

Archaeomagnetism in the Levant and Mesopotamia reveals the largest changes in the geomagnetic field

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Key Points:

- Archaeomagnetic intensity data from 23 groups of pottery collected from 18 consecutive radiocarbon-dated strata in Tel Megiddo (Israel).
- The Levantine Archaeomagnetic Curve (LAC): a Bayesian radiocarbon-calibrated archaeointensity curve of the Levant and Mesopotamia.
- Four geomagnetic spikes between 1050–600 BCE define new constraints for maximum field intensity and secular variation rates.

Abstract

Our understanding of geomagnetic field intensity prior to the era of direct instrumental measurements relies on paleointensity analysis of rocks and archaeological materials that serve as magnetic recorders. Only in rare cases absolute paleointensity datasets are continuous over millennial timescales, in sub-centennial resolution, and directly dated using radiocarbon. As a result, fundamental properties of the geomagnetic field, such as its maximal intensity and change rate have remained a subject of lively discussion. Here, we place firm constraints on these two quantities using Bayesian modeling of well-dated archaeomagnetic intensity data from the Levant and Upper Mesopotamia. We report new data from 23 groups of pottery collected from 18 consecutive radiocarbon-dated archaeological strata from Tel Megiddo, Israel. In the Near East, the period of 1700–550 BCE is represented by 84 groups of archaeological artifacts, 55 of which were dated using radiocarbon or a direct link to clear historically-dated events, providing unprecedented sub-century resolution. Moreover, stratigraphic relationships between samples collected from multi-layered sites enable further refinement of the data ages. The Bayesian curve shows four geomagnetic spikes between 1050 and 600 BCE, with virtual axial dipole moment (VADM) reaching values of 155–162 ZAm^2 – much higher than any prediction from geomagnetic field models. Rates of change associated with the four spikes are ~ 0.35 – 0.55 $\mu\text{T}/\text{year}$ (~ 0.7 – 1.1 ZAm^2/year), at least twice the maximum rate inferred from direct observations spanning the past 190 years. The increase from 1750 BCE to 1030 BCE (73 to 161 ZAm^2) depicts the Holocene's largest change in field intensity.

Plain Language Summary

The strength of Earth's magnetic field is changing in an unpredictable manner. Understanding these changes requires precise information on how the field has changed in the past. Direct instrumental measurements of magnetic field intensity began in the 1840s, providing only a short time window into past intensity changes. Here, we explore the more ancient field by analyzing a rare collection of radiocarbon-dated archaeological materials from stratified archaeological settlements and historically-dated burnt structures in the Levant and Mesopotamia. We use new data from Tel Megiddo (Armageddon) to construct a continuous curve of geomagnetic field intensity spanning 2500 years, with unprecedented detail and resolution. The curve depicts the evolution of a high-intensity anomaly, the largest change in intensity observed during the Holocene. Between 1750 and 1050 BCE, the field rapidly increased to values greater than twice those of today, much higher than any prediction derived from available geomagnetic field models. Subsequent oscillations between 1050 and 550 BCE, with extreme peaks, namely ‘geomagnetic spikes’, reveal change rates of at least twice as fast as the fastest change observed since the advent of direct measurements. Levantine archaeomagnetic data represent a case study in which archaeology provides crucial constraints on the geomagnetic field behavior.

1. Introduction

The absolute intensity of the geomagnetic field was first measured by Carl Friedrich Gauss in 1832 (Courtillet & Le Mouel, 2007). Subsequent measurements with improved precision and spatial resolution have provided quantitative estimates of the amplitude and rate of geomagnetic field intensity changes, but only across the past two centuries. Information from periods preceding observational measurements, fundamental for understanding the magnetic field

behavior, is derived from ancient materials that acquired thermoremanent magnetization upon cooling from high temperatures. For the past several millennia, archaeological materials have been the primary source for this information (e.g., Arneitz et al., 2017; Brown et al., 2021), providing most of the data for late Holocene geomagnetic models (Arneitz et al., 2019; Campuzano et al., 2019; Constable et al., 2016; Nilsson et al., 2014; Nilsson et al., 2022; Panovska et al., 2019; Pavon-Carrasco et al., 2014). However, global models are smoothed by design, as they require a tradeoff between model complexity and fit to heterogeneous data. Regional intensity curves provide important insights into field behavior (e.g., Cai et al., 2017; Garcia et al., 2021; Genevey et al., 2016, 2021; Rivero-Montero et al., 2021; Schnepf et al., 2020), but they depend on the quality of the underlying source data. One of the most significant limiting factors for both global and regional field modeling is the limited precision and accuracy of the published ages. Only ~12% of the global published data from the past 10ky data are directly dated with radiocarbon, and in many cases, the exact nature and context of the dated material are not documented. Instead, most archaeomagnetic ages are based on assignment to regional archaeological chronologies, which may have differing interpretations, can be poorly tied to absolute ages, or have large age ranges (e.g., Shaar et al., 2020). Furthermore, unlike paleomagnetic field direction reconstructions, which use stratigraphic constraints in sedimentary sequences to obtain continuous time series, archaeomagnetic intensity datasets are mostly sporadic in time and space. Given the overall uncertainties in the available paleomagnetic and archaeomagnetic data, some of the most fundamental properties of the geomagnetic field, such as its maximum intensity and maximum possible change rate, have remained elusive and the subject of a lively and fruitful debate (Davies & Constable, 2018; Korte & Constable, 2018; Livermore et al., 2014; Livermore et al., 2021).

One way to improve the resolution of archaeomagnetic data is to focus efforts on large multi-layered archaeological sites, which are composed of distinct consecutive strata, and can provide data in stratigraphic order, e.g., Mari (Tell Hariri), Tell Atij and Tell Gueda in Upper Mesopotamia (Gallet & Butterlin, 2015; Gallet et al., 2020; Gallet et al., 2006; Gallet et al., 2008), Ebla (Tell Mardikh) in Northern Levant (Gallet et al., 2014; Gallet et al., 2008), and Tel Hazor in Southern Levant (Shaar et al., 2020; Shaar et al., 2016). Although these are key sites for Near Eastern archaeology, most of their archaeomagnetic data are not radiocarbon-dated. From this perspective, Tel Megiddo (Israel), with a radiocarbon-based age model covering timespan of 3000-735 BCE time-span, is unique, providing an unprecedented opportunity to construct a stratigraphically constrained radiocarbon-calibrated archaeomagnetic time series.

In the following, we report the data obtained from Tel Megiddo, which to date, is the largest archaeomagnetic intensity dataset available from a single site. We compile the new data with other archaeomagnetic data from the Levant and Mesopotamia that pass our selection criteria. The temporal resolution of the combined data between the 18th and the 6th century BCE is a century or less, as most of the archaeomagnetic ages in this period are derived from radiocarbon-dated contexts and historically-dated burnt structures. Using this high-precision compilation, we develop the Levantine Archaeomagnetic intensity Curve (LAC), utilizing a Bayesian methodology. The LAC elucidates the details of the largest geomagnetic change in the Holocene, associated with the Levantine Iron Age Anomaly (Shaar et al., 2016) and the occurrence of geomagnetic spikes (Ben-Yosef et al., 2009; Shaar et al., 2011; Shaar et al., 2016). We use the LAC to enhance knowledge of the number, duration, and intensity of geomagnetic spikes, which define new robust upper limit constraints for both maximum field intensity and change rate.

2. Archaeomagnetic intensity stratigraphy of Tel Megiddo

2.1. Background

Tel Megiddo (32.585N, 35.185E, Fig. 1) is a world-heritage archaeological site located on the western margins of the Jezreel Valley in northern Israel. Owing to its strategic location on the international route which connected Egypt with Mesopotamia, Megiddo was a central city and an important administrative center throughout the Bronze and Iron Ages (ca. 3500 – 600 BCE). Extensive excavations of the mound have revealed more than thirty Bronze and Iron Age superimposed settlements, with several destruction layers indicating violent endings in military campaigns (Finkelstein, 2009). The chronology of the entire Megiddo sequence was established from Bayesian analyses of ca. 150 radiocarbon samples (the total number of radiocarbon samples at Megiddo is 185) carefully collected from nearly all strata (Boaretto, 2022; Martin et al., 2020; Regev et al., 2014; Toffolo et al., 2014). Special care was taken in assembling the radiocarbon model from mostly short-lived organic materials strongly linked to the archaeological findings. The exceptionally large radiocarbon data from a detailed, continuous, and well-established stratigraphy, along with the intensive ceramic record of Megiddo that defines the relative dating of the region (e.g., late Iron I, early Iron IIA), provides a robust absolute chronology for Near Eastern archaeology.

The archaeomagnetic stratigraphy of Tel Megiddo is based on twenty-three different contexts recovered from 18 layers, excavated in six excavation areas (Fig. 2). Ten contexts (S-3, H-15, K-6, H-11, K-4/H-9/Q-7, Q-4, H-3/Q-2) are destruction layers with distinct boundaries and clear marks of their ending. Megiddo's final destruction by the Assyrian Tiglath-Pileser III is conclusively dated to 732 BCE, based on multiple historical documents. We sampled fragments of indicative pottery from each context, with emphasis on local domestic material. We preferred, when possible, complete or cured vessels that were photographed and documented in the excavation reports. In three contexts (Q-4, Q-5, K-9), we also sampled fragments of cooking ovens (tabuns). The fragments (termed hereafter ‘samples’ for consistency with previous publications) discussed here include new data, as well as data already published in Shaar et al. (2016) and Shaar et al. (2020), which reported the initial archaeomagnetic stratigraphy of Megiddo. Supplementary Table S1 lists the archaeological details of all the materials analyzed in this study.

2.2. Archaeointensity experiments

Thellier-IZZI-MagIC paleointensity experiments were conducted in the shielded paleomagnetic laboratory at the Institute of Earth Sciences, the Hebrew University of Jerusalem, using two modified ASC TD-48 ovens and a 2G-RAPID superconducting rock magnetometer (SRM). Specimens were prepared by gluing small pieces of pottery inside non-magnetic $22 \times 22 \times 20$ mm square alumina crucibles. The protocol followed the IZZI method (Tauxe & Staudigel, 2004; Yu et al., 2004) with routine pTRM checks at every second temperature step using an oven field of 40, 50, or 60 μ T. Heating time ranged from 40 to 65 minutes, depending on the target temperature. In total, each IZZI experiment included 31 or 33 heating steps at 13 or 14 temperature intervals between 100°C to 590°C or 600°C. All specimens were subjected to anisotropy of thermoremanent magnetization (ATRM) experiments, which consisted of eight heating steps at 590°C or 600°C: a baseline zero-field step, six infield steps at orthogonal

directions, and an additional alteration check. ATRM alteration parameter was calculated following Shaar et al. (2015) (Table 1). For specimens with ATRM alteration checks > 6%, anisotropy of anhysteretic remanent magnetization (AARM) was measured at a 100mT AC field in 0.1mT DC bias field, at six orthogonal directions, after thermal demagnetization of the specimens. All specimens were subjected to cooling rate correction experiments, which consisted of 4–5 cooling steps from 590 °C or 600 °C to room temperature, following the protocol described in Shaar et al. (2020). Archaeointensity values were calculated with the *Thellier-GUI* program (Shaar & Tauxe, 2013), incorporated into the *PmagPy* software package (Tauxe et al., 2016), using *Thellier Auto Interpreter* algorithm and the acceptance criteria listed in Table 1. Sample results were calculated by averaging at least 3 specimens per sample using the STDEV-OPT algorithm of the *Thellier-GUI* program and the ‘extended error bounds’ approach (Shaar & Tauxe, 2013; Shaar et al., 2016). When averaging sample data in ‘groups’ (see section 2.3), we calculated a simple mean of the STDEV-OPT values of the samples. A detailed description of the methods can be found in Shaar et al. (2016) and Shaar et al. (2020). All measurement data are available in the MagIC database (earthref.org/MagIC/19395).

2.3. Results

The archaeomagnetic data from Tel Megiddo, including the data already published in Shaar et al. (2016) and Shaar et al. (2020), include 763 specimens from 175 samples. In this study, we analyze 288 specimens from 85 newly collected samples. In total, 583 specimens and 132 samples pass the criteria listed in Table 1, where archaeointensities obtained at the sample level are calculated from a minimum of 3 specimens. Fig. 3 shows representative cases of a successful specimen and interpretations failing criteria. The importance of the anisotropy and cooling rate corrections is illustrated in Fig. 4. Typically, the bias due to anisotropy and cooling rate effects is 5%–15%; in some cases, the combined corrections exceed 20%. Table S2 (Supplementary Material) lists specimens results, anisotropy and cooling rate correction factors, and values of paleointensity statistics listed in Table 1.

Fig. 5 displays sample data with error bars calculated using the “extended error bounds” approach (Shaar & Tauxe, 2013). In general, samples collected from the same archaeological context (termed hereafter ‘group’) show good agreement with only two outliers in groups K-4 and Q-4. Levels H-9 and H-3 exhibit a large dispersion of data, with distinctively different values. As the ceramics in each context represent production during a time interval rather than a singular point in time, this probably indicates fast changes in the field during the interval represented in the ceramic assemblage. Thus, we tentatively split the results from these two contexts into two subgroups. The mean archaeointensity of each group is calculated by averaging the sample means. Detailed sample data are provided in Table 2 and Supplementary Table S3. The archaeomagnetic stratigraphy presented in Fig. 5 (Table 2) shows exceptionally large amplitude changes – between ~39 and ~90 μ T. In the following sections, we explore in detail this amplitude change depicted in the Megiddo data.

The ages of the groups are based on the Bayesian age model of Megiddo, which is assembled from ~150 radiocarbon samples and takes into account the stratigraphic relationships between strata and correlation between levels excavated at different areas (e.g. Fig. 2a). The published 68.2% and 95.4 % probability age intervals (Boaretto, 2022; Martin et al., 2020; Regev et al.,

2014; Toffolo et al., 2014) are shown in Table 2. The actual age ranges we use for the archaeomagnetic analysis overlap the radiocarbon age ranges, but are not identical. This distinction is performed to optimally represent the age range of the ceramic assemblage in each archaeological context, considering the entire archaeological and historical evidence. Thus, the age range of a group may be as short as 50 years for sequences of short-lived phases punctuated by well-dated destruction layers (e.g. Q-4,Q-5,Q-6,Q-7) or 250–300 years for less-constrained strata (e.g. J-4,J-5,J-6).

3. The Levantine Archaeomagnetic Curve (LAC.v.1.0)

The Levantine Archaeomagnetic Curve (LAC) is designed to enable both the statistical analysis of secular variation properties and archaeomagnetic dating. In this article, we focus on the geomagnetic implications of the curve, whereas in a sister article (Vaknin et al., in press), we demonstrate the applications of the curve for correlating ancient historical events. Preliminary versions of the LAC, using an identical methodology as in this study but spanning different time intervals, were published in Shaar et al. (2020) and Gallet et al. (2020). Here, we provide a short description of the different datasets used to construct the LAC and briefly outline the underlying methodology of the Bayesian analysis. A more detailed description of the archaeomagnetic methods, selection criteria, and our approach for sorting and organizing the published data can be found in Shaar et al. (2020). A complete description of the Bayesian method is given in Livermore et al. (2018).

3.1. Data compilation: experimental guidelines

Archaeomagnetic studies in the Levant and Mesopotamia, in the area extending from Egypt in the south to southern Turkey in the north, yielded an incredibly large archaeomagnetic dataset, which includes more than 722 archaeointensity estimates published between 1969 and 2021 and available in the GEOMAGIA50 database (Brown et al., 2015; Brown et al., 2021), the MagIC database (Tauxe et al., 2016), and the ArcheoInt compilation (Genevey et al., 2008). Yet, when these data are simply stacked together, significant discrepancies are evident (Supplementary Material, Fig. S1). From an experimental perspective, screening out the most robust data and assigning a consistent archaeointensity uncertainty to the overall data are not trivial tasks due to large differences in laboratory methods, data analysis approaches, and selection criteria. Moreover, not all of these data were published along with the raw measurement data, preventing a rigorous and identical calculation of the experimental uncertainty. We, therefore, adopt an approach that utilizes only methods that were tested against each other in different laboratories and shown to yield statistically indistinguishable results at the sample level:

- Thellier-IZZI-MagIC: This method incorporates the Thellier-IZZI protocol (Yu et al., 2004) exactly as applied in this study. The automatic interpretation procedure follows the STDEV-OPT algorithm (Shaar & Tauxe, 2013) with the ‘LAC criteria’ provided in Table 1. All measurement data are available in the MagIC database (<https://www.earthref.org/MagIC>) and can be re-interpreted using any set of alternative selection criteria.
- Triaxe (Le Goff & Gallet, 2004) or the Triaxe + Coe (Gallet & Le Goff, 2006): The Triaxe method was tested against the Thellier-IZZI-MagIC in a blind test in Shaar et al. (2020) and yielded indistinguishable results. The Triaxe + Coe includes groups of specimens, which

were analyzed using both the Thellier-Coe method (Coe et al., 1978) by Genevey et al. (2003) and the Triaxe method by Gallet and Le Goff (2006), and found to be equivalent.

Other data will be included in future LAC compilations if the measurement data can be analyzed using identical procedures and selection criteria as the rest of the LAC data.

Another aspect of the LAC compilation is associated with data hierarchy. A portion of the published archaeomagnetic data was reported as averages of specimens prepared from the same mother sample, while another portion was reported as averages of specimens or samples collected from the same archaeological context. In the LAC compilation, we use the latter approach to avoid over (under)-representation of contexts with more (fewer) samples and to ensure that uncertainties are calculated consistently. Thus, each datum in the LAC compilation represents a ‘group’, where a group can be, for example, a collection of indicative pottery from a specific stratum, fired mud bricks from a burnt structure, a layer in a slag mound, or storage jars with identical stamp types. The archaeointensity value of a group is calculated as a simple average of the samples’ means after screening out outliers (e.g., K-4 and Q-4 in figure 5). Two exceptions to this rule are related to destruction layers: a kiln from Horvat Tevet that had gone out of use when the site had been destroyed (Vaknin et al., in press) and a clay-made floor burnt during the historically-dated Babylonian destruction (Vaknin et al., 2020); in these cases, a large number of specimens collected from the same thermal unit are averaged.

Fig. 6 displays 142 groups between 3000 – 500 BCE passing the experimental criteria. In order to minimize effects related to spatial variability of the field, we constrain the geographic distribution of the data to the region that extends between southern Israel, Cyprus, northern Syria, and eastern Syria (Fig. 1), comprising a circle with a radius of ~500km. Data are displayed in terms of virtual axial dipole moment (VADM) – a transformation from local, latitude-dependent field intensity measurement to the equivalent geomagnetic axial dipole moment. Data from Syria (analyzed using the Triaxe or Triaxe+Coe methods) representing Mesopotamia and northern Levant (Gallet & Al-Maqdissi, 2010; Gallet & Butterlin, 2015; Gallet et al., 2014; Gallet et al., 2020; Gallet et al., 2006; Gallet et al., 2008; Genevey et al., 2003; Livermore et al., 2021), were reported as groups and displayed as published with few minor updates on the ages of some fragment groups (Supplementary Text S4 and the mentioned references for all the data obtained at the specimen/fragment levels). Data from Timna-30 slag mound (Shaar et al., 2011), Tel Hazor (Shaar et al., 2016), and the Judean stamped jars (Ben-Yosef et al., 2017), which were published as samples, are averaged to represent group means (Supplementary Tables S4–S9).

3.2. Data compilation: age estimation guidelines

We distinguish between two sets of data with essentially different approaches for age estimation. The ages of the first set, marked in gray in Fig. 6, were assigned by the excavators of the sites using a complex body of archaeological evidence that does not include absolute radiocarbon ages directly associated with the archaeointensity data. This raises two problems for paleomagnetists and modelers. First, tracing back the considerations used to determine the ages requires specific archaeological expertise, and, therefore, the quality, precision, and robustness of these ages cannot be easily assessed without a detailed description of the age data. Second, in many cases, the archaeological time scales, based on ceramic typology and cultural changes, might be loosely linked to an absolute age scale and may have different age interpretations. Our approach in these

cases is to make as few changes as possible to the archaeological ages assigned by the excavators but rather to use wide age range that considers all possible correlations to the absolute age scale.

The ages of the color-coded datasets in Fig. 6 are assigned using radiocarbon or direct associations to historical events whose ages are considered consensuses by most of the archaeological community. The latter are significant for the time interval associated with the Hallstatt Plateau (ca. 800–400 BCE) in the radiocarbon calibration curve (Reimer et al., 2020). From the 8th to 6th century BCE, the ages are based on a correlation to two precisely-dated historical military campaigns described in the Hebrew Bible and other Mesopotamian texts - Assyrian (733–701 BCE) and Babylonian (600 – 586 BCE) – rather than on radiocarbon. In addition, the Aramean occupation (845–815 BCE), which is dated using both radiocarbon and historical constraints, is also used as a useful chronological tie point. The groups with age ranges tied to absolute ages are marked by four different colors in Fig. 6, and include the following datasets:

- The radiocarbon-dated stratigraphy of Tel Megiddo described in this study, which ended in the Assyrian destruction of the city.
- Two radiocarbon-dated layers from Tel Hazor (Stratum XVIII, Stratum XII) and a sequence of three stratigraphically ordered, short-lived phases from strata V-VI that ended in the Assyrian destruction of Hazor (Text S1, Tables S4-S5, supplementary material).
- A radiocarbon-dated sequence of ten slag layers from Timna-30. The Bayesian age model of the mound (Shaar et al., 2011), which was originally established using a magnetostratigraphic correlation with Khirbet en-Nahas (Ben-Yosef et al., 2009), is revised here to include only radiocarbon samples collected from Timna-30 (Text S2, Tables S6-S8, Supplementary Material)
- Materials dated by association to the Assyrian and Babylonian occupations or to the radiocarbon-dated Aramean campaign. This dataset includes 19 burnt structures (Vaknin et al., 2020; Vaknin et al., in press) and three groups of indicative ceramics that can be dated by association with cultural changes related to the occupations. The fired mud-brick structures found in the burnt level are crucial tie points for two reasons. First, as mentioned above, the uncertainty in radiocarbon dating in this period is in the order of 200–400 years due to the plateau in the calibration curve, while the dates of the historical campaigns are unique in their precision. Second, the burnt bricks record a single event, as the fire during the destruction resets their magnetization, in contrast to pottery groups that provide data over a time interval representing a production period.

Supplementary Table S10 lists the VADM and the age range of 142 groups included in the LAC.v.1.0 compilation. We stress that none of the ages in the LAC data compilation were determined or constrained using archaeomagnetism in order to avoid circular reasoning.

3.3. Bayesian modeling

With the data described in Section 3.2 and listed in Supplementary Table S10, we calculate a Bayesian curve with its corresponding 95% credible envelope (Fig. 6, Supplementary Table S11). We term this curve ‘Levantine Archaeomagnetic Curve version 1.0’, or LAC.v.1.0. The LAC is calculated using the age hyperparameter reverse-jump Monte Carlo Markov Chain (AH-RJMCMC) algorithm developed by Livermore et al. (Livermore et al., 2018) (<https://github.com/plivermore/AH-RJMCMC1>). The algorithm is based on a piece-wise linear interpolation of the data between vertices drawn in a random-walk-like perturbation within a space allowed by the acceptance criteria. The prior assumptions of the model are: i) the allowed range of vertices’ VADM values is set to between 60 and 200 ZAm²; ii) the allowed number of vertices (K) is between $K_{\min} = 1$ and $K_{\max} = 150$; iii) ages in all contexts are uniformly distributed, except the ages of Timna, which were modeled as a normal distribution; and iv) group means and standard deviations define a normal distribution of the archaeointensity data. In addition, Supplementary Table S10 defines a stratigraphic order for contexts collected from the multi-layered sites (Tel Megiddo, Tel Hazor, Tell Atij, Tell Gudeda, Timna) and few mutual constraints between groups in Megiddo and Hazor. The AH-RJMCMC procedure takes into account all of these temporal relationships. The model uses the parameters $\sigma_{\text{move}} = 30$ yrs, $\sigma_{\text{change}} = 10$ Z Am², and $\sigma_{\text{birth}} = 10$ Z Am², which define the random perturbation of a vertex in age, in intensity, and that of the linearly interpolated intensity value of a new vertex based on the current vertex distribution respectively. The age of a single datum is perturbed per age-resampling step ($\text{num_age_changes} = 1$); chain length is $2 \cdot 10^8$.

The sub-centennial resolution of the curve from 1100 to 550 BCE (encompassing the Levantine Iron Age anomaly) is achieved through several unique features of the combined datasets. Firstly, we obtained radiocarbon-dated contexts with age uncertainties of approximately a century and, in several cases, even less. Secondly, the stratigraphic relationships in Timna, Megiddo, and Hazor define constraints to the Bayesian model that lead to a reduction in the posterior age ranges. Lastly, we included data obtained from historically well-dated burnt levels and used a dense dataset with a large number of groups during the spike period.

4. Discussion

4.1 New constraints on the highest geomagnetic field intensity

Considering all the published paleointensity estimates from individual samples (i.e., not group means) from the past 5 My available in the GEOMAGIA50 v.3.3 (Brown et al., 2015) and PINT v.8.1.0 (Biggin et al., 2009; Bono et al., 2022) databases, only 1% of the data, which are sporadically scattered in time and space, show VADM > 150 ZAm². As such, VADM values calculated from global geomagnetic models do not exceed 140 ZAm² (e.g., Arneitz et al., 2019; Constable et al., 2016; Korte & Constable, 2018; Panovska et al., 2019; Pavon-Carrasco et al., 2014). The only exception is the time interval between the end of the 2nd millennium BCE and the middle of the 1st millennium BCE, where a number of archaeomagnetic observations point to high field values (>150 ZAm²) at several locations: the Levant (Ben-Yosef et al., 2017; Ertepinar et al., 2012; Shaar et al., 2011; Shaar et al., 2016; Vaknin et al., 2020), Caucasus (Shaar et al., 2017), China (Cai et al., 2017), Bulgaria (Kovacheva et al., 2014), Spain (Osete et al., 2020), Canary Islands (Kissel et al., 2015), Azores (Di Chiara et al., 2014) and Hawaii (Pressling et al., 2006). All these observations suggest short duration for the episodes of high intensity values. This behavior is probably associated with a more complex field structure than today’s (Korte &

Constable, 2018; Osete et al., 2020; Rivero-Montero et al., 2021) and, presumably, with a local high field anomaly in the Near East, termed the ‘Levantine Iron Age Anomaly’ (LIAA) (Shaar et al., 2018; Shaar et al., 2017; Shaar et al., 2016).

The highest VADM values during the climax of the LIAA were termed ‘geomagnetic spikes’ by Ben-Yosef et al. (2009) and Shaar et al. (2011). Note that we use the term ‘spike’ hereafter in a dual sense: first, to describe a short time interval (about a century long) with rates of intensity change far exceeding those observed in the modern era (1840-2020), which would potentially allow for other spikes to be observed outside of the LIAA, and second, to describe short-lived intensity peaks exceeding the typical values in the geological record. Livermore et al. (2021) questioned the robustness of the spikes and stated that the number of spikes and their values strongly depend on the archaeomagnetic data used, particularly the experimental errors and the averaging scheme adopted (i.e., sample groups versus individual samples). Here, we address the issues raised by Livermore et al. (2021) and assemble a much denser dataset based solely on group averages. This way, each data point in our compilation represents exactly the same quantity and gains the same weight in the Bayesian calculation process. The new curve shows the occurrence of four spikes with peak VADM of 155–162 Z Am² around 1030, 840, 740, and 600 BCE. Each spike is represented by several coeval or nearly coeval groups, where overall, fourteen groups have VADM > 150 Z Am². Considering that each group represents a time average of several samples, we suggest a value of 155 Z Am² as a robust and conservative upper limit for the maximum field value. Yet, based on sample data, higher values may have occurred for short time intervals.

Fig. 7 demonstrates that spike-like values are rare in the paleomagnetic record, showing all the published absolute paleointensity data with ages older than 1500 BCE from the GEOMAGIA50 and PINT databases. Only 49 data points out of 6816 have field values higher than 155 Z Am², none of which passes the rather strict statistical tests applied in the LAC. Given the specific conditions associated with the Levantine geomagnetic spikes, the difficulty in detecting similar high-paleointensity values in the global paleointensity record is understandable. First, our dense dataset, including ten archaeointensity groups on average (each consisting of at least two samples) per century during the LIAA interval (Fig. 6a) shows that the duration of the peaks is around a century. Thus, if an average of few samples is required to obtain a robust paleointensity estimate, a large dataset, such as the LAC compilation, is required to detect spikes. Second, global geomagnetic models indicate that the geomagnetic dipole during the LIAA is most likely the highest in the Holocene (Constable et al., 2016; Pavon-Carrasco et al., 2014; Schanner et al., 2022). Thus, the likelihood of detecting spikes may depend on the likelihood that the ancient paleomagnetic dipole was similarly high. Third, the spikes are a regional feature associated with a local geomagnetic anomaly, expressed not only by high field values but also by directional deviations from a dipole field (Osete et al., 2020; Shaar et al., 2018; Shaar et al., 2016). Consequently, there are low chances that the scattered and sparse paleointensity database spanning the geological record can reveal short-lived spikes. Moreover, from the comparison with the global paleointensity database, we can conclude that spikes represent the highest value the geomagnetic field can reach and can serve as a robust upper boundary for the maximum strength of the geomagnetic field at a given location.

4.2 New constraints to maximum secular variation rates

The rates associated with the spikes range between $\sim 0.35 - 0.55 \mu\text{T}/\text{year}$ or $0.7 - 1.1 \text{ Z Am}^2/\text{year}$ in VADM values (Fig. 6c). To place these values within the context of the global geomagnetic field behavior, we calculate in Fig. 8 the maximum rate in today's field by observing the difference between IGRF models epochs 2015 and 2020 (Alken et al., 2021). For most of Earth's surface, the rates do not exceed $0.1 \mu\text{T}/\text{year}$ ($\sim 0.2 \text{ Z Am}^2/\text{year}$), and a maximum rate of $0.12 \mu\text{T}/\text{year}$ occurs only in limited areas (Fig. 8a). VADM transformation accounting for the latitudinal dependency of the field yields a maximum rate of $0.25 \text{ Z Am}^2/\text{year}$ (Fig. 8b). We expand the calculation back to 1840 using the *gufm1* model (Jackson et al., 2000) for 1840–1945 and the IGRF models for 1950–2020 (Alken et al., 2021) by looking at the maximum difference in field intensity every five years at any point on Earth's surface. In our calculation we ignored the interval 1945–1950 as we noticed an abrupt and very time-limited change in the rate at few locations, which is likely an artifact caused by problematic global coverage of the data in the 1945 model. The maximum change rate in the past 190 years is not significantly different from today, i.e., $0.18 \mu\text{T}/\text{year}$ or $0.33 \text{ Z Am}^2/\text{year}$. These rates are considerably lower than those observed during the time intervals associated with geomagnetic spikes. Hence, LAC.v.1.0 also places new robust constraints on how fast local field intensity can change.

We note that the rates calculated using LAC.v.1.0 appear more moderate than previously considered (Ben-Yosef et al., 2009; Shaar et al., 2011), which had raised questions regarding the dynamo processes behind them (Livermore et al., 2014; Troyano et al., 2020). Although the spikes are now within the range of variations permitted by our current understanding of the geodynamo (Davies & Constable, 2017; Davies & Constable, 2018; Livermore et al., 2014), the fluctuations associated with the spikes remain unprecedented in their amplitude and rate.

4.3. Long-term evolution of the Levantine Iron Age Anomaly

The archeo-magnetostratigraphy of Tel Megiddo reveals the Holocene's greatest amplitude change at multi-century scales between 1750 BCE and 1030 BCE. The increase began with a minimum of 73 Z Am^2 in the 18th century BCE, a period characterized by the lowest intensities in the Near East over the last five millennia, comparable to the low values at the beginning of the 3rd millennium BCE (Fig. 6b). The 700 year-long increase is nearly continuous, punctuated by a century-scale peak around 1500 BCE. The spike period may, therefore, be a climax of a long-term evolution in the geomagnetic field intensity in the Near East, which could result from the occurrence of intense and rapidly evolving flux bundles at the core-mantle boundary. It also contrasts with the period spanning from the 3rd millennium BCE to the 18th century BCE, which shows intensity peaks of moderate amplitudes associated with lower change rates of less than $0.2 \mu\text{T}/\text{year}$ (Figs. 6b,c; see also Gallet et al. (2020)). The significant increase in intensity would then result in spikes instead of intensity peaks, with the first events also being shorter and apparently more frequent, which could reflect a remarkable change in core dynamics.

5. Conclusions

We report here the largest archaeomagnetic intensity dataset currently available from a single site, i.e. Tel Megiddo, with 23 sample groups collected from 18 consecutive radiocarbon-dated archaeological strata.

We assemble a new archaeomagnetic compilation of the Levant and Upper Mesopotamia between 3000 BCE to 550 BCE with 142 different groups of samples. The interval from 1700 BCE to 550 BCE is based mostly on contexts directly dated using radiocarbon and clear associations with well-dated historical military campaigns, providing an unprecedented sub-century resolution. We use this compilation to calculate the Levantine Archaeomagnetic Curve (LAC.v.1.0), a Bayesian regional curve for high-precision archaeomagnetic dating.

The LAC depicts four geomagnetic spikes between 1050 BCE and 600 BCE, each lasting about a century, defining new upper limits on both the maximum local field values and change rate. Considering the overall uncertainty, we suggest 155 ZAm^2 and $0.5 \text{ } \mu\text{T/year}$ ($1.0 \text{ ZAm}^2/\text{year}$ in VADM values) as conservative upper boundaries for these quantities.

As a concluding remark, we highlight the challenge in constructing a robust, continuous geomagnetic intensity curve at a sub-centennial resolution over a large millennial time scales. This is made possible in this study by exploiting the advantage of the Near East's abundance of data that allows acquisition of large multiple archaeomagnetic datasets with precise dating, stratigraphic constraints, and cross-correlations between sites. In this respect, Tel Megiddo is an exemplary case study demonstrating a strong link between archaeology, radiocarbon, and geomagnetism.

Acknowledgments

This project has received funding from the European Research Council (ERC) under the European Union's Horizon 2020 research and innovation programme (grant agreement No 804490) to RS. The study was partly supported by the Israel Science Foundation (ISF) grant 1364/15 to RS. It was partly financed by the INSU-CNRS program PNP to YG. Work on the Megiddo samples was supported by the Dan David Foundation and grants from Mark Weissman and Jacques Chahine to IF. We thank Phil Livermore for assisting with the AH-RJMCMC analysis. We thank Maxwell Brown and the anonymous reviewer whose constructive comments helped improve the manuscript.

Open Research

All measurement data are available in the MagIC database (earthref.org/MagIC/19395).

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709 **Table 1: Acceptance criteria, LAC.v.1.0**

Criteria group *	Statistic	Threshold value	Description	Reference †
Specimen	FRAC	0.79	Fraction parameter	Shaar and Tauxe (2013)
	B	0.1	Scatter parameter	Coe et al. (1978); Selkin and Tauxe (2000)
	SCAT	True	Scatter parameter	Shaar and Tauxe (2013)
	GAP-MAX	0.5	Maximum gap	Shaar and Tauxe (2013)
	N _{pTRM}	2	Number of pTRM checks	
	N	5	Number of data points	
	MAD	5	Maximum Angular Deviation of the zero field steps	Kirschvink (1980)
	DANG	10	Deviation Angle	Tauxe and Staudigel (2004)
	Alteration check (correction)	6%	Alteration check in TRM anisotropy and cooling rate experiments	Shaar et al. (2015)
Sample (pottery vessel, furnace, brick, slag)	N _{min}	3	Number of specimens	
	N _{min_aniso_corr}	at least half of the specimens	Minimum number of specimens with anisotropy correction	
	N _{min_cr_corr}	1	Minimum number of specimens with cooling rate correction	
	Σ	$\sigma < 3 \mu\text{T}$ OR $\sigma \% < 8\%$	Standard deviation of the sample mean	
	Anisotropy sample test	6%	If the mean anisotropy correction of all the specimens from the same sample (fragment) is higher than this value, specimens without anisotropy correction are discarded	

* For a complete description and definitions of paleointensity statistics, see Paterson et al. (2014).

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712 **Table 2: Archaeointensity results in Tel Megiddo.**

Megiddo Group	Name in LAC.v.1.0*	Published radiocarbon age range 68.2% probability interval (95.4 % probability interval) (BCE)†	Age range in LAC.v.1.0 compilation (BCE) ‡	N samples	n specimens	B (μT)	B σ (μT)	VADM (ZAm ²)	VADM σ (ZAm ²)
Q-2	mgq02	801-756 (805-735); Assyrian destruction level	820-732	4	15	74.6	3.6	140.3	6.9
H-3-low	mgh03-low	-	820-732	7	37	76.7	3.7	144.3	6.9
H-3-high	mgh03-high			4	18	89.6	2.7	168.6	5.1
Q-4	mgq04	897-821 (901-809)	890-840	3	14	76.4	1.7	143.7	3.1
Q-5	mgq05	956-894 (967-848)	925-875	6	30	69	3.8	129.8	7.1
H-7	Mgh07	930-900 (945-860)	925-875	3	16	71.8	3.7	135.1	6.9
Q-6	mgq06	979-911 (989-876)	950-900	4	16	64.6	3.2	121.5	6
Q-7	mgq07	1047-975 (1052-946)	1050-950	6	18	71	2.7	133.5	5.1
H-9-low	mgh09-low	1038-976 (1056-936)	1050-950	3	15	69.1	4	129.9	7.5
H-9-high	mgh09-high			2	9	83.2	0.4	156.5	0.7
K-4	mgk04	1037-951 (1053-908)	1050-950	4	15	75.5	4.9	142	9.1
H-10	mgh10	1068-1031 (1087-1023)	1100-1025	5	15	69.7	4.3	131	8.1
H-11	mgh11	1105-1051 (1115-1041)	1125-1075	10	30	69.7	4.8	131.1	9.1
K-6	mgk06	1148-1123 (1168-1104)	1175-1125	7	31	64.2	2.6	120.8	4.9
K-8	mgk08	1238-1178 (1268-1158)	1250-1175	8	33	57.3	3.2	107.8	6
K-9	mgk09	1323-1230 (1381-1201)	1400-1250	6	20	54.2	2.2	102	4.1

F-10	mgf10	1545-1354 (1561-1313)	1550-1400	3	14	50.6	1.3	95.2	2.5
H-15	mgh15	1557-1509 (1572-1463)	1550-1475	5	25	54.9	4.3	103.3	8
K-10	mgk10	1581-1545 (1596-1535)	1600-1550	8	35	52.2	5.7	98.2	10.7
K-11	mgk11	1626-1579 (1643-1561)	1650-1600	4	21	52.2	4.5	98.2	8.4
F-13	mgf13	-	1900-1700	4	25	42	2.5	78.9	4.6
S-3	mgs03	1942-1902 (1965-1886)	1950-1900	4	12	38.8	3.5	72.9	6.6
J-6	mgj06	2860-2540 (2880-2450)	2850-2500	4	15	39	0.9	73.4	1.8
J-5	mgj05	2920-2720 (2970-2670)	2900-2800	4	12	44.2	4.3	83.2	8
J-4 §	mgj04	3060-2880 (3180-2830)	3100-2900	8	46	41.1	7.1	77.3	13.3
J-4a §	mgj04a	3060-2880 (3180-2830)	3100-2900	3	10	40.1	1.6	75.4	3.1

* Name in model data (Supplementary Table S10)

† Radiocarbon data from Regev et al. (2014); Toffolo et al. (2014); Martin et al. (2020); Boaretto (2022); There are no radiocarbon data from H-03, and the age is linked to the Assyrian destruction level. Ages from S-3 are preliminary. There are no radiocarbon data from F-13, and the age is inferred from correlation to strata in areas K and S.

‡ Age range used in the archaeomagnetic compilation and in Bayesian modelling.

§ J-4 and J-4a are two groups from the same level. J-4a fragments were collected from loci representing the final days of the temple, whereas the J-4 items originated from a fill context.

Figure 1: Map showing Tel Megiddo and other sites in the Levant and Western Upper Mesopotamia used to construct the Levantine Archaeomagnetic Curve (LAC.v.1.0) shown in Fig. 6. Color code is as in Fig. 6.

Figure 2: Tel Megiddo. a) Aerial photo of the mound displaying the excavation areas discussed in the text. b) Tel Megiddo stratigraphy showing all the contexts analyzed for archaeointensity. The shaded cells mark destruction layers.

Figure 3: Representative results of specimen analysis conducted in this study. (a–d) Red (blue) circles, and triangles in the main Arai plots are ZI steps, IZ steps, and pTRM checks, respectively. Heating temperatures ($^{\circ}\text{C}$) are displayed near the symbols. Blue (red) squares in the inset Zijdeveld plots are x-y (x-z) projections of the remaining NRM, where the x-axis is rotated to the direction of the NRM. The green line is the best fit. a) Specimen passing all criteria. b–d) Interpretations failing the SCAT (b), FRAC (c), and MAD+DANG (d) criteria. e) A successful cooling rate experiment. Blue circles are four measurements at three different cooling rates, and the red square is a projection of the ancient cooling rate on the best-fit (dashed line).

Figure 4: Histograms of anisotropy and cooling rate corrections.

Figure 5: Archaeomagnetic stratigraphy of Tel Megiddo constructed from 132 samples. Full circles (red squares) represent the archaeointensity of samples (group means). Number of samples used to calculate the group means is indicated above each error bar. Vertical lines represent chrono-stratigraphic division. Fragment (sample) groups are plotted according to their relative age.

Figure 6: An archaeointensity curve. a, b) Levantine Archaeomagnetic Intensity Curve (LAC.v.1.0). Colored symbols are groups of samples directly dated with radiocarbon or by clear association with dated historical events. Gray symbols represent groups dated using various archaeological methods. From the 17th to the 6th centuries, there is at least one directly-dated context per century. Curve and shaded area in (a–b) are the average and the 95% credible interval calculated using the AH-RJMCMC algorithm (Livermore et al., 2018), respectively. c) Rate of change. Dashed red and dotted orange lines show the maximum rate for 1840–2020 and the maximum rate in today's field (Fig. 7), respectively. The oscillation pattern revealed in (a) includes four spikes with VADM $>150 \text{ ZAm}^2$ and a change rate of $0.35\text{--}0.55 \mu\text{T/year}$ ($0.7\text{--}1.1 \text{ ZAm}^2$) (c).

Figure 7: Comparison of spikes' paleointensity with the global databases. a) PINT v.8.1.0 database (Bono et al., 2022). b) GEOMAGIA50 v.3.3 (Brown et al., 2015) before 3500 yBP. Horizontal lines show the value of 155 ZAm^2 corresponding to the spike with the lowest intensity maximum according to LAC.v.1.0 (Fig. 6). N is the number of paleointensity estimates below and above 155 ZAm^2 .

Figure 8: Maximum change rate of the geomagnetic field intensity for 1840–2020. a–b) Intensity and VADM rate of change in today's field, calculated using the IGRF13 model (Alken et al., 2021) epochs 2015–2020. c–d) Maximum intensity and VADM rate of change for 1840–2020, calculated using the IGRF models (Alken et al., 2021) between 1950–2020 and the gufm1 model (Jackson et al., 2000) between 1840–1950. e) Field intensity and rate of change at the location

762 with the maximum B change (40°S , 45°E). f) VADM and rate of change at the location with the
763 maximum VADM change (20°N , 65°W).

Figure 1.

a)

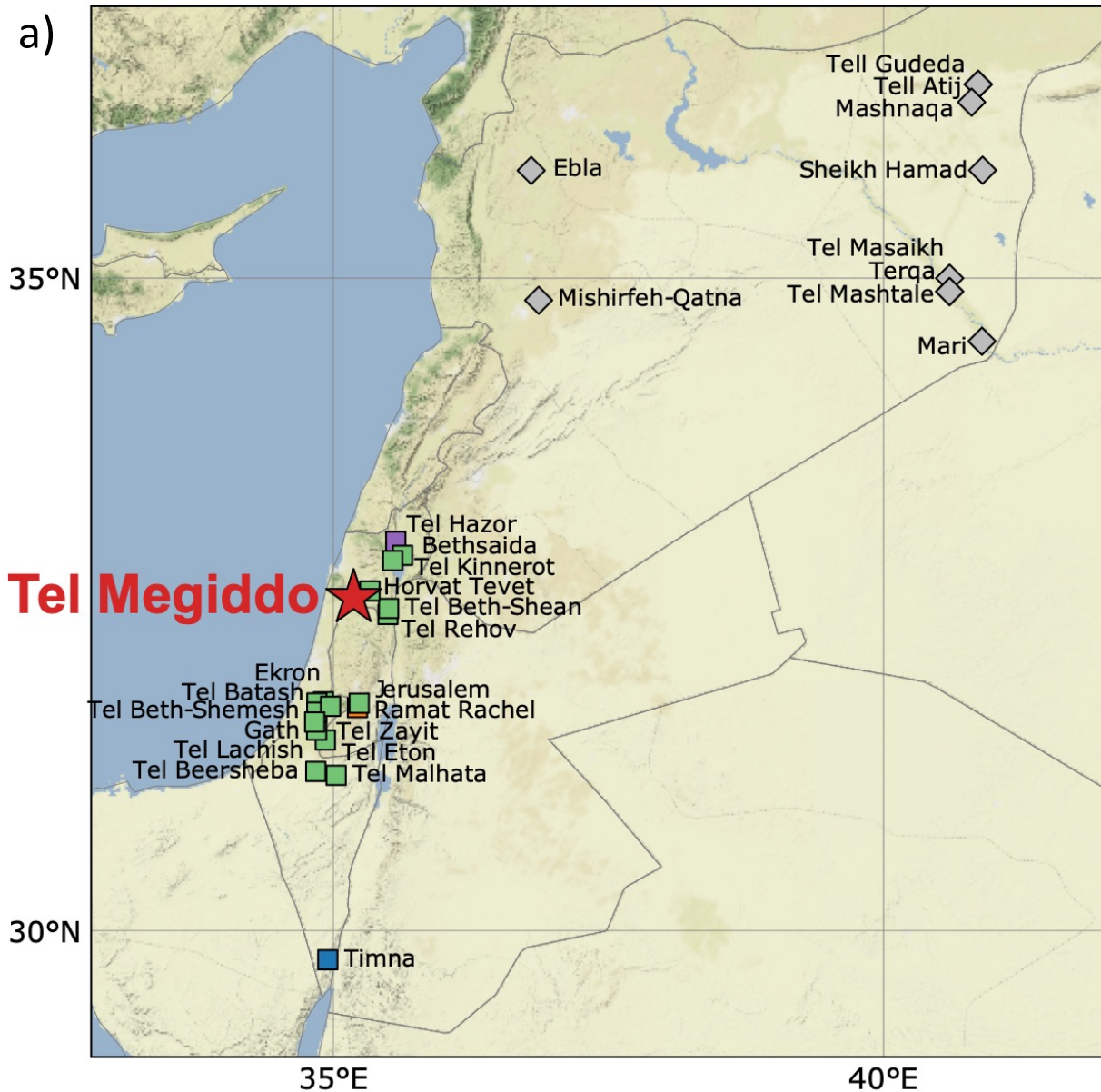
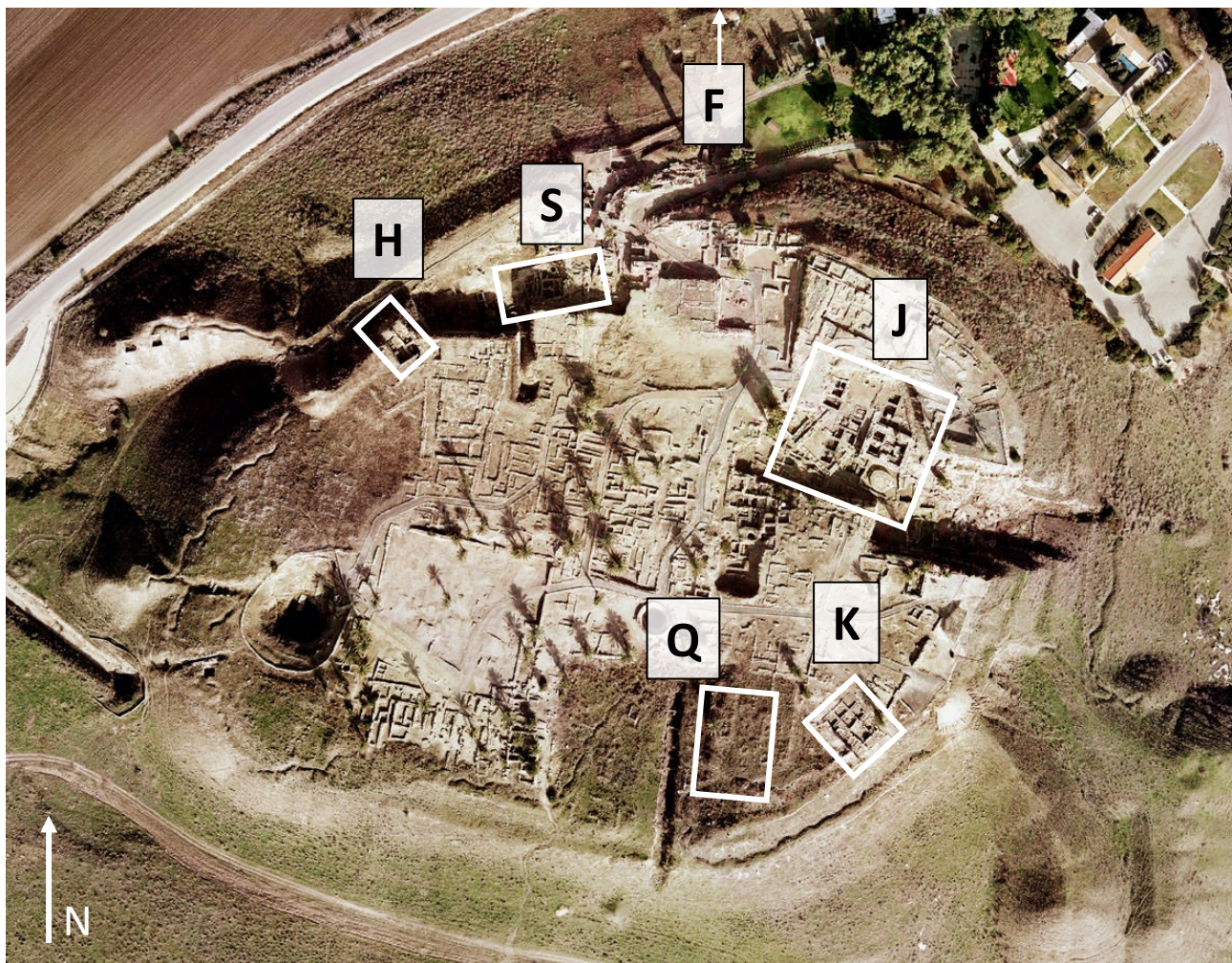


Figure 2.

a)



b)

Area	J	S	F	K	H	Q
Period						
Early Iron IIB					H-3	Q-2
Late Iron IIA						Q-4
Middle Iron IIA					H-7	Q-5
Early Iron IIA						Q-6
Late Iron I				K-4	H-9	Q-7
Middle Iron I					H-10	
Early Iron I					H-11	
Late Bronze III				K-6		
Late Bronze IIB				K-8		
Late Bronze IIA-IIB				K-9		
Late Bronze I			F-10		H-15	
Middle Bronze III				K-10		
Middle Bronze III				K-11		
Middle Bronze II			F-13			
Middle Bronze I		S-3				
Early Bronze III	J-6					
Early Bronze III	J-5					
Early Bronze IB	J-4					

Figure 3.

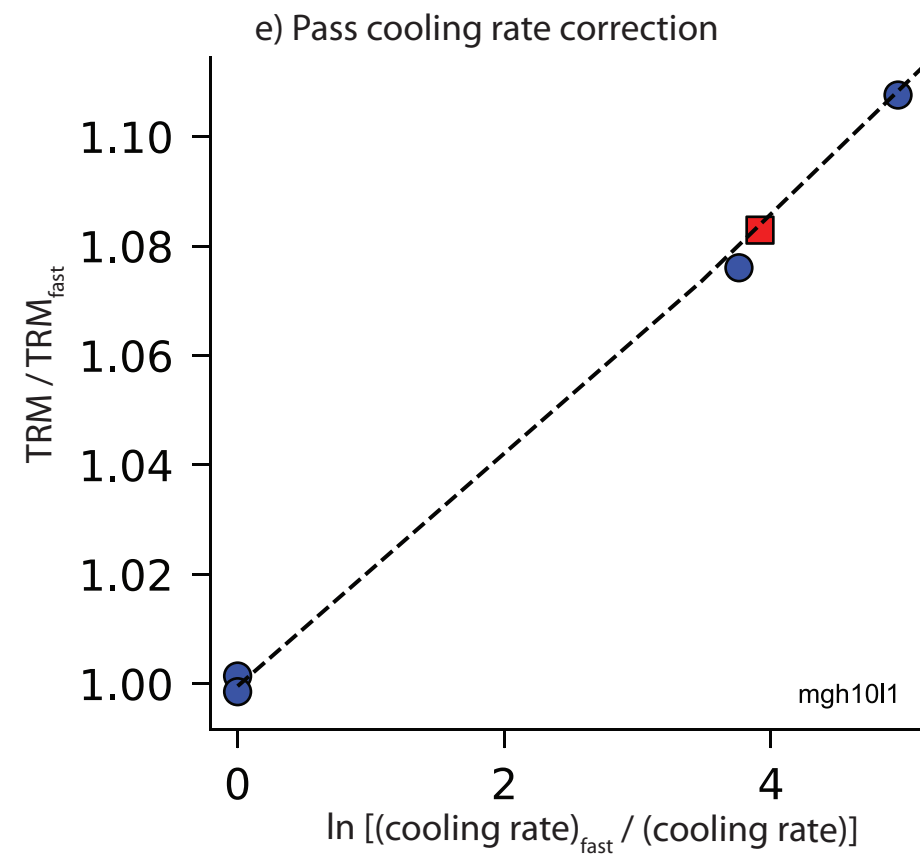
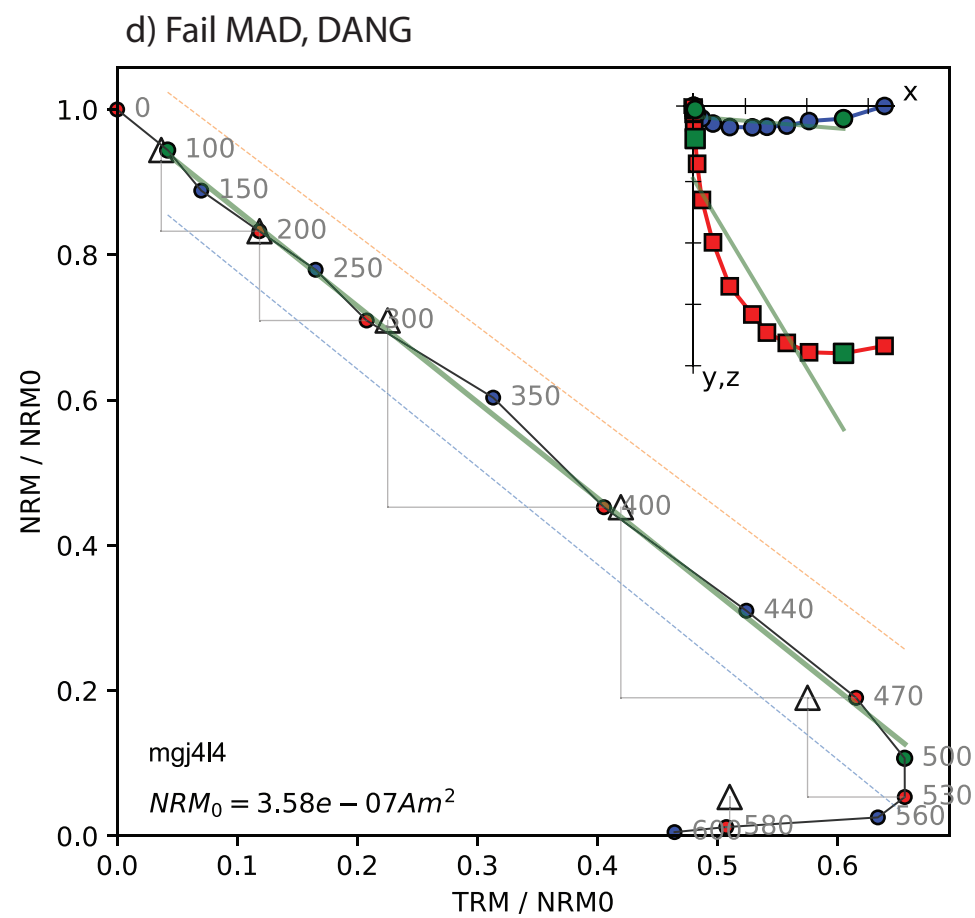
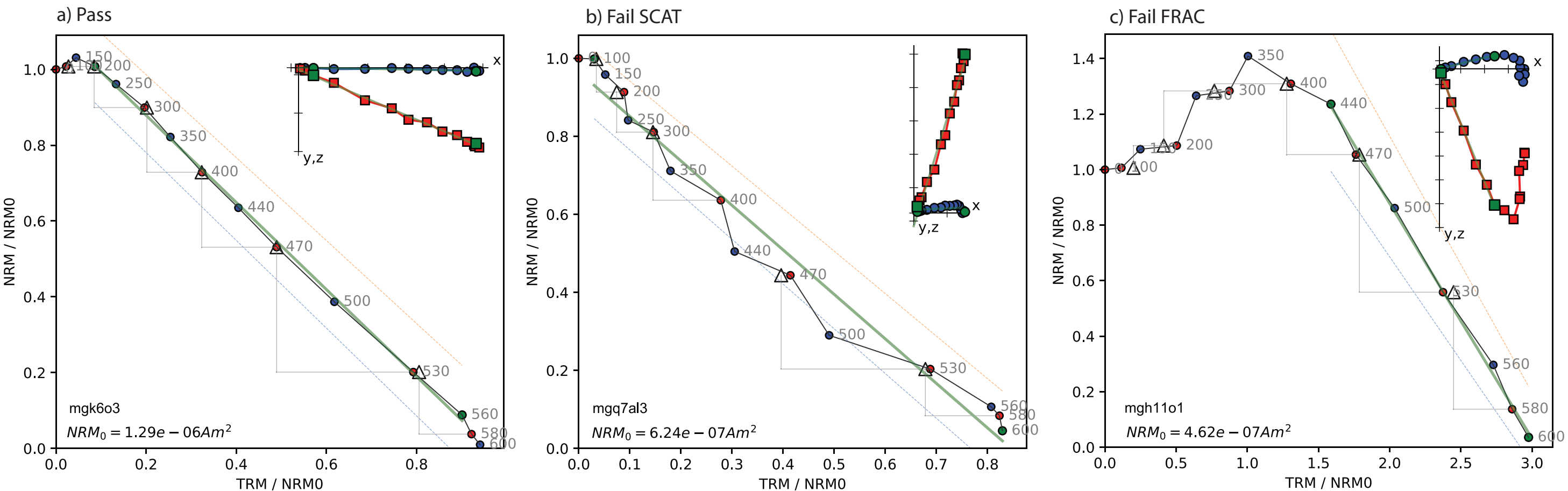
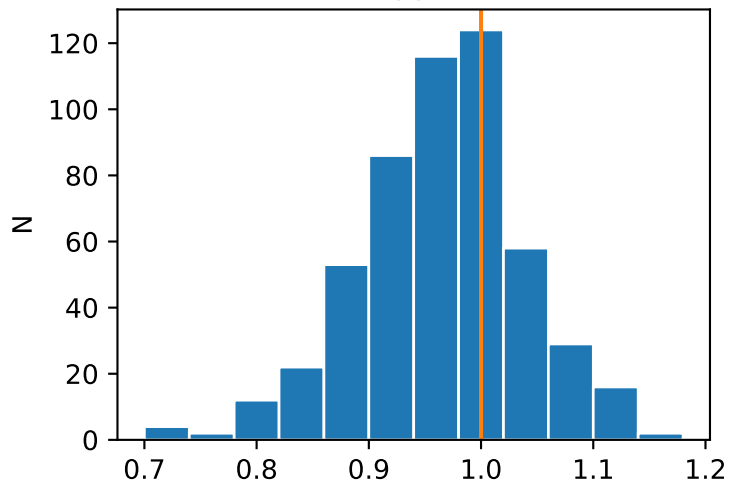
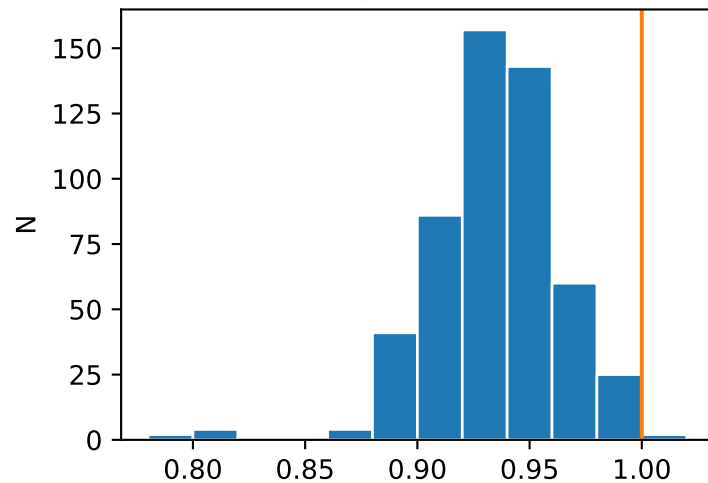


Figure 4.

a. Anisotropy correction



b. Cooling rate correction



c. Anisotropy + cooling rate corrections

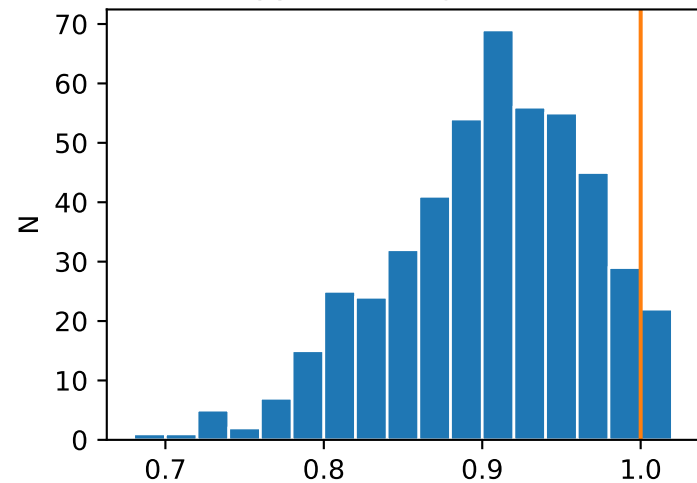


Figure 5.

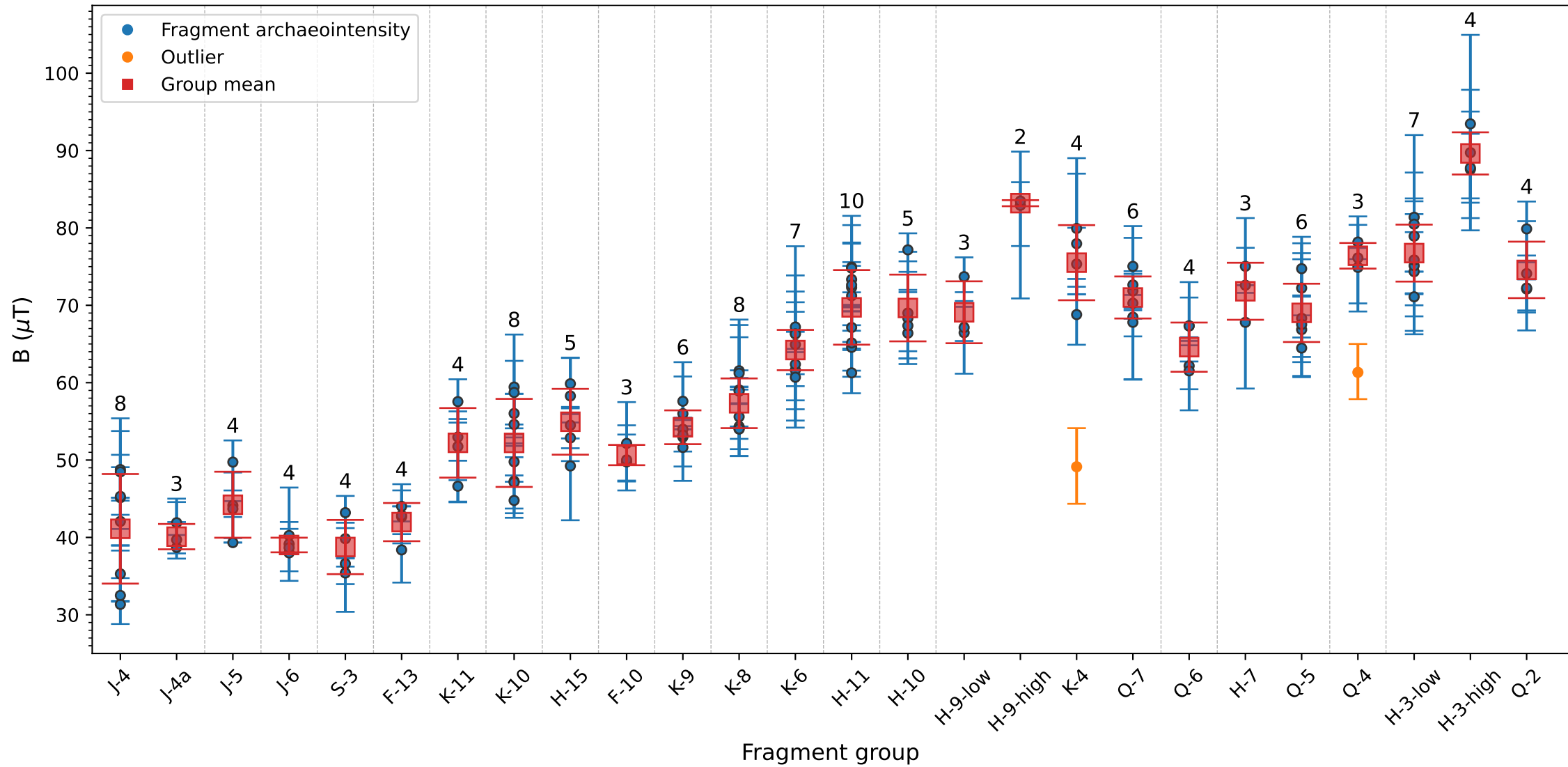


Figure 6.

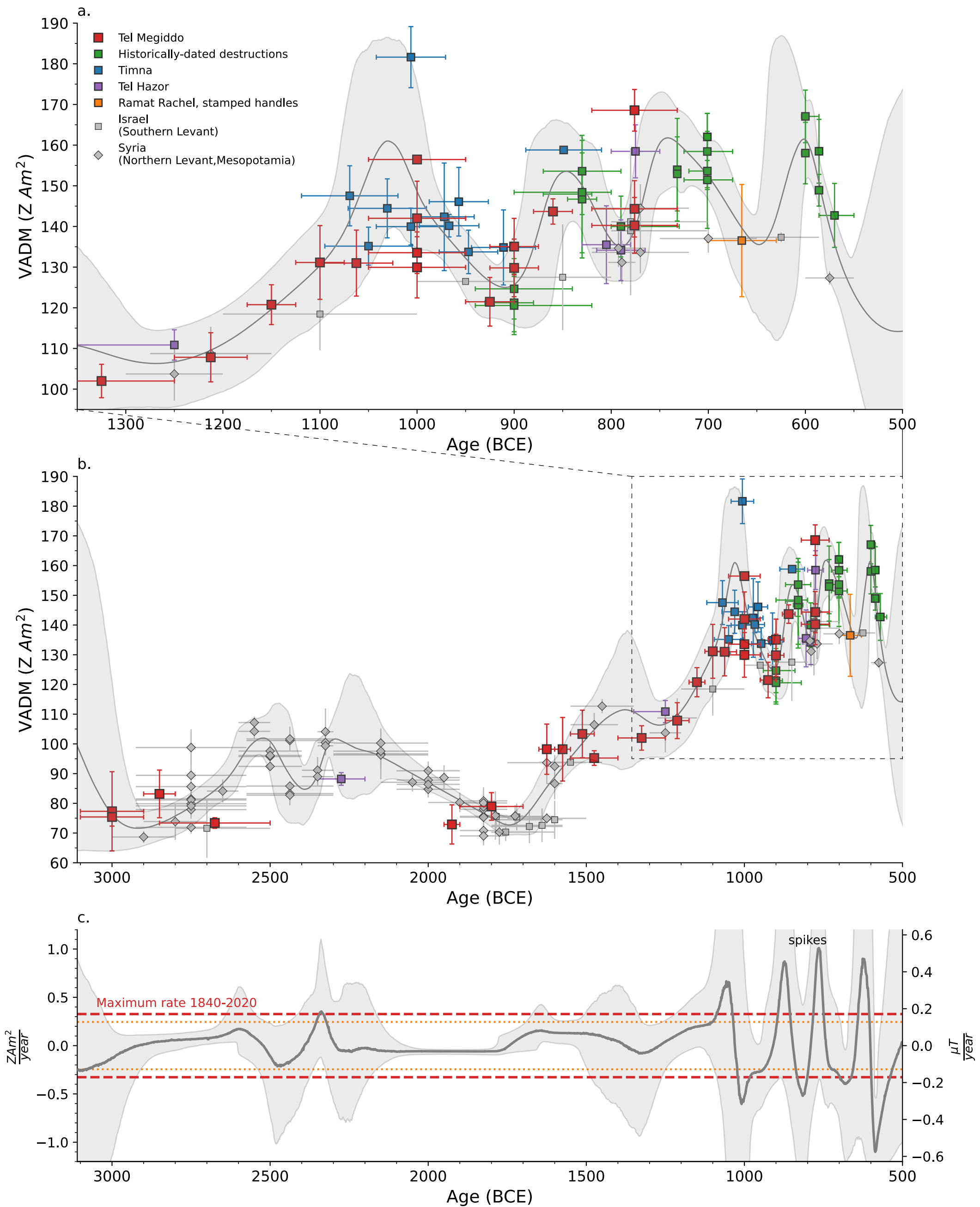
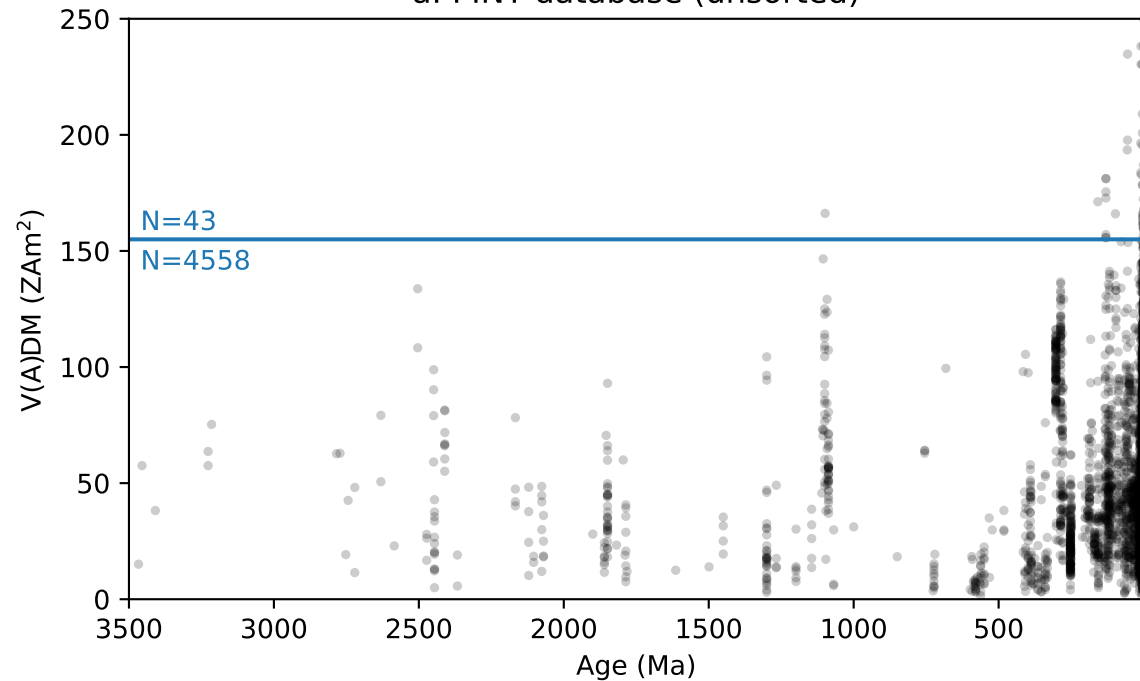


Figure 7.

a. PINT database (unsorted)



b. GEOMAGIA database (unsorted)

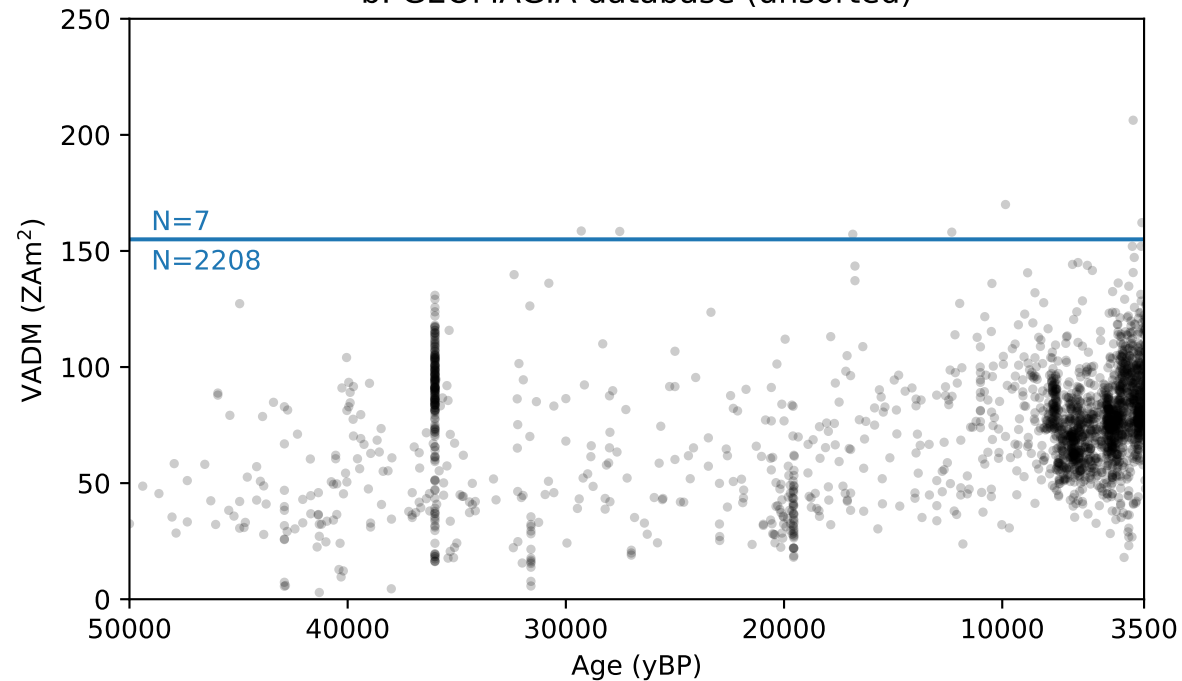
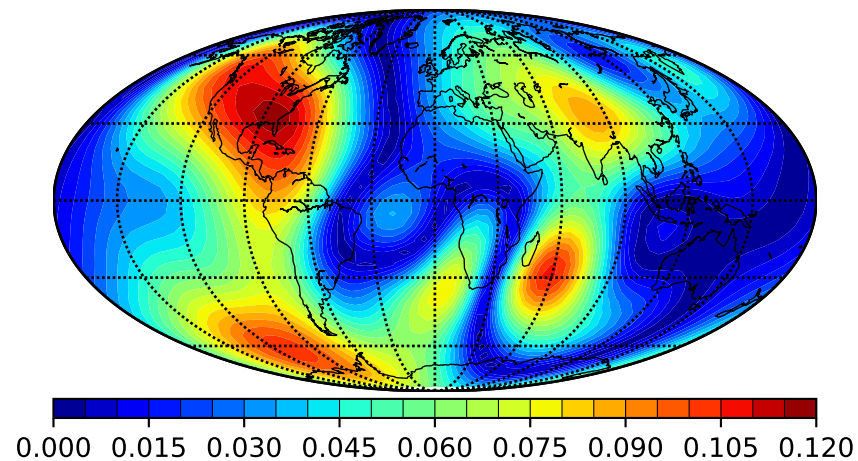
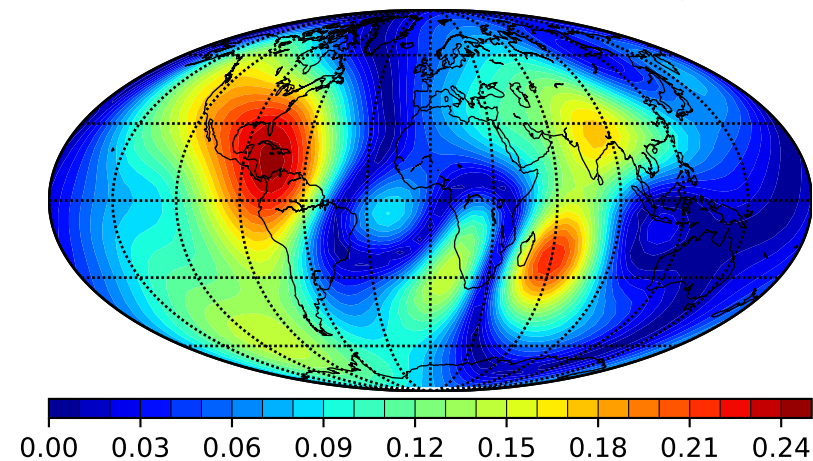


Figure 8.

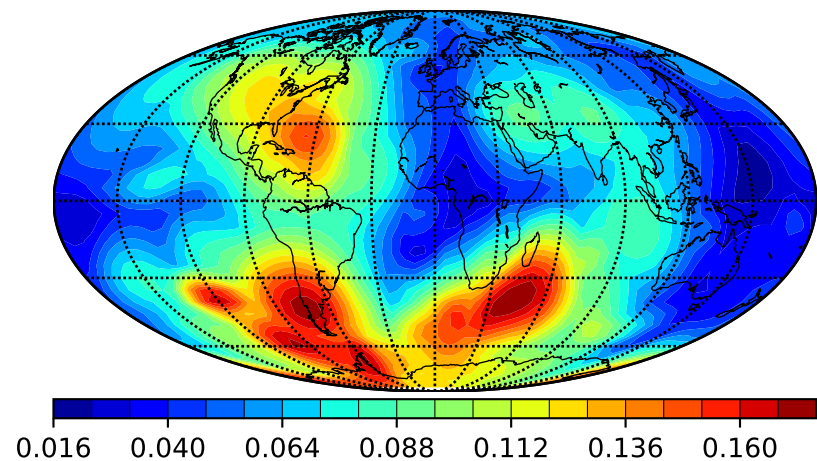
a. Maximum B rate of change ($\mu\text{T}/\text{year}$) 2020



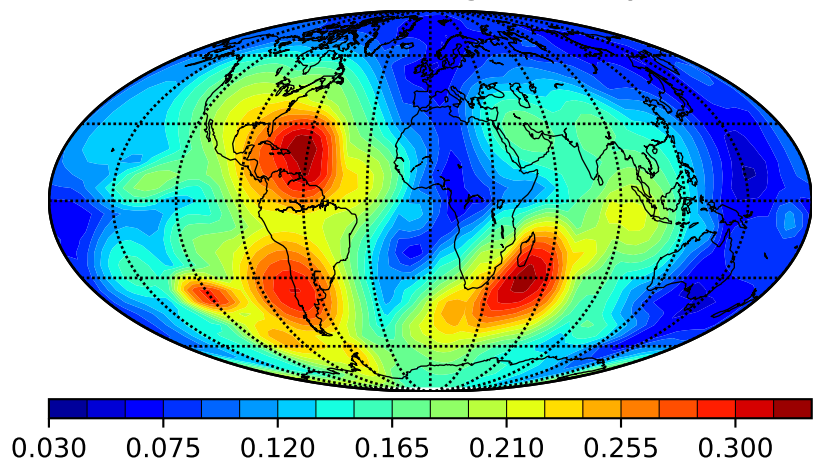
b. Maximum VADM rate of change (ZAm^2/year) 2020



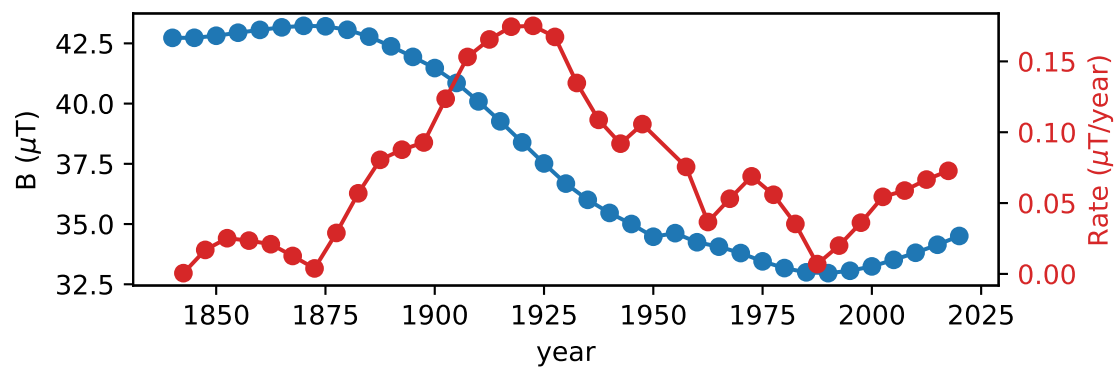
c. Maximum B rate of change ($\mu\text{T}/\text{year}$) 1840-2020



d. Maximum VADM rate of change (ZAm^2/year) 1840-2020



e.



f.

