

# Statistics of bubble plumes generated by breaking surface waves

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## Key Points:

- Bubble plumes generated during ocean surface wave breaking are observed with echosounders on drifting buoys.
- Bubble plume depths are well correlated with whitecap coverage, wind speed, and spectral wave steepness.
- Bubble plumes persist for many wave periods and exceed the persistence of visible surface foam.

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## Abstract

We examine the dependence of the penetration depth and fractional surface area (e.g., whitecap coverage) of bubble plumes generated by breaking surface waves on various wind and wave parameters over a wide range of sea state conditions in the North Pacific Ocean, including storms with sustained winds up to  $22 \text{ ms}^{-1}$  and significant wave heights up to 10 m. Observations include arrays of freely drifting SWIFT buoys together with shipboard wind and optical video systems, which enabled concurrent high-resolution measurements of wind, waves, bubble plumes, and turbulence. We estimate bubble plume penetration depth from echograms that extend to more than 30 m depth in a surface-following reference frame collected by downward-looking echosounders integrated onboard the buoys. Our observations indicate that the mean and maximum bubble plume penetration depths exceed 10 m and 30 m beneath the surface at high winds, respectively, with a plume residence time of many wave periods. Bubble plume depths are well correlated with wind speeds, spectral wave steepness, and whitecap coverage. Plume depths scaled by total significant wave height are strongly linearly correlated with the inverse of wave age. Plume depths scaled by either wind sea or total significant wave height vary non-monotonically with increasing wind speeds. Dependencies of the combined observations on various non-dimensional predictors used for whitecap coverage estimation are also explored. This study provides first field evidence of a direct relation between bubble plume penetration depth and whitecap coverage, suggesting that the volume of bubble plumes could be estimated by remote sensing.

## Plain Language Summary

Quantifying the statistics of bubble plumes generated during ocean surface wave breaking is essential to understand the exchange between the ocean and the atmosphere. Bubble plumes also cause important variations in underwater acoustics and optics. Recent studies also suggest that the statistics of bubble plumes are skillful predictors for total energy loss during wave breaking, which is an essential quantity for accurate wave forecasting. Here we examine the dependence of bubble plume statistics on various wind and wave parameters over a wide range of sea state conditions, including storms. Echosounders integrated onboard drifting buoys are used to detect bubbles and estimate their penetration depth below the ocean surface. Visible surface area of these bubble plumes is also observed using shipboard optical video systems. We successfully provide multiple empirical relationships that predict the observed variability of the penetration depth and surface area of bubble plumes as a function of simple wind and wave statistics (available from existing forecast models or typical ocean buoys). Our results indicate that the penetration depth of bubble plumes is correlated with their visible surface area, suggesting that the volume of bubble plumes could be estimated by observing the ocean surface from above.

## 1 Introduction

Air-entraining breaking surface waves play a significant role in air-sea exchanges of mass, heat, energy, and momentum [Melville, 1996; Sullivan and McWilliams, 2010; Deike, 2022], and are also crucial in various technical applications, such as the design of marine structures and underwater communications. Breaking waves inject a relatively large volume of air into the water column as bubbles which then form intermittent bubble clouds at a wide range of spatial scales, hereafter referred to as bubble plumes. The entrained bubbles change optical properties of the water column [Terrill et al., 2001; Al-Lashi et al., 2016] and generate acoustic noise [Felizardo and Melville, 1995; Manasseh et al., 2006], especially during the active breaking period.

Quantifying the statistics of these bubble plumes (e.g., void fractions, size distributions, penetration depth, surface area, and volume of bubble plumes averaged over many waves) is essential to obtain robust parameterizations of fluxes at the ocean-atmosphere interface and variations in underwater acoustics and optics. Recent studies, including the present observations, also show that the statistics of bubble plume that represent the overall size of bubble plumes are strongly correlated with total wave breaking dissipation [Schwendeman and Thomson, 2015a; Callaghan et al., 2016; Callaghan, 2018; Derakhti et al., 2020a]. This suggests that such bubble plume statis-

66 tics are skillful predictors for the corresponding energy and momentum exchange between the  
67 ocean and atmosphere, especially in high sea states.

68 The statistics that represent the overall size of bubble plumes for a given sea state may be  
69 defined, in a wave-averaged sense, as the long-time (several minutes) average of the surface area  
70 and the penetration depth of individual bubble clouds. The former may be directly approximated  
71 from whitecap coverage  $W$ , which represents the average visible surface area of bubble plumes  
72 and surface foam patches per unit sea surface area.  $W$  is a reasonably easily measurable quan-  
73 tity using optical video systems. Estimation of bubble plume depth is, however, challenging and  
74 rare, especially during active wave breaking period. This study provides concurrent observations  
75 of  $W$  and bubble plume penetration depth in various sea states.

76 Many previous studies have examined the dependence of  $W$  on wind speeds and sea states  
77 [*Monahan and Muirheartaigh*, 1980; *Callaghan et al.*, 2008; *Kleiss and Melville*, 2010; *Schwen-*  
78 *deman and Thomson*, 2015a; *Brumer et al.*, 2017; *Malila et al.*, 2022]. Despite large scatter in  
79 the data, particularly for wind speeds less than  $10 \text{ ms}^{-1}$ , these recent field studies have established  
80 fairly consistent empirical formulations that provide estimates of  $W$  based on given wind and/or  
81 sea state parameters.

82 Fewer previous studies reported mean values of the penetration depth of bubble plumes,  
83  $\overline{D}_{bp}$ , for a range of wind speeds using upward-looking sonars moored to the sea bed or a plat-  
84 form [*Thorpe*, 1982, 1986; *Dahl and Jessup*, 1995; *Vagle et al.*, 2010; *Wang et al.*, 2016; *Strand*  
85 *et al.*, 2020]. These observations show that  $\overline{D}_{bp}$  increases with increasing wind speed and varies  
86 from [1–5] m at low winds to [7–25] m during storms. However, the dependence of the statis-  
87 tics of  $D_{bp}$  on wind and sea state parameters is not well understood.

88 The main objective of this study is to understand and quantify the statistics that character-  
89 ize the size of the bubble plumes (averaged over many waves,  $O(\text{minutes})$ ) generated by break-  
90 ing surface waves in the open ocean. Our observations include arrays of freely drifting, surface-  
91 following SWIFT buoys together with shipboard wind and optical video systems, which enabled  
92 concurrent high-resolution measurements of wind, waves, whitecap coverage, bubble plumes,  
93 and turbulence over a wide range of sea state conditions in the North Pacific Ocean, including  
94 storms with sustained winds up to  $22 \text{ ms}^{-1}$  and significant wave heights up to 10 m. We estimate  
95 bubble plume penetration depth by using echograms that extend to more than 30 m depth in a surface-  
96 following reference frame using downward-looking echosounders integrated onboard the buoys.

97 We focus on examining the dependence of the statistics of the penetration depth of bub-  
98 ble plumes  $D_{bp}$  on various wind and wave parameters and the relation between  $D_{bp}$  statistics  
99 and  $W$ . Further, we comment on the role of wind history on  $W$  values. In a planned companion  
100 paper, we also investigate dynamic relationships between these bubble plume statistics and total  
101 wave breaking dissipation using our synchronized observations of bubble plumes and dissipa-  
102 tion rates.

103 The rest of this paper is organized as follows: §2 describes the observed environmental condi-  
104 tions and our analysis for estimating bubble plume penetration depths. §3 describes the depen-  
105 dency of the bubble plume statistics on various wind and sea state parameters. Discussion and  
106 a summary of the main findings are provided in §4 and §5, respectively.

## 107 2 Methods

### 108 2.1 Data

109 The present data set includes observations of wind, waves, air and sea temperature, near-  
110 surface turbulence, time-depth images of acoustic backscatter (referred to as echograms), above-  
111 and sub-surface optical imagery by freely drifting surface following SWIFT buoys [*Thomson*, 2012;  
112 *Thomson et al.*, 2019], as well as concurrent shipboard measurements of wind, temperature, and  
113 whitecap coverage. The data were collected during an 18-day research cruise in the North Pa-

cific Ocean (Figure A.1a) in December 2019. The primary objective of the cruise was to carry out concurrent observations of breaking surface gravity waves and the associated bubble plume statistics. The secondary objective was the replacement of a long-term moored wave buoy at Ocean Station PAPA ( $50^\circ$  N,  $145^\circ$  W), which reports as CDIP 166 and NDBC 46246. Hereafter we refer to the present data set and cruise with the abbreviation PAPA.

The PAPA cruise, aboard the R/V *Sikuliaq*, departed Dutch Harbor, AK, on 5 December 2019 and ended in Seattle, WA, on 23 December 2019. Arrays of SWIFT buoys were deployed from the ship early in the morning and usually recovered later the same day. Most of the shipboard and autonomous measurements were carried out during local daylight hours, while the eastward transits were continued overnight. Figure A.1a shows the PAPA cruise track and average location of SWIFT buoys during each deployment along the transit. Figures A.1b, A.1c, and A.1d show that the PAPA data set includes a wide range of sea state conditions;  $U_{10N}$  ( $0.8\text{--}22\text{ ms}^{-1}$ ),  $H_s$  ( $2.2\text{--}10.0\text{ m}$ ),  $T_m = f_m^{-1}$  ( $6.6\text{--}11.6\text{ s}$ ),  $T_p$  ( $6.5\text{--}14.6\text{ s}$ ),  $T_{air} - T_{sea}$  ( $-4.4\text{ to }1.2^\circ\text{ C}$ ),  $c_m/U_{10N}$  ( $0.6\text{--}17.5$ ),  $dU_{10N}/dt$  ( $-10.2\text{ to }6.9\text{ ms}^{-1}/\text{hr}$ ); including a storm in the vicinity of Station PAPA with sustained wind speeds up to  $22\text{ ms}^{-1}$  and significant wave heights up to 10m. We note that a significant portion of the data was collected in the presence of persistent rain (though rain rates were not measured).

Raw SWIFT data were collected at sampling rates of  $0.5\text{--}5\text{ Hz}$  in bursts lasting 512 seconds at intervals of 12 minutes. Processed SWIFT data, such as wave spectra and bubble plume statistics, are produced for each burst for each buoy, and then concurrent bursts are averaged among the buoys (typically 4 of them). During the cruise, more than 2000 bursts of data were collected by arrays of two to six SWIFT buoys, and 543 processed data points are obtained at intervals of 12 minutes spread over 14 daylight deployments. Statistics from the shipboard measurements, such as wind speeds and whitecap coverage, represent 10-minute average values and are obtained at the same time that the processed SWIFT data points are produced.

Two versions of SWIFT buoys have been concurrently used here, the version 3 buoys have uplooking Nortek Aquadopp Doppler sonars [Thomson, 2012], and the version 4 buoys have down-looking Nortek Signature1000 Doppler sonars which enable synchronous measurements of acoustic backscatter (i.e., echograms), broadband Doppler velocity profiles, and high-resolution (HR) turbulence profiles through the near-surface layer [Thomson *et al.*, 2019]. This new SWIFT capability allows us to quantify penetration depths of bubble plumes in a surface-following reference frame, with raw data that captures the time evolution within individual waves (i.e., phase-resolved).

The methodologies that we use to process echogram data and obtain bubble plume statistics are described in detail in this section. The instrumentation and methods that are used to obtain the rest of the relevant environmental variables and statistics, such as wind speeds, wave spectra, and whitecap coverage, are described in several previous observational studies [Thomson, 2012; Schwendeman and Thomson, 2015a; Thomson *et al.*, 2016, 2018], and will be briefly summarized here for convenience.

## 2.2 Wind Statistics

We calculate the neutral 10-m wind speed  $U_{10N}$  (Figure A.1b) following Hsu [2003] from wind speed measurements at 10 Hz, corrected for ship motion and airflow distortion, by three shipboard sonic anemometers (Metek Omni-3) at approximately 16.5 m height above the sea surface. The mean  $U_{10N}$  values are obtained over 10-minute bursts of raw data. We note that the atmospheric stability ( $T_{air} - T_{sea}$ ) effect is often neglected in the estimation of the 10-m wind speed, or  $U_{10N}$  is simply approximated using the mean wind profile power law given by  $U_{10}^{PL} = U_z (10/z)^{1/7}$ . Figure A.1b shows the observed range of the shipboard measurements of  $U_{10}^{PL} = U_{16.5} (10/16.5)^{1/7}$  (solid line) and the estimated  $U_{10N}$  values (circles) for the times that the processed SWIFT data are produced.

During the PAPA cruise, the atmospheric stability was negative most of the time ( $T_{air} - T_{sea}$  ranged between  $-4.4$  °C and  $1.2$  °C as shown in Figure A.1d) indicating unstable atmospheric boundary layer conditions. Figure A.2a shows that  $U_{10N}$  values are greater than  $U_{10}^{PL}$  in unstable atmospheric conditions by between 2% and 30%, where the differences decrease with increasing wind speed or  $T_{air} - T_{sea}$  values. Figure A.2a also shows that the differences between  $U_{10N}$  and  $U_{10}^{PL}$  values are within 2% for stable atmospheric conditions (i.e.,  $T_{air} - T_{sea} > 0$ ).

The friction velocity  $u_*$  of the airflow is readily estimated from a modified logarithmic mean wind profile [Hsu, 2003], which accounts for atmospheric stability effects. The air-side friction velocity is also independently estimated using the inertial dissipation method and assuming neutral atmospheric stability as described in Thomson *et al.* [2018]; Yelland *et al.* [1994]. However, robust estimates of  $u_*$  are only achieved for a fraction of the time due to the strict requirements that the ship's heading is within 60 degrees of the wind and that the turbulent wind spectra match an expected frequency<sup>-5/3</sup> shape. Figure A.2b shows the two estimates of  $u_*$  against  $U_{10N}$  during the PAPA cruise. The mean  $u_*$  values are obtained over 10-minute bursts. The corresponding data from Schwendeman and Thomson [2015a], in which  $u_*$  values were estimated using the inertial dissipation method, are also compiled in Figure A.2b. Here we use the  $u_*$  values obtained from a modified logarithmic mean wind profile [Hsu, 2003] for all relevant analyses.

### 2.3 Wave Statistics

Wave spectral information, including the wave power spectral density  $E(f)$  ( $m^2s$ ) and frequency-dependent directional spread  $\Delta\theta(f)$ , are obtained by combining GPS and IMU measurements (collected by the SWIFT buoys) over the frequency range (0.01–0.49) Hz with a 0.012 Hz resolution as described in Schwendeman and Thomson [2015a]; Thomson *et al.* [2018]. As detailed below, several bulk and spectral wave parameters are then calculated using  $E(f)$  and  $\Delta\theta(f)$ .

Figure A.2c shows examples of the observed  $E(f)$ , colored by  $U_{10N}$ , for  $U_{10N} > 10 \text{ ms}^{-1}$ . The two vertical dotted lines in Figure A.2c show the equilibrium range  $\sqrt{2}f_m$  to  $\sqrt{5}f_m$ , defined by Schwendeman and Thomson [2015a], over which the spectra approximately decay as  $f^{-4}$  consistent with the observations of Schwendeman and Thomson [2015a]. Here,  $f_m$  is the spectrally-weighted mean frequency given by

$$f_m = \frac{\int f E(f) df}{\int E(f) df}. \quad (1)$$

Figure A.2d shows the observed range of two commonly used alternatives for a characteristic wave period  $T$ , the peak wave period  $T_p = f_p^{-1}$  and the mean wave period  $T_m = f_m^{-1}$  (Eq. 1), as a function of  $U_{10N}$ . Figure A.2d also shows the wind sea mean wave period  $T_m^{ws} = (f_m^{ws})^{-1}$ , where  $f_m^{ws}$  calculated as given by Eq. 1 but over the wind sea portion of the observed wave spectra  $E^{ws}(f)$ . Here  $E^{ws}(f)$  is estimated using a 1D wave spectral partitioning technique following Portilla *et al.* [2009]. The solid lines in Figure A.2d represent the  $T_m$  and  $T_p$  values predicted by the Pierson-Moskowitz spectrum, a representative spectrum of fully developed wind-driven seas.

Figure A.2e shows the observed range of several characteristic wave heights as a function of  $U_{10N}$ , with  $H_s = 4(\int E(f) df)^{1/2}$  the total significant wave height,  $H_p = 4(\int_{0.7f_p}^{1.3f_p} E(f) df)^{1/2}$  a peak wave height (after Banner *et al.* [2000]), and  $H_s^{ws} = 4(\int E^{ws}(f) df)^{1/2}$  the wind sea significant wave height. Two estimates of the significant wave height of fully developed seas  $H_{s,fd}$  (solid lines) given by Carter [1982] and Chen *et al.* [2002] are also plotted in Figure A.2e. Results shown in Figures A.2d and A.2e indicate that significant swell is present at moderate and calm winds in the PAPA data.

Several estimates of the corresponding wave age are presented in Figure A.2f, where  $c_p$  and  $c_m$  are the wave phase speeds corresponding to  $f_p$  and  $f_m$ , respectively. These results show that a significant portion of the PAPA data at high winds ( $U_{10N} \geq 15 \text{ ms}^{-1}$ ) are characterized as developing seas ( $c_p/u_* < 30$  or  $c_p/U_{10N} < 1.2$ ), and that equilibrium seas ( $c_p/u_* \approx 30$  or  $c_p/U_{10N} \approx 1.2$ ) are mostly observed at moderate winds.

210 It is generally accepted that the wave steepness (or slope), defined as  $S = Hk/2$  with  $H$   
 211 and  $k$  are a characteristic wave height and wavenumber, is the most relevant local geometric wave  
 212 parameter to characterize surface gravity wave breaking and related processes in deep water [*Per-*  
 213 *lin et al.*, 2013]. Several formulations have been proposed to quantify a representative wave steep-  
 214 ness in a wave-averaged sense which are either defined based on wave spectral information [*Ban-*  
 215 *ner et al.*, 2002] or bulk wave parameters [*Banner et al.*, 2000].

216 A measure of mean square slope ( $mss$ ) over a frequency range  $f_1 \leq f \leq f_2$ , as proposed  
 217 by *Banner et al.* [2002], is calculated as

$$mss = \int_{f_1}^{f_2} k^2 E(f) df = \int_{f_1}^{f_2} \frac{(2\pi f)^4}{g^2} E(f) df, \quad (2)$$

218 and is shown to be a skillful spectral steepness parameter for predicting wave breaking statistics  
 219 in the open ocean [*Schwendeman and Thomson*, 2015a; *Brumer et al.*, 2017]. Many field obser-  
 220 vations of the speed of visible breaking wave crests [*Phillips et al.*, 2001; *Melville and Matusov*,  
 221 2002; *Gemmrich et al.*, 2008; *Thomson and Jessup*, 2009; *Kleiss and Melville*, 2010; *Sutherland*  
 222 *and Melville*, 2013; *Schwendeman et al.*, 2014] have shown that most of surface gravity wave break-  
 223 ing occurs at frequencies noticeably greater than the frequency at the peak of  $E(f)$ ,  $f_p$ , with most  
 224 frequent breaking occurring at  $\approx 2f_p$ . We note that  $f_m/f_p$  varies between 0.9 and 1.6 in the PAPA  
 225 data (Figure A.2d) where most of the  $f_m/f_p$  values are within a range (1.1–1.4), and that the  
 226 Pierson-Moskowitz spectrum gives  $f_m/f_p \approx 1.30$ . Following *Schwendeman and Thomson* [2015a],  
 227 here we take an equilibrium range  $mss$  calculated over a frequency range  $\sqrt{2}f_m \leq f \leq \sqrt{5}f_m$   
 228 ( $2k_m \leq k \leq 5k_m$ ,  $c_m/\sqrt{5} \leq c \leq c_m/\sqrt{2}$ ), which is related to an average spectral steepness of  
 229 a significant portion of visible breaking waves, especially in developed and equilibrium sea states.

230 Figures A.2g and A.2h show the variation of the equilibrium range  $mss$  and  $mss/\Delta f$  ( $\Delta f =$   
 231  $(\sqrt{5}-\sqrt{2})f_m$ ) against  $U_{10N}$ , all colored by wind accelerations  $dU_{10N}/dt$  defined as the rate of  
 232 change of  $U_{10N}$  over 1.5 hr, in the PAPA data together with the corresponding data from *Schwen-*  
 233 *deman and Thomson* [2015a]. Figures A.2g and A.2h also show the corresponding values that  
 234 are obtained from the Pierson-Moskowitz spectrum, which is a representative spectrum of a fully  
 235 developed sea under constant wind ( $dU_{10N}/dt = 0$ ), given by  $[mss]_{PM} \approx 0.436\alpha$  ( $\alpha = 8.1 \times$   
 236  $10^{-3}$ ) and  $[mss/\Delta f]_{PM} \approx \pi\alpha g^{-1}U_{10N}$ . Figures A.2g also shows that the observed equilibrium  
 237 range  $mss$  in equilibrium, developing, and old seas are, on average, consistent with, greater, and  
 238 smaller than those predicted by the Pierson-Moskowitz spectrum, respectively. Further, our ob-  
 239 servations corroborate the analytical relations obtained from the Pierson-Moskowitz spectrum,  
 240 i.e., equilibrium range  $mss$  is independent of wind speeds and  $mss/\Delta f \propto U_{10N}$  in fully devel-  
 241 oped seas with constant winds. Further, Figure A.2i shows the corresponding wind sea  $mss^{ws}/\Delta f$   
 242 values where  $mss^{ws}$  is calculated as given by Eq. 2 but using  $E^{ws}(f)$  over a frequency range  $\sqrt{2}f_m \leq$   
 243  $f \leq \sqrt{5}f_m$ .

244 *Schwendeman and Thomson* [2015a] and *Brumer et al.* [2017] used a normalized  $mss$  pa-  
 245 rameter,  $mss/\Delta f \Delta \theta$ , where  $\Delta \theta$  is the average of  $\Delta \theta(f)$  over  $\sqrt{2}f_m \leq f \leq \sqrt{5}f_m$  and reported  
 246 a decrease of data scatter in their plots of whitecap coverage against  $mss/\Delta f \Delta \theta$  compared to  $mss$ .  
 247 At any given wind speed, the  $mss/\Delta f \Delta \theta$  values in the present data are, on average, greater than  
 248 those in *Schwendeman and Thomson* [2015a] despite consistent  $mss$  and  $mss/\Delta f$  values in both  
 249 data sets. We note that  $mss/\Delta f \Delta \theta$  can not be defined in a long-crested wavefield or from a 1D  
 250 wave spectrum. We further note that  $\Delta \theta$  is sensitive to the type of buoy and method of process-  
 251 ing [*Donelan et al.*, 2015], such that values may not be directly comparable between data sets.  
 252 Here we avoid the directional normalization and choose the equilibrium range  $mss/\Delta f$  as a rep-  
 253 resentative measure of spectral steepness of dominant breaking waves.

254 The observed range of several bulk steepness parameters, including the significant spec-  
 255 tral peak steepness  $H_p k_p/2$  (after by *Banner et al.* [2000]) and the significant wave steepness  $H_s k_p/2$ ,  
 256 against  $mss/\Delta f$  are shown in Figures A.2j and A.2k. Here the peak  $k_p$  and mean  $k_m$  wave num-  
 257 bers are obtained from the linear gravity wave dispersion relation given by  $k = (2\pi)^2 g^{-1} T^{-2}$ .  
 258 Consistent with the literature, we consider these bulk steepness parameters here.

259 Finally, several dimensionless bulk parameters with general forms of

$$R_H = u_* H / \nu_w, \quad (3)$$

260 and

$$R_B = u_*^2 / (2\pi T^{-1} \nu_w), \quad (4)$$

261 where  $\nu_w \approx 1.4 \times 10^{-6} \text{ m}^2 \text{ s}^{-1}$  is the kinematic viscosity of seawater for  $T_w \approx 9\text{C}^\circ$ , are consid-  
 262 ered. These parameters represent combined effects of wind forcing and wave field and are shown  
 263 to have skills in predicting oceanic whitecap coverage [Zhao and Toba, 2001; Scanlon and Ward,  
 264 2016; Brumer et al., 2017]. Figure A.21 shows the variation of  $R_{H_{eq}} = u_* H_{eq} / \nu_w$  and  $R_B^m =$   
 265  $u_*^2 / (2\pi T_m^{-1} \nu_w)$  parameters as a function of the equilibrium range  $mss / \Delta f$  in the PAPA data. Here  
 266  $H_{eq} = 4[\int_{\sqrt{2}f_m}^{\sqrt{5}f_m} E(f)df]^{1/2}$  and  $T_m = f_m^{-1}$  are taken as a characteristic wave height  $H$  and  
 267 period  $T$ , respectively.

## 268 2.4 Whitecap Processing

269 The whitecap coverage data set in this study is the same as the North Pacific whitecap cov-  
 270 erage data set described in the recent study by Malila et al. [2022]. This section provides a sum-  
 271 mmary of the acquisition and processing of the data set, much of which is equal or similar in terms  
 272 of hardware and software to the study by Schwendeman and Thomson [2015a].

273 Visual images of the sea surface were collected from shipboard video camera systems on  
 274 the port and starboard sides of the vessel. The cameras, of model PointGrey Flea2 equipped with  
 275 2.8 mm focal-length lenses, recorded at 5–7.5 frames per second during daylight hours. A total  
 276 of 60 hours of image data were collected while the ship was stationary, mostly coincident with  
 277 the SWIFT buoy deployments and recoveries. The video acquisitions varied in length between  
 278 5 and 60 minutes, but the final mean whitecap coverage  $W$  values were obtained over 10–20-minutes  
 279 bursts. Each  $W$  value represents a 10-minute average of consecutive frames.

280 The image processing of the grayscale video frames for whitecap coverage estimation fol-  
 281 lowed the approach of Schwendeman and Thomson [2015a], in which the ship motion due to waves  
 282 (i.e., pitch and roll) was corrected for using a slightly modified version of the horizon tracking  
 283 algorithm of Schwendeman and Thomson [2015b]. The stabilized images were subsequently geo-  
 284 rectified and gridded to regular grids with 0.8 m grid resolution. The whitecap-related foam was  
 285 isolated from the stabilized, geo-rectified, and gridded frames using the pixel intensity thresh-  
 286 olding algorithm of Kleiss and Melville [2011]. The frame-wise fractional whitecap coverage was  
 287 then computed as the ratio of pixels detected as belonging to whitecaps (given a value of one)  
 288 to the total number of pixels in the frame. A subset of the original and thresholded frames in each  
 289 burst was visually quality controlled for satisfactory image exposure and lens contamination (e.g.,  
 290 raindrops or sea spray). Only image sequences with consistent lighting conditions and minimal  
 291 lens contamination were used in the final data set.

## 292 2.5 Echogram Processing

293 Acoustic backscattering data were obtained using the echosounding capabilities of the downward-  
 294 looking beam of the Nortek Signature1000 acoustic Doppler current profiler (ADCP) mounted  
 295 on the version 4 SWIFT buoys. During the PAPA cruise manufacturer firmware version 2205 was  
 296 used. Sampling frequencies and pulse repetition rates for the echosounder were 1 MHz and one  
 297 second, respectively. A transmit pulse duration of 500  $\mu\text{s}$  was used. The vertical sampling res-  
 298 olution provided by the instrument is 1 cm and is presented to depths from  $0.3 \text{ m} \leq z_w \leq 30.3\text{m}$ ,  
 299 where  $z_w$  is positive downward and  $z_w = 0$  represents the instantaneous free surface level af-  
 300 ter accounting for the depth of the unit on the SWIFTS. The echosounder mode was operated in  
 301 512-s bursts gathered in the surface-following reference frame, from which echograms are pre-  
 302 sented. Based on the size of the transducer and the operational frequency, we estimate that the  
 303 echosounders' acoustic near-field based on the definition provided in Medwin and Clay [1998]  
 304 is less than 1 m. To limit the potential impacts of the acoustic near-field, we present data only from  
 305 ranges greater than 1 m from the transducer face (i.e., depth range from  $1.3 \text{ m} \leq z_w \leq 30.3\text{m}$ ).

306 As detailed below, penetration depths of bubble plumes are estimated from the volume backscat-  
 307 tering strength. The volume backscattering strength  $S_v$  [dB re  $1\text{m}^{-1}$ ] is the logarithmic form of  
 308 the backscattering cross section per unit volume  $M_v$  as is given by [Vagle *et al.*, 2012]. When the  
 309 signal is dominated by the presence of bubbles, as is the focus on this manuscript, this is described  
 310 by

$$\begin{aligned}
 S_v &= 10 \log_{10} M_v = 10 \log_{10} \int_0^\infty \sigma_s(a_b) N(a_b) da_b \\
 &= 10 \log_{10} (10^{\frac{Pr}{10}} - 10^{\frac{Nt}{10}}) + 20 \log_{10} r + 2\alpha r + G_{cal} - 10 \log_{10} \left( \frac{c\tau}{2} \right) - \phi,
 \end{aligned}
 \tag{5}$$

311 where  $\sigma_s(a_b) = 4\pi a_b^2 / [(f_R/f)^2 - 1]^2 + \delta^2$  [ $\text{m}^2$ ] is the scattering cross section for a bubble  
 312 with radius  $a_b$  [m] and  $N(a_b)$  is the bubble size distribution. The right-hand side of the equa-  
 313 tion is an implementation of the sonar equation where  $Pr$  is the received signal including noise,  
 314  $Nt$  is the noise threshold,  $r$  is the range from the transducer,  $\alpha$  is the attenuation coefficient,  $c$   
 315 is the speed of sound in the water,  $\tau$  is the transmit pulse duration,  $\phi$  is the equivalent beam angle,  
 316 and  $G_{cal}$  is a gain factor for a configured transmit power level for the transducer (see Appendix  
 317 A).  $G_{cal}$  was determined by using standard calibration techniques for echosounders [Demer *et al.*,  
 318 2015]. We note that we identified issues with the saturation of the signals associated with sys-  
 319 tem gains during calibration. This results in saturated signals at short ranges when measured backscat-  
 320 tering intensity is high, thereby truncating the dynamic range of the system at the upper end. A  
 321 longer discussion of this is also included in Appendix A.

322 To estimate the average noise level of the transducer, we calculate burst-averaged  $Pr$  val-  
 323 ues at large ranges at low sea states at which the measured signal, not compensated for range and  
 324 attenuation, does not vary with depth. At these ranges, we assume that due to transmission losses  
 325 and the weak scattering in the water column the system is simply measuring its own electrical  
 326 noise and that increases in  $S_v$  are driven primarily by the addition of the time-varying gain com-  
 327 ponents in Eq. 5. This approach is consistent with those often applied in fisheries acoustics ap-  
 328 plications (e.g., De Robertis and Higginbottom [2007]). Here we found the average noise level  
 329 of approximately 22 [dB] and set  $Nt = 26$  [dB], i.e., only echogram data values with  $Pr > Nt$   
 330 are considered for the bubble statistics analysis. We note that subsequent firmware revisions and  
 331 different internal or internal processing parameters are expected to result in different noise thresh-  
 332 olds and calibration gains.

333 To estimate the local penetration depth of entrained bubbles, we first need to identify a thresh-  
 334 old  $S_v^{th}$  below which the backscatter signal indicates the absence of signals associated with en-  
 335 trained bubbles exceeding the background conditions. These background conditions may be driven  
 336 by populations of tiny residual bubbles, biological backscattering, or microstructure in the up-  
 337 per water column. Note that the mixed layer depth was always greater than 40 m in areas sam-  
 338 pled during the PAPA cruise; thus, acoustic scattering from stratification can be neglected.

339 The local penetration depth of entrained bubbles is then defined relative to the instantaneous  
 340 free surface level ( $z_w = 0$ ) at the vertical level  $Z_b$ , in the surface-following reference frame, at  
 341 which  $S_v > S_v^{th}$  for  $z_w \leq Z_b$ ; otherwise  $Z_b = \text{NaN}$  (Not-a-Number). We note that this thresh-  
 342 olding technique to estimate bubble penetration depth is analogous to the pixel intensity thresh-  
 343 olding commonly used for whitecap coverage estimations (see §2.4). Similar thresholding tech-  
 344 niques have been used by previous studies [Thorpe, 1986; Dahl and Jessup, 1995; Trevorrow,  
 345 2003; Vagle *et al.*, 2010; Wang *et al.*, 2016] with empirical  $S_v^{th}$  values ranging from -70 dB re 1/m  
 346 to -50 dB re 1/m using sonars with operating frequencies ranging between  $\approx 20\text{kHz}$  and  $\approx 200\text{kHz}$ .  
 347 Hereafter we refer to this bubble detection method as BDM1.

348 We identified the time between 18:00 and 19:00 UTC Dec 16 as a period with relatively  
 349 calm sea surface conditions and minimal whitecapping during which no visible bubbles and sur-  
 350 face foam were observed in the above-surface, and sub-surface images collected by the cameras  
 351 integrated on SWIFT buoys along with the images from the shipboard cameras. Further, Figure A.1b  
 352 shows that the wind speeds just before the SWIFTS deployment on Dec 16 were less than  $3\text{ms}^{-1}$   
 353 for several hours. Figure A.1b also shows that although wind speed was increasing during the

354 rest of the day in the presence of steady rain, it remained below  $5 \text{ ms}^{-1}$  between 18:00 and 19:00  
 355 UTC. These observations suggest this is a suitable period for establishing baseline levels for near-  
 356 surface backscattering with negligible contributions of bubbles injected by active breaking at the  
 357 surface.

358 The baseline can be established by using statistical averages of the  $S_v$  from this relatively  
 359 calm period with low levels of observed volume backscattering. Figure A.3a shows an example  
 360 echogram, above-surface image, and vertical profiles of burst-averaged and top 10%-averaged  
 361 of  $S_v$  values just after the low backscattering conditions on Dec 16 described above. The echogram  
 362 data during low-backscattering conditions reveals that significant portions of the corresponding  
 363  $S_v$  values vary between  $-90 \text{ dB re 1/m}$  and  $-75 \text{ dB re 1/m}$  with the burst-averaged values,  $\bar{S}_v$ , less  
 364 than  $-80 \text{ dB re 1/m}$ . We also found that  $\bar{S}_v < -80 \text{ dB re 1/m}$  holds for the rest of calm sea state  
 365 conditions ( $U_{10N} < 3 \text{ ms}^{-1}$ ,  $dU_{10N}/dt < 1 \text{ ms}^{-1}$ ) within the PAPA data. We take  $S_v^{th} = -70$   
 366  $\text{dB re 1/m}$  (as in *Vagle et al.* [2010]) to distinguish between regions with and without the pres-  
 367 ence of recently entrained bubbles in the water column.

368 Even very low bubble void fractions,  $O(10^{-7})$  or less, can result in  $S_v$  values greater than  
 369  $S_v^{th}$  due to the relatively strong acoustic backscattering response of bubbles [*Dahl and Jessup*,  
 370 1995], even when they are sampled well above resonance. For reference, at 1 MHz bubble radii  
 371 from approximately  $3 \mu\text{m}$  to  $7 \mu\text{m}$  would be resonant in the upper water column [*Medwin and*  
 372 *Clay*, 1998; *Vagle and Farmer*, 1998]. We assume that these smaller bubbles dissolve rapidly ( $<$   
 373 10 seconds), even when the upper water column is supersaturated, as suggested by *Blanchard and*  
 374 *Woodcock* [1957]. Thus, the measured backscattering reflects backscattering from an unknown  
 375 and evolving population of bubbles that are dissolving and slowly transported by their own buoy-  
 376 ancy and/or local currents and turbulence.

377 We define another estimate of the local penetration depth of entrained bubbles as the depth  
 378  $z_b (\leq Z_b)$  at which  $S_v > S_v^{th}$  for  $z_w \leq z_b$  and  $S_v > S_v^{th} + 20 \text{ dB}$  for  $z_b/2 \leq z_w \leq z_b$ ; otherwise  
 379  $z_b = \text{NaN}$ . That is, the penetration depth is defined by the depth at which the volume backscat-  
 380 tering signal continuously exceeds the defined threshold at the surface, and  $S_v$  values deeper in  
 381 the water column exceed background thresholds by at least 20 dB. Hereafter we refer to this bub-  
 382 ble detection method as BDM2.

383 Figure A.3 shows examples of echogram data and the corresponding  $Z_b$  (obtained from BDM1,  
 384 dotted-dashed lines) and  $z_b$  (obtained from BDM2, solid lines) values during a developing sea  
 385 on Dec 16 just after the relatively bubble-free condition described above (panels *a* and *b*) and dur-  
 386 ing a storm with sustained wind speeds of greater than  $18 \text{ ms}^{-1}$  on Dec 11 (panels *c* and *d*). Fig-  
 387 ure A.3 also shows examples of subsurface optical images, collected at times when  $S_v < S_v^{th}$   
 388 for  $1.3 \text{ m} \leq z_w$  (panel *e*), portions of  $S_v$  values are greater than  $S_v^{th}$  but remain below  $S_v^{th} + 20$   
 389 dB (panels *f* and *g*), and a portion of  $S_v$  values is greater than  $S_v > S_v^{th} + 20$  (panels *h*, *i* and  
 390 *j*). These images qualitatively demonstrate that the entrained surface bubbles at times at which  
 391 both BDM1 and BDM2 are satisfied, i.e.,  $Z_b \neq \text{NaN}$  and  $z_b \neq \text{NaN}$ , have significantly more sub-  
 392 surface visible optical signature than those at times at which  $Z_b \neq \text{NaN}$  but  $z_b = \text{NaN}$ . Com-  
 393 paring all available concurrent subsurface images and echogram data, we conclude that a simi-  
 394 lar trend exists across all the PAPA data.

395 Although we cannot ultimately constrain the differences in void fractions or bubble pop-  
 396 ulations using our sampling method, we can confidently state that our second bubble detection  
 397 criterion (BDM2) laid out above identifies periods where void fractions increase by a minimum  
 398 of two orders of magnitude compared to the first bubble detection criterion (BDM1). Under the  
 399 simplest conditions where bubble size distribution remains constant, a 20 dB increase in backscat-  
 400 tering would correspond to an increase in the void fraction of two orders of magnitude. This is  
 401 driven by a linear relationship between backscattering and the number of scatterers so long as  
 402 the distribution has not changed or has not been attenuated by high bubble volumes. (Eq. 5). Fur-  
 403 thermore, the high bubble void fractions following breaking waves may result in significant ex-  
 404 cess attenuation of the signals, which is not accounted for in our analysis here [*Vagle and Farmer*,  
 405 1998; *Deane et al.*, 2016; *Bassett and Lavery*, 2021]. Such observations have been reported at

406 lower frequencies, where extinction cross sections for resonant bubbles are much larger, but we  
 407 expect that the high void fractions following a breaking event will also have a temporary impact  
 408 on measured acoustic backscatter. The result of this is that increases in volume backscattering  
 409 following localized breaking events likely understate the increase in scattering that would oth-  
 410 erwise be observed from the bubble populations, given the transducer’s location near the surface.

411 Overall,  $z_b$  values represent the local penetration depths of entrained bubbles that have sig-  
 412 nificantly more void fraction and visible optical signature than those that reach  $Z_b$ . This is con-  
 413 sistent with a broad range of prior observations measuring bubbles in the upper ocean, which show  
 414 significant decreases in bubble densities with depth [Vagle and Farmer, 1998; Medwin, 1977].

### 415 3 Results

416 In this section, we present the observations of the residence time (§3.1) and the penetra-  
 417 tion depth (§3.2) of bubble plumes and whitecap coverage (§3.3) as a function of various wind  
 418 and sea state parameters defined in §2. Estimations of the volume of bubble plumes from the mea-  
 419 sured whitecap coverage and plume penetration depths are discussed in the next section.

#### 420 3.1 Bubble Plume Residence Time

421 Figure A.4a shows a schematic of a SWIFT track drifting across an intermittent field of sat-  
 422 urated (with visible optical surface signature) and diffused (without visible optical surface sig-  
 423 nature) bubble clouds during a 512-s burst of data along which echogram data are collected in  
 424 a surface following reference frame. The buoy has a ”wind slip” velocity relative to the surface  
 425 water  $U_{slip} \approx 0.01U_{10N}$  that is caused by wind drag on the portion of the buoy above the sur-  
 426 face [Iyer *et al.*, 2022]. Note that the example SWIFT track shown here is calculated with respect  
 427 to the earth frame, so the example includes both the true surface current and the wind slip of the  
 428 buoy (which combine together to make the observed drift velocity of the buoy, typically  $U_{drift} \approx$   
 429  $0.04U_{10N}$ ). Thus apparent residence time of detectable bubble clouds in echogram data could  
 430 be shorter than their true residence time due to the relative drift of the buoys.

431 As illustrated in Figure A.4a, the apparent residence time of each bubble cloud in echogram  
 432 data is directly related to the way the buoy crosses the bubble cloud with respect to its main axis.  
 433 To minimize this potential sampling bias, here we define the residence time of bubble plumes as  
 434 an average of the highest one-third of the apparent residence time of bubble clouds detected in  
 435 all concurrent bursts of the echogram data.

436 Figure A.4b shows the variation of the bubble plume residence times  $T_{bp}$  and  $T_{bp,v}$  scaled  
 437 by the wind sea mean wave period  $T_m^{ws}$  (defined in §2.3) for wind speeds greater than  $6 \text{ ms}^{-1}$ .  
 438 Hereafter the statistics of bubble plumes obtained from the bubble detection methods BDM1 and  
 439 BDM2 (described in §2.5) are denoted by  $(\ )_{bp}$  and  $(\ )_{bp,v}$ , respectively. Results indicate that the  
 440 bubble plumes, especially those detected by BDM1, persist in the water column much longer than  
 441 the corresponding dominant active breaking period, which is expected to be a fraction of  $T_m^{ws}$ .

442 Figure A.5 shows the sub-surface visible signature of an example evolving bubble plume  
 443 at several instances during (panels (a1) to (a3)) and after (panels (a4) to (a8)) active breaking  
 444 collected by a GoPro camera on a SWIFT buoy looking from behind (upwave) the breaking event  
 445 in an old sea with moderate wind speeds of  $U_{10N} \approx 11 \text{ ms}^{-1}$  and  $T_m^{ws} \approx 6 \text{ s}$ . Figure A.6 also  
 446 shows example sub-surface images of two evolving bubble plumes during (panels (a – c) and  
 447 (e – f)) and after (panels d and g – h) active breaking during a storm with sustained wind speeds  
 448 of  $U_{10N} > 18 \text{ ms}^{-1}$  and  $T_m^{ws} \approx 10 \text{ s}$ . These images qualitatively show that void fractions in  
 449 the bubble plumes rapidly decrease after the active breaking period and that residual void frac-  
 450 tions persist for many wave periods. These observations are consistent with previous experimen-  
 451 tal [Lamarre and Melville, 1991; Blenkinsopp and Chaplin, 2007; Anguelova and Huq, 2012] and  
 452 numerical [Derakhti and Kirby, 2014, 2016; Derakhti *et al.*, 2018, 2020a,b] studies of laboratory-  
 453 scale breaking waves showing that average void fractions within bubble clouds vary from  $O(10\%)$

454 to  $O(1\%)$  during active breaking, and then, drop rapidly by several orders of magnitude within  
455 a few wave periods.

456 As discussed in detail in §2.5, plume regions with tiny bubble void fractions, e.g., the dif-  
457 fused bubble clouds shown in panels (a7) and (a8) of Figure A.5, are still detectable in our sam-  
458 pling method. Assuming that the scattering is dominated by bubbles with radii less than  $100\ \mu\text{m}$ ,  
459 the low bubble rise velocities (i.e., a few  $\text{cm s}^{-1}$  or less) would yield bubble residence times of  
460  $O(\text{minutes})$  which is consistent with the apparent residence time of the bubble plumes detected  
461 by BDM1 (Figure A.4b), here  $T_{bp} \approx 100\text{s}$  and  $\approx 200\text{s}$  for sea states similar to Figure A.5 and  
462 Figure A.6, respectively. Thus the statistics of the bubble plumes detected by BDM1, referred  
463 to by subscript  $bp$ , correspond to bubble plumes ranging from saturated plumes during active  
464 breaking to highly diffused plumes that may remain in the water column long after active break-  
465 ing (e.g., panel (a8) of Figure A.5). These observations also confirm that the bubble plumes de-  
466 tected by BDM2 in a given sea state represent plumes that have much shorter residence times and  
467 much more visible optical signature than those detected by BDM1 but noticeably exceed the per-  
468 sistence of visible surface foam formed during breaking, here  $T_{bp,v} \approx 12\text{s}$  and  $\approx 40\text{s}$  for sea  
469 states similar to Figure A.5 and Figure A.6, respectively.

### 470 3.2 Bubble Plume Penetration Depth

471 Example sub-surface images of the bubble plume shown in Figure A.5 illustrate that the  
472 average plume penetration depth (and volume) rapidly increases during the initial phase of the  
473 bubble plume evolution (e.g., panels (a1) to (a5), over several seconds). As shown in panels (a6)  
474 to (a8), the overall size of the plume keeps increasing for several wave periods but at rates much  
475 lower than during active breaking. This is consistent with the evolution of bubble plumes, tur-  
476 bulent kinetic energy (TKE), and dye patches in previous numerical and experimental studies of  
477 laboratory-scale isolated breaking focused waves [*Rapp and Melville, 1990; Melville et al., 2002;*  
478 *Derakhti and Kirby, 2014; Derakhti et al., 2018, 2020a*]. Large-scale coherent structures gener-  
479 ated by wave breaking crests are among potential drivers of such slow but persistent transport  
480 of bubbles long after active breaking [*Melville et al., 2002; Derakhti and Kirby, 2014; Derakhti*  
481 *et al., 2016*].

482 We define the mean,  $\bar{D}_{bp}$  and  $\bar{D}_{bp,v}$ , and significant bubble plume depths,  $D_{bp}^{1/3}$  and  $D_{bp,v}^{1/3}$ ,  
483 as

$$\bar{D}_{bp} = \frac{\sum_{i=1}^{N_{Z_b}} Z_b^i}{N_{Z_b}}, \quad \bar{D}_{bp,v} = \frac{\sum_{i=1}^{N_{z_b}} z_b^i}{N_{z_b}}, \quad (6)$$

484 and

$$D_{bp}^{1/3} = \frac{\sum_{i=2N_{Z_b}/3}^{N_{Z_b}} Z_b^i}{N_{Z_b}/3}, \quad D_{bp,v}^{1/3} = \frac{\sum_{i=2N_{z_b}/3}^{N_{z_b}} z_b^i}{N_{z_b}/3}, \quad (7)$$

485 where  $1.3\text{m} \leq Z_b^i \leq Z_b^{i+1} \leq 30.3\text{m}$ ,  $1.3\text{m} \leq z_b^i \leq z_b^{i+1} \leq 30.3\text{m}$  (see Figure A.3), and  $N_{Z_b}$   
486 and  $N_{z_b}$  are the total numbers of the estimated  $Z_b$  (obtained from BDM1) and  $z_b$  (obtained from  
487 BDM2) values over available concurrent (1 to 4) bursts (each burst includes more than 8 min-  
488 utes of data) of echogram data, respectively. The representative mean and significant bubble plume  
489 depths are obtained at 12-minute intervals at which the wind and wave statistics are available.

490 Figure A.7 shows the variation of the mean (Eq. 6) and significant (Eq. 7) bubble plume  
491 depths as a function of wind speed  $U_{10N}$  and equilibrium range  $mss/\Delta f$  (Eq. 2) and the corre-  
492 sponding best fits. All the plume depth measures are well correlated with wind speed and  $mss/\Delta f$   
493 with data scatter smaller than existing whitecap coverage data sets (including the PAPA data set  
494 shown in Figure A.11 below). Because time-dependent bubble depths less than 1.3 m are unavail-  
495 able here, the resultant plume depth statistics are expected to be biased high in low winds. Here-  
496 after the data points with  $U_{10N} < 6\ \text{ms}^{-1}$  are not considered in obtaining the relevant fits and  
497 their statistics. (This is also a typical minimum wind speed for visible whitecaps to occur.)

498 Of the bubble depths defined here (by Eqs. 6 and 7 above),  $\bar{D}_{bp}$  is defined similar to pre-  
499 vious studies [*Vagle et al., 2010; Wang et al., 2016; Strand et al., 2020*]. Our observations, shown

500 in Figure A.7a, indicate that the mean bubble plume depth  $\overline{D}_{bp}$  could be up to 14 m at  $U_{10N} \approx$   
 501  $20 \text{ ms}^{-1}$ . This is in good agreement with the observations of *Vagle et al.* [2010] and *Strand et al.*  
 502 [2020].

503 The black solid line in Figure A.7a represents the best fit to the binned  $\overline{D}_{bp}$  values with  
 504 a power law form given by

$$\overline{D}_{bp} = 0.092 [U_{10N}]^{1.58} \quad (8)$$

505 with  $r^2 = 0.90$  defined as in Eq. 12 below. As shown in Figure A.7a, the linear fit by *Vagle et al.*  
 506 [2010] also well describes the observed variability of  $\overline{D}_{bp}$  for moderate winds. For high winds,  
 507 however, the relationship between  $\overline{D}_{bp}$  and wind speed becomes nonlinear, and the  $\overline{D}_{bp}$  values  
 508 are, on average, greater than those reported by *Vagle et al.* [2010]. Underprediction of  $\overline{D}_{bp}$  at  
 509 high winds in *Vagle et al.* [2010] could be simply due to the linear extrapolation of  $S_v$  at depths  
 510 greater than 8m (see their Figure 3). *Wang et al.* [2016] also found a nonlinear relationship be-  
 511 tween mean bubble depth and wind speed at high winds. However, their mean bubble depths are  
 512 significantly (factor of 1.5-2) higher than the present (and other) observations. We note that the  
 513 averaging time used to obtain  $\overline{D}_{bp}$  at high winds is 8 or 16 minutes (depending on available con-  
 514 current bursts) which is comparable to that in *Wang et al.* [2016].

515 At any given wind speed, individual breaking events could generate bubble clouds with pen-  
 516 etration depths much higher than  $\overline{D}_{bp}$ . For example, Figure A.3c documents an example indi-  
 517 vidual bubble cloud with a penetration depth of  $\approx 30$  m which is approximately three times greater  
 518 than the corresponding average bubble plume depth (e.g., Eq. 8). Figure A.8 shows that the Rayleigh  
 519 distribution could reasonably describe the observed probability distribution function (PDF) of  
 520 the  $D_{bp}$  values at various wind speeds, especially for  $D_{bp} > \overline{D}_{bp}$ . Assuming the Rayleigh dis-  
 521 tribution for  $D_{bp}$ , we obtain the significant bubble depth as  $D_{bp}^{1/3} \approx 1.6\overline{D}_{bp}$  which is consis-  
 522 tent with our observations especially for  $U_{10N} > 10 \text{ ms}^{-1}$ . The best fit to the observed binned  
 523  $D_{bp}^{1/3}$  values with a power law form (black solid line in Figure A.7c) is obtained as

$$D_{bp}^{1/3} = 0.13[U_{10N}]^{1.63}, \quad (9)$$

524 with  $r^2 = 0.92$ . Assuming the Rayleigh distribution for  $D_{bp}$ , the maximum bubble depth is fur-  
 525 ther approximated as

$$D_{bp}^{max} \approx 2D_{bp}^{1/3} \approx 3.2\overline{D}_{bp}. \quad (10)$$

526 As explained in detail in §2.5 and consistent with observations shown in §3.1, at a given  
 527 sea state condition  $D_{bp,v}$  represents penetration depth of bubbles that have, on average, at least  
 528 two orders of magnitude more void fraction and significantly more visible optical signature than  
 529 those reach to  $D_{bp}$ . Figure A.8 shows that the population of the bubble plume depth  $D_{bp,v}$  val-  
 530 ues around their mean is considerably elevated compared to that in  $D_{bp}$  and that the observed  
 531 PDF of  $D_{bp,v}$  is better described by the Gamma distribution. Furthermore, our observations show  
 532 that  $D_{bp,v}^{1/3}/\overline{D}_{bp,v}$  varies, on average, from 1.2 at low winds to 1.5 at high winds, and that, in con-  
 533 trast to  $D_{bp}^{1/3}$ ,  $D_{bp,v}^{1/3}$  has an approximately linear relationship with wind speed. As shown in Fig-  
 534 ure A.7 the ratio  $D_{bp,v}^{1/3}/D_{bp}^{1/3}$  decreases with increasing wind speeds, varies from  $\approx 1$  at low winds  
 535 to  $\approx 0.6$  at high winds.

536 We examine the predictive skill of several wind and wave parameters, commonly used for  
 537 parameterizations of whitecap coverage, for bubble plume depths  $D_{bp}^{1/3}$  and  $D_{bp,v}^{1/3}$ . We quantify  
 538 the skill of each predictor  $\mathcal{X}$  (e.g.,  $U_{10N}$ ,  $u_*$ ,  $mss/\Delta f$ ,  $S$ ,  $R$ ,  $\dots$ , all defined in §2) by calculating  
 539 the best fit with a power law form  $a\mathcal{X}^n$  to the binned  $D_{bp}^{1/3}$  and  $D_{bp,v}^{1/3}$  values, using the least squares  
 540 method and then comparing the corresponding fit statistics obtained over all individual data points  
 541 with  $U_{10N} \geq 6 \text{ ms}^{-1}$ . Bins containing fewer than four bursts of data are not considered for data  
 542 fitting. Here we consider the root-mean-square error (RMSE) and the coefficient of determina-  
 543 tion  $r^2$ , which represent the overall quality of the fits, given by

$$\text{RMSE} = \sqrt{\frac{\sum_{i=1}^{i=N} D_{res,i}^2}{N}}, \quad (11)$$

544 and

$$r^2 = 1 - \frac{\sum_{i=1}^{i=N} D_{res,i}^2}{\sum_{i=1}^{i=N} (D_i - \overline{D_i})^2}, \quad (12)$$

545 where  $D_{res,i} = D_i - [a(\mathcal{X}_i)^n]$ ,  $D_i$  is either  $D_{bp}^{1/3}$  or  $D_{bp,v}^{1/3}$ ,  $N$  is the number of observations,  
 546 and the overbar indicates an average over all the considered data points. Here RMSE, defined in  
 547 linear space, indicates an average deviation from the fit, and  $r^2$  indicates the proportion of the  
 548 observed variability of the bubble plume depths that is predictable from the  $\mathcal{X}$  parameter. The  
 549 perfect fit corresponds to  $RMSE \sim 0$  and  $r^2 \sim 1$ .

550 Table 1 summarizes the coefficients ( $a$  and  $n$ ) and statistics (RMSE,  $r^2$ ) of the best fits,  $a\mathcal{X}^n$ ,  
 551 to the PAPA data for several predictive parameters  $\mathcal{X}$ . Of all the parameters considered here,  $U_{10N}$   
 552 has the highest skill in predicting the observed variability of both  $D_{bp}^{1/3}$  and  $D_{bp,v}^{1/3}$ . Results sum-  
 553 marized in Table 1 also document that the equilibrium range  $mss/\Delta f$  and  $H_s K_m/2$  show the high-  
 554 est skill among the spectral and bulk wave steepness predictors, respectively. For each type of  
 555 the predictors considered here, those that contain either the peak wave height, peak wave num-  
 556 ber, or peak wave period show the least skill. These results also hold for the mean bubble plume  
 557 depths statistics  $\overline{D}_{bp}$  and  $\overline{D}_{bp,v}$ .

558 Next, we examine how the bubble plume penetration depths, defined in Eqs. 6 and 7 above  
 559 and shown in Figure A.7, scaled by significant wave height  $H_s$  and mean wavelength  $L_m = 2\pi/k_m$   
 560 vary in various sea states. Our observations indicate that  $D_{bp}^{1/3}$  (note that  $\overline{D}_{bp} \approx 0.6 D_{bp}^{1/3}$ ) varies  
 561 from  $0.4H_s$  to  $4.8H_s$  and from  $0.01L_m$  to  $0.20L_m$  for wind speeds greater than  $6 \text{ ms}^{-1}$  (Figure A.9).  
 562 This is in good agreement with the observed range of scaled mean bubble depths reported in pre-  
 563 vious field observations [Thorpe, 1986; Wang et al., 2016; Strand et al., 2020].

564 Bulk wave statistics  $H_s$  and  $L_m$  (or  $H_p$  and  $L_p$ ) could be completely uncorrelated with the  
 565 scales of the corresponding wind sea (and dominant breaking waves) in the presence of propor-  
 566 tionally significant swell, e.g., in low and moderate winds ( $U_{10N} < 15 \text{ ms}^{-1}$ ) in the PAPA data  
 567 set as shown in Figures A.2d and A.2e. Thus we also consider the wind sea significant wave height  
 568  $H_s^{ws}$  and mean wavelength  $L_m^{ws}$  as scaling parameters here. Our data show that  $D_{bp}^{1/3}$  varies from  
 569  $1.4H_s^{ws}$  to  $9.2H_s^{ws}$  and from  $0.06L_m^{ws}$  to  $0.33L_m^{ws}$  for wind speeds greater than  $6 \text{ ms}^{-1}$  (Figure A.9).

570 Further, the corresponding binned data indicate that  $D_{bp}^{1/3} \approx [2.4-4.4]H_s^{ws}$ , and  $\approx [0.11-$   
 571  $0.2]L_m^{ws}$  (with  $\overline{D}_{bp} \approx [1.6-2.8]H_s^{ws}$ , and  $\approx [0.07-0.13]L_m^{ws}$ ). The observed range of these  
 572 scaled bubble plume depths is comparable with the scaled penetration depth of TKE and dye patches  
 573 reported in previous numerical and experimental studies of isolated breaking focused waves [Rapp  
 574 and Melville, 1990; Melville et al., 2002; Derakhti and Kirby, 2014; Derakhti et al., 2018, 2020a]  
 575 while the length scales of these laboratory-scale breaking waves are one to two orders of mag-  
 576 nitude smaller than those in the PAPA data sets.

577 Figures A.9 and A.10 show the dependency of some of the scaled plume depths on wind  
 578 speed and wave age. We note that other scaled plume depths considered here show, on average,  
 579 a similar trend with increasing wind speeds and wave age. Our observations indicate that all the  
 580 scaled bubble plume penetration depths considered here vary non-monotonically with increas-  
 581 ing wind speeds. However, they are all, on average, decreasing functions of wave age in devel-  
 582 oping seas (i.e.,  $c_p/U_{10N} < 1.2$ ). In other words, during the early stages of a young sea, i.e.,  
 583  $c_p/U_{10N} \ll 1.2$ , bubble plume penetration depth scaled by either significant wave height or  
 584 mean wavelength is, on average, much (two times or more) larger than those in equilibrium sea  
 585 states, i.e.,  $c_p/U_{10N} \approx 1.2$ . Previous field observations revealed that the former is dominated  
 586 by plunging breaking waves Thorpe [1992], while the dominant breaker type in the latter is ex-  
 587 pected to be spilling breaking. Previous numerical and experimental studies of laboratory-scale  
 588 breaking waves indicate that bubbles (and breaking-generated turbulence) penetrate, on average,  
 589 deeper in the water column beneath a plunger than a spilling breaker with the same length scale,  
 590 especially during active breaking [Rapp and Melville, 1990; Melville et al., 2002; Derakhti and  
 591 Kirby, 2014; Derakhti et al., 2018, 2020a,b]. Thus the observed dependency of scaled bubble

Table 1: Parameterizations of significant bubble plume depths  $D_{bp}^{1/3}$  and  $D_{bp,v}^{1/3}$  represented by the best fits with a power law form  $a \mathcal{X}^n$  as a function of several wind and wave parameters  $\mathcal{X}$  to the binned PAPA data for  $U_{10N} \geq 6 \text{ ms}^{-1}$ . The statistics of each fit are also calculated. The fits and their statistics are computed in linear space.

Plume Depth	Predictor $\mathcal{X}$	Results of the best fit		Statistics of the best fit	
		$a \mathcal{X}^n$		$U_{10N} \geq 6 \text{ ms}^{-1}$	
		$a$	$n$	RMSE	$r^2$
$D_{bp}^{1/3}$	$U_{10N}$	$1.27 \times 10^{-1}$	1.63	1.326	0.921
$D_{bp}^{1/3}$	$u_*$	$1.49 \times 10^1$	1.14	1.417	0.910
$D_{bp}^{1/3}$	$R_{B,m} = \frac{u_*^2}{v_w \omega_m}$	$1.07 \times 10^{-2}$	0.52	1.502	0.899
$D_{bp}^{1/3}$	$R_{B,p} = \frac{u_*^2}{v_w \omega_p}$	$1.12 \times 10^{-2}$	0.51	1.653	0.877
$D_{bp}^{1/3}$	$R_{Heq} = \frac{u_* H_{eq}}{v_w}$	$2.56 \times 10^{-3}$	0.61	1.894	0.839
$D_{bp}^{1/3}$	$R_{H_s} = \frac{u_* H_s}{v_w}$	$1.36 \times 10^{-3}$	0.60	1.986	0.823
$D_{bp}^{1/3}$	$R_{H_p} = \frac{u_* H_p}{v_w}$	$2.05 \times 10^{-3}$	0.59	2.139	0.794
$D_{bp}^{1/3}$	mss	$1.86 \times 10^4$	1.34	2.893	0.619
$D_{bp}^{1/3}$	mss/ $\Delta f$	$7.60 \times 10^2$	1.32	2.419	0.734
$D_{bp}^{1/3}$	mss/ $\Delta f \Delta \theta$	$3.35 \times 10^2$	1.37	2.911	0.614
$D_{bp}^{1/3}$	$H_p k_p / 2$	$9.06 \times 10^1$	0.88	4.055	0.251
$D_{bp}^{1/3}$	$H_s k_p / 2$	$6.33 \times 10^1$	0.83	4.027	0.262
$D_{bp}^{1/3}$	$H_{eq} k_m / 2$	$1.34 \times 10^4$	2.23	3.017	0.586
$D_{bp}^{1/3}$	$H_p k_m / 2$	$2.20 \times 10^3$	2.31	3.211	0.531
$D_{bp}^{1/3}$	$H_s k_m / 2$	$1.29 \times 10^3$	2.34	2.888	0.620
$D_{bp,v}^{1/3}$	$U_{10N}$	$3.78 \times 10^{-1}$	1.10	1.112	0.822
$D_{bp,v}^{1/3}$	$u_*$	$9.55 \times 10^0$	0.83	1.110	0.822
$D_{bp,v}^{1/3}$	$R_{B,m} = \frac{u_*^2}{v_w \omega_m}$	$5.09 \times 10^{-2}$	0.38	1.139	0.813
$D_{bp,v}^{1/3}$	$R_{B,p} = \frac{u_*^2}{v_w \omega_p}$	$4.88 \times 10^{-2}$	0.37	1.197	0.794
$D_{bp,v}^{1/3}$	$R_{Heq} = \frac{u_* H_{eq}}{v_w}$	$1.58 \times 10^{-2}$	0.45	1.290	0.760
$D_{bp,v}^{1/3}$	$R_{H_s} = \frac{u_* H_s}{v_w}$	$9.56 \times 10^{-3}$	0.45	1.318	0.750
$D_{bp,v}^{1/3}$	$R_{H_p} = \frac{u_* H_p}{v_w}$	$1.43 \times 10^{-2}$	0.43	1.383	0.725
$D_{bp,v}^{1/3}$	mss	$1.43 \times 10^3$	0.94	1.917	0.466
$D_{bp,v}^{1/3}$	mss/ $\Delta f$	$1.55 \times 10^2$	0.94	1.589	0.634
$D_{bp,v}^{1/3}$	mss/ $\Delta f \Delta \theta$	$8.62 \times 10^1$	0.96	1.839	0.509
$D_{bp,v}^{1/3}$	$H_p k_p / 2$	$2.63 \times 10^1$	0.50	2.334	0.209
$D_{bp,v}^{1/3}$	$H_s k_p / 2$	$2.11 \times 10^1$	0.46	2.341	0.205
$D_{bp,v}^{1/3}$	$H_{eq} k_m / 2$	$1.25 \times 10^3$	1.59	1.974	0.434
$D_{bp,v}^{1/3}$	$H_p k_m / 2$	$2.09 \times 10^2$	1.44	2.000	0.419
$D_{bp,v}^{1/3}$	$H_s k_m / 2$	$2.15 \times 10^2$	1.63	1.858	0.499

592 plume penetration depths with wave age in developing seas (i.e.,  $c_p/U_{10N} \approx 1.2$ ), shown in Fig-  
 593 ure A.10, is linked to the change in the dominant breaker type.

594 Further, our results show that the bubble plume penetration depths scaled by either  $H_s$  or  
 595  $L_m$  decrease monotonically with increasing wave age over the observed range of sea states in the  
 596 PAPA data set from developing to old seas. In particular, the data indicate that  $D_{bp}^{1/3}/H_s$  has a  
 597 linear relationship with the inverse of wave age, given by

$$\frac{D_{bp}^{1/3}}{H_s} = 2.42 \left[ \frac{c_p}{U_{10N}} \right]^{-0.96}, \quad (13)$$

598 with relatively small data scatter and  $r^2 = 0.77$  (solid line in Figure A.10a). Assuming an ap-  
 599 proximately linear relationship between  $U_{10N}$  and air friction velocity (Figure A.2b), the results  
 600 shown in Figures A.10a and A.10b and Eq. 13 are consistent with the corresponding results re-  
 601 ported in *Wang et al.* [2016].

### 602 3.3 Whitecap Coverage and Its Relation with Bubble Plume Depths

603 Existing parameterizations of oceanic whitecap coverage  $W$  have a general threshold power  
 604 law form of  $W = a(X-b)^n$ , where  $X$  is a selected predictive parameter (e.g.,  $U_{10N}$ ,  $u_*$ ,  $mss/\Delta f$ ,  $S$ ,  $R$ ,  $\dots$ ,  
 605 all defined in §2) and  $a$ ,  $b$  and  $n$  are empirical coefficients obtained from the best fit to a consid-  
 606 ered data set by minimizing the sum of the squares of the log residuals  $W_{res} = \log_{10} W - \log_{10} [a(X-$   
 607  $b)^n]$  to give equal weight to  $W$  data across several orders of magnitude. It is generally accepted  
 608 that several environmental conditions, including surfactants, salinity, wind fetch and duration,  
 609 wind history, surface shear, and rain, are responsible for data scatter in whitecap variability against  
 610 a typical predictive parameter  $X$ . However, these secondary effects on the corresponding mean  
 611  $W$  values are thought to be relatively small. Thus, we obtain the corresponding best fits over the  
 612 binned data as in §3.2 and similar to *Scanlon and Ward* [2016] and *Brumer et al.* [2017]. Bins  
 613 containing fewer than four bursts of data are not considered for data fitting as in §3.2.

614 Figures A.11a and A.11b show the variation of  $W$  against wind speed and air friction ve-  
 615 locity in the PAPA data and in the data set of *Schwendeman and Thomson* [2015a] as well as the  
 616 best fits to the binned PAPA data and several relevant least squares threshold power law fits from  
 617 the recent literature [*Sugihara et al.*, 2007; *Callaghan et al.*, 2008; *Schwendeman and Thomson*,  
 618 2015a; *Scanlon and Ward*, 2016; *Brumer et al.*, 2017]. Consistent with the recent literature, the  
 619 observed  $W(U_{10N})$  values are considerably smaller than those reported in pioneering  $W$  stud-  
 620 ies [e.g., *Monahan and Muircheartaigh*, 1980] using a manual  $W$  extraction method [*Monahan*,  
 621 1969]. Further, the observed range of  $W(U_{10N})$  and  $W(u_*)$  values and their associated data scat-  
 622 ter are consistent with the recent studies in which their experimental methods are comparable to  
 623 those used here (see §2.4).

624 Figure A.11a shows that the observed  $W(U_{10N})$  values and their corresponding best fits  
 625 at high winds are considerably comparable with those in the other data sets, especially those that  
 626 include  $W$  observations at  $U_{10N} > 16 \text{ ms}^{-1}$ ; the solid line section of each fit shown in Figure A.11  
 627 represents the range of data to which the best fit is obtained. The fits, however, tend to diverge  
 628 for  $U_{10N} < 10 \text{ ms}^{-1}$ . We note that the shape of a threshold power law fit at low and moderate  
 629 winds and, in particular, the coefficient  $b$  (which incorporates the threshold behavior of the fit)  
 630 are sensitive to the data at the lower range of  $X$  values. Thus, any systematic bias at the selected  
 631 wind parameter at low wind speeds will be translated into the resulting best fit. Wind speeds in  
 632 several previous studies were not corrected for atmospheric stability, e.g., *Sugihara et al.* [2007]  
 633 and *Schwendeman and Thomson* [2015a], or  $U_{10}^{PL}$  was simply used as a proxy for  $U_{10N}$ , e.g., *Callaghan*  
 634 *et al.* [2008]. As discussed in §2.2, although these simplifications have a relatively small effect  
 635 on estimated wind speeds at high winds, they can lead to considerable errors in estimated wind  
 636 parameters at low winds.

637 Our observations shown in Figures A.11a and A.11b demonstrate that the observed  $W(U_{10N})$   
 638 and  $W(u_*)$  values at rapidly decreasing ( $dU_{10N}/dt \ll 0$ ), low winds ( $U_{10N} < 4 \text{ ms}^{-1}$  or  $u_* <$

639 0.2 ms<sup>-1</sup>) vary between 10<sup>-4</sup> and 2×10<sup>-3</sup> while the best wind-speed-only or  $u_*$ -only fits obtained  
 640 from the remaining of the data points predict no whitecapping ( $W = 0$ ) at those wind condi-  
 641 tions. This suggests that strong wind history could also result in a systematic bias in  $W(U_{10N})$   
 642 and  $W(u_*)$  data at low winds, and thus, may be responsible for a portion of an apparent diver-  
 643 gence in existing wind-speed-only and  $u_*$ -only fits at low and moderate winds.

644 Figures A.11a and A.11b also show that at the same wind forcing, either represented by  
 645  $U_{10N}$  or  $u_*$ , a large portion of  $W$  values in the PAPA data at increasing ( $dU_{10N}/dt > 0$ ) and de-  
 646 creasing ( $dU_{10N}/dt < 0$ ) wind speeds tend to be smaller and greater than the corresponding  
 647 mean  $W$  values provided by the corresponding best fits, respectively. This trend is consistent with  
 648 the observations of *Callaghan et al.* [2008] for wind speeds above approximately 9 ms<sup>-1</sup>. How-  
 649 ever, in contrast to *Callaghan et al.* [2008], our observations clearly show that the same trend ex-  
 650 ists for moderate and low winds if the magnitude of  $dU_{10N}/dt$  is sufficiently large.

651 Next, we examine the predictive skill of several wind and wave parameters for the observed  
 652 range of  $W$  values in the PAPA data similar to the method described in §3.2 but in log<sub>10</sub> space;  
 653 i.e., we evaluate the overall quality of the fits using Eqs. 11 and 12 with  $W_{res,i} = \log_{10} W_i -$   
 654  $\log_{10}[a(\mathcal{X}_i - b)^n]$ . Here RMSE indicates an average order of magnitude deviation from the fit,  
 655 and  $r^2$  indicates the proportion of the observed log<sub>10</sub>  $W$  variability that is predictable from the  
 656  $\mathcal{X}$  parameter. Negative  $r^2$  indicates that the fit performs worse than a horizontal line at the mean  
 657 of the data. As in §3.2, all the fits are obtained from the binned data for  $U_{10N} \geq 6$  ms<sup>-1</sup>. The  
 658 fit statistics are obtained using the individual 10-minute average data points,  $W_i$  ( $i = 1, \dots, N$ ),  
 659 with three conditions: all data ( $N = 165$ ),  $U_{10N} \geq 6$  ms<sup>-1</sup> ( $N = 144$ ), and  $|dU_{10N}/dt| < 2$   
 660 ms<sup>-1</sup>hr<sup>-1</sup> ( $N = 126$ ).

Table 2: Parameterizations of whitecap coverage represented by the best fits with a threshold power law form  $W = a(\mathcal{X} - b)^n$  as a function of several wind and wave parameters  $\mathcal{X}$  to the binned PAPA data for  $U_{10N} \geq 6$  ms<sup>-1</sup>. The statistics of each fit are also calculated for three conditions. The fits and their statistics are computed in log space.

Predictor $\mathcal{X}$	Results of the best fit $W = a(\mathcal{X} - b)^n$			Statistics of the best fit with conditions:					
	$a$	$b$	$n$	$U_{10N} \geq 6$ ms <sup>-1</sup> RMSE	$r^2$	$ dU_{10N}/dt  < 2$ ms <sup>-1</sup> /hr RMSE	$r^2$	all data RMSE	$r^2$
$U_{10N}$	$2.06 \times 10^{-5}$	3.89	2.65	0.412	0.70	0.471	0.60	0.752	0.05
$u_*$	$3.63 \times 10^{-2}$	0.18	2.00	0.394	0.72	0.476	0.59	0.698	0.18
$R_{B,m} = \frac{u_*^2}{v_w \omega_m}$	$3.87 \times 10^{-9}$	$5.81 \times 10^4$	1.14	0.400	0.72	0.646	0.25	0.935	-0.47
$R_{B,p} = \frac{u_*^2}{v_w \omega_p}$	$3.86 \times 10^{-9}$	$7.01 \times 10^4$	1.12	0.424	0.68	0.657	0.22	0.916	-0.41
$R_{Heq} = \frac{u_* H_{eq}}{v_w}$	$3.02 \times 10^{-10}$	$1.50 \times 10^5$	1.31	0.428	0.68	0.415	0.69	0.645	0.30
$R_{H_s} = \frac{u_* H_s}{v_w}$	$2.45 \times 10^{-10}$	$5.07 \times 10^5$	1.23	0.456	0.63	0.434	0.66	0.692	0.20
$R_{H_p} = \frac{u_* H_p}{v_w}$	$1.64 \times 10^{-9}$	$4.05 \times 10^5$	1.12	0.590	0.38	0.589	0.37	0.801	-0.08
mss	$6.50 \times 10^6$	–	3.60	0.565	0.43	0.557	0.44	0.572	0.44
mss/ $\Delta f$	$1.61 \times 10^2$	$6.23 \times 10^{-3}$	2.79	0.487	0.58	0.482	0.58	0.512	0.55
mss/ $\Delta f \Delta \theta$	4.79	$1.72 \times 10^{-2}$	2.16	0.537	0.49	0.534	0.49	0.557	0.47
$H_p k_p / 2$	4.85	–	2.33	0.737	0.03	0.520	0.06	0.778	-0.04
$H_s k_p / 2$	$2.06 \times 10^{-1}$	$3.86 \times 10^{-2}$	0.99	0.766	-0.05	0.795	-0.14	0.837	-0.20
$H_{eq} k_m / 2$	$1.89 \times 10^7$	–	6.58	0.564	0.43	0.550	0.46	0.576	0.43
$H_p k_m / 2$	$3.80 \times 10^2$	$3.12 \times 10^{-2}$	3.87	0.547	0.46	0.550	0.46	0.552	0.48
$H_s k_m / 2$	$5.53 \times 10^2$	$4.56 \times 10^{-2}$	4.27	0.507	0.54	0.502	0.54	0.503	0.56

661 Table 2 summarizes the coefficients ( $a$ ,  $b$ , and  $n$ ) and statistics of the best fits,  $W = a(\mathcal{X} -$   
 662  $b)^n$ , for several predictive parameters  $\mathcal{X}$  to the PAPA data. Of all the considered predictors for  
 663  $W$  at moderate and high winds,  $u_*$  gives the best fit ( $r^2 = 0.72$ , RMSE = 0.394), which is only  
 664 slightly better than the  $U_{10N}$  fit ( $r^2 = 0.70$ , RMSE = 0.412). Our results show that the quality  
 665 of the fits obtained from various forms of the predictors  $R_H$  (Eq. 3) and  $R_B$  (Eq. 4), which com-  
 666 bine  $u_*$  and a characteristic scale of the wave field, are similar or less than the  $u_*$ -only fit. These  
 667 parameterizations cannot reasonably predict  $W$  at rapidly varying wind speeds (i.e., large wind  
 668 accelerations).

669 Our observations shown in Figure A.2 indicate that either the normalized or unnormalized  
 670 equilibrium range  $mss$  values at increasing winds are smaller than those in decreasing winds at  
 671 a given wind speed. This suggests that these spectral parameters might reflect both wind forc-  
 672 ing and wind history effects. Consistently, the results summarized in Table 2 document that the  
 673 parameterizations based on the equilibrium range  $mss$  have similar skill across all sea state con-  
 674 ditions, including those with large wind accelerations. The results indicate that the equilibrium  
 675 range  $mss/\Delta f$  (Figure A.11c) better predicts the observed  $W$  variability compared to the other  
 676 spectral predictors considered here. Of the bulk steepness predictors,  $H_s k_m/2$  has the highest  
 677 skill. Overall, for each type of the predictors considered here, those that contain either the peak  
 678 wave height, peak wave number, or peak wave period show the least skill (Figure A.11d).

679 Figure A.11 shows that the observed  $W(U_{10N})$ ,  $W(u_*)$ , and  $W(mss/\Delta f)$  values in the PAPA  
 680 data at moderate winds (e.g.,  $8 \text{ ms}^{-1} \leq U_{10N} \leq 16 \text{ ms}^{-1}$ ) are generally smaller than in the *Schwen-*  
 681 *deman and Thomson* [2015a] data set. We note that a significant portion of the data at these wind  
 682 speeds was collected in the presence of rain (Figure A.1b). This observation may suggest that  
 683 whitecap activity is suppressed in the presence of rain. Detailed quantification of rain effects on  
 684  $W$  requires rain rates (not measured here) and remains unknown.

685 Finally, Figure A.12 shows that the mean and significant bubble plume penetration depths  
 686 are, on average, correlated and have a nonlinear relation with whitecap coverage given by

$$\overline{D}_{bp} = 29.5 W^{0.33}, \quad D_{bp}^{1/3} = 52.8 W^{0.36}, \quad (14)$$

687 with  $r^2 = 0.60$  (the fit in Figure A.12a) and  $r^2 = 0.62$  (the fit in Figure A.12c), and

$$\overline{D}_{bp,v} = 12.6 W^{0.19}, \quad D_{bp,v}^{1/3} = 21.9 W^{0.24}, \quad (15)$$

688 with  $r^2 = 0.33$  (the fit in Figure A.12b) and  $r^2 = 0.43$  (the fit in Figure A.12d). Here both fits  
 689 are obtained using the binned data as a function of  $U_{10N}$ ; as before, the data with  $U_{10N} < 6 \text{ ms}^{-1}$   
 690 are not considered for data fitting. As explained in detail in §2.5 and consistent with the obser-  
 691 vations shown in §3.1 and §3.2,  $D_{bp,v}$  represents penetration depth of bubbles that have, on av-  
 692 erage, at least two orders of magnitude more void fraction and significantly more visible optical  
 693 signature than those reach to  $D_{bp}$  for a given sea state condition.

694 Assuming a wave field with narrow-banded breaking waves,  $W$  increases approximately  
 695 linearly with increasing rates of breaking waves while the statistics of bubble plume depths are  
 696 not sensitive to breaking rates, and thus they would not be correlated with  $W$ . However, wave break-  
 697 ing typically occurs at a range of scales, and increasing  $W$  results from increasing both the rate  
 698 and scale of breaking waves. This may partially explain the observed relationship between bub-  
 699 ble plume depths and  $W$  shown in Figure A.12, that is, the plume depths increase, on average,  
 700 with increasing  $W$  but at a much lower rate, i.e., the exponents in Eqs. 14 and 15 are positive but  
 701 much less than 1.

#### 702 4 Discussion: Bubble Plumes Volumes

703 Here we define the volume of bubble plumes as a measure of their overall size rather than  
 704 the total volume of bubbles they contain, with the bubble plumes that are identified as regions  
 705 in which volume backscattering strength (somewhat related to bubble void fractions, see §2.5)

706 is above a threshold value. That said, the volume of bubble plumes per unit sea surface area is  
 707 given by

$$\mathcal{V}_{bp} = \mathcal{A}_{bp} \bar{D}_{bp}, \quad \text{and} \quad \mathcal{V}_{bp,v} = \mathcal{A}_{bp,v} \bar{D}_{bp,v}, \quad (16)$$

708 where  $\mathcal{A}$  is the fractional surface area of bubble plumes,  $\bar{D}$  is the mean penetration depth of bub-  
 709 bles within the plumes, and the subscripts  $bp$  and  $bp, v$  denote the statistics corresponding the  
 710 bubble plumes obtained from our bubble detection methods BDM1 and BDM2 (described in §2.5),  
 711 respectively. As discussed in detail in §2.5,  $\bar{D}_{bp,v}$  represents mean penetration depth of bubbles  
 712 where the volume backscattering is at least 20 dB higher at  $\bar{D}_{bp}$  for a given sea state condition  
 713 and note this is expected to reflect comparable increasing in void fraction. Our observations and  
 714 several simple parameterizations of the mean plume depths  $\bar{D}_{bp}$  and  $\bar{D}_{bp,v}$  are presented in §3.

715 We note that  $\mathcal{A}$  represents the fractional surface area, with or without visible surface sig-  
 716 nature, of bubble plumes that significantly exceed the persistence of visible surface foam gen-  
 717 erated during active breaking as discussed in §3.1. Thus both  $\mathcal{A}_{bp}$  and  $\mathcal{A}_{bp,v}$  are expected to  
 718 be noticeably greater than the measured whitecap coverage  $W$ . However,  $\mathcal{A}_{bp}$  and  $\mathcal{A}_{bp,v}$  can  
 719 not be directly quantified from our sampling method. In the following, we introduce a proxy for  
 720  $\mathcal{A}$  and comment on its relation to  $W$ .

721 We define  $P$  as a time fraction of echogram data over concurrent bursts during which bub-  
 722 ble plumes are detected. Assuming the buoys had an approximately constant "wind slip" veloc-  
 723 ity  $U_{slip}$  during each burst,  $A = P^2$  then provides a proxy for  $\mathcal{A}$  if the drifting distance of the  
 724 buoy relative to the surface water  $\approx U_{slip} T_{burst}$  is much greater than the average horizontal length  
 725 of the bubble clouds  $\approx U_{slip} T_{ab}$  or  $U_{slip} T_{ab,v}$  (see §3.1). Further, at least a few bubble clouds  
 726 should be available in a burst to consider that  $\mathcal{A} \approx A$ .

727 Figure A.13a shows the  $A_{bp}$  and  $A_{bp,v}$  values as a function of  $U_{10N}$  where the size of the  
 728 symbols is a function of the number of the bubble clouds detected in a burst, averaged over con-  
 729 current bursts,  $N$ , with  $0.67 \leq N_{bp} \leq 26$  and  $0.5 \leq N_{bp,v} \leq 24$ . Note that  $P$ , and thus  $A =$   
 730  $P^2$ , values that approach one indicate that either the main portion of the surface layer is covered  
 731 by bubble plumes or the net drifting distance of the buoy (relative to the surface water) is smaller  
 732 than the horizontal length of the sampled bubble cloud. As shown in Figure A.4b and A.13a, the  
 733 latter may explain  $A_{bp} \sim 1$  at moderate winds where  $N < 2$  and  $T_{ab}$  values are on the order  
 734 of several hundreds of seconds (comparable to  $T_{burst} = 512$ s). Despite the uncertainties in the  
 735 interpretation of  $A$ , the observations shown in Figure A.13a suggest that  $A_{bp}$  is several times greater  
 736 than  $A_{bp,v}$ , which is qualitatively consistent with the continuous increase of the overall size of  
 737 the bubble plume shown in Figure A.5 and the corresponding residence time results shown in Fig-  
 738 ure A.4b.

739 Figure A.13b shows that both  $A_{bp}$  and  $A_{bp,v}$  are, on average, increase as a function of  $W$   
 740 as

$$A_{bp} = 2.5 W^{0.33} \leq 1, \quad \text{and} \quad A_{bp,v} = 8.4 W^{0.97} \leq 1. \quad (17)$$

741 Note that the data points with  $N < 3$  are neglected in Figure A.13b. Our observations show that  
 742  $A_{bp}$ , which is comparable to a fractional surface area defined in *Thorpe* [1986], is at least an or-  
 743 der of magnitude larger than  $W$ . This is consistent with the semi-empirical plume area analysis  
 744 of *Thorpe* [1986].

745 Finally by substituting Eqs. 14, 15, and 17 into Eq. 16, we obtain

$$\mathcal{V}_{bp} = \mathcal{A}_{bp} \bar{D}_{bp} \approx 74 W^{0.66} \leq 29.5 W^{0.33} \quad [\text{m}^3/\text{m}^2], \quad (18)$$

746 and

$$\mathcal{V}_{bp,v} = \mathcal{A}_{bp,v} \bar{D}_{bp,v} \approx 106 W^{1.16} \leq 12.6 W^{0.19} \quad [\text{m}^3/\text{m}^2], \quad (19)$$

747 assuming that the best fits to the binned data shown in Figure A.13b (Eq. 17) provide a proxy for  
 748  $\mathcal{A}_{bp}$  and  $\mathcal{A}_{bp,v}$ .

749 We emphasize that uncertainty in our estimates of the fractional surface area of bubble plumes  
 750 (and thus plume volumes) increases with decreasing  $W$ , especially at low  $W$  values (e.g.,  $W <$

751  $10^{-3}$ ) because of increasing effect of sparse sampling of intermittent breaking crests on the re-  
 752 sulting statistics [Derakhti *et al.*, 2020a].

## 753 5 Summary

754 The observational results presented here quantify the statistics of penetration depth and frac-  
 755 tional surface area of bubble plumes generated by breaking surface waves as a function of var-  
 756 ious wind and sea state parameters over a wide range of sea state conditions. Bubble plume data  
 757 include concurrent high-resolution (with a 12 min temporal resolution) plume depth statistics and  
 758 whitecap coverage. The former is obtained from the echogram data with 1 cm vertical resolu-  
 759 tion, collected by downward-looking echosounders mounted on arrays of freely drifting SWIFT  
 760 buoys. The latter is obtained from visual images, collected by shipboard cameras operated near  
 761 the buoys.

762 Our observations indicate that the statistics of bubble plume penetration depths are well-  
 763 correlated with wind speed, spectral wave steepness, and whitecap coverage. Results show that  
 764 the mean plume depths exceed 10 m beneath the surface at high winds, with individual bubble  
 765 clouds reach to depths of more than 30 m. Mean plume depths vary, on average, from 1.6 to 2.8  
 766 times wind sea significant wave height  $H_s^{ws}$ . Plume depths scaled by either  $H_s^{ws}$  or total signif-  
 767 icant wave height  $H_s$  vary non-monotonically with increasing wind speeds. Plume depths scaled  
 768 by  $H_s$  are strongly linearly correlated with the inverse of wave age from developing to old seas.  
 769 All scaled plume depths considered here are decreasing functions of wave age in developing seas.  
 770 We successfully provide multiple parameterizations that predict the observed variability of the  
 771 penetration depth and surface area of bubble plumes as a function of simple wind and wave statis-  
 772 tics available from existing forecast models or typical ocean buoys.

773 This study is the first to provide a direct relation between bubble plume penetration depth  
 774 and whitecap coverage, indicating that the penetration depth of bubble plumes is correlated with  
 775 their visible surface area. This result is significant as it advocates the possibility of estimating  
 776 the volume of bubble plumes by remote sensing. This also significantly expands the applicabil-  
 777 ity of the recent theoretical framework introduced by *Callaghan* [2018] on predicting total wave  
 778 breaking dissipation as a function of bubble plume penetration depth and whitecap coverage. In  
 779 a companion paper, we examine dynamic relationships between the bubble plume statistics pre-  
 780 sented here and total wave breaking dissipation using our synchronized observations of bubble  
 781 plumes and dissipation rates.

782 Finally, the parameterizations of bubble plume penetration depth provided in this study may  
 783 also be used for estimating effective vertical transport of other particles, with a rising velocity  
 784 on the order of few  $\text{cm s}^{-1}$  or less, by breaking surface waves. It is possible that the drifting SWIFT  
 785 buoys used in this study aggregate in convergence zones with enhanced downwelling velocities,  
 786 such that there would be a sampling bias in the interpretation of vertical transport [Zippel *et al.*,  
 787 2020]. However, no obvious convergence zones, windrows, or other organized surface fronts were  
 788 observed during the PAPA data collection. Furthermore, the wind slip (1% of wind speed) of the  
 789 buoys tends to cause a quasi-uniform sampling along a drift track even in the presence of surface  
 790 features.

## 791 Open Research

792 SWIFT data are available from [www.apl.washington.edu/swift](http://www.apl.washington.edu/swift), and whitecap coverage data  
 793 are available from <http://hdl.handle.net/1773/48143>.

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## 800 **A: Echosounder calibration**

801 The echosounder was calibrated using standard sphere calibration techniques *Demer et al.*  
802 [2015]. In this approach, a sphere of a known material is suspended below the beam of an echosounder.  
803 Since the sphere's properties are known, an analytical solution for the acoustic target strength can  
804 be calculated. The difference between the measured intensity of the scattering and the known scat-  
805 tering from the sphere at the transmit frequency is the total gain for the system. In post-cruise  
806 testing, a 38.1 mm diameter tungsten-carbide sphere with 6% cobalt binder was suspended 8 m  
807 below the transducers by a bridle connected to the hull of the SWIFTS. The units were then de-  
808 ployed for 30-60 minutes on Lake Washington (Washington, USA), during which the attitude of  
809 the SWIFTS caused the suspended sphere to pass through the beam of the echosounder. The top  
810 1% of targets at the sphere range, which are assumed to be those associated with the sphere be-  
811 ing on-axis within the beam where the combined transmit-receive beam pattern is highest, were  
812 then selected. The gain is then determined by solving for  $G_{cal}$  in the target strength equation us-  
813 ing the known analytical solution for the target strength of the sphere.

814 In practice, a sphere is sized such that its scattering response contains no significant nulls  
815 within the bandwidth [*Demer et al.*, 2015; *Stanton and Chu*, 2008; *Lavery et al.*, 2017]. How-  
816 ever, this is not feasible at 1 MHz since a tiny  $< 1\text{cm}$  sphere would be required. Furthermore,  
817 for such a small sphere, the strands securing the sphere would contribute significantly to scatter-  
818 ing, bias the results [*Renfree et al.*, 2020]. Thus, we chose to use a larger sphere whose response  
819 is quite complex over the relevant frequency range. The pulse-compressed signal has sufficient  
820 bandwidth to clearly resolve the echo from the front interface and subsequent contributions from  
821 circumference waves. We, therefore, assumed that the peak of the pulse compressed signal rep-  
822 represents the partial wave scattering cross-section of the sphere [*Stanton and Chu*, 2008]. This as-  
823 sumption is necessary given that a frequency-dependent calibration cannot be performed given  
824 the only output data product is a scattering intensity measurement representing the average within  
825 the range bin output by the ADCP.

826 At the time of this experiment, the firmware resulted in scattering that saturated the receiver  
827 in the high gain setting and saturated the receiver when using the calibration sphere at a range  
828 of  $\sim 8$  m. There is, therefore, some uncertainty in the calibration gains and the field observations.  
829 We cannot conclusively state the magnitude of this uncertainty, but it is believed to be on the or-  
830 der of a few dB or less from the calibration gain. The justification for this statement is that the  
831 elastic response of the sphere is well resolved with the intensity (impulse response squared) of  
832 the signal from the first Rayleigh wave, approximately 9 dB smaller than the echo from the front  
833 interface of the sphere when the calibrations were performed at the lower gain setting. This is  
834 consistent with expectations based on the impulse response of a 38.1 mm tungsten carbide sphere  
835 [*Demer et al.*, 2015] and the arrival of the signal associated with the first Rayleigh wave. In the  
836 saturated data, the difference in intensity between the first Rayleigh wave and the saturated echo  
837 from the front interface was approximately 3 dB. Given the impulse response of the 38.1 mm sphere,  
838 this suggests that about 6 dB of scattering from the sphere had been clipped. When used in the  
839 high power setting, gains were applied assuming the clipped value was 6 dB. The practical ef-  
840 fect of this uncertainty is to put consistent error bars on the volume scattering coefficients mea-  
841 sured in the data. That is, all data are shifted similarly, making the absolute intensity of the backscat-  
842 tering more uncertain without impacting the relevant ranges between the thresholds.

843 The fact that scattering from the tungsten carbide sphere saturated at 8 m indicates the high  
844 gain setting almost certainly caused widespread saturation of signals in the upper portion ( $\sim 10$   
845 m) of the water column when high densities of bubbles were present. A consequence of this is  
846 that the full dynamic range of volume backscattering is not resolved. Despite these challenges  
847 and uncertainties, we consider it preferable to present backscattering intensities in this approach

848 to backscattering intensities expressed in decibels with reference value ground in physical mea-  
849 surements.

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