

1 **Leveraging Machine Learning Approaches to Predict Organic Carbon Abundance in**
2 **Mars-Analog Hypersaline Lake Sediments**

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4 Floyd Nichols¹, Alexandra Pontefract², Andrew L. Masterson¹, Mia L. Thompson¹, Christopher
5 E. Carr³, Mia T. Tuccillo¹, Magdalena R. Osburn¹

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7 ¹Department of Earth & Planetary Sciences, Northwestern University, Evanston IL, USA

8 ²Space Exploration Sector, Johns Hopkins University Applied Physics Laboratory, Laurel MD,
9 USA

10 ³School of Earth & Atmospheric Sciences, Georgia Institute of Technology, Atlanta, GA, USA

11

12 + corresponding authors [floydnichols2025@u.northwestern.edu; maggie@northwestern.edu]

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14 **Key Points**

- 15 ● Predictive machine learning models have the potential to aid in life detection efforts beyond
16 Earth using organic geochemical datasets.
- 17 ● We predicted organic carbon abundance using only XRF-derived elemental abundances
18 with greater than 80% accuracy.
- 19 ● *Post-hoc* interpretation of our models highlights the importance of elements associated
20 with clays in determining organic carbon concentrations.
- 21 ● We've developed a user-friendly interface to improve the accessibility of our classification
22 model to predict organic carbon from XRF-derived elemental abundances.

23

24 **Abstract**

25 Modern advancements in laboratory and instrumental techniques in astrobiology have
26 improved our life detection capabilities on both Earth and beyond. These advancements have also
27 increased the complexity of data often resulting in datasets that are characterized by complex and
28 non-linear relationships. Machine learning methods are underutilized in astrobiology; however,
29 these methods are extremely effective at revealing structure and patterns in complex datasets when
30 paired with the right algorithms. Here, we employ a series of classification and regression
31 algorithms to predict the abundance of organic carbon (OC) from X-ray fluorescence (XRF) data

32 in dynamic Mars-analog hypersaline lake sediments. More specifically, we constructed models
33 using the random forest (RF), k-nearest neighbors (KNN), support vector machine (SVM), and
34 logistic regression (LR) algorithms. Overall, our trained models showed good performance with
35 predicting the abundance of OC, with accuracies from 80% to 94%. Our results show how applying
36 predictive models to astrobiology datasets can help life detection efforts. Machine learning
37 approaches such as classification and regression algorithms offer insight into complex data while
38 providing agnostic insights, ultimately creating a more efficient search for OC. We applied our
39 trained model on XRF data from Martian soil using PIXL and Odyssey datasets to produce
40 probability predictions of OC abundance. Our predictions show a high probability that OC
41 abundance is low which is comparable to OC data from recently landed missions. These results
42 highlight the potential for machine learning models to be trained on data from analog environments
43 on Earth and then transferred (transfer learning) to extraterrestrial targets.

44

45 **Plain Language Summary**

46 Modern datasets have become large as a by-product of the desire to discover unknown or
47 characterize complex non-linear relationships. Machine learning approaches are extremely
48 valuable for tackling such problems; however, those methods are underutilized in astrobiology and
49 therefore have not been refined for these types of data. Here, we employ machine learning
50 approaches to predict organic carbon abundance from a series of Mars-analog hypersaline lake
51 sediment core and a freshwater lake from X-ray fluorescence-derived elemental abundances.
52 Overall, our models successfully predicted organic carbon concentration, with average accuracies
53 between 80% and 94% and root mean square errors within 1.0 wt% organic carbon. Furthermore,
54 we applied our model to Martian instruments including PIXL and Odyssey. We compute
55 probability predictions that corroborate organic carbon that has been measured on the Martian
56 surface. Our study demonstrates the potential for machine learning methods to be employed to aid
57 in life detection efforts.

58

59 **Introduction**

60 Exploring the potential for ancient or extant life beyond Earth poses many challenges
61 including but not limited to sample selection, sample priority, and the search for an ideal site
62 (Warren-Rhodes *et al.*, 2023; Theiling *et al.*, 2022). These challenges and limitations are often due

63 to a lack of consensus for what establishes favorable features or conditions for life detection
64 beyond Earth (Theiling *et al.*, 2022). Although underutilized in astrobiological research, machine
65 learning methods can be used to mitigate these limitations. Machine learning excels at detecting
66 patterns and structures within large and/or complex datasets (Warren-Rhodes *et al.*, 2023; Theiling
67 *et al.*, 2022; Peuple *et al.*, 2021). As such, machine learning methods can be extremely valuable
68 for using data that is less resource intensive (such non-destructive spectroscopy including X-ray
69 Fluorescence (XRF)) to predict information that is resource intensive (such as organic carbon (OC)
70 or biosignature analysis; Warren-Rhodes *et al.*, 2023; Jacq *et al.*, 2019). For example, the Curiosity
71 Rover can analyze an effectively unlimited number of samples using its remote laser-induced
72 breakdown spectroscopy (LIBS) instrument, ChemCam (Maurice *et al.* 2012), but the Sample
73 Analysis at Mars (SAM) instrument has a limited number of sample cups for analyzing powdered
74 samples via gas chromatography mass spectroscopy (Mahaffy *et al.* 2012). Similarly, the Planetary
75 Instrument for X-ray Lithochemistry (PIXL) and the Scanning Habitable Environments with
76 Raman & Luminescence for Organics & Chemicals (SHERLOC) instruments on the Perseverance
77 Rover can detect elemental abundances using XRF and scan for organics using Raman
78 spectroscopy; however, it is limited in its ability to detect OC beyond the surface of minerals or
79 within inclusions of transparent evaporite minerals (Bhartia *et al.*, 2021).

80 For Earth-based studies, major element analysis via XRF is often used on sediment cores
81 to reconstruct paleohydroclimate (Shea *et al.*, 2022; Puleo *et al.*, 2020; Zhang *et al.*, 2020). In
82 contrast, these analyses applied beyond Earth such as Martian research are mostly limited to
83 exploring the surface to provide insight into the processes that have guided surficial processes and
84 the evolution of the Martian crust-mantle system (Allwood *et al.*, 2015; Hahn *et al.*, 2007). As
85 such, there are rapid and streamlined procedures for the determination of the major element
86 composition of sediments on Earth. XRF can yield abundance information for more than 30
87 elements, ultimately producing relatively large and complex datasets. Due to this complexity, most
88 researchers select a few of the most important elements to probe based on geologic relevance
89 (Puleo *et al.*, 2020; Evans *et al.*, 2019; Zhang *et al.*, 2019; Rothwell and Croudace, 2015).

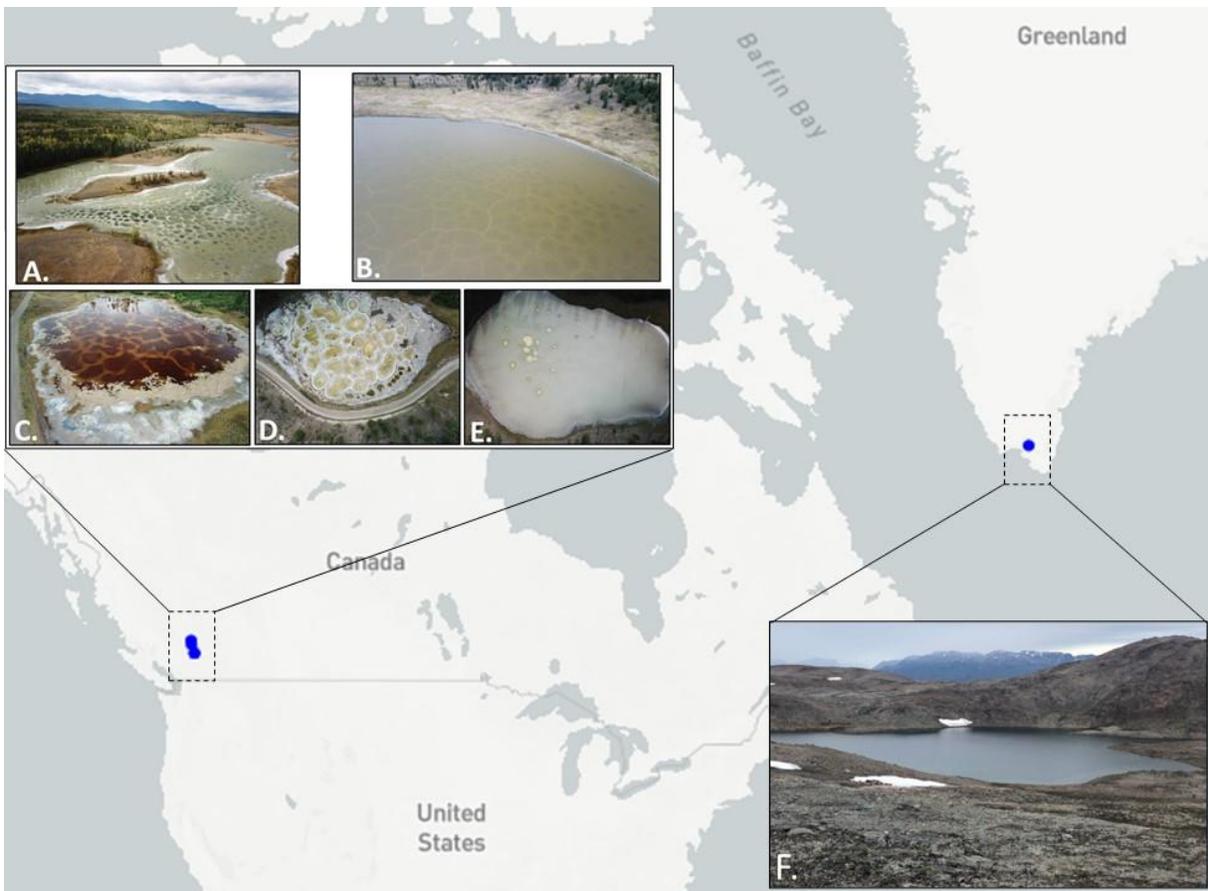
90 Our greater understanding of Earth processes compared to Mars allows for selection of a
91 few elements to explore; however, it is imperative that human biases are eliminated for studies
92 beyond Earth. Thus, similarly to the maximum entropy principle, all pieces of available
93 information should be utilized for exploration to provide agnostic insight into interpretation of the

94 data (Uffink, 1995). Ultimately, viewing data holistically reintroduces the challenges inherent to
95 large and complex datasets. Here, we employ machine learning approaches to reveal patterns and
96 make rapid and agnostic predictions of OC abundance in Holocene-aged Mars-analog hypersaline
97 lake sediment cores from major XRF-derived element abundances. We utilize both unsupervised
98 and supervised learning approaches to understand the structure of the data and make predictions
99 of OC abundance. More specifically, we use unsupervised learning for exploratory data analysis.
100 We then build classification models for broad scale characterization of sediments and then
101 construct regression models for granular predictions of OC abundance.

102

103 **Study Area**

104 We targeted a series of hypersaline lakes located within the Cariboo Plateau of South-
105 Central Interior British Columbia, Canada including Salt Lake, Last Chance Lake, and the Basque
106 Lakes. The chemistry and geography of these lakes are described in greater detail in Nichols *et al.*,
107 2023. In short, these lakes are closed basins and situated within a rain-shadow (~300 mm
108 precipitation per year) contributing to their hypersaline nature. Despite the aridity, heavily
109 vegetated catchment areas surround many of these lakes. Additionally, these lakes feature unusual
110 chemistries dominated by high concentrations of magnesium sulfate (Salt Lake and the Basque
111 Lakes) and of sodium carbonate (Last Chance Lake) ions making them ideal analogs to Mars. To
112 improve the generalization of our model performance we also include a freshwater lake in
113 Greenland near Narsarsuaq, informally named Mel3. Unlike the lakes from the Cariboo Plateau,
114 Mel3 is open-basin, oligotrophic and surrounded by a sparsely vegetated catchment, with the
115 closest meteorological station recording ~650 mm precipitation per year.



116

117 **Figure 1. Location of Study Sites. A.)** Last Chance Lake: 51°19'40.8" N, 121°38'9.6" W; **B.)** Salt
 118 Lake: 51°04'25.44" N, 121°35'11.244" W; **C.)** Basque Lake #1: 50°36'1.8" N, 121°21'32.4" W; **D.)**
 119 Basque Lake #2: 50°35'36.6" N, 121°20'58.2" W; **E.)** Basque Lake #4: 50°35'20.304" N,
 120 121°20'34.397" W.; **F.)** Mel3: 61°07'46.81" N, 45°20'10.68" W. Photo Credit: Mitchell Barklage,
 121 PhD (Canadian Lakes) and Pete J.K. Puleo (Greenland Lake).

122

123 **Methods**

124 *Geochemical Sediment Analysis*

125 In British Columbia, Canada, we sampled nine sediment cores in the summer of 2018 and
 126 the summer of 2019 using a Unicoring device (2018) or an SDI Vibecore mini adapted to 3 inch
 127 diameter polycarbonate pipe (2019). Sediment water interfaces were stabilized with Gelzan and
 128 packing material. Sediment cores were capped and kept cold within 6 hours of collection,
 129 transported to the lab, and stored at 4°C until further processing. The Mel3 sediment core was
 130 collected in August 2022 using an Aquatic Research Instruments “Universal” check-valve and

131 percussion coring system. The core was sealed, transported, and stored at 4°C until analysis. Each
132 sediment core was split using a GeoTek core splitter for bulk analysis and XRF analysis,
133 respectively. The sediment elemental abundance (Ag, Al, As, Bi, Ca, Cd, Co, Cr, Cu, Fe, Hg, K,
134 Light Elements (LE), Mg, Mn, Mo, Nb, Ni, P, Pb, Rb, S, Sb, Se, Si, Sn, Sr, Th, U, V, W, Y, Zn,
135 Zr, Ti), magnetic susceptibility (MS), and core imaging was determined using a Geotek Multi-
136 Sensor Core Logger (MSCL-S) paired with an Olympus Delta X-ray fluorescence (XRF) analyzer,
137 a Bartington MS2E magnetic susceptibility meter, and a 50mm Canon camera of the core split. LE
138 have weaker X-ray energies, thus, are harder to resolve individually. Accordingly, the LE are
139 grouped into a single category. After scanning, bulk sediment material was taken throughout the
140 Canadian cores in approximately 10 cm intervals for ¹⁴C-dating and calibration. The sediment core
141 was then sub-sectioned and homogenized into 3 cm increments for further organic analysis. Bulk
142 sediment samples for ¹⁴C dating were sent to the National Ocean Sciences Accelerator Mass
143 Spectrometry facility at Woods Hole Oceanographic Institution.

144 Total organic carbon (TOC) and total organic nitrogen (TON) abundances were measured
145 in the Northwestern Stable Isotope Biogeochemistry Lab with an elemental analyzer isotope ratio
146 mass spectrometer (EA-IRMS; Costech 4010 EA coupled to a Thermo Delta V+ IRMS through a
147 Conflo IV interface). Freeze-dried samples were weighed then were treated with 1M HCl to
148 remove inorganic carbon and acid soluble salts, rinsed with MilliQ water, then freeze-dried and
149 weighed again. Fourier Transform Infrared Spectrometric (FTIR) analysis on Mel3 sediments
150 confirmed no presence of carbonates (So et al., 2020), and thus these sediments were not acidified
151 before analysis. The homogenized samples were loaded into tin capsules for analysis. Standards
152 were run every 10 samples including IU-acetanilide (precision: ± 1.0%) and urea (precision: ±
153 0.1%; Schimmelman *et al.*, 2009). Additionally, TOC and TON values were corrected for the loss
154 of acid soluble material determined gravimetrically.

155

156 *Model Selection and Descriptions*

157 All models in this study were constructed using the Python package Scikit-learn
158 (Pedregosa, 2011). The explanatory variables were XRF-derived elemental abundances from
159 sediments whereas the response variables were OC abundance divided amongst three classes (high,
160 moderate, and low) as defined and justified below in the *Geochemical and Sediment Analysis*
161 section. There are a variety of machine learning models available, but to limit the number of

162 models used in this study, we choose those that are most common and interpretable/explainable.
163 Interpretable and explainable are defined as the degree to which a human can understand the cause
164 of a decision or consistently predict the model’s result and perform well with small datasets (Belle
165 & Papantonis, 2021; Molnar, 2019; Probst *et al.*, 2018; Kim *et al.*, 2016; Pal & Mather, 2005). We
166 employed t-distributed stochastic neighbor embedding (t-SNE; van der Maaten & Hinton, 2008),
167 principal component analysis (PCA; Jolliffe & Cadima, 2016), logistic regression (Maalouf, 2011),
168 k-nearest neighbor (Taunk *et al.*, 2019), random forest (Biau & Scornet, 2016), and support vector
169 machine (Cortes & Vapnik, 1995). Even though deep learning models such as neural networks
170 have become increasingly common due to their state-of-the-art performance, they require large
171 datasets (thousands of training examples) to perform well (Cronin, 2021), and are therefore not
172 appropriate for our study. Below we give a brief description of the models and hyperparameter
173 tuning we applied; however, a more detailed description of the mathematics involved can be found
174 in the references above.

175

176 *Unsupervised Learning: Dimensionality Reduction*

177 We used unsupervised learning approaches to visualize information such as patterns
178 exclusively from unlabeled data. There are a variety of such algorithms, but we chose two distinct
179 yet common approaches to do so. This included t-distributed stochastic neighbor embedding (t-
180 SNE) and principal component analysis (PCA). The t-SNE method is a non-linear dimensionality
181 reduction data visualization technique that preserves the local structure of data by minimizing the
182 Kullback-Leibler divergence between the two distributions with respect to the locations of the
183 points in the map (van der Maaten & Hinton, 2008). t-SNE excels at revealing structure at many
184 different scales which is very important for high-dimensional data (van der Maaten & Hinton,
185 2008). Conversely, PCA is a linear dimensionality reduction data visualization method that
186 preserves the global structure of the data. To do so, PCA implements an orthogonal transformation,
187 resulting in a number of components equal to or less to the number of original variables (Platzer,
188 2013). As such, using these two methods in tandem allows for the visualization of the data
189 separability as a first-order analysis of the structure of the data.

190

191 *Supervised Learning*

192 We employed four classification algorithms to make broad scale classifications of OC
193 abundance and three regression algorithms to make granular predictions of OC abundance. More
194 specifically, we use logistic regression (LR), k-nearest neighbor (KNN), support vector machine
195 (SVM), and random forest (RF) due to their commonality and explainability. Explainability can
196 be defined as simulatability (ability to be simulated by a human), decomposability (ability to break
197 down a model into parts), and algorithmic transparency (ability to understand the procedure the
198 model goes through to generate its output; Belle & Papantonis, 2021; Lipton, 2016). We applied a
199 Synthetic Minority Oversampling Technique (SMOTE) to address issues of imbalance between
200 classes prior to model construction. The parameters for each model constructed were determined
201 through optimization via an iterative run of parameters. Although by definition SVM and RF are
202 not considered explainable algorithms, *post-hoc* analysis such as feature importance extraction
203 was used to improve their explainability including mean decrease in impurity, permutation
204 importance, and Shapley additive explanations.

205 We calculate a variety of metrics including accuracy, precision, recall, and F1 score to
206 evaluate the errors of each ML model. These metrics were chosen due to their popularity and
207 interpretability. In short, each metric can be defined as follows: *accuracy* is the measure of the
208 fraction of predictions the model got correct:

209 Eq. 1:
$$\left(\frac{\text{number of correct predictions}}{\text{total number of predictions}}\right)$$

210 *precision* is the measure of the proportion of positive identifications correctly identified:

211 Eq. 2:
$$\left(\frac{\text{true positive}}{\text{true positive} + \text{false positive}}\right)$$

212 *recall* is the measure of the proportion of actual positives identified correctly:

213 Eq. 3:
$$\left(\frac{\text{true positive}}{\text{true positive} + \text{false negatives}}\right)$$

214 and *F1 score* is a combined measurement of recall and precision that computes how many times a
215 model made a correct prediction across the entire dataset:

216 Eq. 4:
$$\left(2 \times \frac{\text{Precision} \times \text{Recall}}{\text{Precision} + \text{Recall}}\right)$$

217 In addition to evaluation of errors, we performed a two-step process to prevent and evaluate
218 overfitting. First, we employed a cross-validation technique to check each model's ability to
219 generalize the data. Specifically, we used a repeated stratified k-fold method for cross-validation.
220 The model performance associated with each split was then averaged, allowing for a more accurate
221 estimate of the model's performance. Additionally, since training machine learning models on

222 small datasets can cause the model to memorize all training examples, in turn leading to overfitting,
223 our second step to prevent and evaluate overfitting was introducing varying levels of gaussian
224 noise to the input data. Gaussian noise can lead to an improvement in the generalization of the
225 model performance as it adds structured noise to the input data that is consistent with natural
226 perturbations. Furthermore, when different levels are added, the robustness of each model can be
227 evaluated by comparing the accuracy of the models with respect to each noise level.

228

229 **Results**

230 *Age-Depth Modelling*

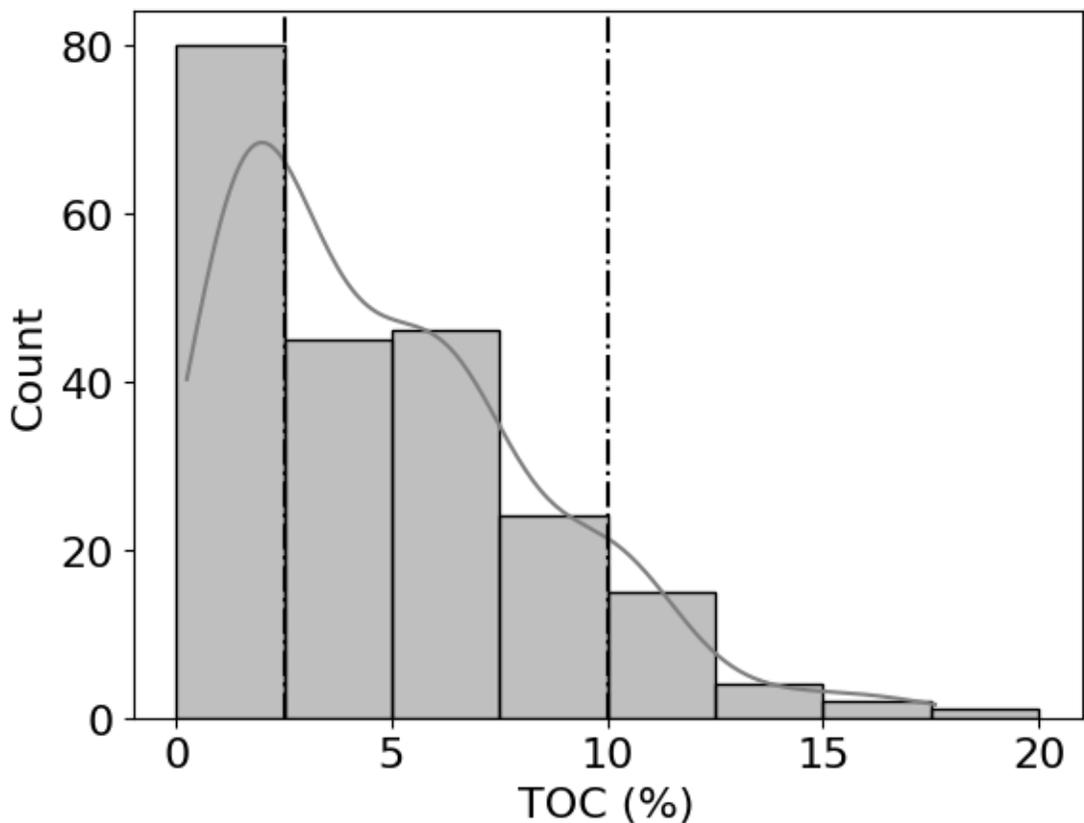
231 The age of the sediment cores was calculated using Bayesian statistics age-modelling with
232 the R Bacon package (Supporting Information; Blaauw & Christen, 2011). Due to a paucity of
233 plant debris in the sediment cores, the age model was calculated using radiocarbon ages derived
234 from bulk OM from the sediments. All sediments were Holocene-aged with ages ranging from
235 present-day to ~6.5 ka with the oldest sediments found at the base of the Salt Lake core. A coherent
236 age model was found for Salt Lake, Basque Lake #1, and Basque Lake #2 whereas significant
237 reversals were noted in the Last Chance Lake cores. Disrupted ages in the upper sediments were
238 likely caused by mixing during coring or through mixing via salt growth and ice growth throughout
239 the year. The radiocarbon age model of the Mel3 core is not reported here; however, preliminary
240 chronologies derived from terrestrial plant macrofossils throughout the sediments indicate the
241 record spans roughly 4.2 ka to present.

242

243 *Geochemical Sediment Analysis*

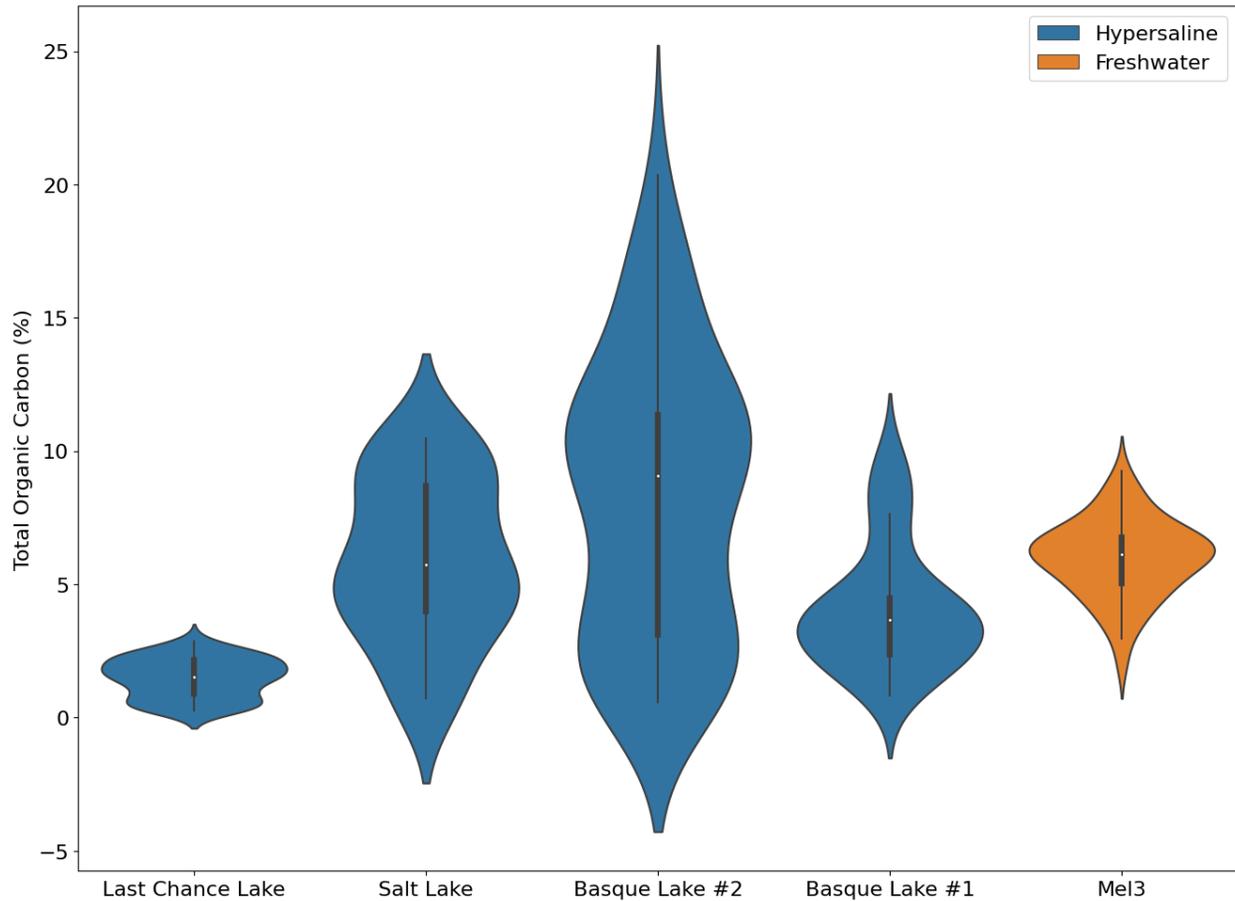
244 The TOC for our samples showed a frequency distribution that had a high density around
245 2.5 wt% TOC that decreased with increasing wt% TOC (Figure 2). Generally, sediments are
246 classified as low in OC if the TOC is less than 1.0 wt% (Fox *et al.*, 2017). Considering our
247 distribution and sparsity of samples with less than 1.0 wt% TOC we considered samples less than
248 2.5 wt% as low. Conversely, sediments are generally classified as organic rich when the TOC is
249 greater than 10.0 wt% (Fuller *et al.*, 2021; Fox *et al.*, 2017). We used the distribution from our
250 data and previously defined boundaries to determine our boundary conditions for the three-class
251 model where low concentration is less than 2.5 wt% TOC, moderate concentrations are between
252 2.5 and 10.0 wt%, and high concentrations are greater than 10.0 wt% TOC.

253 The TOC of our samples ranged from 0.3 wt% to 20.4 wt% (Figure 3). Overall, we
 254 observed the highest average TOC from Basque Lake #2 of 8.2 wt% with minimum and maximum
 255 values of 0.6 wt% and 20.4 wt%, respectively. Conversely, Last Chance Lake had the lowest
 256 average TOC of 1.5 wt% with minimum and maximum values of 0.3 wt% and 2.9 wt%,
 257 respectively. The remaining lakes including Basque Lake #1, Salt Lake, and Mel3 had similar
 258 average TOC values of 4.1 wt%, 5.0 wt%, and 6.0 wt%, respectively. We observed minimum TOC
 259 values of 0.8 wt%, 0.7 wt%, and 2.0 wt% for Basque Lake #1, Salt Lake, and Mel3, respectively.
 260 Conversely, we observe maximum TOC values of 9.8 wt%, 10.5 wt%, and 9.2 wt%.



261
 262 **Figure 2. Histogram and Kernel Density Distribution of Organic Carbon Abundance.** For
 263 this study, we divided OC into three classes: low, moderate, and high (dashed lines). This
 264 classification was based on a combination of the distribution of OC in our sediments and general
 265 classification of sediment OC concentration (Fuller *et al.*, 2021; Fox *et al.*, 2017) where low < 2.5
 266 wt%, 2.5 wt% < moderate < 10.0 wt%, and 10.0 wt% < High. We employed a Synthetic Minority
 267 Oversampling Technique (SMOTE) to address the issue of class imbalance prior to model
 268 construction.

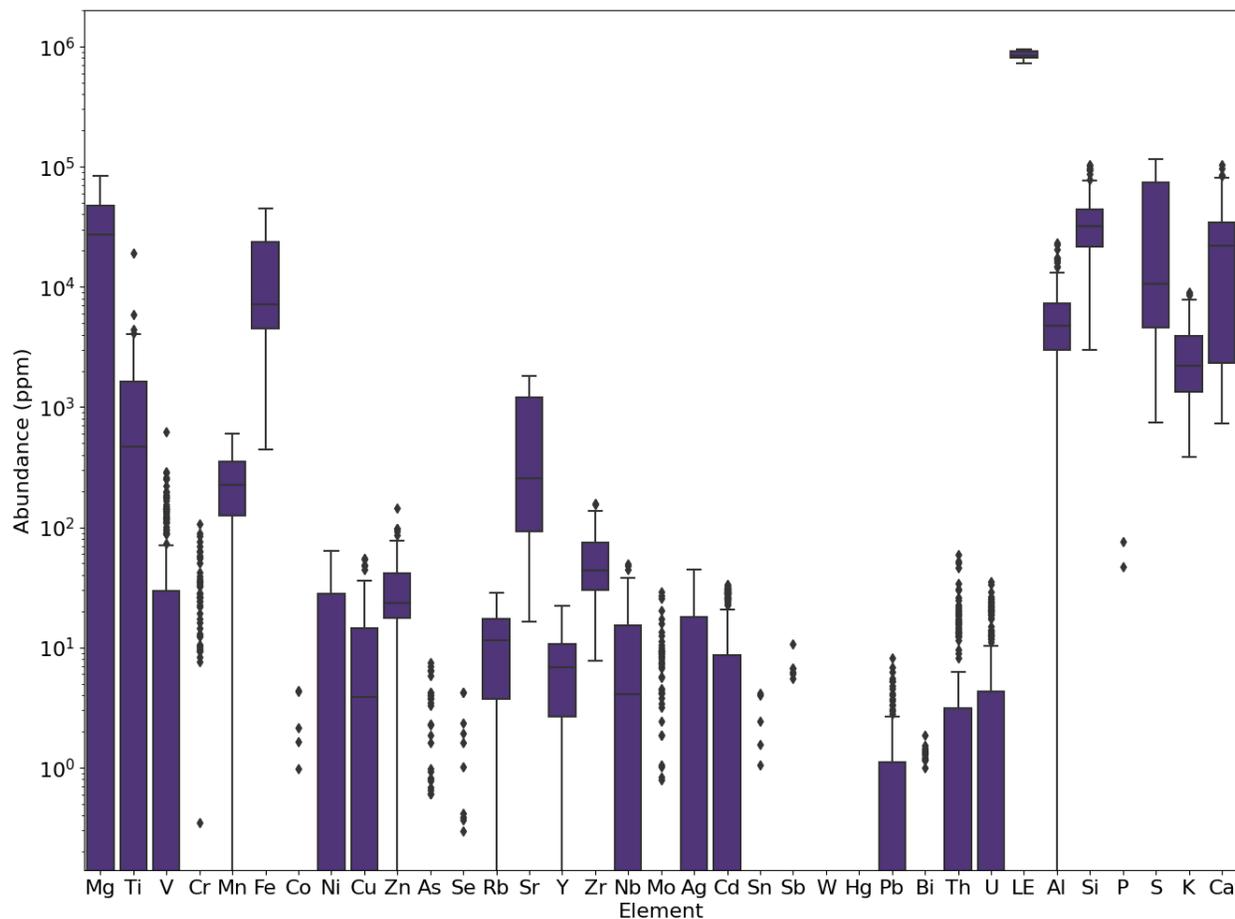
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271
272 **Figure 3. Violin plot of the distribution of TOC by lake.** Blue represents hypersaline Mars-
273 analog systems and orange represents the freshwater system.

274
275 Our XRF analysis shows that the most abundant elements on average were the light
276 elements (LE, atomic mass < Mg); however, of the individual elements that can be resolved, the
277 five most abundant were Mg, Fe, Si, S, and Ca (Figure 4). The average concentrations of the most
278 abundant elements were 2.7 wt%, 1.3 wt%, 3.4 wt%, 3.8 wt%, and 2.4 wt%, respectively. The
279 majority of the other elements had concentrations less than 0.01 wt%.

280



281
 282 **Figure 4. Box and whisker plot of the distribution of detected elements from XRF in our**
 283 **samples with respect to their concentrations in parts per million (ppm).** Points represent
 284 outliers (points farther than $1.5 \times$ interquartile range). The y-axis is plotted on a \log_{10} scale.
 285 Additionally, there is an absence of a boxplot in a few elements due to extremely low abundances
 286 and little variance.

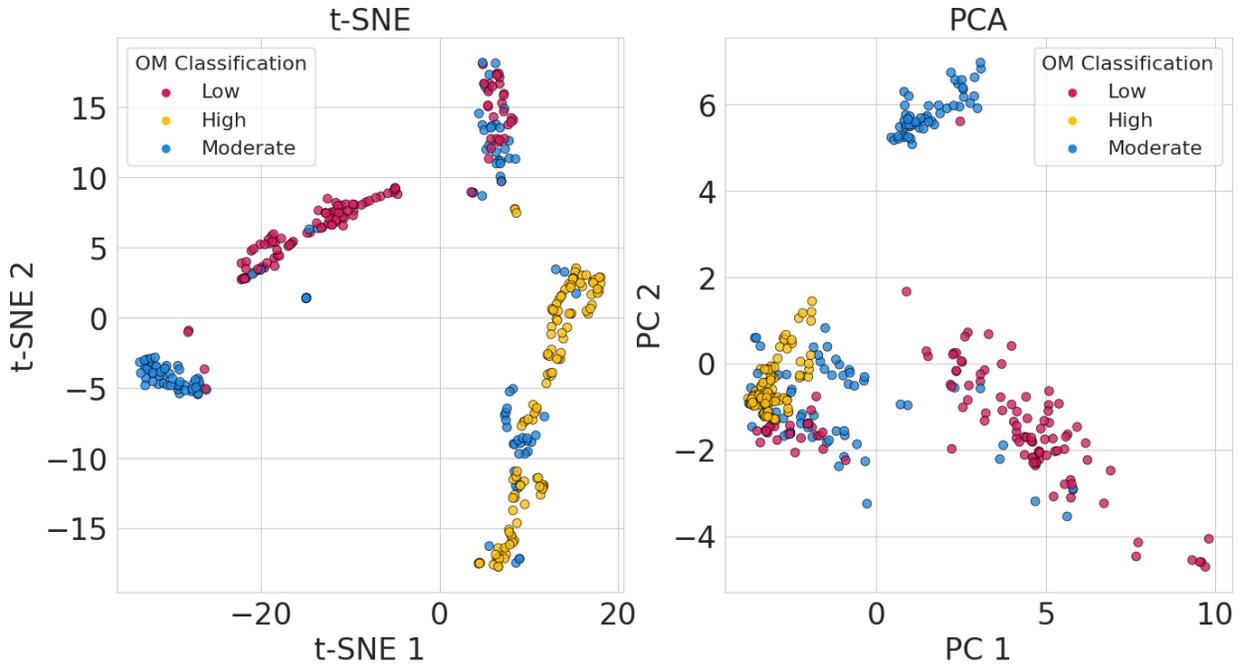
287
 288 *Machine Learning*

289 *Unsupervised Learning*

290 We first employed unsupervised learning approaches to visualize the structure of the data.
 291 The two main approaches used were t-SNE and PCA. In our analysis, we observed that t-SNE
 292 showed 4 distinct clusters while PCA showed 3 distinct clusters (Figure 5). More specifically, in
 293 the t-SNE approach, each OC classification generally formed its own cluster. Conversely, with
 294 PCA, there are clear clusters of low and moderate OC classes, but there is more overlap with

295 respect to the high OC class. Overall, both methods were able to identify distinct clusters within
296 the data, but t-SNE showed better structure in the clusters.

297



298

299 **Figure 5. Unsupervised Learning: Dimensionality Reduction.** We applied two unsupervised
300 learning methods including t-distributed stochastic neighbor embedding (t-SNE; *left*) and principal
301 component analysis (PCA; *right*). Both methods clustered the data to some capacity; however, t-
302 SNE revealed more structure in the data as shown by the four distinct clusters representing the
303 different OC classifications.

304

305 *Supervised Learning*

306 The accuracies of our supervised models ranged from 80% to 94% (Table 1). More
307 specifically, KNN performed with an accuracy of 90%, precision of 89%, recall of 90% and F1
308 score of 89%. Random Forest (RF) performed with an accuracy of 94%, precision of 94%, recall
309 of 95% and F1 score of 94%. SVM performed with an accuracy of 80%, precision of 81%, recall
310 of 95%, and F1 score of 80%. Logistic regression performed with an accuracy of 91%, precision
311 of 91%, recall of 91% and F1 score of 91%. We used cross-validation to evaluate the generalization
312 capabilities of our models. The cross-validation accuracies of our RF, KNN, SVM, and LR were
313 91%, 89%, 84%, and 88%, respectively.

314

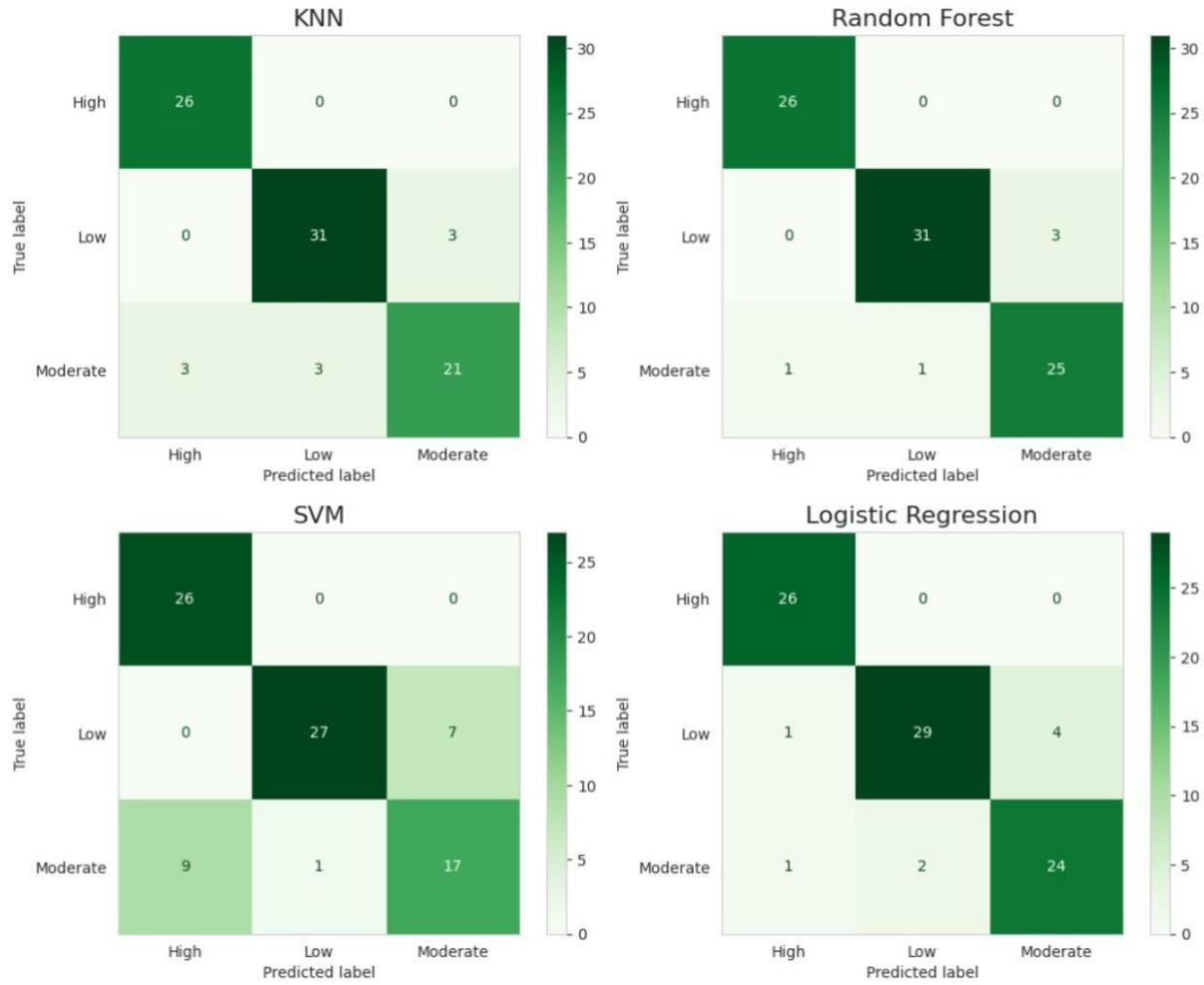
315 **Table 1. Model Performance Metrics.**

Model	Training Accuracy	Precision	Recall	F1	CV
KNN	90	89	90	89	88
RF	94	94	95	94	91
SVM	80	81	81	80	84
LR	91	91	91	91	88

316

317 To further interrogate the performance of our models, we computed a confusion matrix for
318 each model (Figure 6). In short, a confusion matrix allows for the visualization of correctly labeled
319 and mislabeled classifications. Generally, all models correctly labeled high and low OC abundance
320 with high accuracy (~80%). Conversely, all models correctly labeled moderate OC abundance with
321 a lower accuracy. The SVM and KNN models were more likely to misclassify moderate OC
322 abundance (54% and 70%, respectively) than the RF and LR (79%). We computed feature
323 importance calculations using mean decrease in impurity, permutation importance, and Shapley
324 additive explanations to extract additional information from our models, (Figure 7). These
325 calculations highlight that the top elemental abundances contributing to the classification of OC
326 were Al, Fe, Si, Ti, U, and Zn.

327



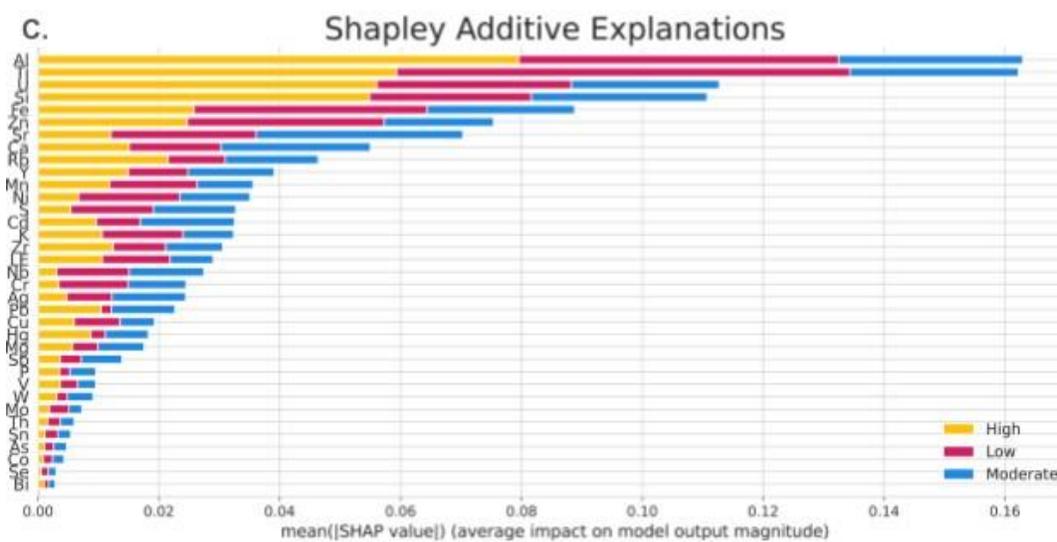
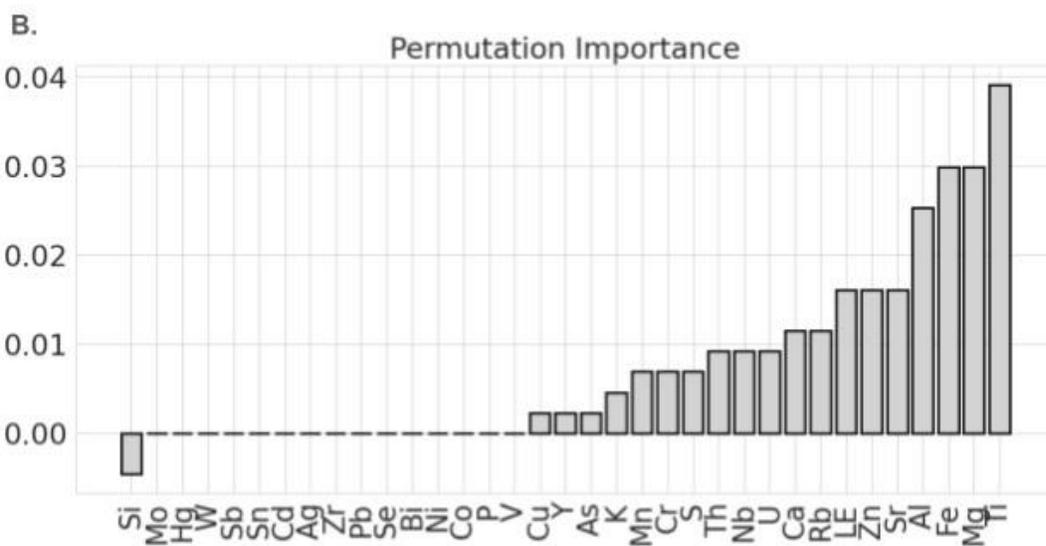
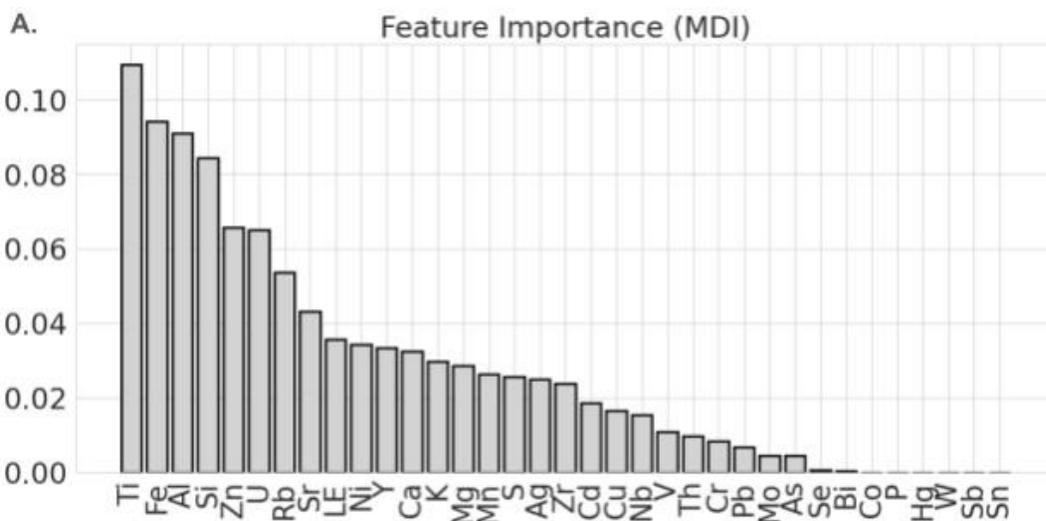
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330 **Figure 6. Confusion Matrix.** Visual representation of correctly labeled classes against mislabeled
 331 classes.

332

333



335 **Figure 7. Interpretation of Model Results.** A.) Feature Importance Mean Decrease in Impurity
 336 (MDI), B.) Permutation Importance, C.) Shapley Additive Explanations.

337

338 We also introduced three levels of gaussian noise (1%, 5%, and 10%) to improve
 339 generalization and evaluate overfitting for each model (Table 2). We observe that in general, the
 340 model performance decreases slightly with added gaussian noise. More specifically, there is a
 341 gradual decrease in overall model performance including accuracy, precision, recall, and F1 scores
 342 with each gaussian noise percentage step. Despite the performance of the models decreasing with
 343 respect to the original models with no added noise, we still obtain accuracies above 80% even at
 344 10% added gaussian noise.

345

346 **Table 2. Model Performances with Varying Levels of Gaussian Noise.**

Gaussian Noise	Model	Accuracy	Precision	Recall	F1
1%	KNN	87	88	87	87
	RF	84	84	84	83
	SVM	85	86	85	85
	LR	90	90	89	90
5%	KNN	86	87	86	86
	RF	84	84	84	83
	SVM	82	83	81	81
	LR	89	89	88	88
10%	KNN	84	86	85	83
	RF	85	87	86	84
	SVM	83	85	84	82
	LR	84	84	85	84

347

348 In addition to constructing a classification model, we built three regression models
 349 including random forest regression (RF), support vector regression (SVR), and k-nearest neighbor
 350 regression (KNN) to examine OC prediction at a more granular level. Overall, our models show
 351 good performance across all metrics including coefficient of determination (r^2), root mean squared
 352 error (RMSE), mean absolute error (MAE), and mean absolute percent error (MAPE; Table 3). Of
 353 the models, RF had the best metrics, followed by KNN, and then SVR (Table 3). We compute a
 354 moderate average r^2 value of 0.63 across all models. Additionally, we compute an average RMSE
 355 value of 1.05 across all models.

356 **Table 3. Statistic metrics for our regression models.**

Model	r ²	RMSE	MAE	MAPE
RF	0.80	0.80	0.61	0.17
SVR	0.51	1.24	0.95	0.37
KNN	0.60	1.12	0.85	0.29

357

358 **Discussion**

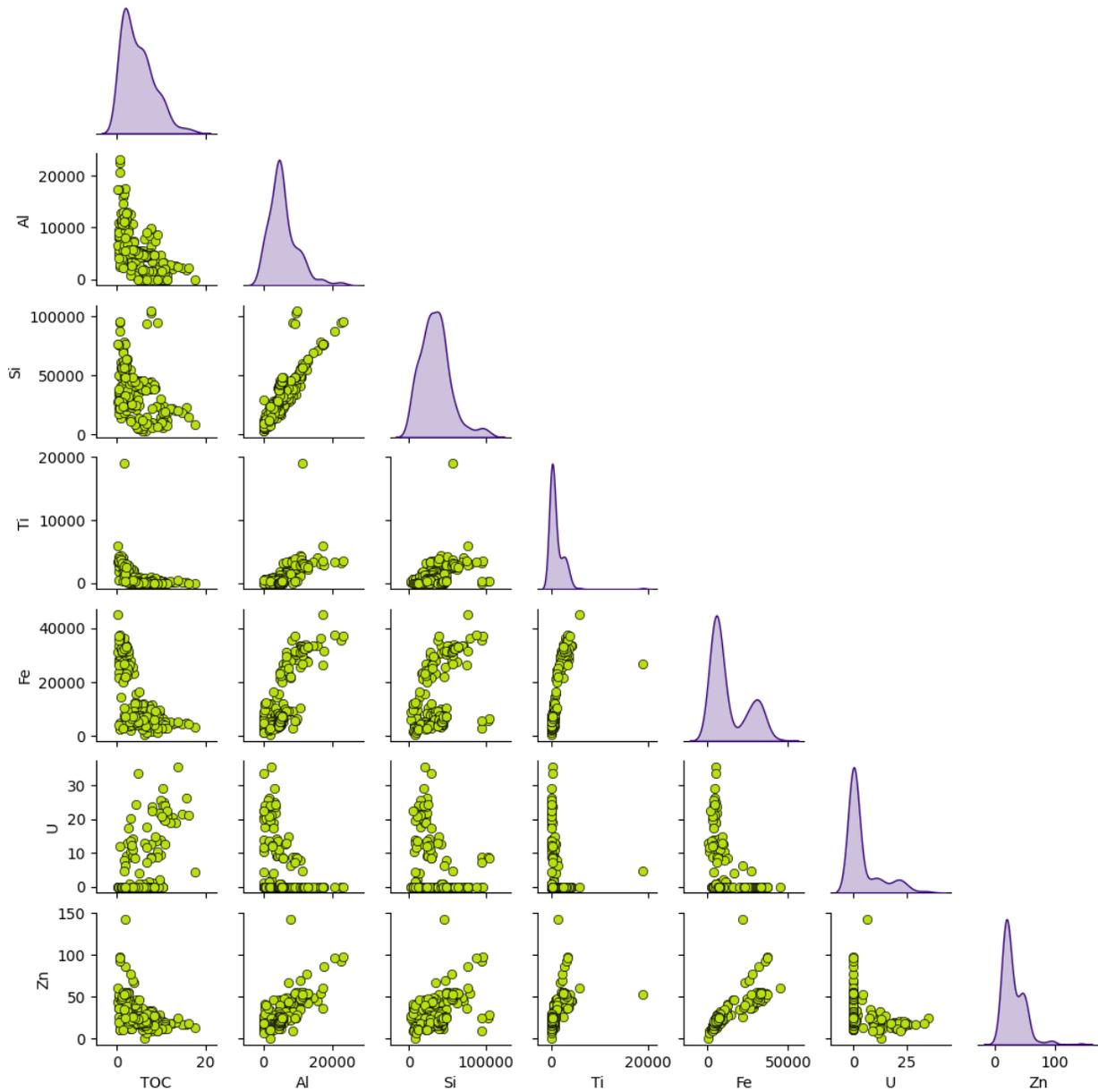
359 *Sediment Geochemistry*

360 On Earth, several competing mechanisms control the abundance of OC in sediments
 361 including mineral protection from clays and salts, selective preservation of refractory
 362 biomolecules, chemical speciation of trace metals and their affinity towards organic compounds,
 363 and redox conditions (Hemingway et al, 2019; Burdige, 2007; Aubrey *et al.*, 2006; Bilali *et al.*,
 364 2002; Hedges et al, 2001). Our feature extraction analysis converged on common elements that
 365 had the greatest effect on the model predictions for OC abundance. These elements included Al,
 366 Fe, Si, Ti, U, and Zn. The abundance of OC in sedimentary systems can be indirectly linked to
 367 elemental abundances (Evans *et al.*, 2019; Bilali *et al.*, 2002). For instance, the concentration of
 368 OC in lake systems often depends on aridity, salinity, and other hydrologic variations which can
 369 be captured by elements linked to phyllosilicates such as Al, Fe, and Si or detrital elements such
 370 as Ti and Zn (Evans *et al.*, 2019; Zhang *et al.*, 2019; Rothwell and Croudace, 2015; Bilali *et al.*,
 371 2002). Conversely, OC may be directly linked to elements such as trace metals including U in
 372 which OC influences its mobilization within sediments (Bone *et al.*, 2019; Bone *et al.*, 2017;
 373 Cumberland *et al.*, 2016; Bilali *et al.*, 2002). These relationships can be complex and challenging
 374 to unravel in lake sedimentary systems as these elements and OC can show varying degrees of
 375 correlations due to their potential to become bound to the sediment or released into the water and
 376 removed from the system (Frings *et al.*, 2014; Makinen *et al.*, 2005; Bilali *et al.*, 2002). Improving
 377 our understanding of these complex mechanisms will be key to understanding areas of interest for
 378 OC on the Martian landscape.

379 Appropriately, we capture a wide range of OC concentrations from lean samples (<1 wt%
 380 TOC) to organic rich samples (>10 wt% TOC) in addition to varying levels of XRF-derived
 381 elemental abundances. The varying levels of TOC and elemental abundances, especially within an
 382 individual lake, highlight the dynamic nature of these lakes. In our analysis of the relationship
 383 between TOC and the top elements from our feature extraction analysis, we show that Al, Fe, Si,

384 Ti, and Zn have very strong positive correlations between each other (Figure 8). This strongly
385 suggests that Al, Fe, Si, Ti, and Zn inputs are linked to weathering as they are generally associated
386 with phyllosilicates and detrital elements (Evans *et al.*, 2019; Zhang *et al.*, 2019; Rothwell and
387 Croudace, 2015). Conversely, TOC shows moderately positive correlations with U. This
388 relationship between OC and U is potentially due to organic ligands forming stable complexes
389 with U (Bone *et al.*, 2019; Bone *et al.*, 2017; Bilali *et al.*, 2002). Bone *et al.*, 2017 show that as
390 OC increases, U increases proportionally as the U has a higher potential to adsorb to the surface
391 of organic matter. Another potential mechanism controlling this relationship between OC and U
392 is due to a redox effect. It is known that oxidized U (VI) is soluble whereas reduced U (IV) is
393 insoluble. Organic matter acts as the reductant that immobilizes U through its reduction
394 (Cumberland *et al.*, 2016). Ultimately, both of these potential mechanisms influence the fate of U
395 in subsurface sediments with the same behavior that we observe in the sediments in this study.

396 On the other hand, we observe moderately negative correlations with Al, Fe, Si, Ti, and Zn
397 (Figure 8). We suggest the following mechanisms for the observed correlation. As previously
398 mentioned, the concentrations of Al, Fe, Si, Ti and Zn are controlled by hydrologic processes
399 including weathering and aridity due to their association with phyllosilicates and detrital material.
400 As weathering increases and aridity decreases, the loading of these elements to the sedimentary
401 system increases as described by the chemical index of alteration (Wang *et al.*, 2020). Considering
402 the negative correlation that we observe with OC abundance with respect to Al, Fe, Si, Ti, and Zn
403 we suggest that periods of increased lake desiccation and aridity are the primary driving forces for
404 increased OC. It has been shown that the increasing desiccation of lakes drives an increase of
405 nutrients and ions to the system, ultimately increasing OC productivity (Sarkar *et al.*, 2023; Duarte
406 *et al.*, 2008; Jones and Decampo, 2003). Lake desiccation can also drive evaporite mineral dilution
407 effects such that as the sediments fluctuate between low and high weight percentage of soluble
408 evaporite minerals, the signal of the elements associated with phyllosilicates and detrital material
409 are diluted out. This effect would also influence the negative correlation that we observe with OC.
410 Additionally, an increase in salinity is known to slow remineralization rates of OC, resulting in a
411 positive correlation of OC abundance with drier periods (Jellison *et al.*, 1996).



412

413 **Figure 8. Pairplot of relationship between TOC and the top elemental abundance features.**

414 Using the top five features from our model, we compute the pairwise relationship to evaluate their
 415 correlation with OM. Abundance is in ppm.

416

417 *Model Performance and Validation*

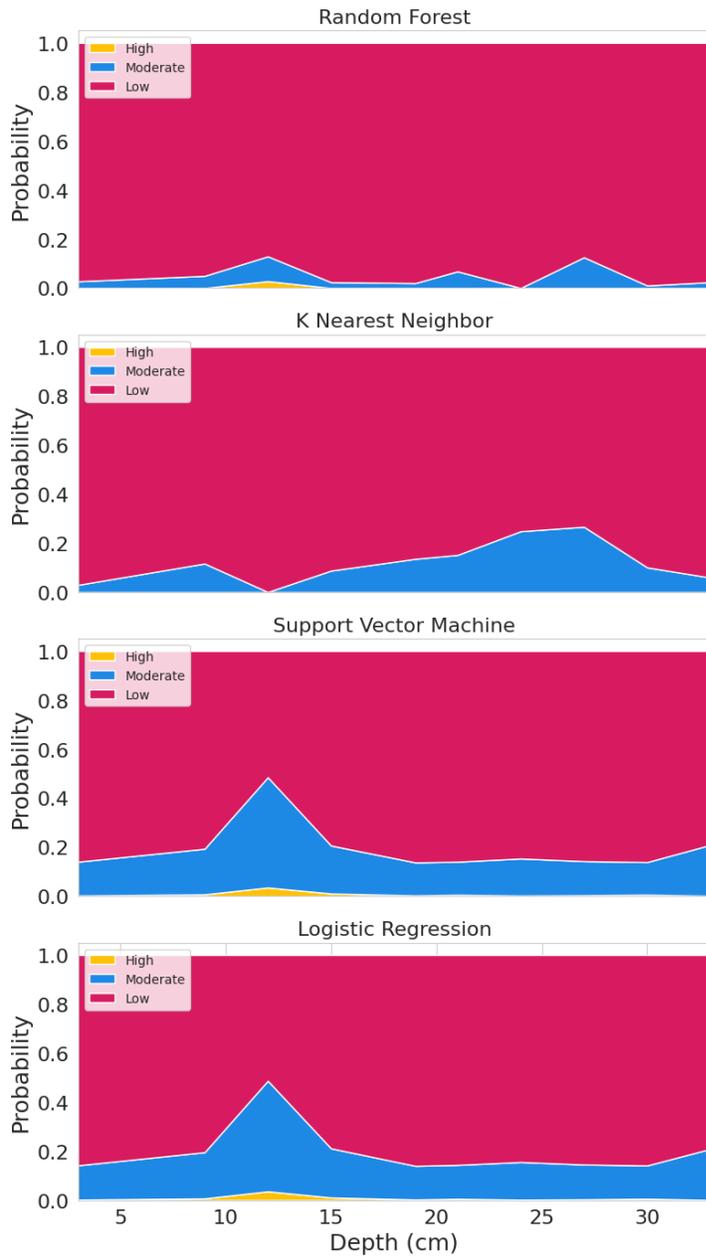
418 PCA is widely used for visualization of high dimensional data and data pre-processing;
 419 however, while robust in some scenarios such as handling linearly separable data, it does not
 420 always perform well where the data has a non-linear structure. Additionally, the clustering of

421 points in a PCA is highly affected by outliers. In contrast, t-SNE has the ability to reveal more
422 structure, underlying patterns, assess data separability, and better handles outliers. Despite these
423 advantages, t-SNE is also limited in its use as a critical drawback of its approach compared to PCA
424 is its inability to preserve the global structure of the data, thus, making it useful only for data
425 visualization and not pre-preprocessing. As such, the tandem use of these methods allows for
426 holistic exploratory data analysis.

427 In this study, we show that t-SNE clustered the data more distinctly with respect to PCA.
428 These results suggest that our data is complex and has a non-linear structure, as highlighted by our
429 OC and elemental abundance data, rather than a linear structure. The supervised learning
430 algorithms that we employed validates our results from the unsupervised learning as the highest
431 performing algorithms for both classification and regression analysis were RF and KNN.
432 Similarly, to t-SNE, RF and KNN are generally better with handling data with non-linearities
433 compared to SVM or LR (Acito, 2023; Chauhan *et al.*, 2019; Auret & Aldrich, 2012; Maalouf,
434 2011). It is important to note that SVM can overcome non-linear challenges when different kernels
435 are employed such as the radial basis function (RBF); however, it often does not perform as well
436 algorithms specifically designed to handle non-linearities (Dong *et al.*, 2014). Our results from our
437 unsupervised and supervised learning approaches are promising as it is very likely when
438 considering life detection beyond Earth, specifically Mars, the environments for which life may
439 have thrived were likely in dynamic and non-linear states as the planet evolved. Subsequently,
440 choosing the correct algorithms for predictions will require knowledge of the environment.

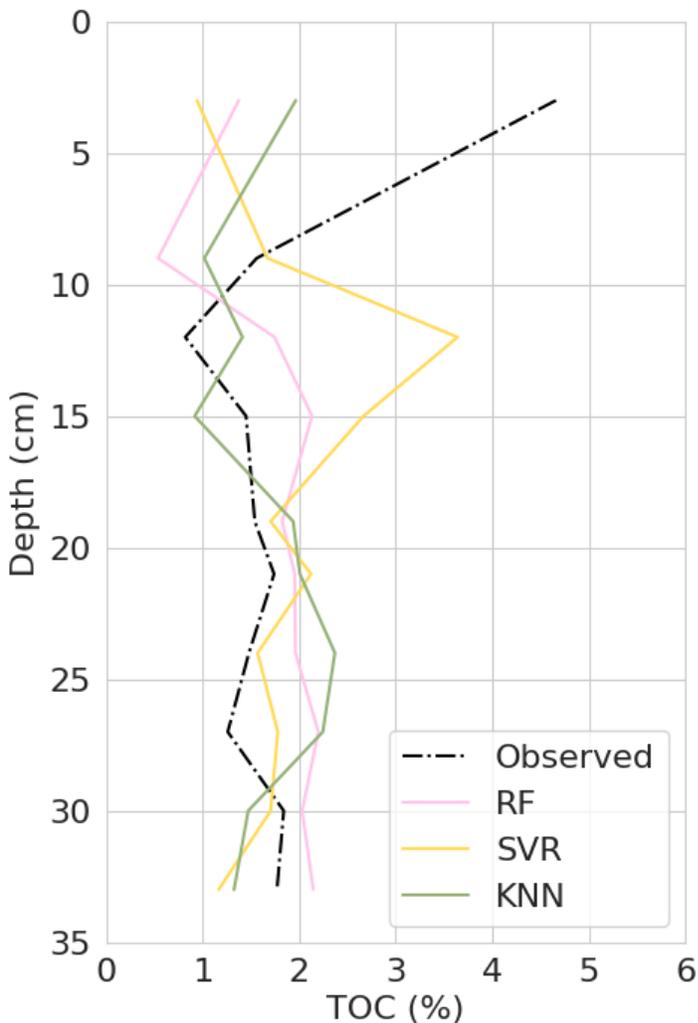
441 Additionally, we used our trained models to predict the OC abundance of a new sediment
442 core to validate the models and produce a probability map of the OC abundance within that core
443 (Figure 9). This sediment core (Last Chance Lake 5) from British Columbia was relatively low in
444 OC with the exception of the topmost point which was greater than 2.5% OC. All classification
445 models showed good performance (~80% accuracy) with predicting the actual TOC classification
446 of the sediment core. We then examine the predictive capability of our regression models for the
447 same sediment core (Figure 10). Similarly, to our classification models, all three regression models
448 were generally able to predict the TOC to within 1.0 wt% of the observed data. More specifically,
449 ~80% of the predicted samples were within 1.0 wt% TOC of the observed data. In general, RF and
450 KNN outperformed SVR as they predicted TOC more closely to the real data. It is worth
451 mentioning that all three of our models generally predicted slightly higher than the observed

452 values. This behavior is likely due to the tested core (Last Chance Lake 5) being exceptionally
453 lean in TOC (mean = 1.8 wt%) with respect to the training data, thus overpredicting values.



454
455 **Figure 9. Model Validation. Area Plot of Classification Probabilities.** To validate the models,
456 we apply new data that the models have never seen. Similar to the training sets, all models
457 classified the sediment organic carbon abundance with ~80% accuracy.

458



459
 460 **Figure 10. Regression Model Comparison to Real Data.** We plot our modelled predictions
 461 (solid-colored lines) with respect to the observed data of Last Chance Lake 5 (black dashed line)
 462 to compare the accuracy of their predictions. Generally, all models predicted within 1.0% of the
 463 actual data. Additionally, RF and KNN outperformed SVR.

464
 465 *Graphical User Interface: Organic Matter Abundance Predictor (OMAP)*

466 To improve accessibility of the model and provide rapid prediction of OC abundance from
 467 XRF-derived elemental abundances, we have developed an open-source graphical user interface:
 468 Organic Matter Abundance Predictor (OMAP; Nichols, 2023). This application is an interactive
 469 data visualization tool and predictor for the constructed model (Figure 11). Due to the random
 470 forest algorithm having the best performance for our testing and validation, we use it as the primary
 471 algorithm for the application. The application consists of three main components including model

472 data and performance visualization, organic carbon probability predictor, and geographic
473 distribution of samples that the model is based on. Additionally, this application serves as an open-
474 source database for others to add OC and XRF-derived elemental abundances from other
475 sedimentary systems to improve and expand upon the model.

A.

Organic Matter Abundance Predictor (OMAP)

Information **Model Data & Performance** Make Prediction Geographic Distribution

Done! (using st.cache_data)

Show Model Data

Select elements (default is all):

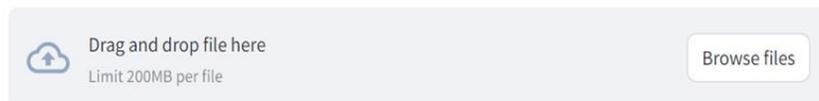


B.

Information Model Data & Performance **Make Prediction** Geographic Distribution

Upload a Data File for Organic Matter Prediction

Choose a file



476
477 **Figure 11. Layout of the Graphical User Interface: Organic Matter Abundance Predictor**
478 **(OMAP).** A.) Model data and performance tab which allows the user to selectively choose the
479 parameters to make OC predictions including specific elements and boundary conditions for OC
480 classification. Additionally, once the elements and boundary conditions are chosen, the data
481 dimensionality reduction and model performance metrics will be printed and B.) The make a
482 prediction tab allows for the user to make a probability prediction of OC abundance based on the

483 elements and OC boundary conditions chosen in the Model data and performance data via
484 uploading a CSV that includes a sample column and elements of interest.

485

486 *Transfer Learning: Mars Regolith*

487 As a proof-of-concept, we applied our model to the average elemental abundance
488 composition from Mars regolith samples as determined by Perseverance’s Planetary Instrument
489 for X-ray Lithochemistry (PIXL; Christian *et al.*, 2023) (Table 4). Our model computes a 73.5%
490 probability that the OC abundance is low ([OC] < 2.5 wt% TOC), a 17.8% probability that it is
491 moderately abundant in OC (2.5 wt% < [OC] < 10.0 wt%), and an 8.2 % probability that OC is
492 high in abundance (10.0 wt% < [OC]). These predictions corroborate OC analyses that have been
493 done on Martian soils using resource intensive combustion methods from Curiosity in which they
494 determined the sediment to contain ~0.1% OC (Stern *et al.*, 2022). Importantly, this calculation is
495 only for the average composition of Mars regolith. In addition, we also applied our model to
496 elemental abundances from the Mars Odyssey Orbiter (which uses gamma-ray spectroscopy rather
497 than XRF to determine elemental abundances) to make predictions for OC abundance. Unlike
498 PIXL, Mars Odyssey determines elemental abundances at a much lower resolution. This lower
499 resolution allows it to cover broader areas of the Martian surface. Hahn *et al.*, 2007 determined
500 the elemental abundance average for three different aged soils/rocks including Noachian,
501 Hesperian, and Amazonian. We use this data to determine the probability of OC abundance as a
502 comparison to PIXL. Interestingly, we compute very similar probabilities to those determined from
503 PIXL elemental abundances (Table 4). On average, we calculate an 80.3% probability that OC
504 abundance is low, a 19.7% probability that OC abundance is moderate, and a 0% probability that
505 OC abundance is high.

506

507 **Table 4. Classification probability of Martian soils.**

Instrument/Sample	High	Moderate	Low
PIXL	8.2	18.2	73.5
Odyssey (Noachian)	0	17.7	82.2
Odyssey (Hesperian)	0	17.8	82.3
Odyssey (Amazonian)	0	23.7	76.3

508 Despite the overall good performance of our model, it is important to note that we cannot
509 assume that OC behaves exactly the same on Mars as it does on Earth. As such, our model serves
510 as a basis for transfer learning which in this instance is an algorithm that is trained using Earth-
511 based data sets after which that knowledge is “transferred” and applied to Mars (Theiling *et al.*,
512 2022). An advantage to the transfer learning concept is that the vast amounts of datasets on Earth
513 can serve as a starting point to eventually adapt algorithms for Mars, which is currently necessary
514 considering the paucity of data on Mars compared to Earth (Theiling *et al.*, 2022).

515

516 *Potential Earth-Based Application*

517 Although our model was constructed with the goal of life detection beyond Earth, there are
518 potential applications for Earth-based analysis. For instance, carbon flux through time in lake and
519 marine sediments is an active area of research (Sarkar *et al.*, 2023; Lee *et al.*, 2019; Zhang *et al.*,
520 2018; Leach *et al.*, 2008). More specifically, high quality, continuous, and high-resolution records
521 are sought after (Leach *et al.*, 2008). One of the limitations to continuous and high-resolution
522 studies is that due to the time intensive nature of OC analysis, it is more challenging to achieve the
523 same level of resolution as elemental abundances from XRF, ultimately resulting in a sparsity of
524 OC data (Lee *et al.*, 2019; Zhang *et al.*, 2018). Our model potentially provides a useful screening
525 tool to evaluate where to target biomarker work and/or assess necessary analytical amounts for
526 TOC and biomarker work.

527 Our model has the potential to be especially useful for hypersaline lakes as they are widely
528 distributed across the globe; however, they are often neglected in climate models as they are
529 assumed to be smaller in number and sparsely distributed (Sarkar *et al.*, 2023; Marce *et al.*, 2019).
530 Contrary to this belief, hypersaline lakes contribute significantly to the global lake volume (85 x
531 10³ km³ to the total 190 x 10³ km³; Sarkar *et al.*, 2023; Williams, 2002). This assumption has
532 created a blind spot in the global carbon cycle, which is estimated by the amount of primary
533 production and long-term sediment storage of OC (Marce *et al.*, 2019). Considering this,
534 hypersaline environments have a great potential to influence the global carbon budget and
535 ultimately global climate. Given their potential to strongly influence global climate, it is crucial
536 that we identify a rapid way to identify the amount of potential carbon flux into the atmosphere.

537

538 **Conclusion**

539 This work provides a proof-of-concept application for leveraging machine learning models
540 to aid in life detection efforts. More specifically, we use data that is relatively easy to collect (XRF-
541 derived elemental abundances) to make predictions about data that is more resource intensive (OC
542 analysis). All models constructed in this study showed good performance (>80% accuracy) for
543 predicting OC abundance in lake sediments from XRF data; however, the RF algorithm
544 outperformed the others for both classification and regression predictions. Warren-Rhodes *et al.*,
545 2023 showed that a random biosignature search yielded only a 9.2% probability of detecting
546 biosignatures. As such, our models improve the probability of detecting OC with varying levels
547 by >70% compared to random searches. Ultimately, our model has the potential to be used to make
548 predictions about OC abundance prior to sample analysis on high-resolution but narrow field of
549 view Martian rovers to save time and resources for life detection. Additionally, it can be used to
550 create a prediction map at the global scale using orbiters such as Mars Odyssey which are low-
551 resolution but have a broad field of view.

552 Despite the overall good performance of our model, we recognize that a model is only as
553 strong as its training set, and this model is most adapted to hypersaline, lacustrine sediments.
554 Further efforts for broader classification tools will require more data from diverse environments
555 and older sediments or rocks. Additionally, our model lacks an abundance of samples that are very
556 lean in OC (<1 wt% [OC]), thus, it can also benefit from an addition of more samples lean in
557 organic material. While the models constructed in this work have showed good performance,
558 future work will focus on adding more data from diverse environments as described above in
559 addition to extending the explanatory variables to mineralogy. Our work is expected to be critical
560 for better understanding which samples and/or sites on Mars are most likely to harbor an
561 abundance of OC and will further propel future life detection missions.

562

563 **Open Research**

564 All data and Python scripts written for data analysis, visualization, and model construction
565 are hosted at <https://zenodo.org/records/10433398> and can be cited as “Nichols, Floyd. (2024).
566 FloydNichols97/OMAP: OMAP. Release (1.0) [Dataset]. Zenodo.
567 <https://zenodo.org/records/10433398>”

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