

1 **Point-source moment tensor inversion via a Bayesian hierarchical**
2 **inversion with 2D-structure uncertainty: Implications for the 2009-2017**
3 **DPRK nuclear tests**

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9 **Key Points:**

- 10 • A new seismic moment tensor inversion with Bayesian approach incorporates 2D
11 structural uncertainty along specific source-station paths.
- 12 • Effective affine-invariant ensemble samplers mitigate the ISO-CLVD tradeoff that
13 impedes resolving shallow explosive sources.
- 14 • The newly developed inversion method reveals similar explosive-source mechanisms of
15 five DPRK underground nuclear explosions.
16

17 **Abstract**

18 Determining the seismic moment tensors (MT) from the observed waveforms, known as full-
19 waveform seismic MT inversion, remains challenging for small to moderate-size earthquakes at
20 regional scales. Firstly, there is an intrinsic difficulty due to a tradeoff between the isotropic
21 (ISO) and compensated linear vector dipole (CLVD) components of MT that impedes resolving
22 shallow explosive sources, e.g., underground nuclear explosions. It is caused by the similarity of
23 long-period waveforms radiated by ISO and CLVD at regional distances. Secondly, regional
24 scales usually bear complex geologic structures; thus, inaccurate knowledge of Earth's structure
25 should be considered a theoretical error in the MT inversion. However, this has been a
26 challenging problem. So far, only the uncertainty of the 1D Earth model (1D structural error),
27 apart from data errors, has been explored in the source studies. Here, we utilize a hierarchical
28 Bayesian MT inversion to address the above problems. Our approach takes advantage of affine-
29 invariant ensemble samplers to explore the ISO-CLVD tradeoff space thoroughly and
30 effectively. Furthermore, we invert for station-specific time shifts to treat the structural errors
31 along specific source-station paths (2D structural errors). We present synthetic experiments
32 demonstrating the method's advantage in resolving the ISO components. The application to
33 nuclear explosions conducted by the Democratic People's Republic of Korea (DPRK) shows
34 highly similar source mechanisms, dominated by a high ISO, significant CLVD components, and
35 a small DC component. The recovered station-specific time shifts from the nuclear explosions
36 present a consistent pattern, which agrees well with the geological setting surrounding the event
37 location.

38 **Plain Language Summary**

39 The seismic sources, including underground faulting, volcanic processes, and manufactured
40 underground explosions, can be represented by a point-source moment tensor (MT), which is an
41 equivalent force system at a point in space and time. Inferring the seismic MT from the observed
42 seismic waveforms is an MT inverse problem. This study designs a new Bayesian inference
43 method to solve this inverse problem by considering two challenging issues: (a) estimating the
44 uncertainty for theory error due to the assumption of 1D Earth's model for the true 3D Earth, and
45 (b) mitigating the theoretical tradeoff between nondouble couple source types at a shallow depth.
46 Here, we determine the MTs of five underground nuclear explosions conducted by the
47 Democratic People's Republic of Korea (DPRK) by fixing their sources at a realistic burial depth
48 of 0.5 km. The robustness of these MT solutions is demonstrated through a series of simulation
49 experiments. Comparisons with previous studies reveal a typical explosive nature of the
50 manmade seismic sources. The recovered theory error is consistent among five explosions,
51 providing a meaningful interpretation of the regional geological setting.

52 **1 Introduction**

53 The seismic moment tensor (MT, a symmetric 3×3 matrix) is a generalized mathematical
54 representation for various seismic sources, including tectonic earthquakes and non-tectonic
55 events, such as manufactured underground explosions and volcanic processes, including
56 eruptions. The point source assumption must hold to use MT, which is generally valid for small-
57 to-medium-size earthquakes (Aki & Richards, 2002). The seismic MT introduces source
58 components beyond a double-couple (DC) force system, which only describes slip on a planar
59 fault (Gilbert, 1971). One convenient way is to decompose an MT into double-couple (DC) and

60 non-double-couple (NDC) components consisting of isotropic (ISO) and compensated linear
 61 vector dipole (CLVD) components, which was proposed by Knopoff and Randall (1970), then
 62 further developed by others (e.g., Jost & Herrmann, 1989; Julian et al., 1998; Sipkin, 1986;
 63 Vavryčuk, 2015). This decomposition of MT has specific physical properties. DC part depicts
 64 the shear faulting, which is the focal mechanism of most tectonic earthquakes. The ISO
 65 represents the explosion/collapse and involves volumetric changes. Even though an MT only
 66 including a pure CLVD does not correspond to any simple seismic sources, its combination with
 67 ISO can explain the tensile or compressive faulting (Vavryčuk, 2001, 2011, 2015). Besides,
 68 shear faulting on a non-planar fault can be represented by the combination of DC and CLVD,
 69 referred to as deviatoric MT, assuming zero ISO. A ring fault was proposed to explain the
 70 teleseismic and regional long-period waveforms of the 1996 Bárðarbunga earthquake (e.g.,
 71 Konstantinou et al., 2003; Nettles & Ekström, 1998; Tkalčić et al., 2009).

72 The NDC sources have been found in various geologic settings. At the early stage of
 73 seismology, some minor departures from the DC mechanism were considered artifacts of the
 74 inversion, e.g., data noise or theory error. As the instruments and methods are developed, the
 75 NDC components are confirmed to correspond to the source processes. They are found in
 76 various geological settings but are most common in volcanic environments (e.g., Dreger et al.,
 77 2000; Duputel & Rivera, 2019; Julian, 1983; Mustać & Tkalčić, 2016; Nettles & Ekström, 1998;
 78 Saraò et al., 2001; Tkalčić et al., 2009), and geothermal environments (e.g., Johnson, 2014;
 79 Martínez-Garzón et al., 2017; Mustać et al., 2018; Mustać & Tkalčić, 2017; Ross et al., 1996),
 80 and underground explosions (e.g., Alvizuri et al., 2018; Chiang et al., 2014; Dreger et al., 2021;
 81 Ford et al., 2009; Mustać et al., 2020). Julian et al. (1998) and Miller et al. (1998)
 82 comprehensively reviewed the NDC sources in theory and applications. The relative significance
 83 of the NDC component is a critical indicator in discriminating between tectonic earthquakes and
 84 non-tectonic events (e.g., volcanic or explosive events). Therefore, the resolvability of MT,
 85 especially the NDC components, plays an essential role in seismic source studies, which relies on
 86 the seismic MT inversion.

87 Utilizing seismological observations to determine the MT comprises a recurring and
 88 broad central theme of modern seismology, which refers to seismic MT inversion. There are four
 89 groups of MT inversion methods based on the used observations. The first group of MT
 90 inversion uses the P-wave first motion polarities recorded at various directions to determine the
 91 fault geometry, i.e., the focal mechanism (e.g., Dillinger et al., 1972; Eaton & Mahani, 2015;
 92 Hardebeck, 2002; Julian, 1986; Reasenberg & Oppenheimer, 1985). The second group fits P-
 93 and S-wave amplitude or their ratio. For example, the absolute P and S amplitudes were used by
 94 Ebel and Bonjer (1990), Rögnvaldsson and Slunga (1993), and Stanek et al. (2014). The third
 95 group of MT inversion uses hybrids of various observations, including the first-motion polarity
 96 and amplitude ratios (e.g., Julian & Foulger, 1996; Shang & Tkalčić, 2020). The fourth group
 97 takes advantage of the full waveforms, which contain much more information than the body-
 98 wave polarity and amplitude ratio. However, it can be readily applied only to $M_w > 4.0$
 99 earthquakes. Based on the different implementations, it is divided into two main categories: The
 100 time-domain full-waveform MT inversion (e.g., Dreger et al., 2000; Dziewonski et al., 1981;
 101 Minson & Dreger, 2008; Pasyanos et al., 1996; Romanowicz et al., 1993), and the frequency-
 102 domain full-waveform MT inversion (e.g., Cesca et al., 2006; Dahm et al., 1999; Nakano et al.,
 103 2008; Romanowicz, 1982; Stump & Johnson, 1977). Cesca et al. (2010) and Vavryčuk and
 104 Kühn (2012) combined the time and frequency domain inversions. Future discussions about the

105 advantages and disadvantages of each method and their categories can be found in Shang and
106 Tkalčić (2020).

107 Rigorous uncertainty estimate has been one of the frontiers in seismic MT inversion. A
108 complete uncertainty treatment should consider both data noise mainly involved in the data
109 acquisition/processing and theoretical error primarily caused by the imperfect knowledge of
110 Earth's structure (i.e., structural error). Data noise has been estimated with different noise
111 models, such as a Gaussian or an exponentially decaying noise model (e.g., Bodin et al., 2012;
112 Duputel et al., 2012), empirical noise model from data residuals (e.g., Dettmer et al., 2007;
113 Mustać et al., 2020), from synthetic noise series (e.g., Gouveia & Scales, 1998; Piana Agostinetti
114 & Malinverno, 2010; Sambridge, 1999), or model with approximating the pre-event ambient
115 noise with two-attenuated cosine functions (Mustać et al., 2018; Mustać & Tkalčić, 2016).
116 Incorporating structural uncertainty has been conducted in the case of 1D Earth's structure by
117 assuming a Gaussian noise distribution for teleseismic Green's functions (Yagi & Fukahata,
118 2011), by estimating a covariance matrix from linear perturbation of Green's functions (Duputel
119 et al., 2014), or evaluating a covariance matrix from synthetically generated Green's functions
120 with randomly perturbed Earth's models (e.g., Hallo & Gallovič, 2016). These studies made
121 remarkable efforts to handle data noise and theoretical error separately. Recent advancements
122 treating data noise and theoretical errors jointly have been made. Vasyura-Bathke et al. (2021)
123 analyzed different combinations of covariance matrixes for data noise and structural uncertainty.
124 Pham and Tkalčić (2021) constructed a combined covariance matrix for data noise and structural
125 error. Namely, an explicit covariance matrix of structural error is obtained by the Monte Carlo
126 method from linear perturbations of the 1D-Earth model. These works provide a pathway to
127 estimating 1D structural error considering the overall structural effect averaged for all stations.

128 Constraining the source parameters better relies on possessing the accurate Earth
129 structure model. The MT inversion using the 1D Earth model has earned many successes by
130 using long-period waveforms, which are not sensitive to the small-size 3D heterogeneity (e.g.,
131 Dziewonski et al., 1981; Ekström et al., 2012). Moreover, the MT inversion has been advanced
132 further by incorporating the 1D Earth structural uncertainty, as discussed above. At the same
133 time, we recognize that an accurate knowledge of 3D anisotropic, heterogeneous Earth would
134 constrain source parameters significantly better. Multiple studies have addressed this issue,
135 concluding that the 3D Earth model can improve the source resolvability (e.g., Donner et al.,
136 2020; Fichtner & Tkalčić, 2010; Gallovič et al., 2010; Hejrani et al., 2017; Hingee et al., 2011;
137 Kim et al., 2011; Wang & Zhan, 2020). However, due to high computational demand, treating
138 uncertainty from the imperfection of 3D Earth structures (3D structural error) remains
139 challenging. Therefore, in this study, we explore a transitional solution before progressing the
140 uncertainty quantification from 1D to 3D structural errors.

141 Apart from the above aspect, an inherent ambiguity of the NDC components exists in
142 seismic source inversion for shallow sources. The resolvability of MT becomes more difficult as
143 the point-source focus becomes shallower (Dziewonski et al., 1981; Kanamori & Given, 1982;
144 Kawakatsu, 1996). Hejrani & Tkalčić (2020) analyzed two main challenges in conjunction with
145 the shallow-source inversion: an unbalanced range of amplitudes from a vertical dip-slip
146 mechanism in various frequency bands and the tradeoff between ISO and CLVD. They
147 addressed the first problem by utilizing high-frequency waveforms (>0.025 Hz), a possible
148 approach for a relatively simple geologic setting. However, the intrinsic difficulty in analyzing
149 shallow explosive sources such as underground nuclear explosions remains due to the similarity

150 of long-period waveforms at regional distances. Unless short periods (high frequencies) can be
151 utilized, many different MTs can fit the regional observed waveforms equally well, leading to
152 considerable uncertainty in MT solutions. Even though the problem can be mitigated by extra
153 constraints such as adding the first motion polarities of the teleseismic P-waves (e.g., Chiang et
154 al., 2014; Dreger et al., 2021; Ford et al., 2012), there is still an urgent need for advanced
155 inversion algorithms to avoid the local optimal solution traps and explore the solution space
156 thoroughly.

157 In this study, we develop an MT inversion within a hierarchical Bayesian framework to
158 address the abovementioned problems. Tkalčić et al. (2009) and Hallo & Gallovič (2016) noted
159 that the significant source of long-period Green's functions uncertainty is due to the
160 misalignment between predicted waveforms and observations when using a 1D layered model to
161 present the medium between the source and receivers. Therefore, we propose a scheme to treat
162 the structural error along specific source-station paths when assuming a 1D Earth model (i.e., 2D
163 structural error) as a transition from 1D structural error to 3D structural error, which uses station-
164 specific time shifts between the observed and predicted waveforms. The station-specific time
165 shifts are set as free parameters and determined simultaneously with MT parameters during the
166 inversion, which is the hierarchical aspect of the inversion problem. Treating the time shifts as a
167 part of the inversion is different from the widely used practices, where a grid search with
168 repeating inversions usually determines time shifts (e.g., Mustač et al., 2020), or cross-
169 correlations match the synthetics with observed waveforms (e.g., Alvizuri et al., 2018; Dreger et
170 al., 2021).

171 Secondly, to mitigate the ISO-CLVD tradeoff, we apply an advanced sampling algorithm
172 for Bayesian MT inversion to explore the parameter space thoroughly and effectively. This
173 sampling method is named “effective affine-invariant ensemble samplers” and was proposed by
174 Goodman & Weare (2010) and well implemented with Python (Foreman-Mackey et al., 2013).
175 The ensemble samplers work simultaneously and efficiently to sample the posterior distribution
176 of the parameter model, compared with other traditional sampling algorithms such as the
177 Metropolis-Hastings algorithm (MHA, Hastings, 1970; Metropolis et al., 1953), which applies
178 only one sampler. Its performance is not strongly affected by the linear dependence between MT
179 parameters caused by the ISO-CLVD tradeoff, which makes it more suitable for MT inversion
180 for shallow seismic events.

181 The rest of the paper is as follows. In section 2, we introduce the methodology
182 development of the proposed hierarchical Bayesian MT inversion framework, i.e., 2D structural
183 error treated by the station-specific time shift and the advanced sampling method with effective
184 affine-invariant ensemble samplers. In section 3, we conduct synthetic experiments using an
185 actual configuration of a shallow underground explosion and stations to demonstrate the
186 feasibility of our method. Section 4 is the application to five underground nuclear explosions
187 conducted by the Democratic People's Republic of Korea (DPRK). Finally, in sections 5 and 6,
188 we discuss the MT solutions for real data applications and compare them with previous studies.
189 A brief conclusion is presented at the end.

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193 2 Methodology

194 2.1 Forward modeling of waveforms

195 In the point-source assumption, the synthetic displacement on the Earth's surface can be
 196 expressed as a linear combination of Green's functions (GFs). By following the method
 197 developed initially by Jost and Hermann (1989), then improved by Minson and Dreger (2008),
 198 the displacement of data samples in the direction at a seismic station is written as

$$g_i(\mathbf{m}) = \mathbf{G}_i \mathbf{m}, \quad (1)$$

199 where $\mathbf{G}_i \in \mathbf{R}^{N \times 6}$ is the six-component GFs for a given Earth's structure model, $\mathbf{m} \in \mathbf{R}^6$ is the
 200 seismic MT. This will hold when the source location and origin time are known precisely. This is
 201 a reasonable assumption for manmade seismic sources such as nuclear explosions. The specific
 202 expressions of synthetic displacements, $g_i(\mathbf{m})$ in vertical, radial, and tangential directions for a
 203 full MT, $\mathbf{m} = [M_{xx}, M_{yy}, M_{zz}, M_{xy}, M_{xz}, M_{yz}]^T$, can be found in Minson and Dreger (2008).

204 2.2 Bayesian MT inference

205 The MT can be inferred from the observed seismograms because each synthetic $g_i(\mathbf{m})$
 206 corresponds to an observed seismogram d_i . The Bayesian approach is one of the most powerful
 207 inversion methods because it can explore the solution space thoroughly by using appropriate
 208 samplers and generates an ensemble of solutions instead of only an optimal solution. The spread
 209 of the sampled solutions quantifies solution uncertainty.

210 The MT parameters are treated as random variables in Bayes' theorem (Bayes & Price,
 211 1763), and its posterior distribution can be derived through a likelihood function. The posterior
 212 probability of MT parameters \mathbf{m} given the observation $\mathbf{d} := \{d_i\}$, based on the likelihood
 213 function $p(\mathbf{d}|\mathbf{m})$, a prior distribution $p(\mathbf{m})$, and the evidence of the data $p(\mathbf{d})$, is given as

$$p(\mathbf{m}|\mathbf{d}) = \frac{p(\mathbf{d}|\mathbf{m})p(\mathbf{m})}{p(\mathbf{d})} \quad (2)$$

214 We assume an uninformative prior, $p(\mathbf{m}) = c$, and the evidence $p(\mathbf{d})$ is also an unknown
 215 constant. These two constants, $p(\mathbf{m})$ and $p(\mathbf{d})$, can be omitted without affecting the posterior
 216 distribution's relative landscape but ensuring the algorithm's efficiency. Consequently, the
 217 likelihood function $p(\mathbf{d}|\mathbf{m})$ is used as the posterior probability $p(\mathbf{m}|\mathbf{d})$ in this study. The
 218 posterior probability can be numerically estimated by coordinate distributions obtained by a
 219 Markov chain Monte Carlo (McMC) sampling method (Sambridge & Mosegaard, 2002).

220 The likelihood function includes all information from the data and Earth's structures for
 221 the Bayesian inversion. The widely-used likelihood function has a Gaussian distribution (e.g.,
 222 Dettmer et al., 2007; Duputel et al., 2012; Mustać & Tkalčić, 2016; Phạm & Tkalčić, 2021;
 223 Sambridge et al., 2006)

$$p(d_i|\mathbf{m}) = \frac{1}{\sqrt{(2\pi)^N |C_i|}} \exp\left(-\frac{1}{2} (g_i(\mathbf{m}) - d_i)^T C_i^{-1} (g_i(\mathbf{m}) - d_i)\right), \quad (3)$$

224 C_i and $|C_i|$ are uncertainty covariance matrix and its determinant. The subscript i denote an
 225 individual seismogram component in the observed data. We assume stochastically independent

226 observed components of all stations so that the aggregated likelihood function for $M = n_s \times 3$
 227 (n_s is the number of three-component stations) component seismograms is

$$p(\mathbf{d}|\mathbf{m}) = \prod_{i=1}^M \frac{1}{\sqrt{(2\pi)^N |C_i|}} \exp\left(-\frac{1}{2}(\mathbf{g}_i(\mathbf{m}) - \mathbf{d}_i)^T C_i^{-1}(\mathbf{g}_i(\mathbf{m}) - \mathbf{d}_i)\right). \quad (4)$$

228 It measures the overall waveform fit level between the observed and the predicted seismograms,
 229 which makes it a critical factor in Bayesian seismic source inversion.

230 2.3 Estimating the covariance matrix

231 The covariance matrix C_i in Equation 4 enables the consideration of various sources of
 232 uncertainty in the inversion problem. There are two sources of uncertainty: data noise, the
 233 empirical theory error, or their combination. Firstly, data noise is mainly caused by background
 234 ambient noise at the recording site and instrumental noise in the data acquisition. Secondly, the
 235 theory uncertainties, or uncertainties relating to the forward problem, are any source of errors
 236 due to theoretical approximations in the forward problem. It is reasonable to assume that the
 237 most significant contribution to the theory error is due to our imperfect knowledge of the Earth's
 238 interior structure, also referred to as structural uncertainty in this study.

239 To thoroughly consider the uncertainty in an MT inversion problem, the covariance
 240 matrix should account for both sources of uncertainties. Therefore, a combined covariance
 241 matrix was proposed by Tarantola & Valette (1982) and further explored by other studies (e.g.,
 242 Duputel et al., 2012; Phạm & Tkalčić, 2021; Tarantola, 2005; Vasyura-Bathke et al., 2021),
 243 which is written as

$$C_i = C_i^d + C_i^t, \quad (5)$$

244 where C_i^d and C_i^t are covariance matrices for the data noise and structural error, respectively. The
 245 structural covariance matrix, C_i^t , is estimated empirically by perturbing a 1D Earth model using
 246 the Monte-Carlo simulation. Moreover, Duputel et al. (2012) and Phạm & Tkalčić (2021)
 247 demonstrated the dependency of C_i^t on a prior MT, i.e., $C_i^t(m)$, which is computationally
 248 expensive, especially when 3D Earth is considered. Furthermore, the empirical estimation of the
 249 structural covariance matrix requires subjective choices for scale and parameterization of the
 250 Earth model perturbations, which are currently subjected to future research.

251 Here, we propose a simplified treatment of the structural uncertainty to avoid the
 252 expensive Monte-Carlo simulation, in which the structural errors are treated using station-
 253 specific time shifts (more details to be considered in Section 2.4). The covariance matrix C_i from
 254 Equation 4 only includes uncertainty from data noise. In further simplification, data noise on
 255 each component is assumed to be uncorrelated when signal-to-noise ratios (SNR) of inverted
 256 waveforms are large, which is usually the case for intermediate-large earthquakes. The
 257 covariance matrix C_i becomes diagonal

$$C_i = \sigma_i^2 \mathbf{I}, \quad (6)$$

258 where σ_i^2 is the unknown noise variance of each seismogram. To reduce the number of noise
 259 parameters and avoid the wide range to search for them, we follow the approach proposed by
 260 Phạm & Tkalčić (2021) to parameterize the covariance matrix in Equation 6 as,

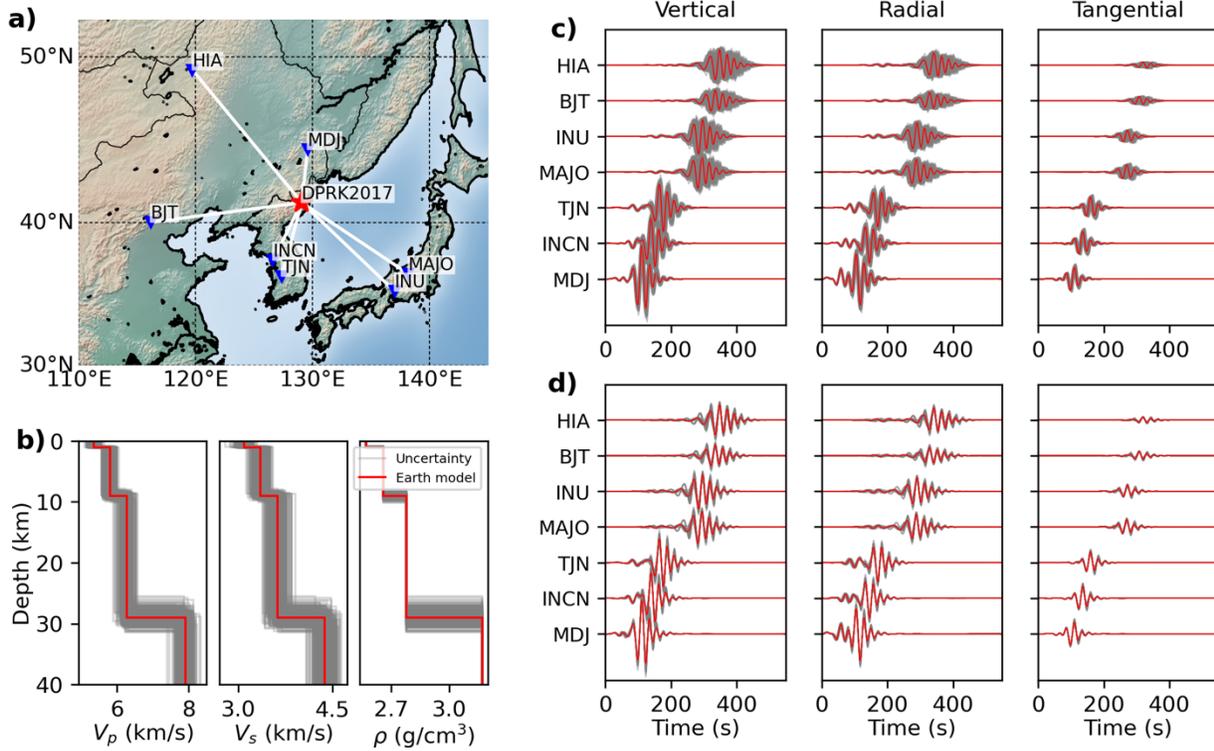
$$C_i = h \cdot (\sigma_i^{ref})^2 \mathbf{I}, \quad (7)$$

261 where σ_i^{ref} is the reference noise strength for each component that is the pre-computed standard
 262 deviation of the 1-hour pre-event ambient noise of three components at each station, and h is the
 263 station-specific noise hyper-parameter. The pre-event noise used to calculate covariance matrix
 264 is pre-processed in the same way as the data used in the inversion.

265 2.4 Accounting for 2D Earth's model uncertainty by station-specific time shifts

266 This study provides a simplified scheme to treat the 2D structural error, i.e., structural
 267 error along specific source-station paths, by inverting for the station-specific time shifts between
 268 predicted waveforms and observations. To demonstrate the validity of this simplification, we
 269 take the DPRK2017 explosion as an example to indicate the misalignment between waveforms
 270 from perturbed 1D Earth models. As Figure 1b shows, a four-layer velocity model (MDJ2,
 271 Ford et al., 2009) is randomly perturbed 300 times given 5% uncertainty (see Pham & Tkalčić,
 272 2021 for the description of 1D model perturbation). An ensemble of waveforms generated by the
 273 same explosive MT in these perturbed 1D models is plotted in Figure 1c. The waveforms at the
 274 same station feature a high degree of similarity in long period band, e.g., 20 – 50 s, used in this
 275 study. At stations MDJ, INCN, and TJN, these 300 waveforms of each component almost
 276 overlap, showing insignificant misalignments in phase and amplitude. However, the
 277 misalignments in phase (referred to as time shift) become more apparent and more significant as
 278 the epicenter distance increases at the other four stations while the amplitudes remain similar.

279 The high order of similarity after waveform alignment confirms the dominance of time
 280 shifts by the model uncertainty in 1D. Specifically, we performed a grid search for the time shift
 281 at each component to achieve the best waveform fit (i.e., the highest variance reduction, VR,
 282 defined in Equation S17b of Pham & Tkalčić, 2021) between the waveforms from the MDJ2
 283 model (red in Figure 1b) and the perturbed MDJ2 model (gray in Figure 1b). The re-aligned
 284 waveforms are shown in Figure 1d. The overall VR of waveform fit is 95.8% after realignment.
 285 Therefore, time shifts dominate the structural error within 5% perturbation uncertainty, providing
 286 a pathway to treat the primary source of the uncertainty from structural errors. Hallo & Gallovič
 287 (2016) derived an approximate covariance matrix by considering these random time shifts in
 288 waveforms. In this study, alternatively, we directly invert the station-specific time shifts
 289 simultaneously with MT parameters, which sets the station-specific time shifts as free parameters
 290 determined by the data to account for the structural error along specific wave propagation paths.



291

292 **Figure 1.** Synthetic scenario to demonstrate the time shifts generated by perturbed 1D velocity
 293 models. (a) Map showing the DPRK2017 explosion location (red star) and seven seismic stations
 294 (blue triangles). (b) The P-wave and S-wave velocity and density of the MDJ2 model (red),
 295 which is a four-layer velocity model (Ford et al., 2009), and its 300 perturbed structures (gray)
 296 given 5% uncertainty. (c) The three-component waveforms for perturbed 1D Earth structures in
 297 (b) and the MT of DPRK2017 explosion from Alvizuri and Tape (2018). All waveforms are
 298 filtered using 20–50 s period band. (d) The re-aligned waveforms from (c) by grid search for the
 299 optimal time shift at each component to obtain the best variance reduction (i.e., 95.8%).

300 Allowing noise amplitudes and time shifts, i.e., the hierarchical aspect of Bayesian
 301 inference, makes the MT inversion non-linear. The noise parameters are already included in the
 302 Bayesian inversion through the likelihood function in Equations 4 and 7. The time-shifting of a
 303 waveform can be described analytically as,

$$g'_i(\mathbf{m}) = F^{-1}[F[g_i(\mathbf{m})] \cdot e^{-i\omega\tau}], \quad (8)$$

304 in which F , F^{-1} denote forward and inverse Fourier transformation, respectively. τ is the station-
 305 specific time-shift parameter, which allows continuous time-shifting values rather than being
 306 restricted by discrete sampling intervals. In this work, the τ is bounded by $[-10, 10]$ to avoid
 307 cycle skipping for waveforms filtered between 20 - 50 s, which is the frequency band we used in
 308 this study. Therefore, the complete parameter model to invert for is defined as $[\mathbf{m}, \mathbf{h}, \boldsymbol{\tau}]$ where
 309 $\mathbf{m} = [M_{xx}, M_{yy}, M_{zz}, M_{xy}, M_{xz}, M_{yz}]^T$ parameterizes a full MT, $\mathbf{h} = [h_1, h_2 \dots h_{n_s}]$
 310 parameterizes station-specific data noise strengths, and $\boldsymbol{\tau} = [\tau_1, \tau_2 \dots \tau_{n_s}]$ are the station-specific
 311 time shifts. Finally, the likelihood function in Equation 4 is rewritten as

$$p(\mathbf{d}|\mathbf{m}, \mathbf{h}, \boldsymbol{\tau}) = \prod_{i=1}^M \frac{1}{\sqrt{(2\pi)^N |C_i|}} \exp\left(-\frac{1}{2}(g'_i(\mathbf{m}) - d_i)^T C_i^{-1}(g'_i(\mathbf{m}) - d_i)\right). \quad (9)$$

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2.5 Exploring the parameter space using affine-invariant ensemble samplers

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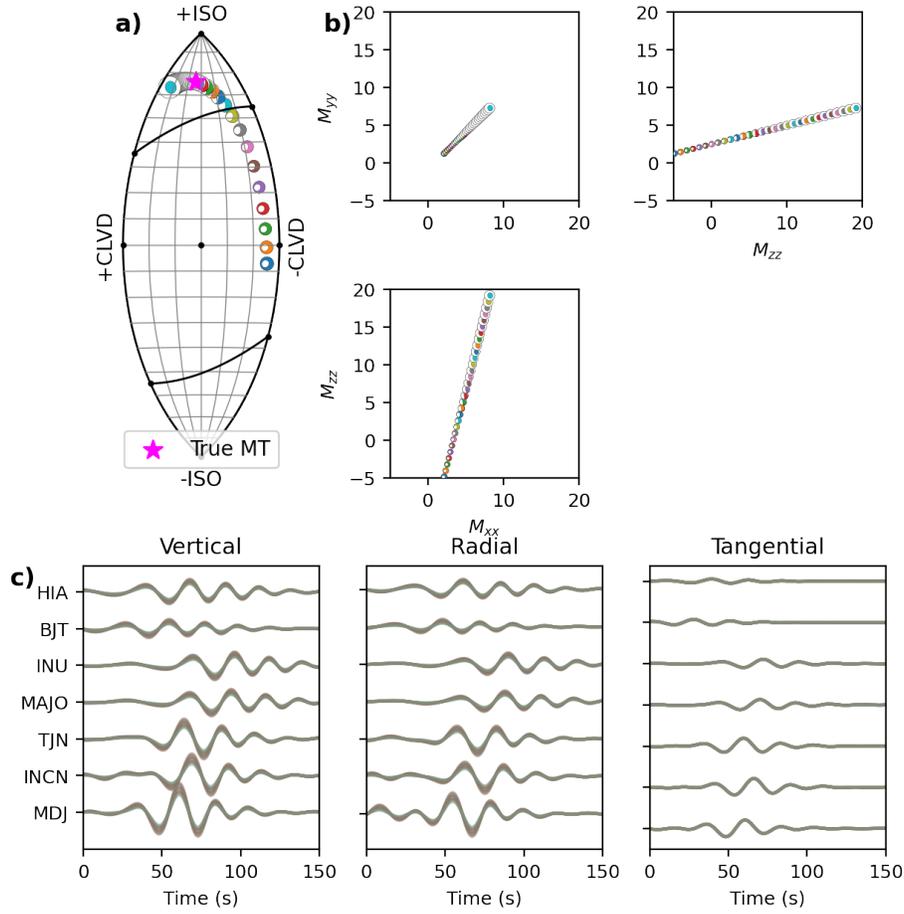
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The inherent ambiguity between pure ISO and vertical CLVD is a significant challenge in MT inversion for shallow seismic sources using long-period regional waveforms. At the shallow depths, seismic waveforms recorded by regional stations (< 1200 km) are dominated by surface waves, which have minimal sensitivities to the vertical force couple. This explains the high similarity between waveforms in Figure 2c generated by various ISO-dominating and vertical-CLVD-dominating sources in Figure 2a at 0.5 km depth, which is meant to reproduce the comparison by Kawakatsu (1996). The waveform similarity leads to the severe tradeoff between ISO and CLVD when resolving for NDC components of the shallow sources, e.g., manmade underground explosions. In parameter space, this ISO-CLVD tradeoff presents a strong linear dependence among three diagonal elements of an MT, i.e., M_{xx} , M_{yy} , and M_{zz} , as shown in Figures 2b. It is challenging to thoroughly sample this type of parameter distribution in Bayesian MT inversion using sampling algorithms such as the Metropolis-Hastings algorithm (MHA, Hastings, 1970; Metropolis et al., 1953). Here, we promote using the affine-invariant ensemble samplers (Goodman & Weare, 2010) for this MT inverse problem to effectively sample the MT solution spaces to mitigate the challenge caused by the shallow source depths.



329

330 **Figure 2.** The ambiguity of non-double-couple components of the shallow seismic source. (a)
 331 Various inverted seismic MTs (shown as focal mechanisms in different colors) yield almost
 332 identical seismic waveforms. The magenta star is the input MT from Alvizuri and Tape (2018).
 333 (b) The linear relationship between three pairs of MT parameters, i.e., M_{xx} and M_{yy} , M_{xx} and
 334 M_{zz} , and M_{yy} and M_{zz} . (c) The synthetic three-component waveforms at seven stations (Figure
 335 1a) produced by the MTs shown in (a).

336 This approach of ensemble samplers employs K walkers in a coordinated manner by
 337 exchanging their current coordinates to explore the N -dimensional unknown model space.
 338 Goodman & Weare (2010) proposed the ‘stretch move’ proposal scheme, in which the next
 339 move of a walker \mathbf{m}_i is proposed in two steps, as in Figure 3. First, a random partner is chosen
 340 from the complementary walkers in the ensemble, say \mathbf{m}_j . Then, the proposed move is drawn
 341 randomly along the line connecting the two walkers,

$$\mathbf{m}'_i = \mathbf{m}_j + Z \cdot (\mathbf{m}_i - \mathbf{m}_j). \quad (10)$$

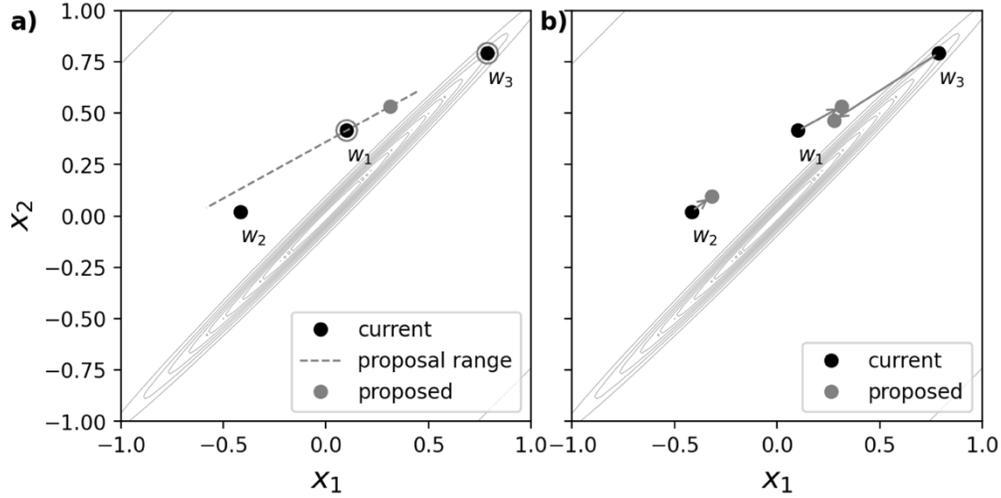
342 In Equation 10, Z is a random, positive number drawn from a probability distribution $g(z)$ in the
 343 $[1/a, a]$ interval,

$$g(z) \propto \begin{cases} \frac{1}{\sqrt{z}} & \text{if } z \in [1/a, a] \\ 0 & \text{otherwise} \end{cases} \quad (11)$$

344 The parameter a , where $a > 1$, is the only parameter to adjust the performance of the ‘stretch
 345 move’ scheme. Furthermore, $a = 2$ has empirically been found to be an optimal choice in many
 346 large-scale inverse problems (Foreman-Mackey et al., 2013; Goodman & Weare, 2010). This
 347 proposed move of the walker \mathbf{m}_i is accepted based on a probability involving the probabilities of
 348 the current coordinate and the proposed move,

$$q = \min\left(1, Z^{N-1} \frac{p(\mathbf{d}|\mathbf{m}'_i)}{p(\mathbf{d}|\mathbf{m}_i)}\right). \quad (12)$$

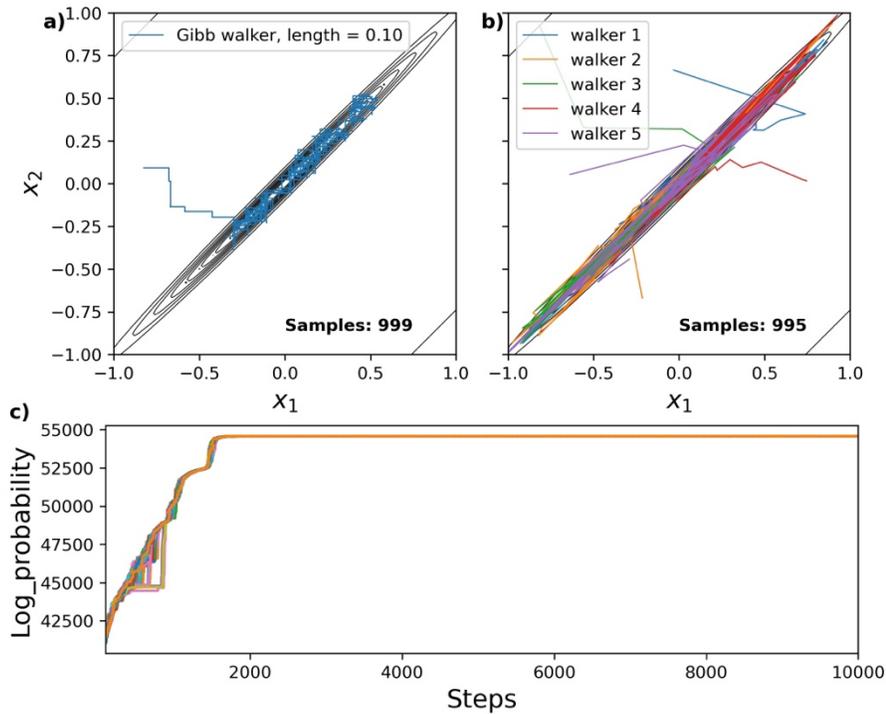
349 The stretch move is iterated for other walkers in the ensemble before proceeding to the next
 350 iteration. The ensemble samplers are implemented in a lightweight, well-tested Python package,
 351 emcee (Foreman-Mackey et al., 2013).



352

353 **Figure 3.** Schematic demonstration in two-dimensional parameter space of the stretched move
 354 used in the affine-invariant MCMC (Goodman & Weare, 2010). The background shows the
 355 contours of the probabilistic distribution to be sampled. In (a), black dots mark the current
 356 positions of three walkers. Grey dot is a proposed move for the walker w_1 , with a randomly
 357 chosen partner w_3 . The dashed gray line shows the range of proposals for the next move of w_1 .
 358 In (b), gray dots are proposed to move all three walkers from their current positions, which will
 359 be accepted or rejected randomly.

360 The ensemble samplers, designed as above, possess the affine invariant property, whose
 361 performance is not affected by an affine transformation of the coordinates. Such transformations
 362 are often caused by the linear dependence between parameters, which leads to a highly
 363 anisotropic probability distribution, as demonstrated in Figure 2b. However, the affine-invariant
 364 ensemble samplers can thoroughly and effectively sample this type of distribution compared to
 365 traditional sampling algorithms. As the example in Figures 4a and 4b shows, with the same
 366 number of sampling steps, i.e., 1000, Gibb’s sampler only samples part of the target distribution,
 367 while the ensemble samplers of 5 walkers with 200 steps each explore the whole target
 368 distribution. This property makes it more suitable for MT inversion for shallow sources. In the
 369 following numerical experiments and applications to real data, we will demonstrate the
 370 advantages of the ensemble samplers for the MT inversion problem of non-double-couple
 371 components in shallow seismic sources.



372 **Figure 4.** Comparison of sampling efficacy between (a) the traditional Metropolis-Hasting
 373 method and (b) the ensemble samplers with stretched moves (Goodman & Weare, 2010). The
 374 background contours show the target probability distribution. Each colored trace represents the
 375 trajectory of a walker. There are 1000 random samples drawn in both cases. (c) Posterior
 376 probability varying with the inversion step during the proposed Bayesian MT inversion using
 377 affine-invariant ensemble samplers. Color-coded lines are for different 512 walkers during
 378 10,000 iterations.

379 3 Synthetic Experiment

380 3.1 Experiment configuration

381 We design numerical experiments having a realistic source-receiver configuration to
 382 demonstrate the feasibility of this approach on the MT inversion for resolving NDC components
 383 of shallow seismic sources. Figure 1 shows the event location and seven stations providing good
 384 azimuthal coverage to the interested event located at the DPRK nuclear test site. Epicentral
 385 distances from the stations range from 370 km up to 1100 km. The four-layer 1D velocity model
 386 MDJ2 (Ford et al., 2009) simulates synthetic waveforms. An explosive event is fixed at 0.5 km
 387 depth, and its input MT is the solution of the DPRK2017 event from Alvizuri & Tape (2018),
 388 which includes 63.7% ISO, 6.4% CLVD, and 29.8% DC, with a moment magnitude $M_w = 5.21$.

389 The “noisy” synthetic waveforms are calculated with data and structural uncertainties.
 390 Noise-free waveforms are band-passed filtered between 20–50 second periods. First, three-
 391 component real recorded ambient noise before the origin time of DPRK2017 explosion, pre-
 392 processed in the same way as noise-free waveforms, are added to corresponding three-
 393 component noise-free waveforms at the sites to represent the data noise. The reference noise

394 strengths, σ_i^{ref} , are pre-computed from the 1-hour pre-event ambient noise (Equation 7) and the
 395 input relative noise levels, $h_1, h_2 \dots h_7$, are set to unity. Secondly, to introduce the structural
 396 uncertainty, we shift the data with station-specific times (Table 1). Waveforms are shifted
 397 forward, corresponding to positive time shifts for three stations in China and South Korea, and
 398 backward, corresponding to negative time shifts for two stations in Japan. The signs of the shifts
 399 simulate the actual difference between the MDJ2 model and slower continental crust toward the
 400 western sites and faster oceanic crust toward the eastern sites. The time shifts are the only source
 401 of structural uncertainty introduced in synthetic waveforms.

402 **Table 1.** True station-specific time shifts (unit: second), used for the numerical experiment of MT
 403 inversion for the DPRK2017 test.

Explosion	IC.MDJ	IC.BJT	IC.HIA	IU.INCN	KG.TJN	IU.MAJO	G.INU
DPRK2017	4.0	3.7	4.0	2.0	1.5	-4.5	-5.5

404

405 3.2 Inversion results for a synthetic, shallow-source explosion

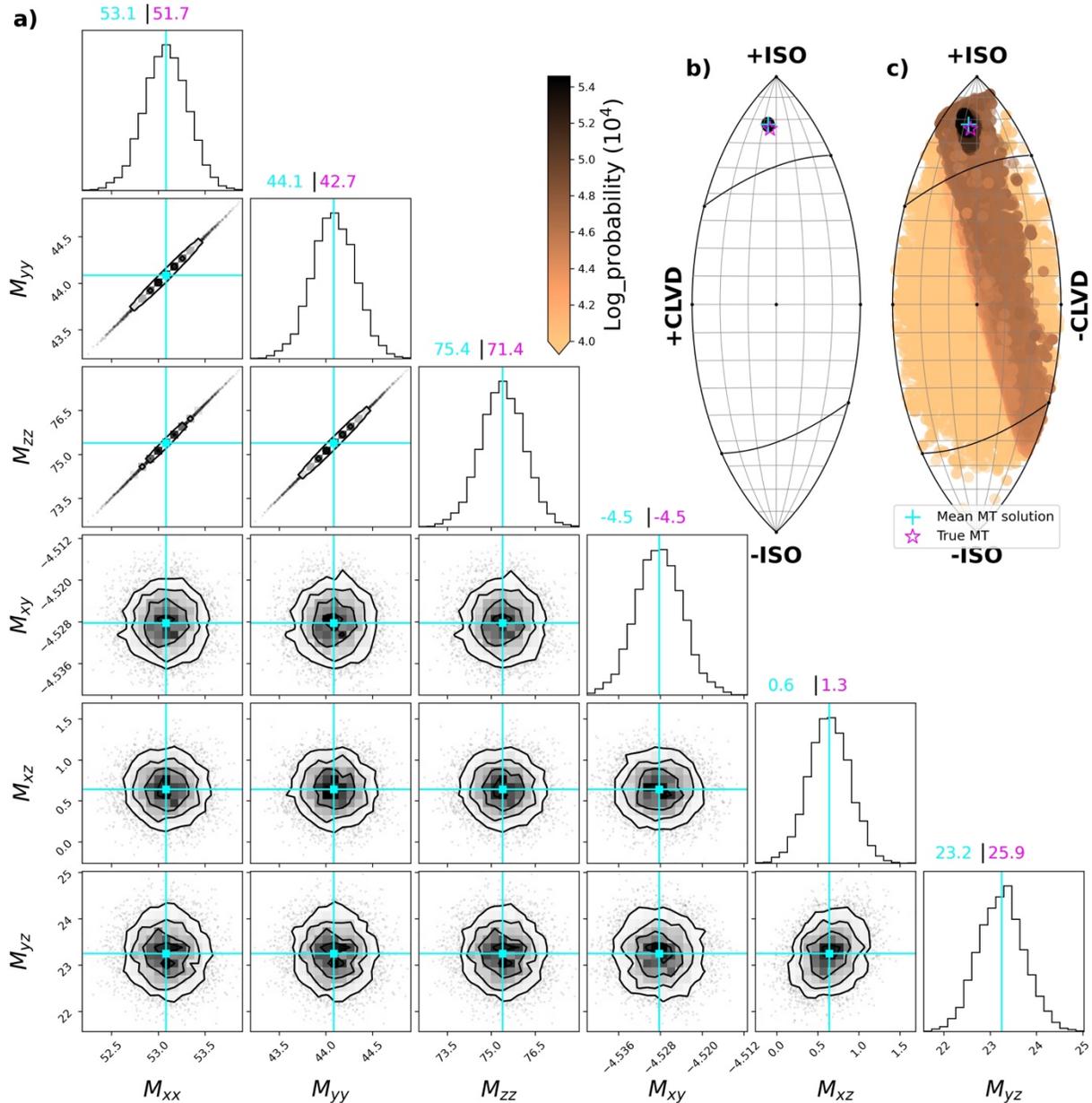
406 The affine-invariant ensemble samplers introduced for the seismic MT inversion in this
 407 study (Section 2.5) perform excellently in terms of efficiency and effectiveness. We used 512
 408 walkers and 10,000 iterations in all inversions presented in this study. The samples from each
 409 walker are not independent. The emcee follows Goodman & Weare (2010) and uses the
 410 autocorrelation time τ_f , i.e., the number of steps before producing independent samples of the
 411 target distribution, to estimate the effective number of independent samples. Running with a
 412 large number of walkers is beneficial to obtain more independent samples and a higher
 413 acceptance rate, that is, the fraction of proposed steps to be accepted (Foreman-Mackey et al.,
 414 2013; Goodman & Weare, 2010). Finally, the first several times τ_f of samples of each walker are
 415 discarded as the burn-in stage. The number of discarded samples is determined via tests prior to
 416 the inversion to make sure the remaining samples have reached the convergence, where all
 417 walkers fluctuate around the similar highest probability. The samples in the convergence stage
 418 are thinned by half the autocorrelation time and flattened across the walkers to obtain the
 419 solution ensemble. In this study, we discard the first half of 10,000 iterations of each walker that
 420 is about 10 times of the maximum τ_f of all walkers. The remaining half of 10,000 iterations are
 421 used as the convergence stage. The probability varying with the inversion step for all walkers is
 422 plotted in Figure 4c with different colors. As one can see, in the burn-in stage, the probability
 423 from each walker increases quickly before reaching the convergence stage. The inversion takes
 424 4.5 minutes on a personal computer (3.1 GHz 6-Core Intel Core i5) for this numerical
 425 experiment.

426 This proposed Bayesian MT inversion successfully recovers the shallow explosive source
 427 using affine-invariant ensemble samplers. The inversion results are summarized in Figures 5, 6
 428 and 7. According to the lune source-type diagram (Tape & Tape, 2012) shown in Figure 5c, the
 429 algorithm with ensemble samplers effectively explores the parameter space. Initially, a wide
 430 variety of source types is explored (copper dots). Then the samplers go through a stripe in the
 431 lune diagram to explore the ISO-CLVD tradeoff with higher posterior probabilities (dark brown
 432 dots). The samplers eventually converge to a small area corresponding to the highest posterior

433 probability (black dots; also plotted in Figure 5b for clarity), where the cyan cross denotes their
 434 mean. As can be seen in Figures 5b and 5c, the mean MT solution is close to the true MT
 435 (represented by the magenta star) in the lune source-type diagram. The decomposition of the
 436 mean MT solution (Figure 6a) gives 65.5%ISO, 8.4%CLVD, and 26.2%DC, which agrees with
 437 63.7% ISO, 6.4% CLVD, and 29.8% DC of the true MT. Its moment magnitude is $M_w=5.22$,
 438 which is close to the input $M_w=5.21$.

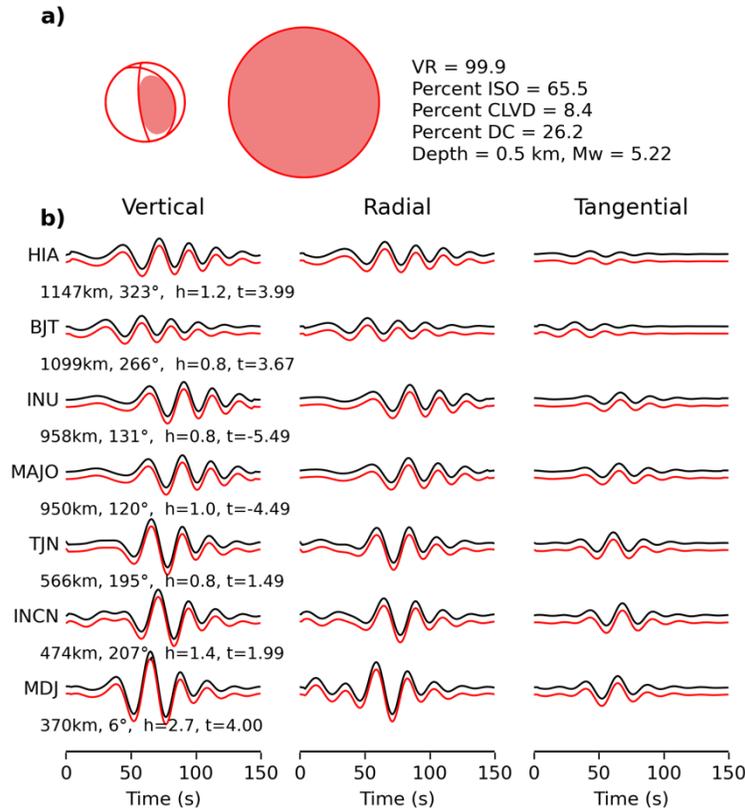
439 The evolution of MT solutions from low to high probability demonstrates the
 440 effectiveness of the employed search engine. The plot of the posterior probability in Figure 5c is
 441 consistent with the contour plot of variance reduction shown in Alvizuri & Tape (2018) by grid
 442 search over source types to achieve the best waveform fit. Moreover, based on the posterior
 443 probability, our method avoids most MTs in the ISO-CLVD tradeoff area and shows smaller MT
 444 uncertainty in the converging stage. The posterior distribution of each MT parameter is near
 445 Gaussian, as shown in Figure 5a, consistent with the assumption made when deriving the
 446 likelihood function in Section 2.2. The linear correlation between M_{xx} , M_{yy} and M_{zz} is a result
 447 of the tradeoff between pure ISO and vertical-CLVD components for shallow sources, as
 448 discussed in Section 2.5.

449 Apart from the MT parameters, the station-specific noise levels (Figure 7a) and time
 450 shifts (Figure 7b) are also recovered by the ensemble samplers. As mentioned before, all noise
 451 levels are fixed to a single value (1.0) in the current numerical experiment. The recovered mean
 452 noise levels for all stations are generally close to the input value. Besides, the recovered time
 453 shifts are also close to the input time shifts (Table 1). The posterior distributions of station-
 454 specific noise and time shift parameters show a Gaussian character. An excellent waveform fit
 455 (VR>99%) between the observed (black) and predicted waveforms (red) using the mean MT and
 456 time shifts is obtained in Figure 6b. Therefore, we conclude that the inversion framework using
 457 regional stations is successful.



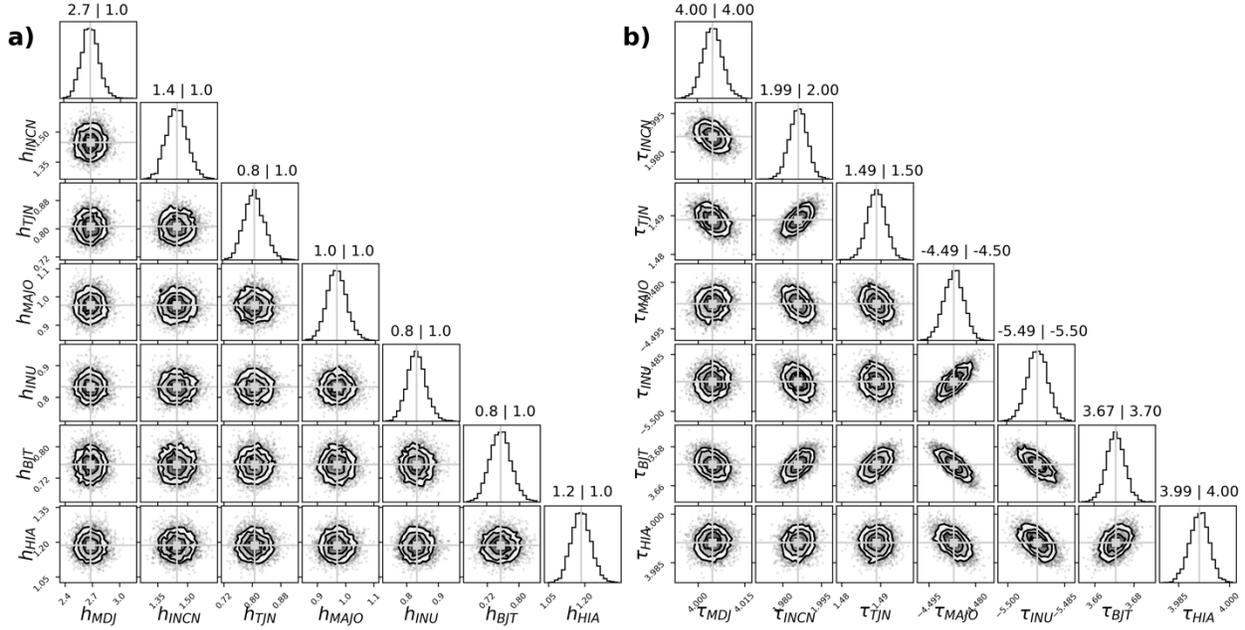
458

459 **Figure 5.** The synthetic scenario MT inversion considering uncorrelated data noise and 2D
 460 structural error within a hierarchical Bayesian inversion framework. The source depth is 0.5 km.
 461 Synthetic waveforms are filtered in the 20-50 s period band. (a) Each sub-panel shows a pair of
 462 the MT parameters in the convergence stage of the inversion. For a definition of the convergence
 463 stage, see the main text. The unit of MT parameters is 10^{15} Nm. The cyan lines are the MT
 464 parameters' means which are also indicated by the cyan numbers above each column, separated
 465 from the true (input) values (magenta numbers) by a vertical bar. (b) The lune diagram with the
 466 converging MT solution from (a). The magenta star shows the source type of the true MT input.
 467 The cyan cross shows the mean MT solution of the convergence stage. The color bar is used to
 468 display log probability. (c) The Lune source-type diagram shows the evolution of every 2 MT
 469 solutions during the entire inversion stage.



470

471 **Figure 6.** MT decomposition and waveform fit for the synthetic scenario. (a) Decomposition of
 472 MT solution into deviatoric (left) and isotropic (right) parts. The beachball sizes are proportional
 473 to the MT component percentages. (b) Waveform fit between ‘observed’ (black) and predicted
 474 (red) waveforms from the MT solution shown in (a), measured by the variance reduction. The
 475 waveforms are offset vertically for clarity. The numbers shown beneath the waveforms are
 476 source-receiver distance, azimuth, recovered station-specific noise parameter and time shift.



477

478 **Figure 7.** Recovered station-specific noise parameters (a) and time shifts (b) for the synthetic
 479 scenario. Each sub-panel shows a pair of parameters in the convergence stage of the inversion.
 480 The two numbers above each column are each parameter’s mean and the true (input) values,
 481 respectively, which are separated by a vertical bar. The light gray lines show the mean values.

482 3.3 Sensitivity tests

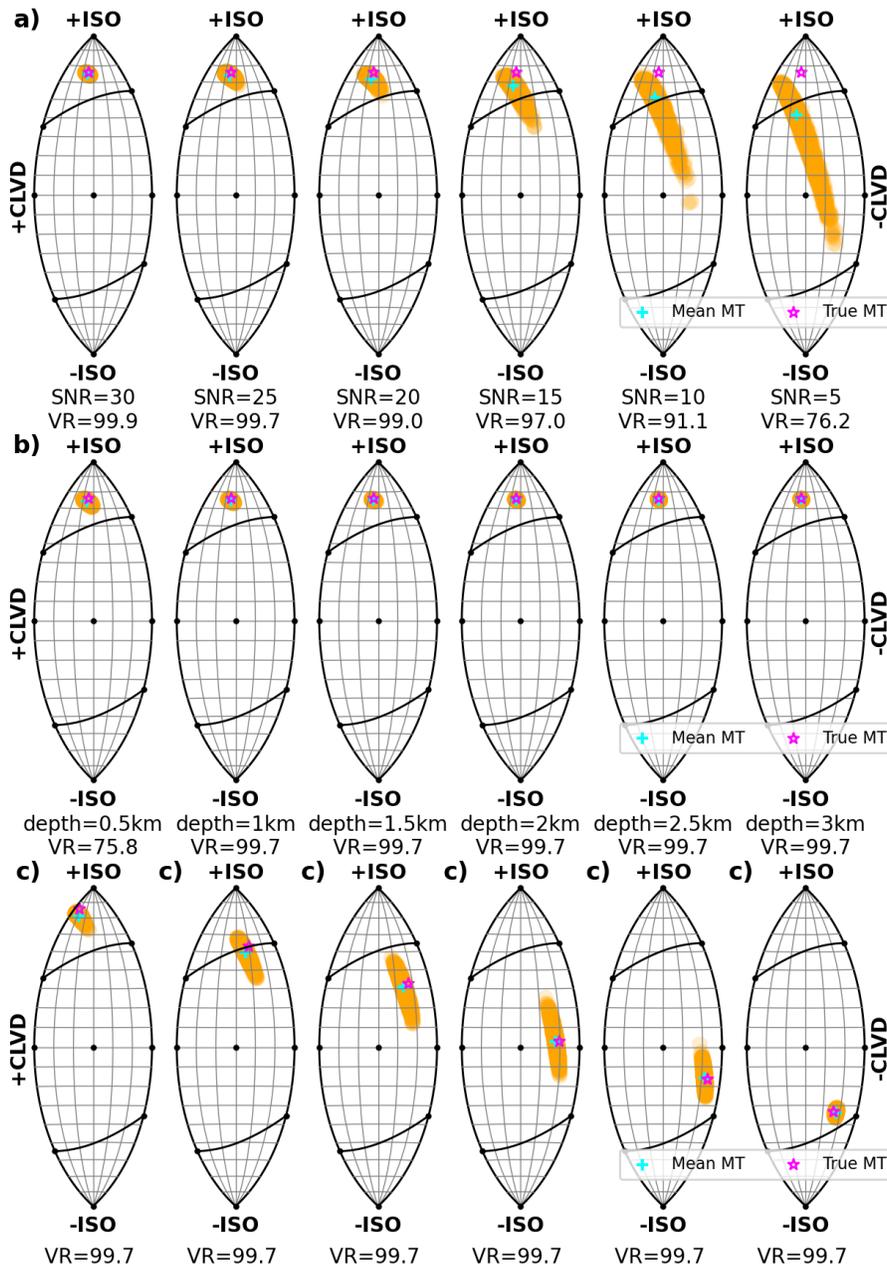
483 Given that the inversion solution is sensitive to the presence and the way of treating the
 484 data noise, we consider its sensitivity against several scenarios, including different datasets
 485 corresponding to high, intermediate, and low SNR, different source depths, and different source
 486 types. The SNR is defined by

$$SNR = 20 \log_{10} \left(\frac{A_s}{C \cdot A_n} \right), \quad (13)$$

487 where A_s and A_n are the root mean square of the simulated waveform and 1-hour pre-event
 488 ambient noise amplitude. C is a component-based coefficient multiplying with the ambient noise
 489 to generate waveforms of specific SNR. We conducted six datasets of different SNRs from 5 to
 490 30, with increments of five units. The real recorded data noise is correlated, and its correlated
 491 property should be considered in the noise model in an inversion problem; however, we argue
 492 that assuming uncorrelated noise is reasonable when the SNR is high.

493 The assumption of uncorrelated noise is reasonable in the cases of high SNR, while it
 494 may fail in the cases of low SNR. As shown in Figure 8a, the shallow source can be recovered in
 495 the case of high SNR ($SNR = 30$). The MT converges to a small area in orange, which is close
 496 to the true source (magenta star), with small uncertainty. As the correlated noise becomes more
 497 significant (i.e., $SNR=25$ or 20), the solution uncertainty also becomes more significant, and the
 498 theoretical tradeoff due to shallow depths becomes more challenging to mitigate. However, there
 499 is still a chance to retrieve the source parameters by only considering uncorrelated noise for
 500 intermediate-size earthquakes whose data SNR is usually above 20. For a typical SNR, i.e., 25,

501 this inversion method works for the same MT sources at depths varying from 0.5 to 3.0 km, as
502 shown in Figure 8b. Besides, six different non-DC sources, including ISO-dominated and
503 CLVD-dominated sources at the same depth of 0.5 km (Figure 8c), are also recovered with the
504 uncorrelated noise model. However, in the case of low SNR data ($SNR = 10$ or 5), our
505 algorithm, assuming uncorrelated noise, cannot reasonably recover the input MT. The solution
506 uncertainty is substantial, as shown by the orange dots in the last two panels of Figure 8a, and the
507 mean MT is far away from the true one. Besides, the theoretical tradeoff between ISO and
508 CLVD remains unresolved due to the inappropriate noise estimate. This happens whenever noisy
509 stations are involved or the earthquake is small.



510

511 **Figure 8.** Source-type lune diagrams for recovered MT solutions in the following scenarios: (a)
 512 varying signal-to-noise ratios (SNR) from 30 to 5, with decrements by five units from left to
 513 rights, for the true source depth of 0.5 km; (b) varying true source depths from 0.5 to 3.0 km,
 514 with increments by 0.5 km, for the waveforms with SNR = 25; and (c) varying true source-types
 515 at the depth = 0.5 km and SNR = 25. In each scenario, the source depth is treated as known. A
 516 magenta star represents the true MT in each panel. Overlapped orange dots are MT solutions in
 517 the convergency stage. A cyan cross marks their mean MT. The variance reduction between
 518 ‘observed’ and predicted waveforms from mean MT is shown beneath each panel. The noise in
 519 the simulated waveform is the pre-event noise multiplied by different factors to obtain “noisy
 520 waveforms” with given SNR.

521 4 Application for DPRK nuclear tests

522 4.1 Data preparation

523 Using lessons from the synthetic experiments, we now apply the developed MT inversion
 524 framework to the five DPRK nuclear tests between 2009 and 2017. The DPRK2006 test is not
 525 included in this study due to poor data quality. When possible, we use the same set of stations for
 526 all events to cross-check the recovered time shifts besides the recovered MT solutions. We
 527 choose five standard stations (i.e., MDJ, MAJO, INU, BJT, and HIA, as shown in Figure 1a)
 528 with sufficient SNR for each nuclear explosion. To fill the azimuth coverage gap in South Korea,
 529 the station INCN is added for the DPRK2009 test, the stations CHNB and YNCB for the
 530 DPRK2013 test, and the stations INCN and TJN for the three tests in 2016–2017. Finally, we
 531 used six stations for the DPRK2009 and seven for the DPRK2013–2017 tests. The recorded 3-
 532 component waveforms are corrected for the instrumental response to obtain displacements and
 533 filtered in the 20–50 second period band using a 4-corner acausal Butterworth bandpass filter.
 534 The waveforms are then incised into 150 s-windows starting at manually picked delay times after
 535 the origin times which are 50 s for stations MDJ, CHNB and YNCB, 70 s for INCN, 100 s for
 536 TJN, 200 s for MAJO and INU, and 280 s for BJT and HIA, respectively. The epicenter location
 537 and origin time used in this study are from Table 1 of Alvizuri and Tape (2018). GFs are
 538 calculated using the MDJ2 model (Ford et al., 2009) with a fixed depth of 0.5 km. The
 539 configuration of ensemble samplers is the same as used in synthetic experiments.

540 4.2 MT inversion results of DPRK2009–2017 tests

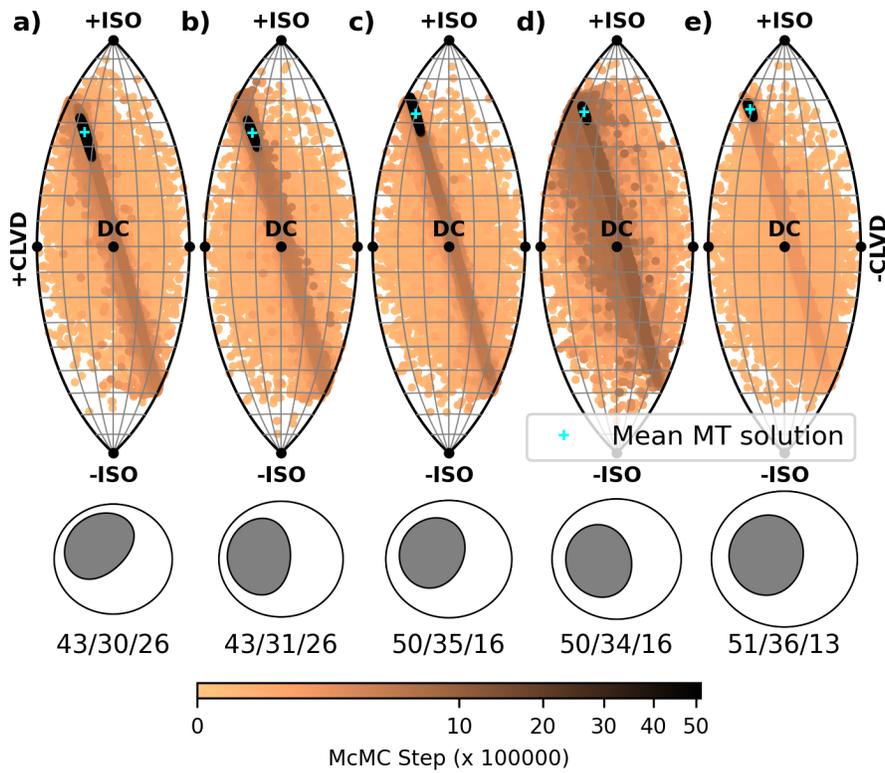
541 Figure 9 presents the entire evolution of the Monte-Carlo chains during the sampling for
 542 all five explosions. Like in the synthetic case, starting with randomly chosen MTs, our inversion
 543 method with ensemble samplers explores a wide variety of source types, including the ISO-
 544 CLVD tradeoff area (the darker stripe in each sub-panel) with a higher posterior probability.
 545 Finally, the chains converge to a small area with the highest posterior probability (consisting of
 546 black dots in each sub-panel in Figure 9). The evolution patterns of MTs are consistent among
 547 the five explosions, which, to some extent, agrees with the patterns obtained by grid search over
 548 source types to achieve the best waveform fit for the DPRK tests by Chiang et al. (2018) and
 549 Alvizuri & Tape (2018). Moreover, by accounting for the station-specific data noise and time
 550 shifts between predictions and observations (i.e., 2D structural error), our inversion method skips
 551 most MTs in the ISO-CLVD tradeoff area and shows smaller uncertainty of the MT solution in
 552 the convergence stage. The mean MT solution of each explosion, i.e., the cyan cross in each sub-
 553 panel, is calculated by averaging the MTs in this convergence stage. Figure 10 shows the
 554 excellent fit of the predicted waveforms corresponding to the mean MTs and the observed
 555 waveforms.

556 The source mechanisms recovered from the five DPRK explosions in 2009–2017 exhibit
 557 similar explosive nature. Large ISO components dominate their MT solutions, i.e., 43% in the
 558 DPRK2009 test and DPRK2013 test, and 50% in three DPRK2016–2017 tests, respectively,
 559 which indicates their explosive nature of sources. The three diagonal elements of mean MT
 560 solutions, M_{xx} , M_{yy} , and M_{zz} , are all positive and larger than off-diagonal elements, M_{xy} , M_{xz} ,
 561 and M_{yz} . Furthermore, M_{xx} and M_{yy} are almost equal and smaller than M_{zz} , which indicates
 562 these five explosions are close to a crack source. The results also show significant CLVD
 563 components required in these five explosions ($\geq 30\%$) and small DC components, e.g., 13% of

564 DC for the 2017 explosion. The high degree of similarity among these five explosions, i.e., near
565 the ISO pole and close to the crack source in the source-type lune diagram, has already been
566 pointed out by Liu et al. (2018) using a unique dataset that includes more broadband stations on
567 the China side. Their similar long-period waveforms are responsible for this source similarity.
568 However, the crack source mechanism for underground nuclear explosions remains unclear.
569 Interestingly, our results coincide with the MTs of nuclear explosions at Nevada National
570 Security Site obtained by Pasyanos & Chiang (2021) using MT inversion for 130 nuclear
571 explosions from 1970 to 1996, which are also distributed around the crack source. Compared
572 with other studies (e.g., Alvizuri & Tape, 2018; Chiang et al., 2018), we report slightly higher
573 moment magnitudes, i.e., $M_w = 4.69$, $M_w = 4.93$, $M_w = 5.0$, $M_w = 5.13$, and $M_w = 5.79$,
574 respectively. The values obtained are closer to the moment magnitudes that Liu et al. (2018)
575 obtained.

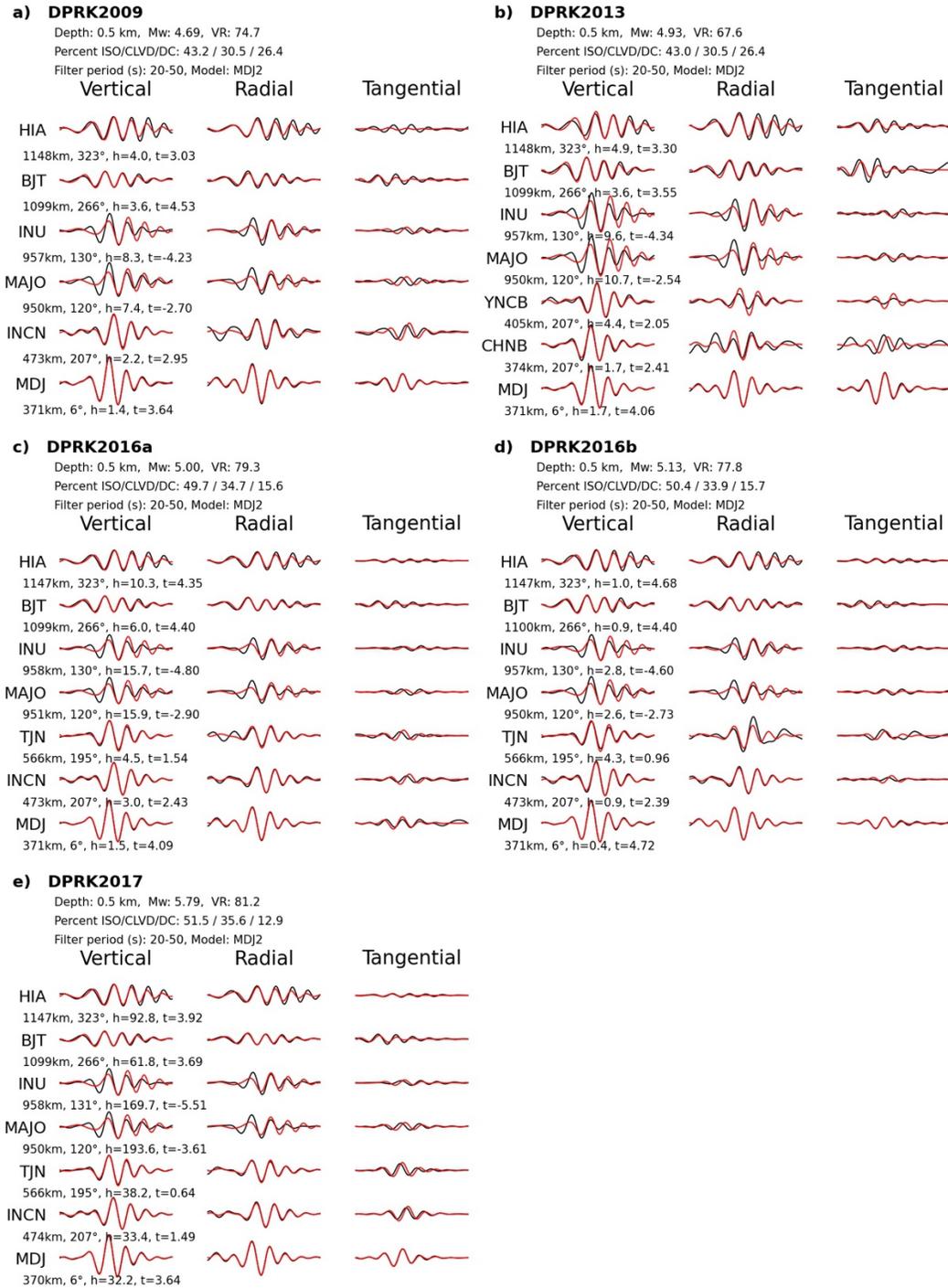
576 The station-specific uncorrelated noise levels and time shifts are recovered as free
577 parameters in the inversion. The noise parameter is relative to the standard deviation calculated
578 from 1-hour pre-event ambient noise records. As shown in Figure 10, the noise parameter of
579 MDJ is the smallest for all explosions. At the same time, MAJO and INU stations have the most
580 significant noise parameters. This result agrees with the perfect waveform fit at MDJ and the
581 poorer waveform fit at MAJO and INU stations. Note that the contribution of each station is
582 quantified by the likelihood function instead of only data noise strength because the data noise C_i
583 in Equation 9 has two competing effects on the likelihood function (Bodin et al., 2012). The
584 resulting likelihood reflects the importance of each station (Shang & Tkalčić, 2020).

585 A visual comparison of individual station contributions reveals their relative significance
586 in the overall solution. For example, Figure 11 shows the logarithm of the likelihood (log-
587 likelihood) for all stations used in the inversion for DPRK2017 (plots for the other four
588 explosions can be found in Figure S1), and the station MDJ plays the most critical role because it
589 presents the highest likelihood. The MDJ is the closest station to the sources and has a high SNR.
590 Overall, MDJ, INCN, and BJT are the most important stations that drive the DPRK2017 MT
591 inversion, while stations MAJO and INU on the Japanese side only have least contributions.



592

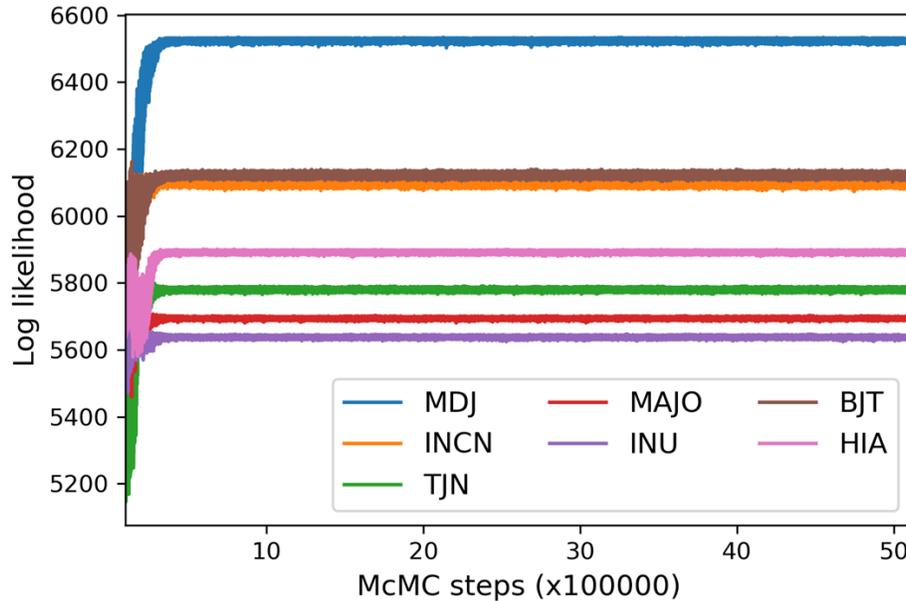
593 **Figure 9.** Source type lune diagrams for the five DPRK tests shown chronologically from 2009
 594 to 2017: (a) DPRK2009, (b) DPRK2013, (c) DPRK2016a (6 Jan 2016), (d) DPRK2016b (9 Sep
 595 2016), and (e) DPRK2017. The color bar indicates the equivalent inversion steps with the power
 596 law normalization of $2/5$ for clearer viewing of the convergence. In each panel, the overlapping
 597 color-coded dots show the MT evolution as the inversion step increases. The cyan cross is the
 598 mean MT of the convergency stage for each explosion. The resulting mechanisms are shown by
 599 the beachballs. The size of each beachball is proportional to its moment magnitude. The numbers
 600 below each beachball are a percentage of ISO, CLVD, and DC, respectively.



601

602 **Figure 10.** Fits between observed (black) and predicted (red) waveforms for the five DPRK
 603 explosions shown chronologically: (a) DPRK2009, (b) DPRK2013, (c) DPRK2016a (6 Jan
 604 2016), (d) DPRK2016b (9 Sep 2016), and (e) DPRK2017. The same 4-corner acausal bandpass
 605 (20–50 s) filter was used for each explosion. The numbers shown beneath each station are the
 606 source-station distance, azimuth, the recovered station-specific noise parameter and time shift in
 607 seconds.

608



609

610 **Figure 11.** Log-likelihood for each station in the DRPK2017 MT inversion. Most burn-in steps
 611 are discarded to illustrate the likelihood function in the convergence stage.

612 The recovered station-specific time shifts from five explosions reveal a consistent pattern,
 613 which demonstrates the robustness of our Bayesian MT inversion. Table 2 lists the station-
 614 specific time shifts from five explosions obtained in this study. Firstly, time shifts at the same
 615 stations are similar among the five explosions: three stations in China (MDJ, BJT, and HIA)
 616 have positive time shifts (up to 4.72 s), stations in South Korea (INCN, TJN, CHNB, and
 617 YNCB) have smaller positive time shifts (0.64 – 2.95 s), while two stations in Japan require
 618 negative time shifts (up to -5.51 s). The time shifts at the same station remain of the same sign
 619 even though the actual values vary in different inversions. This is because the possible errors in
 620 event origin times also contribute to the time shifts in the observed data. From the waveform fit
 621 in Figure 10, some small residual time shifts remain on the tangential components, likely due to
 622 ignoring the structures' anisotropy by applying the same time shift for all three components at
 623 each station. Treating the anisotropy using two-time shifts per station, one for vertical/radial
 624 components sensitive to vertically polarized Rayleigh waves and the other for horizontally
 625 polarized Love waves, is the subject of future studies. To summarize the results, we average the
 626 time shifts on each station for various inversions and plot their distribution with respect to the
 627 MDJ2 velocity model in Figure 12.

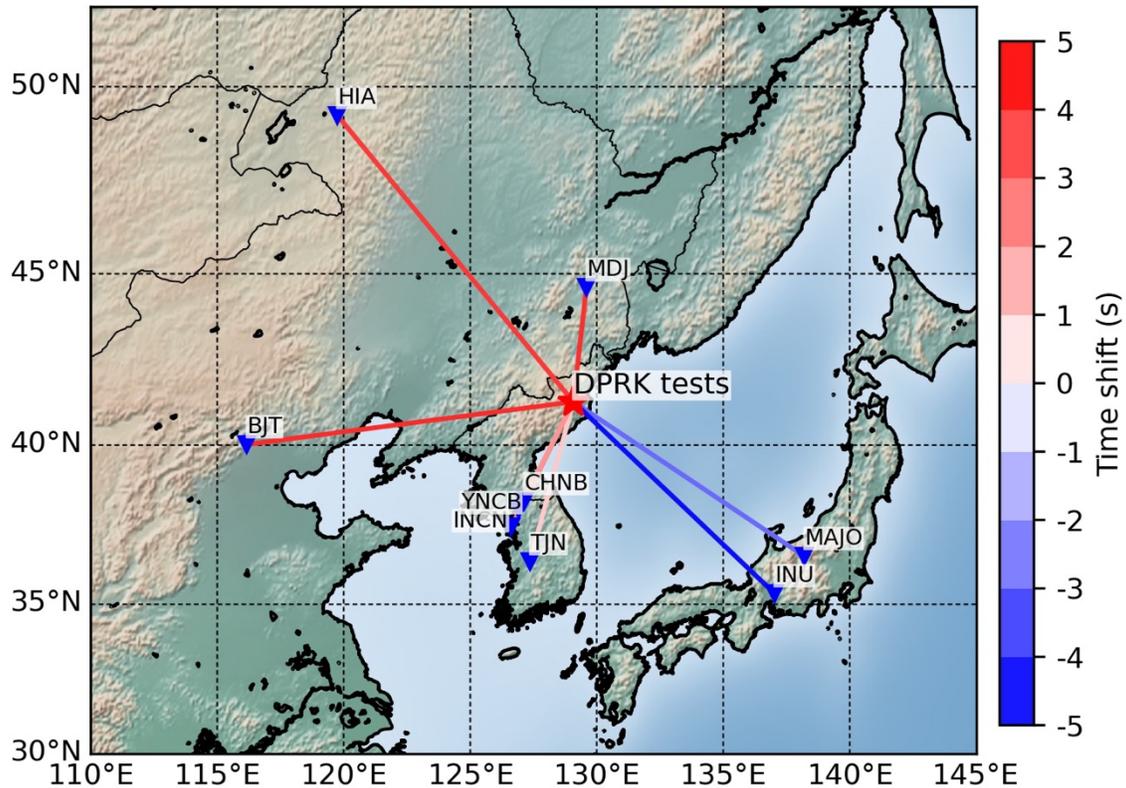
628 The distribution of station-specific time shifts coincides with the regional 2D structures
 629 surrounding the test site. In this study, the station-specific time shift between observations and
 630 predictions accounts for the possible deviation of Earth structure along specific paths with
 631 respect to the assumed 1D Earth model (i.e., MDJ2 model) for the entire study region. Positive
 632 time shifts indicate that the MDJ2 model is faster than the actual Earth's structure along these
 633 paths, while negative time shifts suggest that the MDJ2 model is slower than the actual Earth's
 634 structure. As seen in Figure 12, the Korean Peninsula is at the margin of continental crust to the
 635 west and north and oceanic crust to the east in the Japanese Sea. Therefore, the paths of surface

636 waves to stations in Japan (i.e., MAJO and INU) are sensitive primarily to the high-speed
 637 mantle, which protrudes to shallower depths beneath a thin oceanic crust. Two stations in Japan
 638 hence require negative time shifts because the MDJ2 model is slower. The paths of surface
 639 waves to stations in China (MDJ, BJT, and HHIA) are sensitive to a relatively slower, thick
 640 continental crust. Three stations in China require positive time shifts because the MDJ2 model is
 641 faster. Furthermore, the two stations in South Korea require smaller positive time shifts
 642 compared with the three stations in China. That could be due to the variation of continental crust
 643 thickness along the paths. Thus, overall, the recovered time shifts are consistent with the regional
 644 geological structures of the study region.

645 **Table 2.** Recovered station-specific time shifts (Unit: second) for five DPRK2009-2017 tests. For
 646 the DPRK2013 test, the two stations in South Korea were CHNB and YNCB.

Explosions	IC.MDJ	IC.BJT	IC.HIA	IU.INCN	KG.TJN	IU.MAJO	G.INU
DPRK2009	3.64	4.53	3.03	2.95		-2.7	-4.23
DPRK2013	4.06	3.55	3.3	2.41(CHNB)	2.05(YNCB)	-2.54	-4.34
DPRK2016a	4.09	4.4	4.35	2.43	1.54	-2.9	-4.8
DPRK2016b	4.72	4.4	4.68	2.39	0.96	-2.73	-4.6
DPRK2017	3.64	3.69	3.92	1.49	0.64	-3.61	-5.51

647



648

649 **Figure 12.** The mean time shift at each station. Positive time shifts (red) result from shifting the
 650 predicted waveforms forward, while negative time shifts (blue) stem from shifting the predicted
 651 waveforms backward.

652 4.3 Robustness of the MT inversion

653 Here we discuss the robustness of the proposed Bayesian MT inversion in three aspects.
 654 Firstly, these five DPRK explosions can arguably be considered five repetitive, shallow sources
 655 with different moment magnitudes. We used the same data preprocessing, similar source-station
 656 configuration, and the same 1D Earth model to perform the seismic source inversions. Our
 657 Bayesian MT inversion provides similar results for these five explosions, including MT solutions
 658 and station-specific time shifts.

659 Secondly, as noted above, the two stations in Japan, i.e., MAJO and INU, play a less
 660 important role than the other five stations in the source inversion for the DPRK2017 event.
 661 Therefore, we are motivated to remove these two stations and only use the other five stations in
 662 South Korea and China to invert the DPRK2017 event's MT. The solution is shown in Figure S2
 663 and is close to a crack source mechanism, with 52% ISO, 37% CLVD, 11% DC, and a moment
 664 magnitude of 5.8. It is consistent with the source obtained from seven stations in Figure 9e. The
 665 recovered station-specific time shifts and noise parameters (Figure S2c) also remain stable
 666 compared with those of the seven stations shown in Figure 10e. The variance reduction of
 667 waveform fit improves from 81.2% to 92.2% because two stations with a poorer fit are neglected
 668 in the inversion.

669

670 Thirdly, to demonstrate our approach's robustness, we use another unique dataset from
 671 seven other stations closer to the Punggye-ri test site (Figure S3a) to invert the DPRK2017
 672 event's MT. We apply the same band-pass filter to the waveforms and manually pick 150s-
 673 window waveforms. The inversion result using the same velocity model (i.e., MDJ2) shows a
 674 similar character to the previous dataset in Figure 9e. The source is dominated by an ISO=54%
 675 and is close to the crack source type. The CLVD component is up to 38%, and the DC
 676 component is negligible (only 8%), with a smaller contribution than the result shown in Figure
 677 9e. The pattern of recovered station-specific time shifts (Figure S3a) agrees with Table 2. Four
 678 stations (KSA, CHNB, CHC2, and OKEB) where the surface waves propagated through a
 679 combination of thin oceanic and thick continental crust require a slight positive time shift. Three
 680 stations (NSN, MDJ, and DACB) need more significant time shifts because the surface waves
 681 mainly propagate through the thick continental crust. In addition, these two datasets include a
 682 common station, MDJ. The time shift and noise parameter at this station from two inversions
 683 remain stable, specifically, ~ 3.6 s time shift and ~ 32 for noise parameter. Therefore, we conclude
 684 that our new hierarchical Bayesian MT inversion algorithm is robust under the same assumption
 685 of Earth's structure.

686 5 Discussion

687 5.1 The effect of the uncorrelated noise assumption

688 In this study, we assume the uncorrelated data noise using a diagonal covariance matrix
 689 C_i and focus on another, arguably more critical uncertainty (2D structural error). As
 690 demonstrated in the synthetic experiments (Section 3.3), this assumption of uncorrelated noise
 691 succeeds in the cases of high SNR (25 or larger) while failing in the cases of low SNR. From
 692 Figure 9, the MT solutions of the DPRK2009, DPRK2013, and DPRK2016a events show more
 693 considerable uncertainty than those of the DPRK2016b and DPRK2017 events. Possibly, a more
 694 comprehensive treatment of data noise should be conducted for these three explosions. For
 695 instance, Mustac et al. (2020) accounted for correlated noise with empirical noise covariance
 696 matrices, obtaining a large ISO composition (about 70%) for the DPRK2013 event at the
 697 preferable source depth of 2 km. Here, taking advantage of the affine-invariant ensemble
 698 samplers, we fix the sources at a near-surface depth, i.e., 0.5 km. This is the highlight of the
 699 present study because setting the depth near the surface in the presence of the ISO-CLVD
 700 tradeoffs was a challenging aspect in previous DPRK explosion studies.

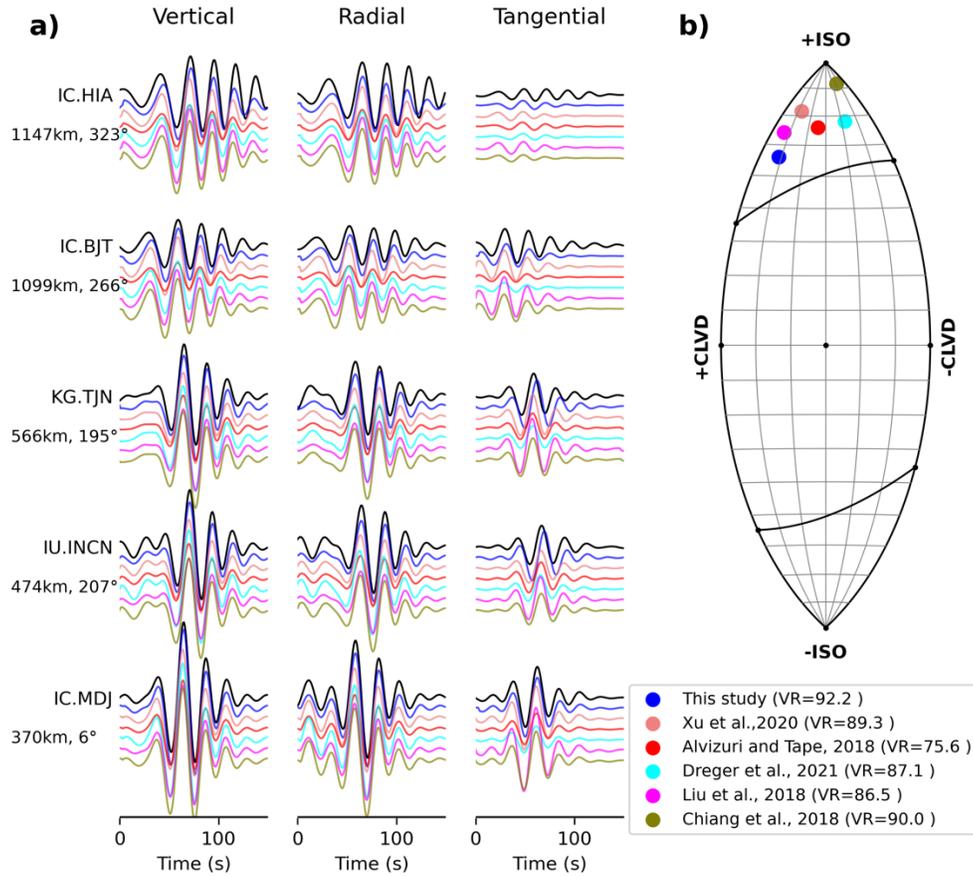
701 We note that the uncorrelated data noise is still a significant aspect of the source
 702 inversion. To illustrate its significance, we fix the noise level at each station to 1.0 instead of
 703 inverting it. This means the noise strength is assumed to be the same as pre-event ambient noise.
 704 The MT inversions for the five explosions are plotted in Figure S4. Relaxing the noise levels as
 705 free parameters increased the ISO components by $\sim 21\%$ for the DPRK2009 event, $\sim 15\%$ for the
 706 DPRK2013 event, $\sim 22\%$ for the DPRK2016a event, and $\sim 6\%$ for the DPRK2016b and
 707 DPRK2017 events. Besides, the recovered noises at different stations do not appear to have a
 708 specific pattern for the five considered explosions. This is explainable given that the ambient
 709 noise at each station could be primarily influenced by instantaneous conditions at recording sites,
 710 e.g., the seasonal variations. These five explosions happened at different times with significant
 711 time gaps.

712 5.2 Uncertainty of MT for shallow explosions

713 Previous MT inversions of the DPRK events confirmed the explosive source nature by
714 recovering a significant ISO component (Alvizuri & Tape, 2018; Chiang et al., 2018; Dreger et
715 al., 2021; Liu et al., 2018; Mustac' et al., 2020; Wang et al., 2018; Xu et al., 2020). However, as
716 we discussed, an MT inversion can suffer severe uncertainty due to several issues. Firstly, there
717 is an ambiguity between ISO and vertical CLVD mechanisms for very shallow source depths.
718 This is because the long-period waveforms at regional stations are most sensitive to the radiated
719 energy along the equator of the focal sphere with large take-off angles, where the pure ISO and
720 vertical CLVD emit similar surface waves at regional distances. Their significant difference in
721 radiation pattern happens only for small take-off angles, meaning teleseismic data are required to
722 distinguish them, as suggested by Ford et al. (2012) and Chiang et al. (2014).

723 Secondly, the region surrounding the Punggye-ri test site comprises a complex structural
724 setting (e.g., Mustac' et al., 2020), located at a margin of the continental crust in the west to the
725 oceanic crust in the east across the Sea of Japan (East Sea). Using a 1D velocity model ignoring
726 this strong 3D structure effect may result in uncertainty to MT inversion. This study uses the
727 station-based time shift between synthetics and observations to treat this significant 3D structural
728 effect on specific source-station paths.

729 Thirdly, data noise can also introduce uncertainty to MT solutions. These effects are
730 barely considered for the DPRK explosions in previous studies. As shown in Figure 13, five
731 previous studies and this study of the DPRK2017 event gave different MTs even though all of
732 them obtained a high ISO content and fit the observed waveforms with high VR, spanning from
733 75% to 95%. The differences testify to and confirm the inversion's non-uniqueness. This study's
734 moment magnitude and MTs results are most similar to those of Liu et al. (2018), using a
735 different 1D velocity model and an independent dataset in the 0.03-0.09 Hz band.



736

737 **Figure 13.** The fits between observed (black) and predicted waveforms (color-coded lines)
 738 obtained from five previous studies (see the legend) and this study for the DPRK2017 test. The
 739 predicted waveforms in this study are shifted using the recovered time shifts. In contrast, the
 740 other five sets of predicted waveforms are shifted using the times that give the highest cross-
 741 correlation coefficient to the observations. The fit levels (i.e., variance reduction) are listed in
 742 panel (b) legend.

743 6 Conclusions

744 In this study, we consider the uncertainty due to data noise involved in the data
 745 acquisition process and structural uncertainty along specific source-station paths due to imperfect
 746 knowledge of Earth structure (i.e., 2D structural error) for full MT inversion within the
 747 hierarchical Bayesian framework. The data noise on each component is assumed to be
 748 uncorrelated and measured by a standard deviation determined by an inversion in a manner of a
 749 free parameter. Besides, we use the station-specific time shifts between observed and predicted
 750 waveforms to address the 2D structural uncertainty. Unlike previous studies, the time shifts are
 751 relaxed as free parameters, determined simultaneously with noise and moment tensor parameters.
 752 We demonstrate the feasibility of this method via well-designed synthetic experiments.

753 Then we perform MT inversions for the five DPRK nuclear explosions from 2009 to
 754 2017. The MT inversion results indicate that the five explosions feature high degrees of
 755 similarity. A significant ISO component dominates their sources, i.e., 43% for the DPRK2009

756 and DPRK2013 events, and 50% for the DPRK2016a, DPRK2016b, and DPRK2017 events,
 757 respectively, which confirms the nature of the explosive source. Additionally, the five events
 758 have significant CLVD components (30%, 31%, 35%, 34%, and 36%). The DC components are
 759 small: 26%, 26%, 16%, 16%, and 13%, respectively. Relaxing the station-based data noise
 760 strength also plays a vital role in the MT inversion for DPRK explosions by increasing the ISO
 761 components. The likelihood function combining the noise and waveform residuals weights
 762 stations' contribution differently. Moreover, the recovered station-based time shifts recover the
 763 2D Earth structure character in the surrounding region of these nuclear events, demonstrating
 764 that our method appropriately accounts for the 2D structural heterogeneities.

765 Rigorously treating structural errors, especially incorporating the effects of 3D structural
 766 heterogeneity, is at leading-edge research in seismic source inversion. This study can be
 767 considered a transitional solution between incorporating the 1D to 3D Earth models in the
 768 regional MT inversion.

769 **Data Availability Statement**

770 Seismic waveform data at seven stations, MDJ, HIA, BJT, MAJO, INU, INCN and TJN
 771 used in this study are freely downloaded from Incorporated Research Institution for Seismology
 772 Data Management Center (IRIS DMC, <http://ds.iris.edu/ds/nodes/dmc/>) using ObsPy software
 773 package (Beyreuther et al., 2010). Seismic waveform data at other stations (e.g., CHNB and
 774 YNCB) come from local networks operated by the Korea Institute of Geoscience and Mineral
 775 Resources (KIGAM) and the Korea Meteorological Administration (KMA).

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 778 the figures are made with Matplotlib (Hunter, 2007).

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