

1       **Compartmentalization of Axial Seamount’s magma reservoir inferred by**  
2       **analytical and numerical deformation modeling with realistic geometry**

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16      **Key Points:**

- 17      • Uniform pressurization of Axial Seamount's seismically imaged magma reservoir does  
18      not adequately fit the observed geodetic data
- 19      • Our models estimate that Axial’s magma reservoir inflated by 0.054-0.060 km<sup>3</sup> during  
20      the inter-eruptive recharge period between 2016-2020
- 21      • Axial's magma reservoir is likely compartmentalized, with magma accumulating in sills  
22      along the western-central edge of the magma reservoir  
23

24 **Abstract**

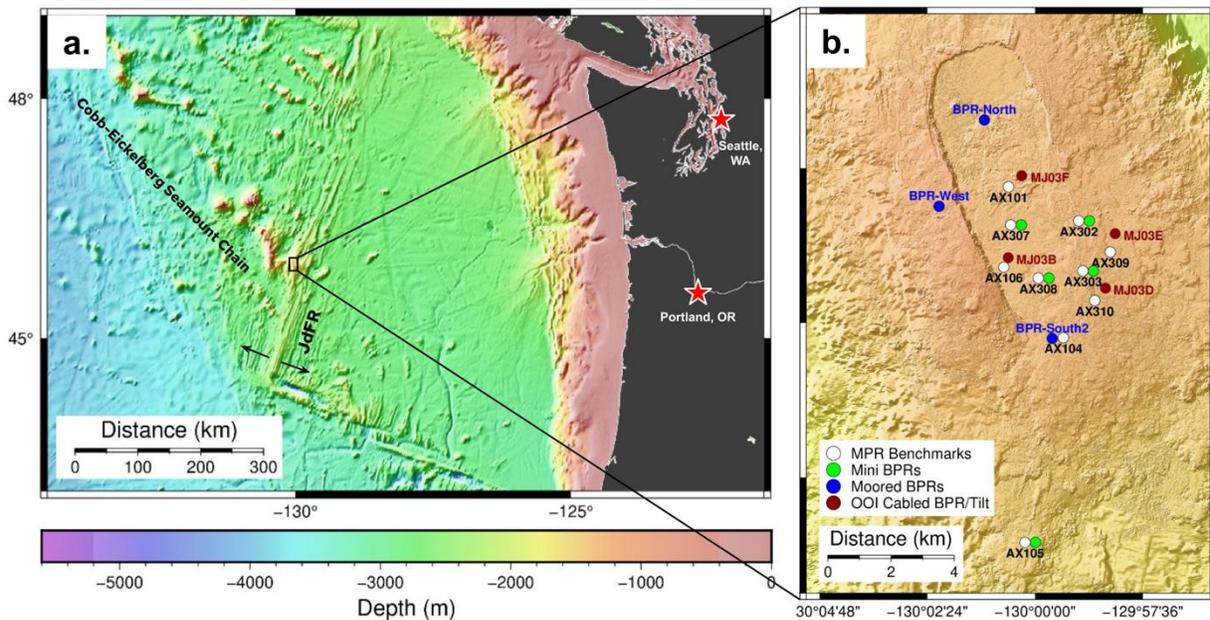
25 Axial Seamount is a submarine volcano on the Juan de Fuca Ridge with enhanced magma supply  
26 from the Cobb Hotspot. Here we compare several deformation model configurations to explore  
27 how the spatial component of Axial's deformation time series relates to magma reservoir  
28 geometry imaged by multi-channel seismic (MCS) surveys. To constrain the models, we use  
29 vertical displacements from pressure sensors at seafloor benchmarks and repeat autonomous  
30 underwater vehicle (AUV) bathymetric surveys covering 2016-2020. We show that  
31 implementing the MCS-derived 3D main magma reservoir (MMR) geometry with uniform  
32 pressure in a finite element model poorly fits the geodetic data. To test the hypothesis that there  
33 is compartmentalization within the MMR that results in heterogeneous pressure distribution, we  
34 compare analytical models using various horizontal sill configurations constrained by the MMR  
35 geometry. Using distributed pressure sources significantly improved the Root Mean Square Error  
36 (RMSE) between the inflation data and the models by an order of magnitude. The RMSE  
37 between the AUV data and the models was not improved as much, likely due to the relatively  
38 larger uncertainty of the AUV data. The models estimate the volume change for the 2016-2020  
39 inter-eruptive inflation period to be between 0.054-0.060 km<sup>3</sup> and suggest that the MMR is  
40 compartmentalized, with most magma accumulating in sill-like bodies embedded in crystal mush  
41 along the western-central edge of the MMR. The results reveal the complexity of Axial's  
42 plumbing system and demonstrate the utility of integrating geodetic data and seismic imagery to  
43 gain deeper insights into magma storage at active volcanoes.  
44

45 **Plain Language Summary**

46 Axial Seamount is a submarine volcano on the Juan de Fuca Ridge (NE Pacific Ocean) with  
47 enhanced magma supply from the Cobb Hotspot. Its frequent activity and long-term deformation  
48 time series covering eruptions in 1998, 2011 and 2015 make it an ideal place to study volcanic  
49 processes. Improved magma reservoir modeling at Axial will aid in understanding how magma  
50 transport and storage are related to surface deformation, seismicity, and eruption timing. Here we  
51 compare several models of Axial's magma reservoir to explore how the spatial component of the  
52 observed deformation at Axial compares to seismically imaged magma reservoir geometry. To  
53 constrain the models, we use vertical displacements covering an inflation period between 2016-  
54 2020, derived from pressure measurements collected at seafloor benchmarks and repeated  
55 bathymetric surveys. The models estimate the volume change for the 2016-2020 inflation period  
56 to be between 0.054-0.060 km<sup>3</sup>. Our results suggest that the Axial's magma reservoir is  
57 compartmentalized, with most magma accumulating in sill-like bodies embedded in crystal  
58 mush. The results reveal the spatial complexity of Axial's plumbing system and demonstrate  
59 how deformation data and seismic imagery can be used together to gain deeper insights into  
60 magma storage at active volcanoes.  
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64 **1. Introduction**

65 Axial Seamount is an active submarine volcano located at the intersection of the Juan de  
66 Fuca Ridge and the Cobb hotspot about 500 km west of the Oregon coast in the NE Pacific  
67 (Figure 1). It has erupted at least 52 times over the last 800 years (Clague et al., 2013), most  
68 recently in 1998, 2011, and 2015. A nearly continuous deformation time series from 1998  
69 through the present covering the past 3 eruptions has revealed that Axial exhibits a relatively  
70 repeatable inflation-deflation cycle, which has allowed for two successful eruption forecasts  
71 (Chadwick et al., 2012; Nooner & Chadwick, 2016). Even though Axial itself does not pose a  
72 direct threat to humans because of its remoteness, insight gleaned from observations made at  
73 Axial contribute to a growing body of knowledge about eruptive precursors that can be applied  
74 to more threatening locations.



75 **Figure 1.** a) Axial Seamount's tectonic setting at the intersection of the Juan de Fuca Ridge  
76 (JdFR) and the Cobb hotspot. b) Zoom-in of Axial's summit caldera with geodetic  
77 instrumentation as of 2020 labeled. White dots are benchmarks where campaign-style mobile  
78 pressure recorder (MPR) measurements are made, green dots are mini bottom pressure recorders  
79 (BPRs), blue dots are moored BPRs, and red dots are BPRs and tilt meters connected to the  
80 Ocean Observatories Initiative (OOI) cabled observatory.  
81  
82

83 Deformation models of Axial have evolved from simple to more complex over the years  
84 as more geodetic data have become available. A point source (Mogi, 1958) was initially used as  
85 the pressure source when few observations were available to constrain models and little was  
86 known about the actual geometry of Axial's magma storage system (Chadwick et al., 2006;  
87 Nooner & Chadwick, 2009). Once more benchmarks for pressure measurements were added and  
88 more analytical model geometries were considered, a steeply dipping prolate spheroid geometry  
89 became the best-fit model (Hefner et al., 2020; Nooner & Chadwick, 2016). The prolate spheroid  
90 model depth, location, and geometry were somewhat consistent with a set of vertically stacked

91 deep sills later imaged by multi-channel seismic (MCS) data and interpreted by Carbotte et al.,  
92 (2020). However, as autonomous underwater vehicle (AUV) repeat bathymetry data (Caress et  
93 al., 2020) has begun to provide more spatial coverage and therefore additional constraints for  
94 deformation modeling than the limited number of point-pressure observations alone, a  
95 rectangular horizontal sill deformation model with about the same outline as the summit caldera  
96 has been found to fit both the AUV and pressure data better than a prolate spheroid (Hefner et  
97 al., 2021).

98  
99 The acquisition of multi-channel seismic (MCS) data at Axial (Arnulf et al., 2014, 2018;  
100 Carbotte et al., 2020) provided a high-resolution view of the magma reservoir geometry beneath  
101 the summit of Axial for the first time. Given the simplicity of the previous analytical deformation  
102 models, a logical next step was to investigate how a more realistic geometry of the magmatic  
103 system relates to deformation observed at the surface, in order to add more physical meaning to  
104 the modeling results. Arnulf et al., 2018 used MCS data to define the 3-D geometry and location  
105 of the main magma reservoir (MMR) beneath the summit caldera at Axial, as well as a secondary  
106 magma reservoir (SMR) to the east. The MMR vertically extends from 1.1-2.8 km depth below  
107 seafloor, is slightly offset from Axial's caldera to the east, and extends beyond the caldera to the  
108 north and south (Figure 2). The deep stacked sills imaged by Carbotte et al., (2020) are located  
109 below the southern half of the MMR between 3-5 km below the seafloor.

110  
111 We constructed deformation models constrained by the MMR geometry in several ways.  
112 First, we directly used the 3D MMR geometry with uniform internal pressure in a finite element  
113 model (FEM), but we found that doing so provides very poor fit to the geodetic data. We then  
114 constructed and considered several analytical deformation models as alternatives, including: 1)  
115 approximating the MMR shape using one rectangular horizontal sill, 2) approximating the MMR  
116 shape using 3 rectangular non-horizontal sills, 3) allowing for non-uniform pressure distribution  
117 in a 2D horizontal sill at the average depth of the MMR roof, and 4) allowing for non-uniform  
118 pressure distributed over the 3D MMR roof. The models are constrained by the observations of  
119 vertical deformation from seafloor pressure data and repeated AUV bathymetric surveys during  
120 Axial's current inter-eruption phase between 2016-2020. Our inversion results suggest that the  
121 MMR is likely compartmentalized, which is consistent with current thinking on magma reservoir  
122 structure.

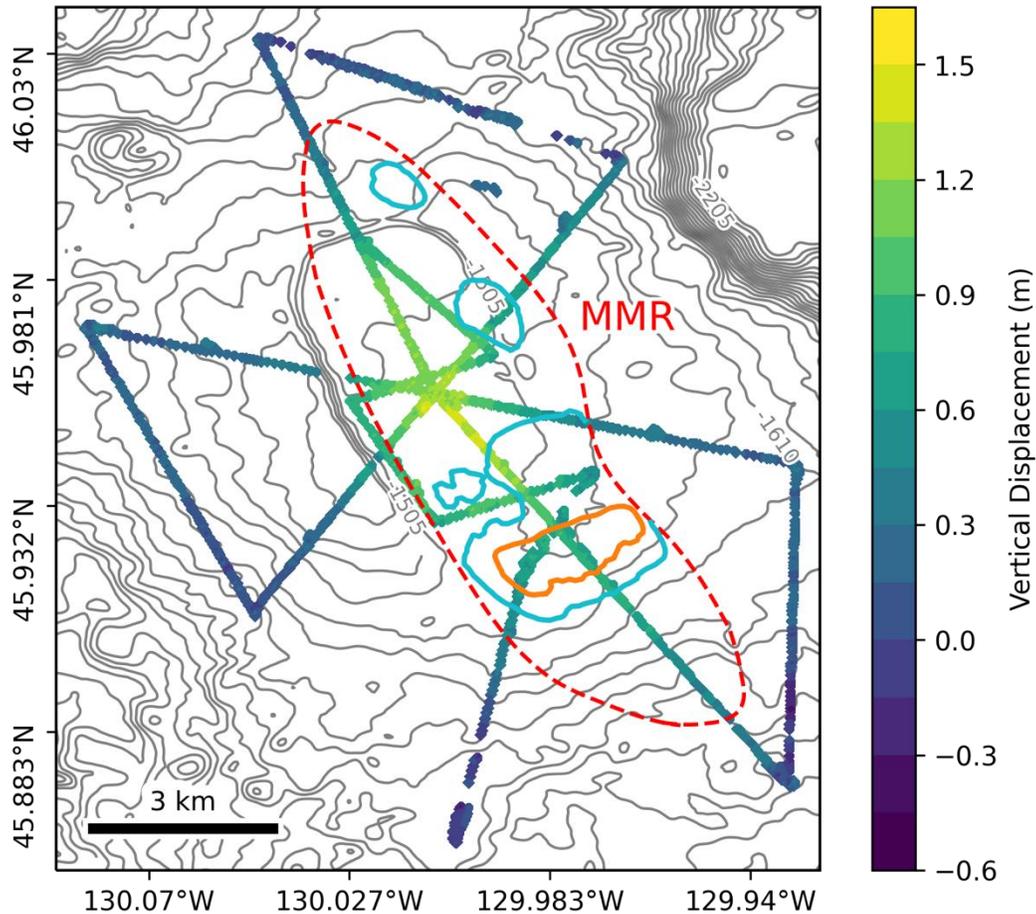
## 123 124 **2. Deformation data**

125 Bottom pressure recorders (BPRs) measure pressure at the seafloor; if the seafloor is  
126 uplifted, there is less water column above it and therefore lower pressure. Similarly, if the  
127 seafloor subsides, the BPR measures higher pressure. The pressure data are converted to depth  
128 after removing tidal signals (Eble et al., 1989). BPRs were deployed at Axial's summit caldera in  
129 1998 when Axial's first observed eruption occurred (Chadwick et al., 2013; Dziak & Fox, 1999;  
130 Embley et al., 1999; Fox, 1999; Fox et al., 2001). After a two-year gap in coverage, the

131 deformation time series resumed in 2000 with an array of seafloor benchmarks and the time  
132 series has been continuous through the present (Figure 1; Chadwick et al., 2006, 2012, 2022;  
133 Nooner & Chadwick, 2009, 2016). Since 2000, BPR measurements have been supplemented by  
134 measurements from mobile pressure recorders (MPRs), which are used in campaign-style  
135 surveys at seafloor benchmarks with a remotely operated vehicle (ROV) every 1-2 years to  
136 correct for the BPRs' long-term drift where the two are co-located (Chadwick et al., 2006). We  
137 used the MPR data for our study instead of BPR data because there were more MPR  
138 measurement locations in 2016-2020 and we are more interested in the spatial component of  
139 deformation than the temporal component.

140  
141 Bathymetric surveys at 1-m scale have been conducted at Axial since 2006 using  
142 multibeam sonar equipped AUVs, first to obtain comprehensive coverage of the volcanic terrain,  
143 and then to measure the extent and thickness of lava flows from the 2011 and 2015 eruptions  
144 through differencing of repeated surveys (Caress et al., 2012; Chadwick et al., 2016). Beginning  
145 after the most recent eruption in 2015, a new sparse pattern of AUV survey lines extending well  
146 outside the caldera (Figure 2) was established to measure vertical surface deformation by  
147 differencing (Caress et al., 2020); this pattern has been repeated each summer since except 2021.  
148 Differencing the repeated components of the surveys reveals vertical surface deformation over a  
149 broader area than from the pressure sensors alone. However, compared to the MPR data which  
150 has an accuracy of  $\pm 1$  cm, the AUV repeat bathymetry data have a lower vertical displacement  
151 accuracy of  $\pm 20$  cm. We used AUV vertical displacement data between two surveys in 2016 and  
152 2020 (Figure 2). An AUV bathymetric survey was also conducted in 2015, but this survey  
153 apparently had higher errors than subsequent surveys, because the AUV depth changes between  
154 2015-2020 poorly match the MPR depth changes from the same time period. Since MPR  
155 measurements were made in 2015 and 2017 (but not in 2016), we estimated the uplift values in  
156 2016 at the MPR benchmarks by interpolating between the 2015 and 2017 MPR measurements  
157 assuming a linear deformation rate. The BPR record shows that deformation at the center of the  
158 caldera during this time period was not entirely linear (Chadwick et al., 2022). The benchmark at  
159 the center of the caldera had uplifted by 55 cm from mid-2015 to mid-2016, about 10 cm  
160 shallower in summer of 2016 than a linear interpolation would predict (Figure S1 in  
161 Supplementary Material). The deformation rate is highest at this benchmark compared to the  
162 other benchmarks, so our interpolation introduces an additional uncertainty of  $\leq 10$  cm in the  
163 estimated 2016 benchmark depths. Nevertheless, the estimated 2016-2020 depth changes at the  
164 benchmarks agree relatively well with the 2016-2020 AUV data (Figure S2 in Supplementary  
165 Material).

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168  
 169 **Figure 2.** AUV repeat bathymetry data covering 2016-2020. Colors represent depth changes  
 170 between AUV surveys. Background bathymetry contour interval is 35 m and select contours are  
 171 labeled. The MMR is outlined with a red dotted line. The shallowest parts of the MMR roof are  
 172 shown with depth contours at -1250m and -1500m (below seafloor) in orange and cyan,  
 173 respectively.  
 174

175 We only used deformation data covering the inflation period from 2016 to 2020 to  
 176 constrain our models because the main objective of our study is to investigate the spatial  
 177 component of the deformation signal and its implications for the underlying magma storage  
 178 system. Previous studies have shown that the spatial pattern of inflation does not vary  
 179 significantly between different time periods (Nooner & Chadwick, 2016), except for major  
 180 episodes of deflation during eruptions when slip on the caldera ring faults may contribute to the  
 181 deformation field (Hefner et al., 2020).  
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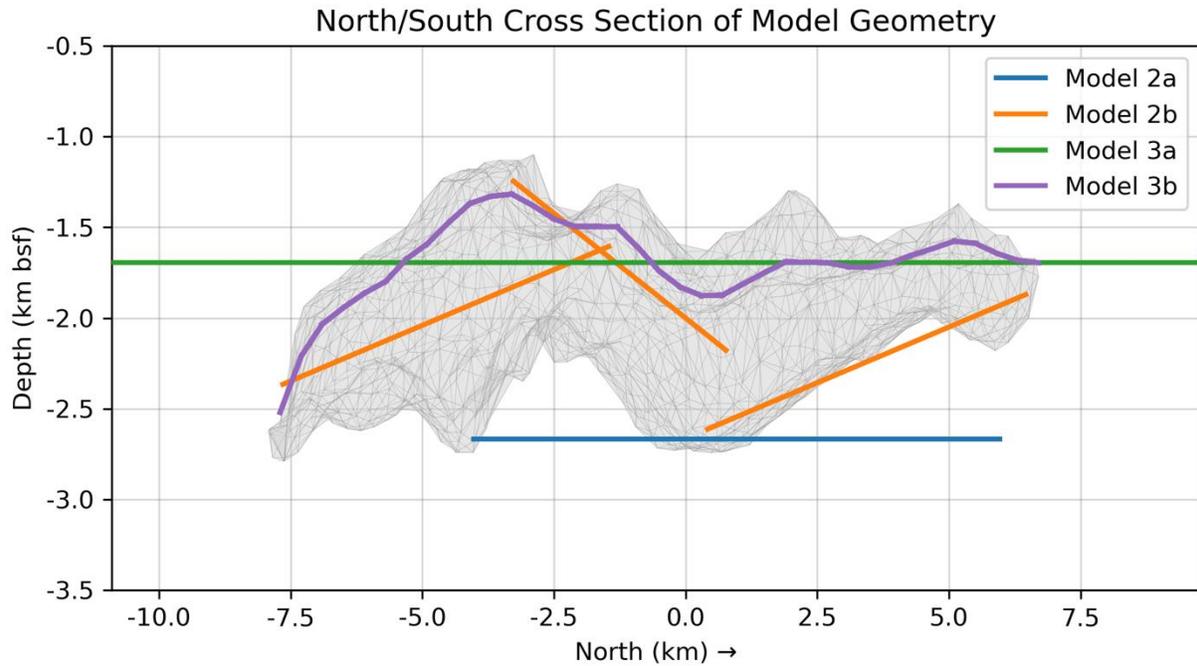
### 183 3. Deformation modeling

184 Our objective was to improve upon previous deformation models by reconciling the  
 185 MMR geometry with the observed spatial deformation pattern. To do this, we constructed a  
 186 series of models with increasing complexity, all constrained/bounded by the MMR. Each is

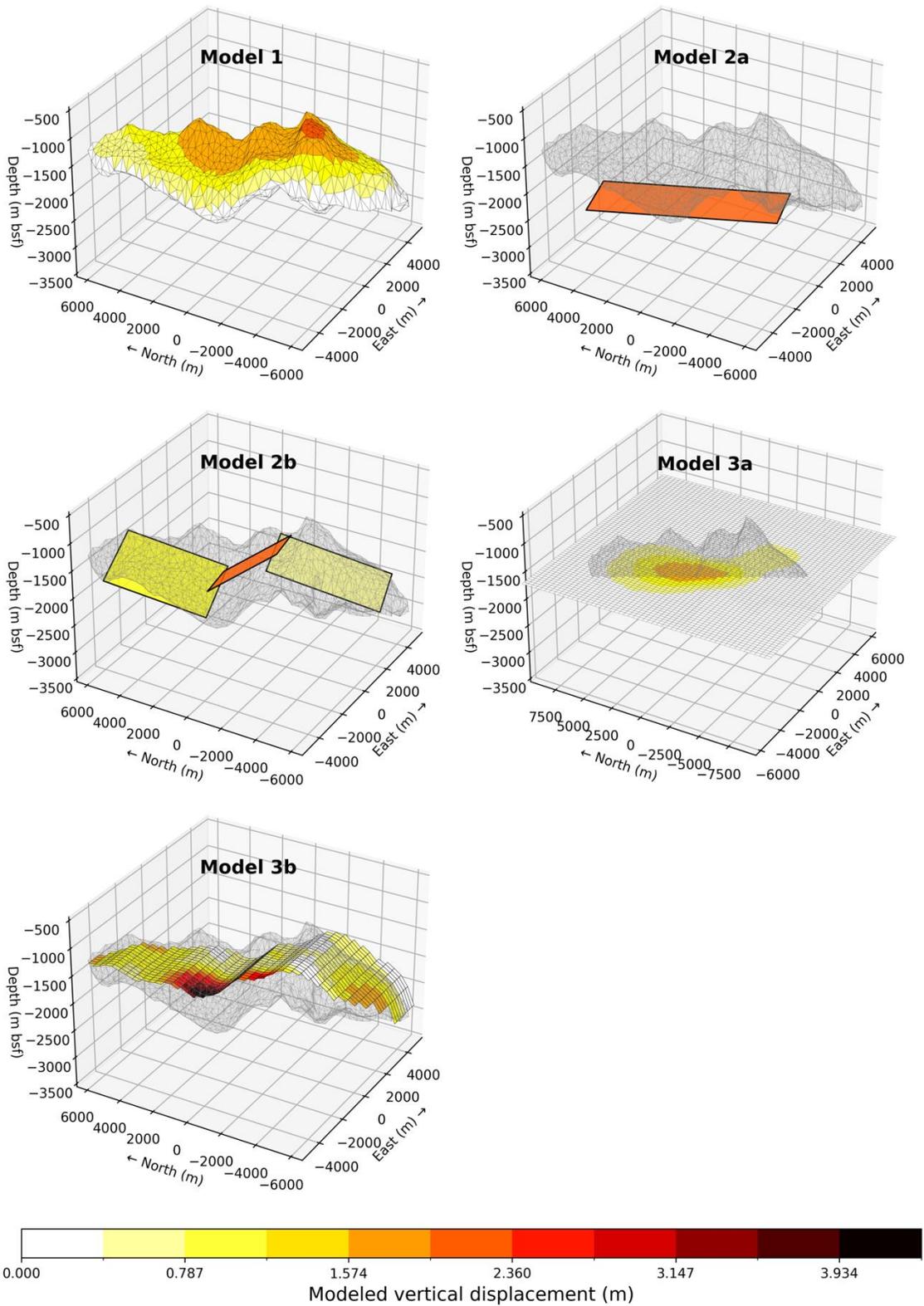
187 discussed in detail below. For all models, typical mechanical properties were used (Poisson's  
 188 ratio = 0.25, shear modulus = 30 GPa, Young's modulus = 70 GPa; Turcotte & Schubert, 2014).  
 189 Table 1 contains a summary of model configurations, inversion methods, and performance. See  
 190 Figures 3 and 4 for a comparison of model geometries.  
 191

	<b>Model configuration</b>	<b>Inversion method</b>	<b>Volume change (km<sup>3</sup>)</b>	<b>RMSE<sub>mpr</sub> (m)</b>	<b>RMSE<sub>auv</sub> (m)</b>
<b>Model 1</b>	FEM, MMR with uniform internal pressure	Parameter search	0.173	0.312	0.254
<b>Model 2a</b>	Analytical, 1 rectangular, horizontal sill	MCMC	0.056	0.059	0.122
<b>Model 2b</b>	Analytical, 3 rectangular, non-horizontal sills	MCMC	0.06	0.047	0.097
<b>Model 3a</b>	Analytical, 2D horizontal grid of Okada sill sources	Least squares regression	0.06	0.009	0.130
<b>Model 3b</b>	Analytical, 3D Okada sill sources draped over MMR roof	Least squares regression	0.054	0.002	0.139

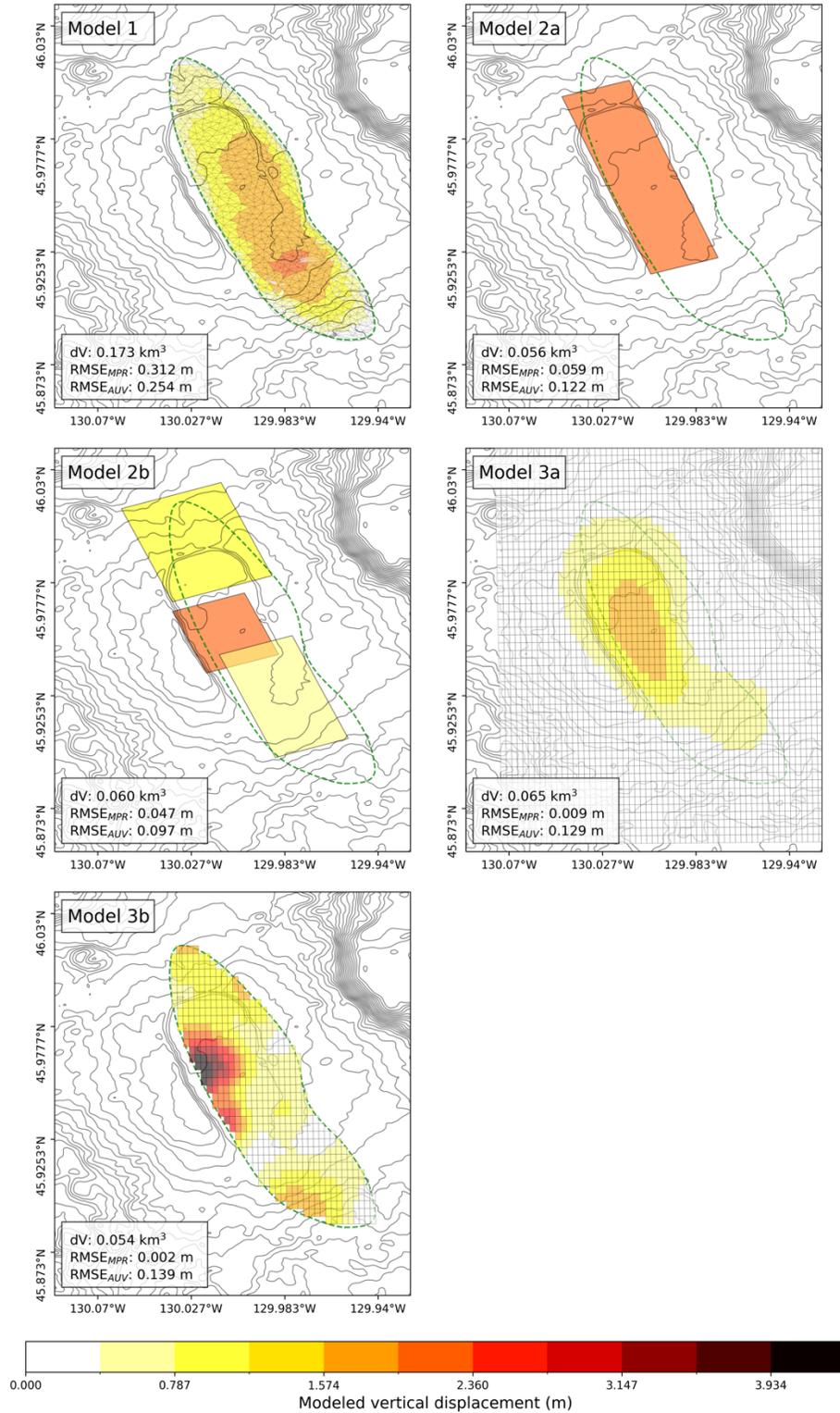
192  
 193 **Table 1.** Summary of model configurations, inversion methods, modeled volume changes, and  
 194 Root Mean Square Error (RMSE) values between each model and the MPR and AUV data.



195  
 196 **Figure 3.** North/South cross section showing deformation model geometries investigated in this  
 197 study. The MMR geometry from Arnulf et al., (2018) is shown as a gray mesh and represents the  
 198 Model 1 FEM source geometry. The other models are single or multiple combination of  
 199 rectangular Okada sills (colored lines) with either uniform or distributed (non-uniform) opening.  
 200 See text for details.



201  
 202 **Figure 4.** Model configurations and modeled openings of each pressure source. For Models 2a-  
 203 3b, the MMR is shown as a transparent gray mesh to provide context for the model geometries.



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**Figure 5.** Model configurations and modeled openings of each pressure source overlain on bathymetric contours (each contour is 35 m). The MMR is outlined in each plot with a dotted green line. Each model's volume change (dV) and RMSE values between the model and the MPR and AUV data are plotted in the lower left corner of each panel.

209 **Model 1a: Finite element model with MMR geometry and uniform pressure**

210 As a first step, we constructed an FEM using the MMR geometry from Arnulf et al., 2018  
211 with a uniform pressure source. We started with a 3D point cloud defining the combined MMR  
212 roof and floor (see Arnulf et al., 2018 for more detail on how the roof and floor boundaries were  
213 defined). A 3D surface was constructed from the point cloud using a ball-pivoting algorithm,  
214 which starts with a seed triangle and creates new triangles by pivoting a ball with user-defined  
215 radius around the edges until it meets new points (Bernardini et al., 1999). This 3D surface was  
216 then loaded into Abaqus/CAE 2020, which we used to carry out the FEM simulations. To  
217 validate the FEM methodology, we compared an analytical prolate spheroid model (Yang et al.,  
218 1988) to an FEM with a pressurized cavity of the same dimensions and verified that both models  
219 predict the same surface deformation (Figure S3 in Supplementary Material).

220  
221 The FEM domain measures 50 km long x 50 km wide x 30 km deep and the boundary  
222 conditions were specified by a free top surface, a roller constraint on the side surfaces, and a  
223 fixed bottom surface. We added bathymetry to the model using GMRT bathymetry data (Ryan et  
224 al., 2009). The effect of gravity was accounted for by adding an additional analysis step (prior to  
225 pressurization of the source) in which gravitational equilibrium is established by adding a pre-  
226 stress defined by hydrostatic equilibrium. This is an ‘initial guess’ which is used as a starting  
227 point to solve for the gravitational force that balances out the pressure force to result in near-zero  
228 ground deformation according to a defined threshold. We tested the effect of ocean loading by  
229 adding a downward hydrostatic pressure applied to the seafloor and found it to be negligible.

230  
231 The MMR was incorporated by subtracting its volume from the domain and applying a  
232 uniform internal pressure on the cavity walls. The pressure was varied over many simulations to  
233 minimize the combined root-mean-squared error (RMSE) between the modeled surface  
234 displacements and the AUV and MPR data.

235

236 **3.1. Models 2a and 2b: Analytical sill models using Bayesian inference**

237 Model 2a is a single rectangular horizontal sill (Okada, 1985) and Model 2b consists of 3  
238 non-horizontal rectangular sills constrained by the MMR geometry. We used the Volcanic and  
239 Seismic Source Modeling (VSM) package (Trasatti, 2022) to conduct joint inversions using  
240 Markov chain Monte Carlo (MCMC) simulations to estimate the source parameters that produce  
241 surface deformation that best fits the AUV and MPR data.

242

243 For Model 2a, all inversion parameters were allowed to vary except for the dip angle of  
244 the sill, which was fixed at zero (horizontal). The sill’s depth was bounded by the minimum and  
245 maximum MMR depth. For Model 2b, the 3-sill geometry was constrained by the MMR  
246 geometry by fixing the strike and dip angles in the inversion to follow the general trend of 3  
247 main MMR segments (Figures 4 and 5). The locations of the sills were allowed to vary within 3

248 defined segments of the MMR volume and the sill opening values were allowed to vary freely.  
 249 See Table 2 for a summary of fixed and best-fit variable parameters for Models 2a and 2b.  
 250

		Centroid Longitude	Centroid Latitude	Centroid depth (m bsf)	Length (m)	Width (m)	Strike	Dip	Opening (m)
<b>Model 2a</b>		-130.0100 ± 258 m	45.9637 ± 110 m	2666 ± 306	2561 ± 1119	9680 ± 253	341° ± 2	0	2.256 ± 0.682
<b>Model 2b</b>	<b>Sill 1</b>	-130.0249 ± 79 m	45.9968 ± 77 m	2241 ± 74	3829 ± 170	5170 ± 115	340°	-7°	1.285 ± 0.041
	<b>Sill 2</b>	-130.0110 ± 83 m	45.9543 ± 97 m	1712 ± 70	2769 ± 166	3475 ± 164	340°	13°	2.071 ± 0.152
	<b>Sill 3</b>	-129.9850 ± 126 m	45.9265 ± 175 m	1985 ± 206	2800 ± 207	5707.36 ± 325	340°	-7°	0.893 ± 0.076

251  
 252 **Table 2.** Summary of fixed and best-fit inverted parameters with standard deviations for Models  
 253 2a and 2b. The strike angle is the orientation of the plane measured clockwise from North  
 254 according to Okada (1985) (i.e., strike = 0 if the plane is oriented North-South and dips to the  
 255 East, strike = 90 if the plane is oriented East-West and dips to the South). Fixed parameters have  
 256 red shading, parameters allowed to vary within the confines of the MMR geometry have yellow  
 257 shading, and parameters allowed to freely vary have green shading.  
 258

### 259 3.2. Models 3a and 3b: 2D and 3D distributed pressure inversions

260 Inverting geodetic data to determine variable slip or opening distribution is a standard  
 261 method for inferring co-seismic slip on faults (e.g., Moreno et al., 2009) and has also been  
 262 applied in volcanic settings (e.g., Grandin et al., 2009). We performed two joint inversions of the  
 263 MPR and AUV data following this approach. For Model 3a, we created a 2D horizontal grid of  
 264 rectangular sill-patches at the average depth of the MMR roof and extending beyond the MMR  
 265 boundary horizontally by 3 km in both the x and y directions. For Model 3b, we gridded the  
 266 MMR roof point cloud into rectangular patches where each patch is defined by its position,  
 267 length, width, strike, and dip. The patches are allowed to dip to the North/South but not  
 268 East/West to create a continuous 3D grid with no gaps; this is appropriate since there is much  
 269 more dip variation along the North/South direction of the MMR than there is along the East/West  
 270 direction. The depths of the patches were defined by the average MMR roof depth at that  
 271 location (Figures 3 and 4).  
 272

273 For both Models 3a and 3b, we treated each patch as a rectangular dislocation (Okada,  
 274 1985) and inverted for the opening value of each patch. Posed as a forward problem, the  
 275 relationship between surface displacements and patch openings can be expressed by the linear  
 276 system:  
 277

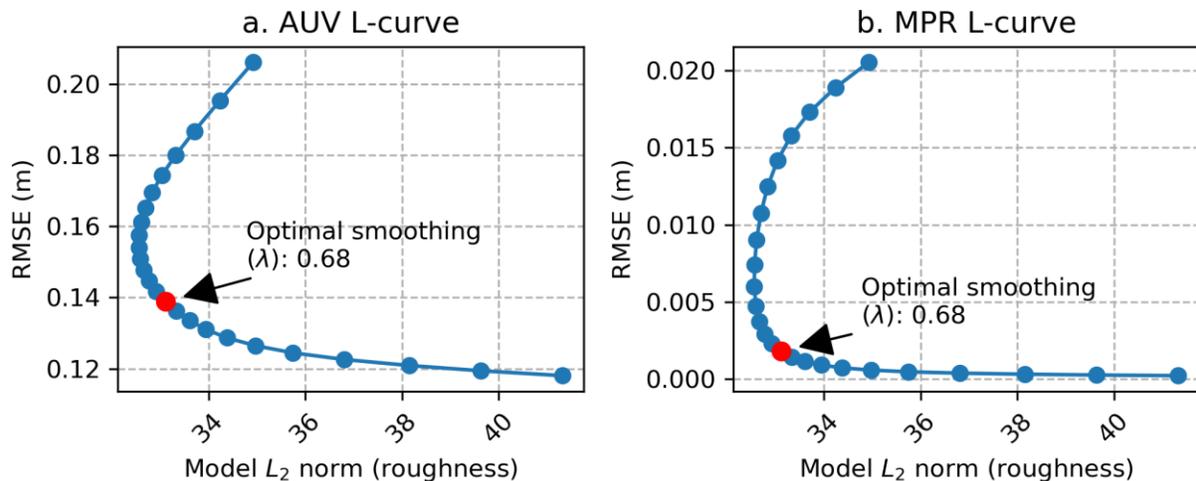
$$d = Gm$$

278

279 where  $d$  is the observation vector composed of vertical surface displacements,  $G$  is the Green's  
 280 function matrix, and  $m$  is the vector of model parameters (patch openings).  $G$  was constructed by  
 281 computing the expected vertical displacement at every observation point for each patch caused  
 282 by a unit opening on that patch. To solve for  $m$ , we used a regularized linear least squares  
 283 method which minimizes the objective function,  $\phi(m)$ :

$$\phi(m) = \|W(G \cdot m - d)\|_2^2 + \lambda^2 \|L \cdot m\|_2^2$$

285  
 286 The first term  $\|W(G \cdot m - d)\|_2^2$  represents weighted misfit, i.e., the squared Euclidean norm  
 287 difference between the observed data and the data predicted by the model, where  $W$  is a diagonal  
 288 weight matrix which normalizes the contribution of the MPR and AUV datasets based on the  
 289 relative uncertainties and the number of relative data points. The second term  $\lambda^2 \|L \cdot m\|_2^2$  is the  
 290 regularization term, where  $\lambda$  is the regularization parameter that controls the smoothness of the  
 291 model, and  $L$  is the regularization matrix. The optimal  $\lambda$  value was chosen using an L-curve,  
 292 where the preferred smoothness is located at the corner of the curve created by plotting  
 293 roughness vs. the L2 norm of misfit (Figure 6).



294  
 295 **Figure 6.** L-curves showing model roughness vs. Root Mean Square Error (RMSE) as a measure  
 296 of misfit between the model and the data for (a) the AUV data and (b) the MPR data. The  
 297 optimal smoothness occurs at the corner of the curve. Example shown is for Model 3b.  
 298

### 299 3.3. Weighing the AUV and MPR data

300 We weighed the AUV and MPR data on a case-by-case basis for each model due to  
 301 differences among inversion methods. For Model 1, since the best-fit model was found by a  
 302 parameter search over uniform pressure values on the MMR surface (all other model parameters  
 303 were fixed), we calculated the AUV RMSE and MPR RMSE for each iteration then normalized  
 304 them by dividing each by the maximum RMSE value across all iterations and by the relative  
 305 uncertainties in the datasets. We then calculated the combined RMSE for each iteration by  
 306 summing the normalized AUV RMSE and MPR RMSE values. The optimal model was chosen

307 as the model with the lowest combined RMSE value. For Models 2a and 2b, we first weighed the  
308 datasets in an MCMC simulation according to their relative uncertainties, then further adjusted  
309 the weights over many MCMC simulations to find the weight combination that minimized the  
310 combined AUV and MPR RMSE values.

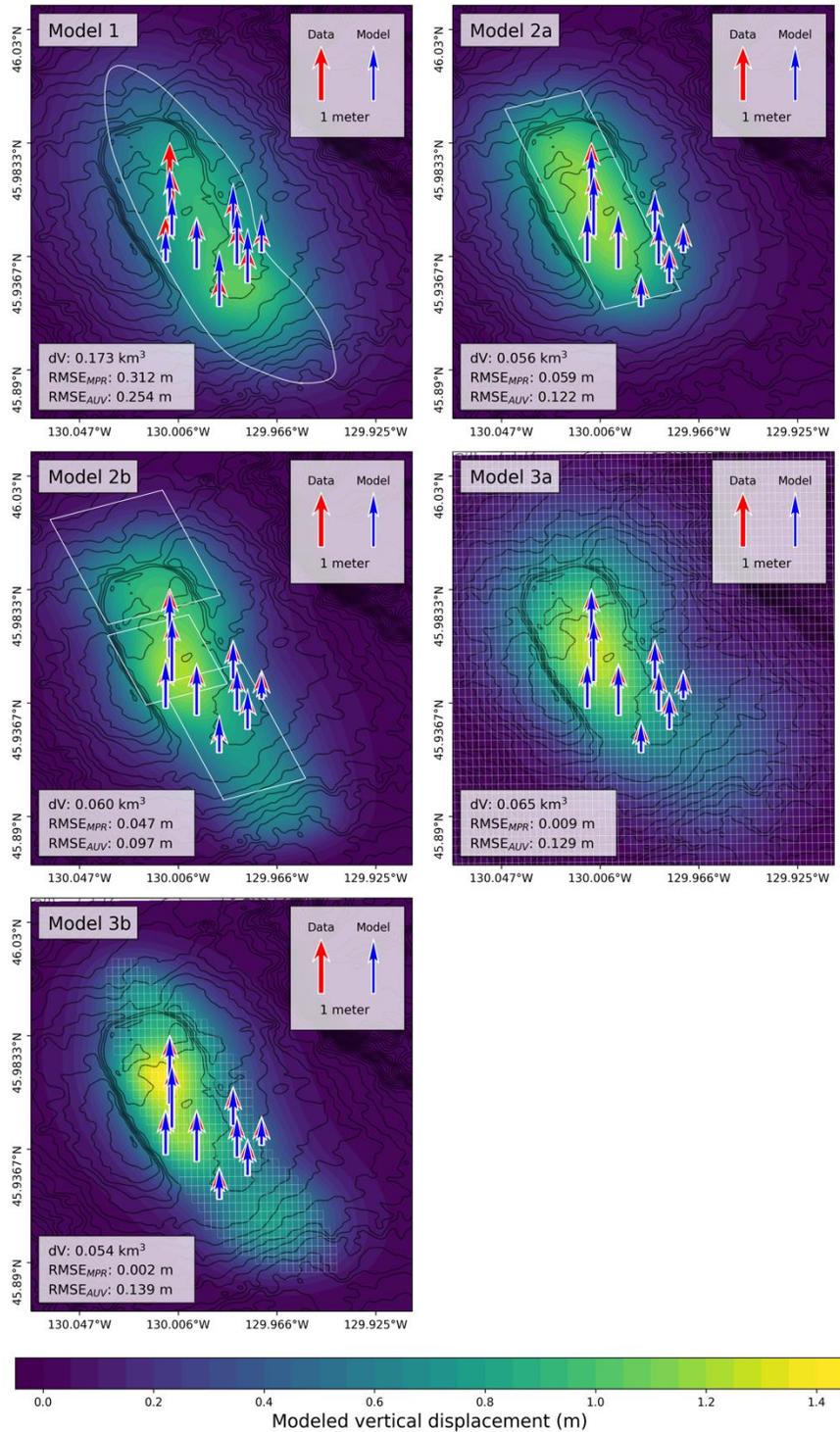
311  
312 For Models 3a and 3b, we found a tradeoff between the regularization parameter  $\lambda$  and  
313 the relative weights, due to higher noise in the AUV data than in the MPR data. Instead of just  
314 normalizing the AUV and MPR datasets using their relative uncertainties, we further normalized  
315 them by the number of data points in each dataset. The  $\lambda$  value was then chosen as described  
316 above in Section 3.3.

#### 317 318 **4. Results**

319 We found that Model 1 (uniform pressurization of the 3-dimensional MMR) did not fit  
320 either the MPR or AUV data well. This was not unexpected, since the MMR geometry is offset  
321 from the caldera to the east while the observed deformation is centered on the caldera. Also, the  
322 shallowest features along the MMR roof are located beneath the SE part of the caldera and  
323 because of this, the model creates the largest surface deformation there, 4-5 km SE of the caldera  
324 center (Figures 2, 4a and 6a). This makes sense intuitively since these shallowest MMR features  
325 have less overburden and therefore uplift more readily under uniform pressurization. This result  
326 tells us that the observed deformation cannot be simply produced by uniform pressure within the  
327 entire MMR, which suggests that perhaps the MMR is compartmentalized with isolated melt  
328 pockets that are not well connected. The other four models, which were developed to test this  
329 idea, showed increasing improvement of fit to the MPR data as more parameters were added, and  
330 the AUV RMSE values were also improved, but not as much and varied from model to model  
331 (Table 1). We suspect that this is because of the higher uncertainty associated with the AUV  
332 data, which was factored into how the datasets were weighed.

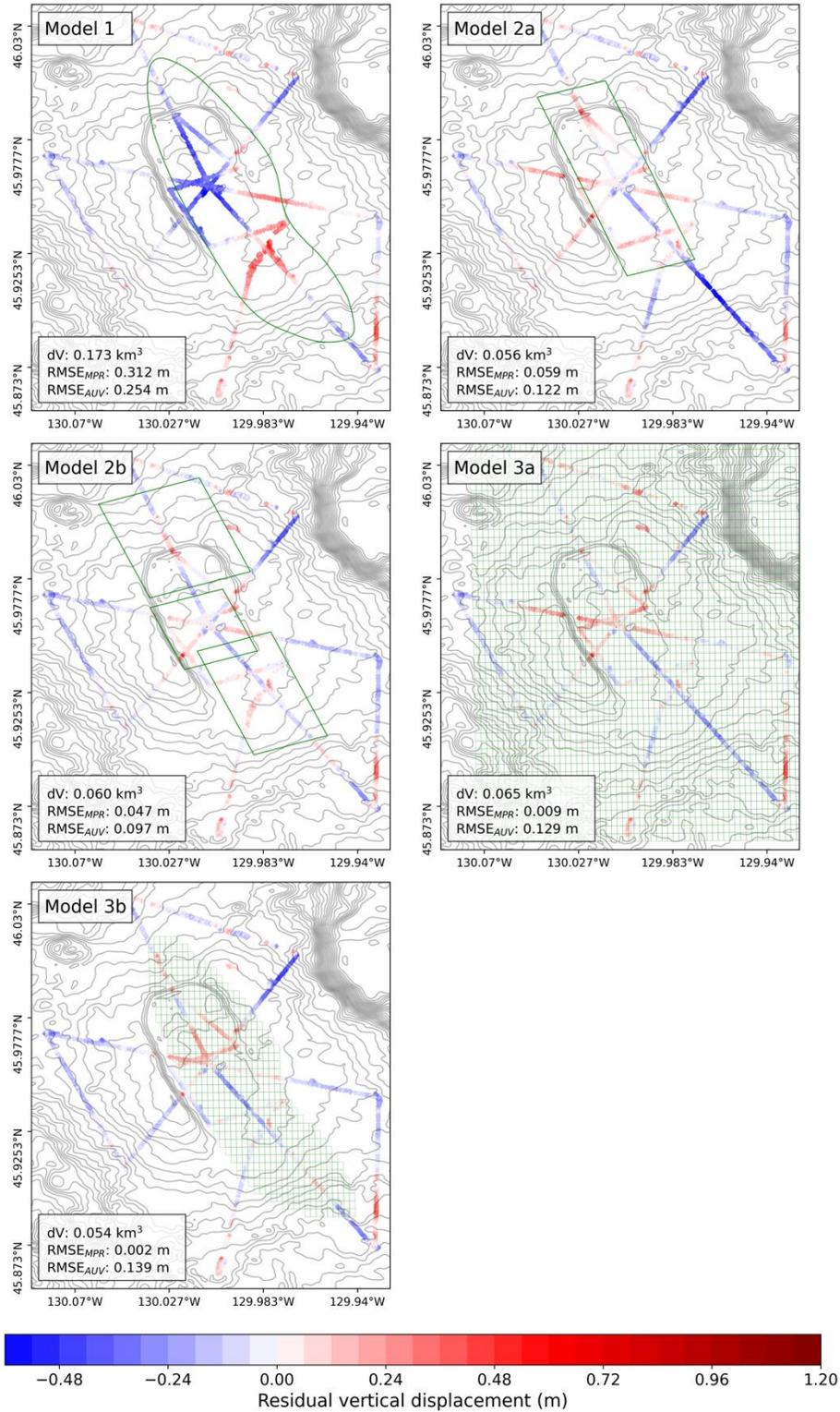
333  
334 Despite differences among model geometries, the models consistently estimated a best-fit  
335 volume change of between 0.054-0.060 km<sup>3</sup>, except for Model 1 which estimated 0.173 km<sup>3</sup>.  
336 Modeled deformation and fit to the MPR data are plotted in Figure 7 and AUV repeat  
337 bathymetry residuals are plotted in Figure 8. In Models 3a and 3b where pressure was allowed to  
338 spatially vary, modeled pressure changes were highest along the western-central edge of the  
339 MMR (Figures 4 and 5). There is also a region of positive pressure change in the southern-most  
340 southward dipping region of the MMR due to a long wavelength deformation signal present in  
341 this area in the AUV data.

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**Figure 7.** Predicted surface vertical deformation for all best-fit models with comparison between the MPR data (red arrows) and modeled surface displacements (blue arrows). The surface projection of each model geometry is shown as a white outline. Each model's volume change (dV) and rmse values between the model and the MPR and AUV data are plotted in the lower left corner of each panel.



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**Figure 8.** AUV repeat bathymetry residuals plotted by subtracting the modeled displacements from the AUV data. The surface projection of each model geometry is plotted as a green outline. Each model's volume change (dV) and rmse values between the model and the MPR and AUV data are plotted in the lower left corner of each panel.

## 354 5. Discussion

### 355 5.1. Model assumptions and limitations

356 All our models assume homogeneous and isotropic elastic half spaces (except for Model  
357 1, which includes bathymetry). Masterlark (2007) showed that the presence of layered crustal  
358 material can increase source depth estimates when compared to models assuming elastic half  
359 spaces with uniform properties. Since Axial's volcanic edifice is composed of lava flows  
360 emplaced upon one another over time, there is likely some anisotropy in which stiffness is  
361 different in the vertical and lateral directions, which could cause an underestimation of source  
362 depths. In addition, we found in sensitivity testing that inclusion of Axial's bathymetry in a finite  
363 element model using a prolate spheroid pressure source fixed at a depth of 3.8 km (the best-fit  
364 model of Nooner & Chadwick, 2016) can affect the volume change estimate by up to 27%  
365 (Figure S3 in Supplementary Material). This effect would increase with shallower source depths  
366 (Williams & Wadge, 1998) such as at the depth of the MMR. This result was unexpected  
367 because of Axial's relatively modest bathymetric relief, so more work must be done to better  
368 understand which bathymetric features (e.g., caldera walls vs surrounding bathymetric features)  
369 influence the expression of vertical deformation for a given pressure source geometry.

370  
371 Our assumption of elasticity could also affect the modeling results since there may be  
372 non-elastic or viscoelastic effects unaccounted for in the models. Numerical modeling  
373 implementing viscoelasticity at Mt. Etna has shown that lower pressures can produce the same  
374 deformation as elastic models with higher pressure due to viscoelastic relaxation over time (Del  
375 Negro et al., 2009). Depending on where this region of viscoelasticity is defined (either above or  
376 below the pressure source), this phenomenon could result in either inflation or deflation observed  
377 on the surface (Nooner & Chadwick, 2009). Cabaniss et al., (2020) found that non-temperature-  
378 dependent elastic rheology requires greater reservoir overpressures to reproduce the observed  
379 surface deformation at Axial compared to models that incorporate a temperature-dependent  
380 rheology. Additionally, petrological and tomographic studies increasingly show that magma  
381 reservoirs are likely composed of discrete melt lenses/sills embedded within a crystal-rich  
382 magma mush (Cashman et al., 2017). Magma mush is expected to behave  
383 poroelastically/poroviscoelastically (Gudmundsson, 2012; Liao et al., 2018, 2021). Although  
384 viscoelastic effects and the presence of magma mush would likely not significantly impact the  
385 spatial distribution of modeled pressure changes in our results, it could impact volume change  
386 estimates due to magma compressibility. Modeling viscoelastic effects at Axial would be more  
387 strongly relevant to the temporal component of the deformation time series, for example to test  
388 hypotheses regarding short-term deflation events proposed by Chadwick et al., (2022).

389  
390 We also assumed that Axial's spatial deformation pattern does not vary temporally,  
391 except for during eruptions when there may be slip on the caldera ring faults (Hefner et al., 2020;  
392 Levy et al., 2018). The degree to which these ring faults are activated during Axial's inter-  
393 eruptive periods is unknown. Although there is little evidence of slip on these faults during the

394 2016-2020 inflation in the AUV repeat bathymetry data (in the form of a sharp offsets along  
395 AUV track lines crossing the faults), the uncertainty in the AUV data ( $\pm 20$  cm) may be equal to  
396 or higher than the amount of expected slip (8-30 cm; Levy et al., 2018). An FEM that includes  
397 bathymetry, spatially variable pressure, and inclusion of ring faults would be most thorough,  
398 although the number of free parameters in a complex model like this may not be constrainable by  
399 the current deformation data. Recent expansions of the geodetic monitoring network at Axial  
400 will be able to better quantify any slip across the caldera faults in the future.

401

## 402 **5.2. Implications for magma storage beneath Axial caldera**

403 A best-fitting deformation model cannot reveal the exact geometry of a magma storage  
404 system and should not be interpreted as such; rather, a deformation model can provide the  
405 approximate location and volume changes of the region(s) where the greatest pressure changes  
406 occurred during inter-eruption magma supply into the storage system. While our results provide  
407 improved horizontal constraints on where magma accumulates between eruptions, there is  
408 inherent non-uniqueness among modeled depths due to the tradeoff between depth and pressure.  
409 While our best-fit horizontal sill (Model 2a) is similar in horizontal geometry to the best-fit  
410 horizontal sill found by Hefner et al. (2021), the depth of our sill is deeper at 2.7 km compared to  
411 0.97-1.24 km, which is likely due to differences in inversion methods and/or the tradeoff  
412 between depth and pressure.

413

414 Despite this tradeoff, the consistent volume change estimates of 0.054-0.060 km<sup>3</sup> among  
415 Models 2a-3b suggests that the volume change is not significantly sensitive to model depths  
416 within the depth range of the MMR. In addition, we tested an FEM model using the prolate  
417 spheroid geometry (best-fit solution from Nooner & Chadwick, 2016) constrained only by the  
418 2015-2020 MPR data, which resulted in a volume change of 0.077 km<sup>3</sup> (Figure S3a in  
419 Supplementary Material). Since this included an extra year's worth of inflation compared to the  
420 2016-2020 models we show in this study, the estimated volume change for the 2016-2020 time  
421 period would be somewhat lower, more or less consistent with the volume change estimates  
422 using geometries constrained by the MMR above. This demonstrates that the estimated volume  
423 change for this inter-eruptive recharge period is not highly sensitive to model geometry, depth, or  
424 location.

425

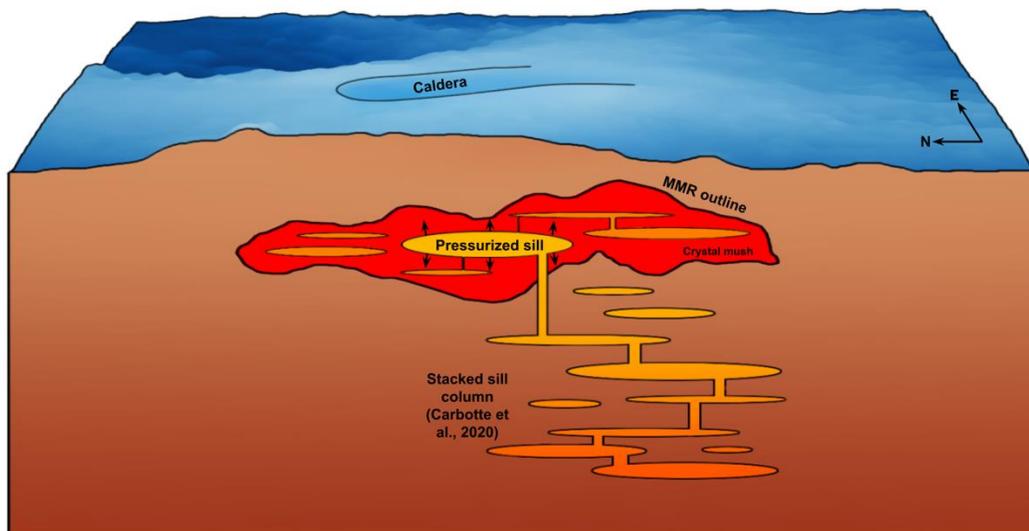
426 The total volume of the shallow magma storage system beneath Axial was estimated by  
427 Arnulf et al. (2014) to be 18-30 km<sup>3</sup> and the modeled co-eruptive volume change associated with  
428 previous eruptions has been estimated to vary between 0.147 – 0.206 km<sup>3</sup> using analytical model  
429 source depths of 3-3.8 km (Chadwick et al., 1999, 2012; Hefner et al., 2020; Nooner &  
430 Chadwick, 2016). Our study models the observed inflation from 2016-2020, during a time when  
431 the magma supply rate was initially high, but then waned with time following the 2015 eruption  
432 (Chadwick et al., 2022). Given that the magma supply rate is estimated to have varied from >0.1

433 km<sup>3</sup>/year to <0.01 km<sup>3</sup>/year during that time period (Chadwick et al., 2022), our volume change  
434 estimates are reasonable.

435

436 Mullet & Segall (2022) demonstrated that as the melt fraction of mush in a magma  
437 reservoir increases, the deformation caused by a mush-dominated magma storage system is  
438 increasingly driven by the overall shape of the mush body, instead of any pressurized melt lens  
439 within the mush. If the melt fraction within the MMR is high enough to cause Axial's  
440 deformation to be driven by the entire mushy body (instead of individual sills) and if we assume  
441 that the MMR is a continuous body, it follows that using the MMR geometry as a pressure  
442 source should fit the deformation data. The poor fit to the data of Model 1 as well as the pattern  
443 of pressure distribution in Models 3a and 3b are instead suggestive of compartmentalization of  
444 melt within the MMR and a relatively low melt fraction in the surrounding mush (Figure 9). In  
445 this context, compartmentalization means that melt bodies within the MMR are not connected  
446 hydraulically, at least on time scales that are relevant to the deformation cycle at Axial.

447



448

449 **Figure 9.** Schematic diagram illustrating possible compartmentalized melt distribution in which  
450 sills are emplaced in crystal mush both within and below the MMR.

451

452 The depth of magma residence estimated by petrological analyses (Dreyer et al., 2013) is  
453 deeper at 3-6 km than the MMR depth range of 1.1-2.8 km, but is consistent with the deeper  
454 system of stacked sills beneath the MMR imaged by Carbotte et al., 2020 extending from 3-5 km  
455 depth below seafloor. Since we did not consider deformation sources in this depth range, we  
456 cannot rule out contribution to the deformation field of a potential pressure source (or multiple  
457 sources) in the stacked sill region. Non-uniqueness among models due to the tradeoff between  
458 depth/pressure would likely hinder efforts to resolve pressurization of multiple vertically stacked  
459 sills or the combination of compartmentalized MMR pressurization with a source representing  
460 the stacked sill region. However, since the stacked sills are exclusively beneath the SSE part of

461 the caldera, they probably cannot produce the observed caldera-centered deformation by  
462 themselves.

463

## 464 **6. Conclusions**

465 The ability to accurately forecast volcanic eruptions is an important goal in hazard  
466 mitigation research. Linking precursory signals like ground deformation to subsurface processes  
467 is therefore essential. With the increase in spatial coverage of Axial's deformation monitoring  
468 due to the application of AUV repeat bathymetric surveys, there is now adequate data to justify  
469 more complex deformation modeling than what has been done previously. We constructed a  
470 suite of numerical and analytical models geometrically constrained by the shape of the  
471 seismically imaged MMR to investigate the role of the MMR in creating the observed surface  
472 deformation and to test the hypothesis that the MMR is compartmentalized. Although our  
473 estimated volume change of 0.054-0.060 km<sup>3</sup> for the inflation period between 2016-2020 is  
474 reasonable considering previous estimates of inflation and eruption volumes, the models make  
475 assumptions (flat seafloor, full elasticity, no ring faults) that could influence the volume change  
476 and/or depth estimates. Nevertheless, the models with spatially varying pressure (Models 3a and  
477 3b) suggest that magma accumulates during Axial's inter-eruptive recharge periods along the  
478 western-central edge of the MMR with some potential additional accumulation in the southern-  
479 most southward dipping region of the MMR. Future modeling efforts with additional complexity  
480 and more parameters will likely require increased data constraints in the form of higher  
481 resolution seismic imagery and/or horizontal deformation measurements.

482

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489 Facility) for conducting the MPR surveys, and the operational teams of AUV Sentry (WHOI  
490 National Deep Submergence Facility) and the MBARI Mapping AUVs for conducting the AUV  
491 surveys.

492

## 493 **Data Availability Statement**

494 The code and data used for this research can be found at <https://zenodo.org/records/10219997>.  
495 Academic licensing for Abaqus software is provided by Simulia, Dassault Systèmes. The VSM  
496 software used for analytical modeling can be found at <https://github.com/EliTras/VSM> (Trasatti,  
497 2022).

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