



## 20 **Abstract**

21 Our understanding of geomagnetic field intensity prior to the era of direct instrumental  
22 measurements relies on paleointensity analysis of rocks and archaeological materials that serve as  
23 magnetic recorders. Only in rare cases absolute paleointensity datasets are continuous over  
24 millennial timescales, provide sub-centennial resolution, and are directly dated using radiocarbon.  
25 As a result, fundamental properties of the geomagnetic field, such as its maximum intensity and  
26 maximum rate of change have remained a subject of lively discussion. Here, we place firm  
27 constraints on these two quantities using Bayesian modelling of well-dated archaeomagnetic  
28 intensity data from the Levant and Upper Mesopotamia. We report new data from 23 groups of  
29 pottery collected from 18 consecutive radiocarbon-dated archaeological strata from Tel Megiddo,  
30 Israel. In the Near East, the period between 1700-550 BCE is now represented by 87 groups of  
31 archaeological artifacts, 57 of which dated using radiocarbon and/or direct association to clear  
32 historically-dated events, providing an unprecedented sub-century resolution. Moreover,  
33 stratigraphic relation between samples collected from multi-layered sited enable further refinement  
34 of the archaeomagnetic ages. The Bayesian curve shows four geomagnetic spikes between 1050  
35 and 600 BCE, with virtual axial dipole moment (VADM) reaching values of 155-162  $ZAm^2$  –  
36 much higher than any prediction from geomagnetic field models. Rates of change associated with  
37 the four spikes are  $\sim 0.35$ - $0.55$   $\mu T/year$  ( $\sim 0.7$ - $1.1$   $ZAm^2/year$ ) – at least twice the maximum rate  
38 inferred from direct observations spanning the past 190 years. Moreover, the increase from 1750  
39 BCE to the first spike depicts the Holocene largest change in field intensity.

40

## 41 **Plain Language Summary (200 words)**

42 The strength of Earth magnetic field is constantly changing in a chaotic and unpredicted manner.  
43 Understanding these changes requires precise information on how the field has changed in the past.  
44 Yet, direct instrumental measurements of field intensity began only in the 1840s, offering a merely  
45 short time-window. Here, we explore the more ancient field by analyzing a rare collection of  
46 radiocarbon-dated archaeological materials from stratified mounds and historically-dated burnt  
47 structures in the Levant and Mesopotamia. Based on new data from Tel Megiddo (Armageddon),  
48 we construct a continuous curve of geomagnetic field intensity spanning 2500 years, with  
49 unprecedented detail and resolution. The curve depicts the evolution of a high-field anomaly,  
50 representing the largest change in the geomagnetic field over the Holocene. Between 1750 and  
51 1050 BCE the field rapidly increased toward values of more than twice those of today - much  
52 higher than any prediction derived from the available geomagnetic field models. Subsequent  
53 oscillations between 1050 and 550 BCE, with extreme peaks namely ‘geomagnetic spikes’, reveal  
54 rates of change that are at least twice faster than the fastest change observed since the advent of  
55 direct measurements. This is an exemplary case-study where archaeology provides crucial  
56 constraints on the geomagnetic field behavior.

## 57 **1 Introduction**

58 The absolute intensity of the geomagnetic field was first measured by Carl Friedrich Gauss in 1832  
59 (Courtillet & Le Mouel, 2007). Subsequent measurements, with improved precision and spatial  
60 resolution, have provided quantitative estimates of the amplitude and rate of geomagnetic field  
61 intensity changes, but only at a limited time interval of less than two centuries. Information from  
62 periods preceding observational measurements, fundamental for understanding field behavior, is

63 derived from ancient materials that acquired thermoremanent magnetization on cooling from high  
64 temperatures. For the past several millennia, archaeological materials are the ultimate source of  
65 information (Arneitz et al., 2017; Brown et al., 2021), providing most of the data for late Holocene  
66 geomagnetic models (Arneitz et al., 2019; Campuzano et al., 2019; Constable et al., 2016; Nilsson  
67 et al., 2014; Panovska et al., 2019; Pavon-Carrasco et al., 2014). Yet, owing to limited precision  
68 and accuracy of the dating in the published archaeomagnetic data, geomagnetic stacks,  
69 compilations and models are inherently scattered and smoothed, hampering our ability to put firm  
70 constraints on geomagnetic field changes. Only ~12% of the published data are directly dated with  
71 radiocarbon and in many cases the exact nature and context of the dated material are not  
72 documented. Instead, most majority of the available archaeomagnetic ages are based on correlation  
73 to regional archaeological chronologies, which can be debated, controversial, and poorly tied to  
74 absolute ages (Shaar et al., 2020). Furthermore, unlike reconstructions of the paleomagnetic field  
75 direction, which use stratigraphic constraints in sedimentary sequences to obtain continuous time-  
76 series, archaeomagnetic intensity datasets are mostly sporadic in time and space. Given the overall  
77 uncertainties in the available paleomagnetic and archaeomagnetic data, some the most  
78 fundamental properties of the geomagnetic field, such as its maximum intensity and maximum  
79 possible rate of change have remained elusive and a subject of a lively and fruitful debate (Davies  
80 & Constable, 2018; Korte & Constable, 2018; Livermore et al., 2014; Livermore et al., 2021).

81 One way to improve the resolution of archaeomagnetic data is focusing efforts on large multi-  
82 layered sites, which are assembled from distinct consecutive strata, and can provide data in a  
83 stratigraphic order. Few examples for this approach in archaeomagnetism are the datasets from  
84 Mari, Tell Atij and Tell Gudeda in Upper Mesopotamia (Gallet & Butterlin, 2015; Gallet et al.,  
85 2020; Gallet et al., 2006; Gallet et al., 2008), Ebla in Northern Levant (Gallet et al., 2014; Gallet  
86 et al., 2008) and Tel Hazor in Southern Levant (Shaar et al., 2020; Shaar et al., 2016). However,  
87 although being key sites in Near Eastern archaeology, most of the data from these sites are not  
88 radiocarbon-dated. In this perspective, Tel Megiddo (Israel), with a radiocarbon-based age model  
89 covering the 3000-800 BCE time-span, is unique, providing unprecedented opportunity for  
90 stratigraphic, radiocarbon-calibrated archaeomagnetic time-series.

91 In the following we report the data of Tel Megiddo, which can be regarded as the largest  
92 archaeomagnetic intensity dataset from a single site. We compile the new data with nearby  
93 archaeomagnetic data derived mostly from radiocarbon-dated stratified mounds and historically-  
94 dated burnt structures to a single coherent continuous dataset that provides a sub-century resolution  
95 over millennial timescale. Based on the high-precision compilation, we calculate the Bayesian  
96 Levantine Archaeomagnetic intensity Curve (LAC). The LAC elucidates the details of the largest  
97 geomagnetic change in the Holocene, associated with the Levantine Iron Age Anomaly (Shaar et  
98 al., 2016) and the occurrence of geomagnetic spikes (Ben-Yosef et al., 2009; Shaar et al., 2011;  
99 Shaar et al., 2016). We use the LAC to enhance knowledge of the number, duration, and intensity  
100 of geomagnetic spikes, which define new robust upper limit constraints for both maximum field  
101 intensity and rate of change.

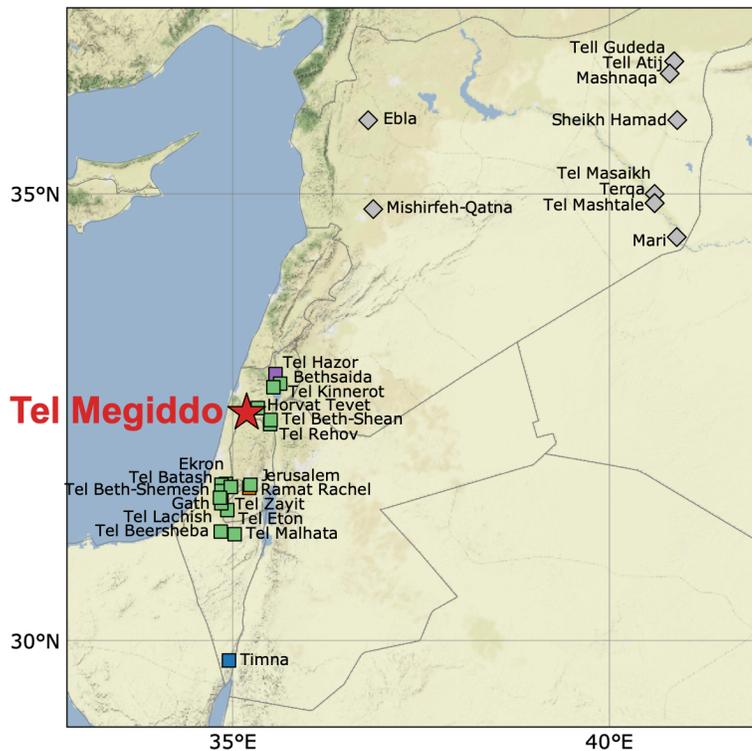
## 102 **2 Materials and Methods**

### 103 **2.1 Tel Megiddo**

104 Tel Megiddo (32.585N, 35.185E, Fig. 1) is a world-heritage archaeological site located on the  
105 western margins of the Jezreel Valley in northern Israel. Owing to its strategic location on the

106 international route which connected Egypt with Mesopotamia, Megiddo was a central city and an  
 107 important administration center through the Bronze and Iron Ages (ca. 3500 – 600 BCE).  
 108 Extensive excavations of the mound have revealed more than thirty Bronze and Iron Age  
 109 superimposed settlements, with several destruction layers marking violent endings in military  
 110 campaigns (Finkelstein, 2009). The chronology of the entire Megiddo sequence was established  
 111 from Bayesian analyses of ca. 150 radiocarbon samples (the total number of radiocarbon samples  
 112 at Megiddo is 185) carefully collected from nearly all strata (Boaretto, 2022; Martin et al., 2020;  
 113 Regev et al., 2014; Toffolo et al., 2014). Special care was taken in assembling the radiocarbon  
 114 model from mostly short-lived organic material securely linked to the archaeological findings. The  
 115 exceptionally large radiocarbon data from a detailed, continuous and well-established stratigraphy,  
 116 along with the intensive ceramic record of Megiddo that defines the relative dating (e.g., late Iron  
 117 I, early Iron IIA), provide a reliable absolute chronology of the Near Eastern archaeology.

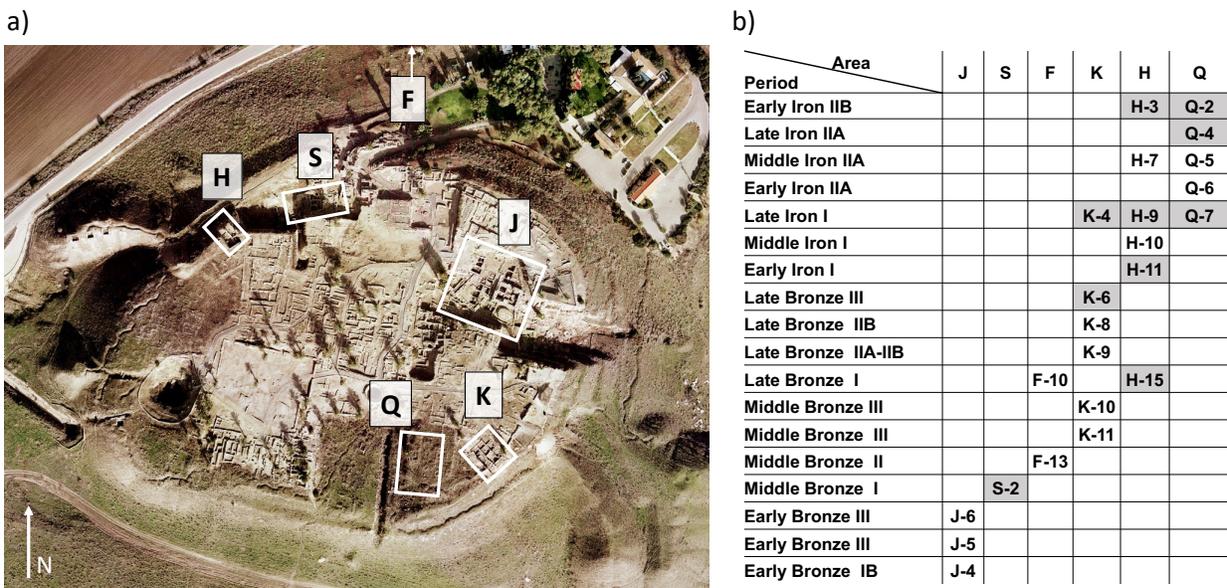
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119  
 120 *Figure 1: Location map showing Tel Megiddo and other sites in the Levant and Western upper*  
 121 *Mesopotamia used to construct the Levantine Archaeomagnetic Curve (LAC.v.1.0) shown in Fig. 6. Color*  
 122 *code is as in Fig. 6.*

123 The archaeomagnetic stratigraphy of Tel Megiddo is based on twenty-three different contexts  
 124 recovered from 18 layers excavated in six excavation areas (Fig. 2). Ten contexts (H-3/Q-2, Q-4,  
 125 K-4/H-9/Q-7, H-11, K-6, H-15, S-2) are destruction layers with distinct boundaries and clear  
 126 marks of their ending. From each context, we sampled fragments of indicative pottery with  
 127 emphasis on local domestic material. We preferred, when possible, complete or cured vessels that  
 128 were photographed and documented in the excavation reports. In three contexts (Q-4, Q-5, K-9)  
 129 we also sampled fragments of cooking ovens (tabun). The samples discussed here include new  
 130 data, as well as data already published in Shaar et al. (2016) and Shaar et al. (2020) who reported

131 the initial archaeomagnetic stratigraphy of Megiddo. Supplementary Table S1 lists the  
 132 archaeological details of all the material analyzed in this study.



133

134 *Figure 2: Tel Megiddo. a) Air photo of the mound showing excavation areas discussed in the text. b) Tel*  
 135 *Megiddo stratigraphy showing all strata analyzed for archaeointensity. Shaded cells mark destruction*  
 136 *layers.*

## 137 2.1 Experimental Design

138 Thellier-IZZI-MagIC paleointensity experiments were carried out in the shielded paleomagnetic  
 139 laboratory at the Institute of Earth Sciences, The Hebrew University of Jerusalem, using two  
 140 modified ASC TD-48 ovens and 2G-RAPID superconducting rock magnetometer (SRM).  
 141 Specimens were prepared by gluing small pieces of pottery inside non-magnetic 22 × 22 × 20 mm  
 142 square alumina crucibles. The protocol followed the IZZI method (Tauxe & Staudigel, 2004; Yu  
 143 et al., 2004) with routine pTRM checks at every second temperature step using an oven field of  
 144 40,50, or 60 μT. Heating time ranged from 40 to 65 minutes depending on the target temperature.  
 145 In total, each IZZI experiment included 31 or 33 heating steps at 13 or 14 temperature intervals  
 146 between 100°C to 590°C or 600°C. All specimens were subjected to anisotropy of thermoremanent  
 147 magnetization (ATRM) experiments, which consisted of eight heating steps at 590°C or 600°C: a  
 148 baseline zero-field step, six infield steps at orthogonal directions, and an additional alteration  
 149 check. ATRM alteration parameter was calculated following Shaar et al. (2015) (Table 1). For  
 150 specimens with ATRM alteration check > 6% anisotropy of anhysteretic remanent magnetization  
 151 (AARM) was measured at 100mT AC field in 0.1mT DC bias field, at six orthogonal directions,  
 152 after thermal demagnetization of the specimens. All specimens were subjected to cooling rate  
 153 correction experiments, which consisted of 4-5 cooling steps from 590 °C or 600 °C to room  
 154 temperature following the protocol described in Shaar et al. (2020). Archaeointensity values were  
 155 calculated with the *Thellier-GUI* program (Shaar & Tauxe, 2013), incorporated into the *PmagPy*  
 156 software package (Tauxe et al., 2016), using *Thellier Auto Interpreter* algorithm and the  
 157 acceptance criteria listed in Table 1. Samples results are calculated by averaging at least 3  
 158 specimens per sample using the STDEV-OPT algorithm of the *Thellier-GUI* program and the

159 ‘extended error bounds’ approach (Shaar & Tauxe, 2013; Shaar et al., 2016). When averaging  
 160 samples data in ‘groups’ (section 3.1) we calculate a simple mean of the STDEV-OPT values of  
 161 the samples. Detailed explanation on the methods can be found in Shaar et al. (2016) and Shaar et  
 162 al. (2020). All measurements data are available in the MagIC databas (earthref.org/MagIC/19395).  
 163 (note for reviewer/editor: The MagIC link will be activated only after publication and receiving  
 164 DOI. Hence, we provide here a private link to the MagIC data, to be used during the review  
 165 process. <https://earthref.org/MagIC/19395/4406c155-3096-42e3-9e73-11ceac8580f2> ).

166

167 **Table 1: Acceptance criteria, LAC.v.1.0**

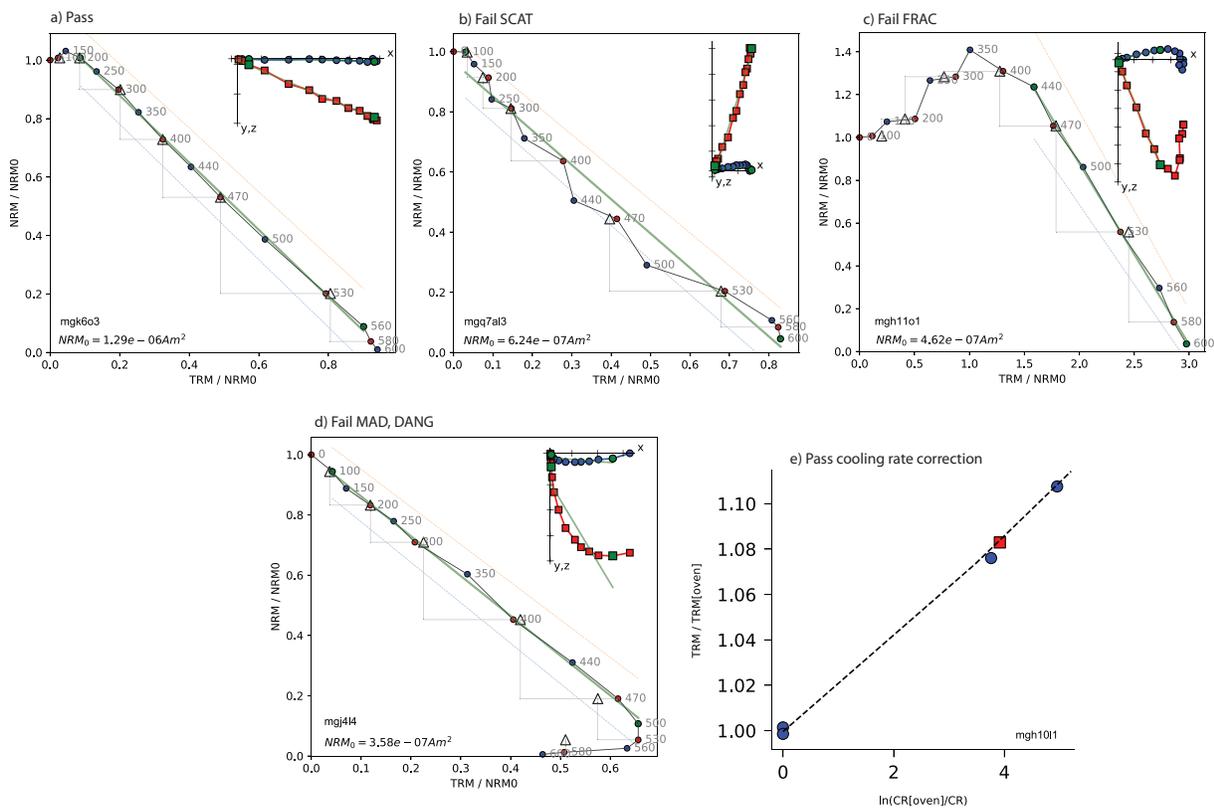
Criteria group *	Statistic	Threshold value	Description	Reference †
Specimen	FRAC	0.79	Fraction parameter	Shaar and Tauxe (2013)
	B	0.1	Scatter parameter	Coe et al. (1978); Selkin and Tauxe (2000)
	SCAT	True	Scatter parameter	Shaar and Tauxe (2013)
	GAP-MAX	0.5	maximum gap	Shaar and Tauxe (2013)
	N <sub>PTRM</sub>	2	Number of pTRM checks	
	N	5	Number of data points	
	MAD	5	Maximum Angular Deviation of the zero field steps	Kirschvink (1980)
	DANG	10	Deviation Angle	Tauxe and Staudigel (2004)
Alteration check (correction)	6%	Alteration check in TRM anisotropy and cooling rate experiments	Shaar et al. (2015)	
Sample (pottery vessel, furnace, brick, slag)	N <sub>min</sub>	3	Number of specimens	
	N <sub>min_aniso_corr</sub>	at least half of the specimens	Minimum number of specimens with anisotropy correction	
	N <sub>min_cr_corr</sub>	1	Minimum number of specimens with cooling rate correction	
	Σ	$\sigma < 3 \mu\text{T}$ OR $\sigma \% < 8\%$	Standard deviation of the sample mean	
	Anisotropy sample test	6%	If the mean anisotropy correction of all the specimens from the same sample (fragment) is higher than this value, specimens without anisotropy correction are discarded	

168 \* For a complete description and definitions of paleointensity statistics see Paterson et al. (2014).

169 **3 Results**170 **3.1 Archaeomagnetic intensity stratigraphy of Tel Megiddo**

171 The archaeomagnetic data from Tel Megiddo, including the data already published in Shaar et al.  
 172 (2016) and Shaar et al. (2020) consists of 763 specimens from 175 samples (pottery fragments).  
 173 In this study, we analyzed 288 specimens from 85 newly collected samples. In total, 583 specimens  
 174 and 132 samples pass the criteria listed in Table 1, where archaeointensities obtained at the sample  
 175 level are calculated from a minimum of 3 specimens. Fig. 3 shows representative cases of a  
 176 successful specimen as well as interpretation failing criteria. The importance of the anisotropy and

177 cooling rate corrections is illustrated in Fig. 4: typically, the bias due to anisotropy and cooling  
 178 rate effects is 5%-15%; in some cases the combined correction exceeds 20%.

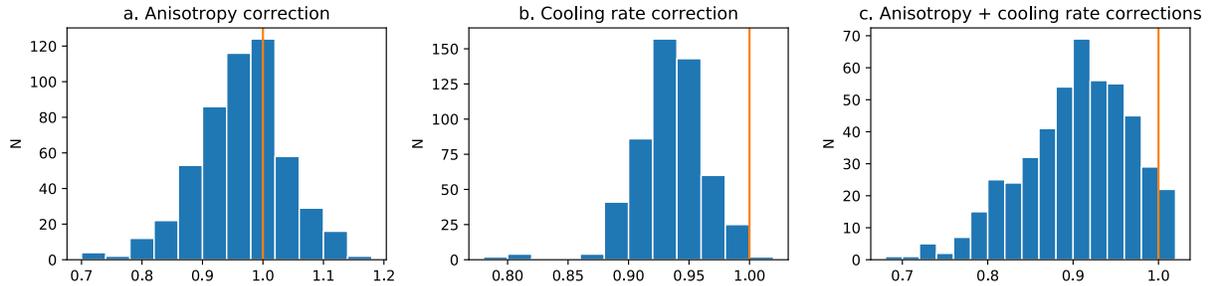


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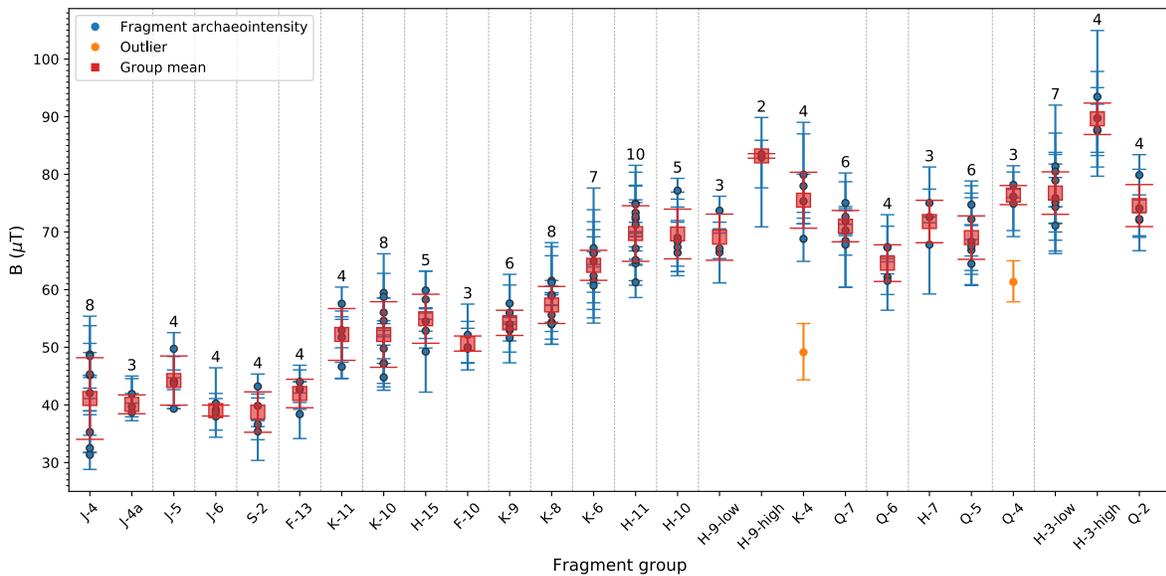
180 *Figure 3: Representative results. (a-d) Red (blue) circles, and triangles in the main Arai plots are ZI steps,*  
 181 *IZ steps, and pTRM checks, respectively. Blue (red) squares in the inset Zijderveld plots are x-y (x-z)*  
 182 *projections of the NRMs remaining, where x axis is rotated to the direction of the NRM. Green line is the*  
 183 *best-fit line. a) specimen passing all criteria. b-d) interpretations failing the SCAT (b), FRAC (c),*  
 184 *MAD+DANG (d) criteria. e) successful cooling rate experiment. Blue circles are four measurements at*  
 185 *three different cooling rates, red square is a projection of the ancient cooling rate on the best-fit line*  
 186 *(dashed line).*

187 Fig. 5 displays samples data with errors bars calculated using the “extended error bounds”  
 188 approach (Shaar & Tauxe, 2013). In general, samples collected from the same archaeological  
 189 context (termed hereafter ‘group’) show good agreement with only two outliers in group K-4 and  
 190 Q-4. Levels H-9 and H-3 exhibit a large dispersion of data, with distinctively different values. As  
 191 the ceramics in each context represent production during a time interval rather than a singular point  
 192 in time, this probably indicates fast change in the field during the interval represented in the  
 193 ceramic assemblage. Thus, we tentatively split the results from these two contexts to two sub  
 194 groups in order to model the varying field. The mean archaeointensity of each group is calculated  
 195 by averaging the sample means. Detailed samples data are given in Supplementary Table S2. Table  
 196 2 summarizes the groups data, including the archaeological affiliations and details on the  
 197 radiocarbon data. We note that the age range we assign in Table 2 to each group (listed as ‘model  
 198 age range’) is different than the radiocarbon age range in order to optimally represent the age range  
 199 of the archaeological context considering the overall archaeological and historical constraints and  
 200 the uncertainty associated with the production of the pottery in each archaeological layer. The

201 archaeomagnetic stratigraphy shown in Fig. 5 (Table 2) shows an exceptionally large amplitude  
 202 changes - between 39 to 90  $\mu\text{T}$  - with few significant oscillations. In the following sections we  
 203 explore in details the extreme amplitude change depicted by Megiddo data.



204  
 205 *Figure 4: Histograms of anisotropy and cooling rate corrections.*



206  
 207 *Figure 5: Archaeomagnetic stratigraphy of Tel Megiddo constructed from 132 samples. Filled circles*  
 208 *(red squares) show archaeointensity of samples (groups means). Number of samples used to calculate*  
 209 *group mean is shown above each error-bar. Vertical lines represent chrono-stratigraphic division.*  
 210 *Fragments (samples) groups are plotted according to the relative age.*

211 **Table 2: Archaeointensity results, Tel Megiddo**

Megiddo Group	name in LAC.v.0.1*	Radiocarbon age BCE		Model Age Range (BCE) ‡	N samples	n specimens	B ( $\mu\text{T}$ )	B $\sigma$ ( $\mu\text{T}$ )	VADM (ZAm <sup>2</sup> )	VADM $\sigma$ (ZAm <sup>2</sup> )
		68.2% interval	probability (95.4% probability interval) †							
Q-2	mgq02	801-756	(805-735); Assyrian destruction level	820-732	4	15	74.6	3.6	140.3	6.9

H-3-low	mgh03-low	-	820-732	7	37	76.7	3.7	144.3	6.9
H-3-high	mgh03-high			4	18	89.6	2.7	168.6	5.1
Q-4	mgq04	897-821 (901-809)	880-840	3	14	76.4	1.7	143.7	3.1
Q-5	mgq05	956-894 (967-848)	925-875	6	30	69	3.8	129.8	7.1
H-7	Mgh07	930-900 (945-860)	925-875	3	16	71.8	3.7	135.1	6.9
Q-6	mgq06	979-911 (989-876)	950-900	4	16	64.6	3.2	121.5	6
Q-7	mgq07	1047-975 (1052-946)	1050-950	6	18	71	2.7	133.5	5.1
H-9-low	mgh09-low	1038-976 (1056-936)	1050-950	3	15	69.1	4	129.9	7.5
H-9-high	mgh09-high			2	9	83.2	0.4	156.5	0.7
K-4	mgk04	1037-951 (1053-908)	1050-950	4	15	75.5	4.9	142	9.1
H-10	mgh10	1068-1031 (1087-1023)	1100-1025	5	15	69.7	4.3	131	8.1
H-11	mgh11	1105-1051 (1115-1041)	1125-1075	10	30	69.7	4.8	131.1	9.1
K-6	mgk06	1148-1123 (1168-1104)	1175-1125	7	31	64.2	2.6	120.8	4.9
K-8	mgk08	1238-1178 (1268-1158)	1250-1175	8	33	57.3	3.2	107.8	6
K-9	mgk09	1323-1230 (1381-1201)	1400-1250	6	20	54.2	2.2	102	4.1
F-10	mgf10	1545-1354 (1561-1313)	1550-1400	3	14	50.6	1.3	95.2	2.5
H-15	mgh15	1557-1509 (1572-1463)	1550-1475	5	25	54.9	4.3	103.3	8
K-10	mgk10	1581-1545 (1596-1535)	1600-1550	8	35	52.2	5.7	98.2	10.7
K-11	mgk11	1626-1579 (1643-1561)	1650-1600	4	21	52.2	4.5	98.2	8.4
F-13	mgf13	-	1900-1700	4	25	42	2.5	78.9	4.6
S-2	mgs02	1942-1902 (1965-1886)	1950-1900	4	12	38.8	3.5	72.9	6.6
J-6	mgj06	2860-2540 (2880-2450)	2850-2500	4	15	39	0.9	73.4	1.8
J-5	mgj05	2920-2720 (2970-2670)	2900-2800	4	12	44.2	4.3	83.2	8
J-4 §	mgj04	3060-2880 (3180-2830)	3100-2900	8	46	41.1	7.1	77.3	13.3
J-4a §	mgj04a	3060-2880 (3180-2830)	3100-2900	3	10	40.1	1.6	75.4	3.1

212 \* Name in model data (Supplementary Table S9)

213 † Radiocarbon date from Regev et al. (2014) ; Toffolo et al. (2014); Martin et al. (2020); Boaretto (2022); There are  
 214 no radiocarbon data from H-03 and the age is linked to the Assyrian destruction level. Ages from S-2 are preliminary.  
 215 There are no radiocarbon data from F-13 and age is inferred from correlation to strata in areas K,S.

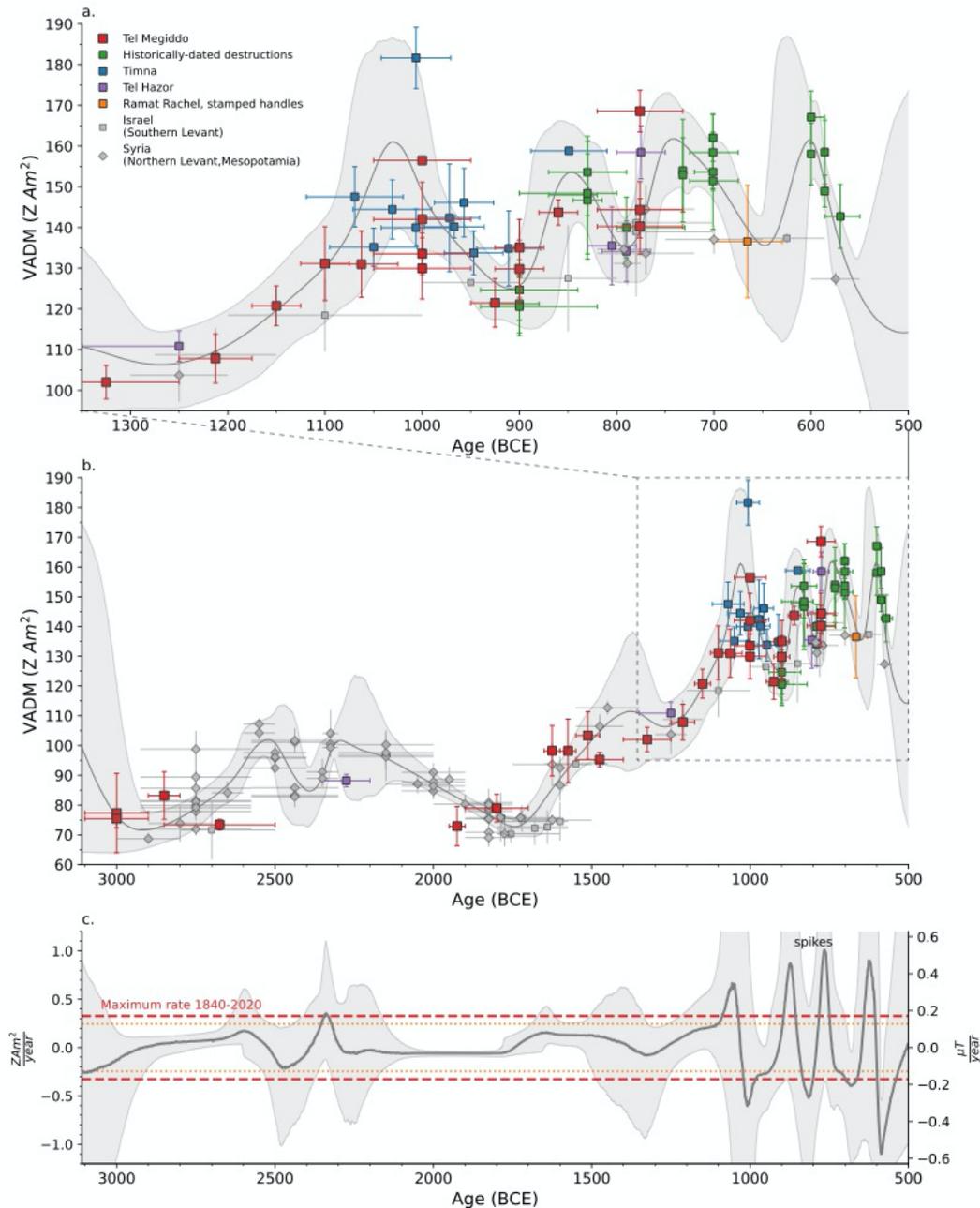
216 ‡ Age range used in LAC.v.1.0

217 § J-4 and J-4a are two groups from the same level. J-4a fragments were collected from loci representing the end-days  
 218 of the temple, whereas the J-4 items originated from a fill context.

### 219 **3.2 Levantine archaeomagnetic curve (LAC.v.1.0)**

220 In an effort to minimize uncertainties associated with differences in laboratory techniques, data  
221 analysis approaches, and selection criteria (see review in Shaar et al. (2020)), our compilation of  
222 the Levant and Upper Mesopotamia is based on data obtained using the Thellier-IZZI-MagIC  
223 method (Shaar & Tauxe, 2013; Tauxe & Staudigel, 2004) with the criteria defined in Shaar et al.  
224 (2016, 2020), and the Triaxe method (Le Goff & Gallet, 2004). These two methods were tested  
225 against each other in a blind test and proved to be equivalent (Shaar et al., 2020). Also included  
226 data obtained by Genevey et al. (2003) that applied the Thellier-Coe method (Coe et al., 1978)  
227 because these data were tested against the Triaxe in Gallet and Le Goff (2006). We include in the  
228 compilation only groups that include at least two samples with overlapping intensity error bars,  
229 after inspection for outliers, where a group is defined as archaeological context, locus, level,  
230 stratum, burnt building, and slag layer. Fig. 6 displays data from 142 groups between 3000 – 500  
231 BCE passing these criteria. Data are displayed in terms of virtual axial dipole moment (VADM) –  
232 a transformation from local, latitude-dependent field intensity measurement to the equivalent  
233 geomagnetic axial dipole moment (Tauxe et al., 2018). The data from Syria, representing  
234 Mesopotamia and Northern Levant (Gallet & Al-Maqdissi, 2010; Gallet & Butterlin, 2015; Gallet  
235 et al., 2014; Gallet et al., 2020; Gallet et al., 2006; Gallet et al., 2008; Genevey et al., 2003;  
236 Livermore et al., 2021), are displayed as published with few updates on the ages of some fragment  
237 groups (see Supplementary Material) . Data from Timna-30 slag mound (Shaar et al., 2011), Tel  
238 Hazor (Shaar et al., 2016), and the Judean stamped jars (Ben-Yosef et al., 2017), which were  
239 published as samples, are shown here as group means (Supplementary Tables S3-S8). We note  
240 few updates on the ages of some of these published data (see Supplementary Material for details):  
241 We arrange the data from strata V-VI in Tel Hazor (Shaar et al., 2016) in three distinct groups  
242 based on sub-division to phases and re-assign their ages accordingly (Supplementary Table S3).  
243 The Bayesian age model of Timna-30 slag mound (Shaar et al., 2011), which was originally  
244 established using magnetostratigraphic correlation with Khirbet En Nahas (Ben-Yosef et al.,  
245 2009) is revised here, and now includes only radiocarbon samples collected from Timna-30

246 (Supplementary Tables S5-S7). Also, the age range of the Judean stamped handles is extended  
 247 from Ben-Yosef et al. (2017) following Vaknin et al. (submitted) (Supplementary Table S8).



248

249 *Figure 6: Archaeointensity curve. a,b) Levantine archaeomagnetic intensity curve (LAC.v.1.0). Colored*  
 250 *symbols are groups of samples directly dated with radiocarbon and/or by clear association to dated*  
 251 *historical events. Gray symbols represent groups dated using various archaeological methods. From the*  
 252 *17<sup>th</sup> to the 6<sup>th</sup> centuries there is at least one directly-dated context per century. Curve and shaded area in*  
 253 *(a-b) are the average and the 95% credible interval calculated using the AH-RJMCMC algorithm*  
 254 *(Livermore et al., 2018). c) Rate of change. Dashed red and dotted orange lines show the maximum rate*  
 255 *between 1840-2020 and the maximum rate in today's field (Fig. 7), respectively. The oscillation pattern*  
 256 *shown in (a) includes four spikes with VADM  $> 150 \text{ Z Am}^2$  and rate of change of  $0.35\text{-}0.55 \mu T / \text{year}$  (c).*

257 The data from Tel Megiddo, obtained from materials that are directly linked to the radiocarbon-  
 258 dated contexts (red symbols in Fig. 6) form the skeleton of the archaeomagnetic dataset with a  
 259 time series of well-established chronological anchors. These data are combined with other  
 260 radiocarbon archaeomagnetic contexts from Tel Hazor and Timna (shown in colored symbols in  
 261 Fig. 6) as well as with a newly published dataset obtained from burnt structures destroyed during  
 262 historically-dated military campaigns (Vaknin et al., 2020; Vaknin et al., submitted). The burnt  
 263 structures are crucial tie points for two reasons. First, the uncertainty in radiocarbon ages during  
 264 the late Iron Age is in the order of 200-400 years due to the plateau in the calibration curve (Reimer  
 265 et al., 2020), while the dates of the historical campaigns are unique in their precision. Second, the  
 266 burnt structures record a single event, as the fire during the destruction resets the magnetization in  
 267 the bricks, compared to the ceramic groups that provide data over a time interval representing a  
 268 production period. Taken together, the period between 1700-550 BCE in the Levant and western  
 269 upper Mesopotamia is now represented by 87 archaeomagnetic contexts, 57 of which are dated  
 270 using radiocarbon and/or direct association with historically-dated secure events. All the directly-  
 271 dated contexts (highlighted in colors in Fig. 6) are spread in such a way that there is at least one  
 272 context per century after 1700 BCE, providing a robust century-scale (or better) resolution.

273 With the data described above, and summarized in Supplementary Table S9, we calculate a  
 274 Bayesian curve with its corresponding 95% credible envelope (Fig. 6, Supplementary Table S10).  
 275 We term this curve ‘Levantine Archaeomagnetic Curve version 1.0’, or LAC.v.1.0. The LAC is  
 276 calculated using the age hyperparameter reverse-jump Monte Carlo Markov Chain (AH-  
 277 RJMCMC) algorithm developed by Livermore et al. (Livermore et al.,  
 278 2018)(<https://github.com/plivermore/AH-RJMCMC1>). The algorithm is based on piece-wise  
 279 linear extrapolation between vertices drawn in a random-walk like perturbation within a space  
 280 allowed by acceptance criteria. The prior assumptions of the model are as follows: i) the allowed  
 281 range of vertices VADM values is between 60 to 200 ZAm<sup>2</sup>; ii) the allowed number of vertices  
 282 (K) is between K<sub>min</sub> = 1 to K<sub>max</sub> = 150; iii) ages in all contexts are uniformly distributed except  
 283 the ages of Timna, which were modeled as normal distribution; iv) groups mean and standard  
 284 deviation define a normal distribution of archaeointensity data. In addition, Supplementary Table  
 285 S9 defines a stratigraphic order for contexts collected from the layered sites (Tel Megiddo, Tel  
 286 Hazor, Tell Atij, Tell Tel Gudeda, and Timna slag mound) as well as for some of the destructions  
 287 levels in Vaknin et al. (submitted). The model uses the following parameters:  $\sigma_{\text{move}} = 30$  yrs,  $\sigma_{\text{change}}$   
 288 = 10 Z Am<sup>2</sup>, and  $\sigma_{\text{birth}} = 10$  Z Am<sup>2</sup>, which define the distributions of a vertex move in age, vertex  
 289 change in intensity, and intensity of a new vertex born with respect to the extrapolated intensity at  
 290 the vertex age; each perturbation includes one age resampling of data in each perturbation  
 291 (num\_age\_changes = 1); chain length is  $2 \cdot 10^8$ .

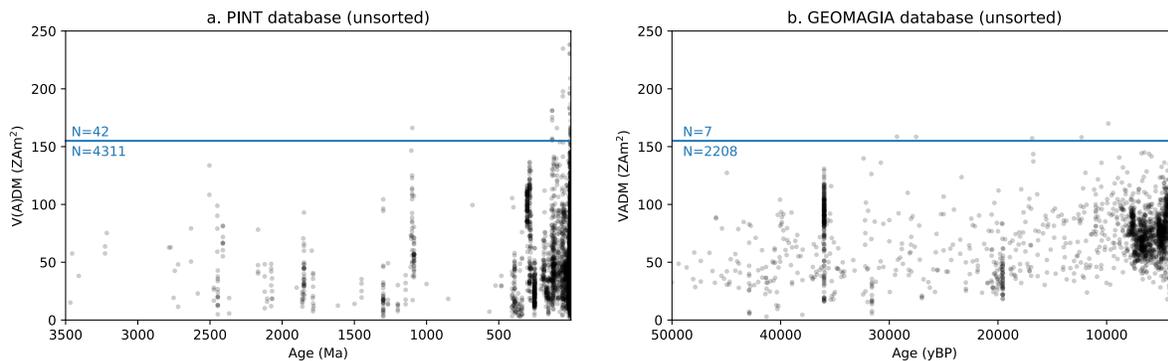
## 292 **4 Discussion**

### 293 **4.1 What is the highest geomagnetic field intensity?**

294 Considering all the published paleointensity estimates from individual samples (i.e. not group  
 295 means) from the past 5 My available in the GEOMAGIA50 (Brown et al., 2015) and PINT (Biggin  
 296 et al., 2009; Bono et al., 2022) databases, only 1% of the data show VADM > 150 ZAm<sup>2</sup> and they  
 297 are sporadically scattered. As such, VADM values calculated from global geomagnetic models do  
 298 not exceed 140 ZAm<sup>2</sup> (Arneitz et al., 2019; Constable et al., 2016; Korte & Constable, 2018;  
 299 Panovska et al., 2019; Pavon-Carrasco et al., 2014). The only exception is the time interval

300 between the end of the 2<sup>nd</sup> millennium BCE and the middle of the 1<sup>st</sup> millennium BCE, where a  
 301 significant number of observations point to unusually high field values at several locations in  
 302 Eurasia (Ben-Yosef et al., 2017; Cai et al., 2017; Di Chiara et al., 2014; Ertepinar et al., 2012;  
 303 Kissel et al., 2015; Molina-Cardin et al., 2018; Shaar et al., 2011; Shaar et al., 2017; Shaar et al.,  
 304 2016; Vaknin et al., 2020). This high field episode is probably associated with a more complex  
 305 field structure than today's (Korte & Constable, 2018; Osete et al., 2020) and presumably with a  
 306 local high field anomaly in the Near East termed 'Levantine Iron Age Anomaly' (LIAA) (Shaar et  
 307 al., 2018; Shaar et al., 2017; Shaar et al., 2016). The highest VADM values, representing the climax  
 308 of the LIAA were termed 'geomagnetic spikes' by Ben-Yosef et al. (2009) and Shaar et al. (2011).  
 309 Livermore et al. (2021) raised the question of the robustness of the spikes and pointed out that the  
 310 number of spikes and their values are strongly dependent on the archaeomagnetic data used, in  
 311 particular the experimental errors and the averaging scheme adopted (i.e. samples groups versus  
 312 individual samples). Here, we address the issues raised by Livermore et al. (2021) and assemble a  
 313 much denser dataset, which is based solely on group averages. This way, each data point in our  
 314 compilation represents exactly the same quantity and gain the same weight in the Bayesian  
 315 calculation process. The new curve shows the occurrence of four spikes with peak VADM of 155-  
 316 162 Z Am<sup>2</sup> around 1000, 850, 750, and 600 BCE. Each spike is represented by several coeval or  
 317 nearly coeval groups, where overall, fourteen groups have VADM > 150 Z Am<sup>2</sup>. It should be noted  
 318 that the same series of four spikes is obtained if only radiocarbon-dated data are used for the  
 319 construction of the curve, as well as those related to historically-dated destruction layers.  
 320 Considering that each group represents a time-average of several samples, we suggest a value of  
 321 155 Z Am<sup>2</sup> as a robust and conservative upper limit for maximum field value. Yet, based on  
 322 samples data, it is still possible that highest values may have occurred for a short time interval.

323



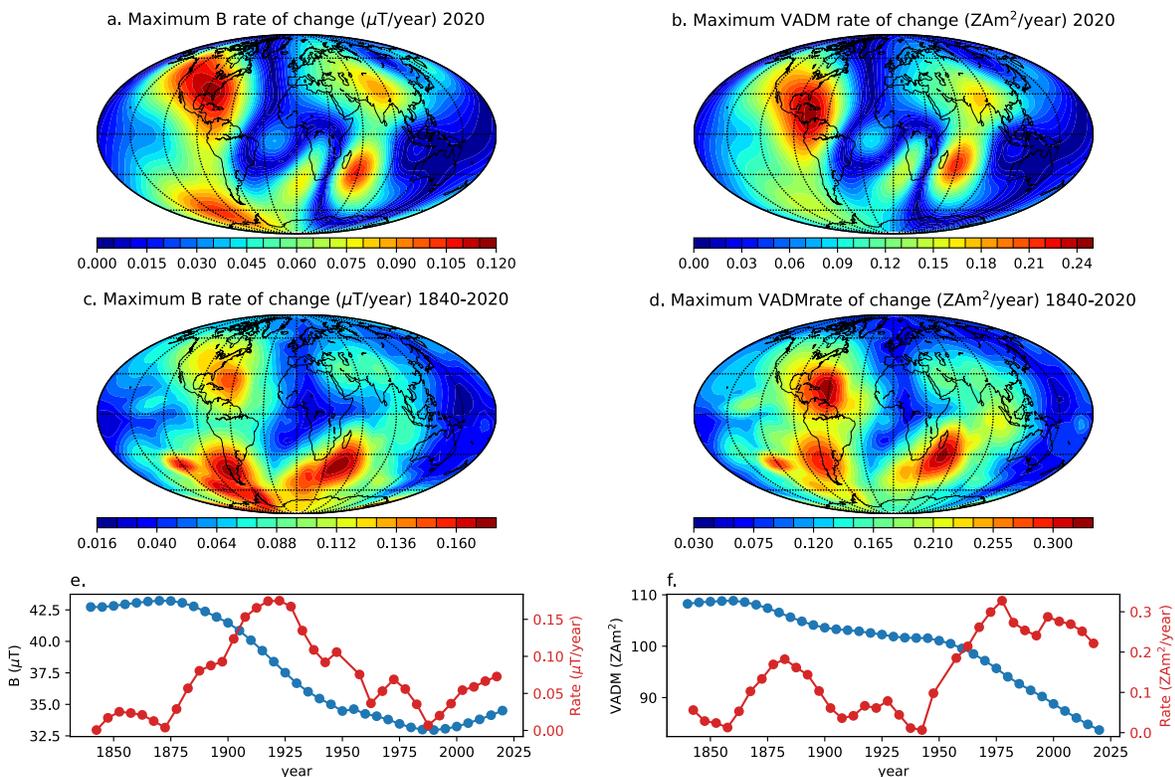
324

325 *Figure 7: Comparison of spikes paleointensity with the global databases. a) PINT database (Bono et al.,*  
 326 *2022). b) GEOMAGIA50 (Brown et al., 2015) before 3500 yBP. Data were not screened out by any criteria*  
 327 *and shown as are. Horizontal lines show the value of 155 ZAm<sup>2</sup> corresponding to the lowest spikes*  
 328 *according to LAC.v.1.0 (Fig. 6). N is the number of paleointensity estimates below and above 155 ZAm<sup>2</sup>.*

329

330 Figure 7 demonstrates that spike-like values are rare in the paleomagnetic record, showing all the  
 331 published absolute paleointensity data with age older than 1500 BCE from the GEOMAGIA50  
 332 and PINT databases. Only 49 data points out of 6568 have field values higher than 155 Z Am<sup>2</sup>,  
 333 none of which pass the rather strict statistical tests applied in the LAC. Given the specific

334 conditions associated with the Levantine geomagnetic spikes, the difficulty in detecting similar  
 335 high paleointensity values in the global paleointensity record is not surprising. First, our dense  
 336 dataset that includes on average 10 archaeointensity groups (each consists of at least 3 samples)  
 337 per century during the high-field interval (Fig. 6a) shows that the duration of the spikes is not more  
 338 than a century. Thus, if an average of several samples is required to obtain a robust paleointensity  
 339 estimate, a large dataset, such as the LAC dataset, is required to detect spikes. Second, global  
 340 geomagnetic models indicate that the geomagnetic dipole during the spikes episode is most likely  
 341 the highest in the Holocene (Constable et al., 2016; Pavon-Carrasco et al., 2014; Schanner et al.,  
 342 2022). Thus, the likelihood of detecting spikes may depend on the likelihood that the ancient  
 343 paleomagnetic dipole was as similarly high. Third, the spikes is a regional feature associated with  
 344 a local geomagnetic anomaly, which is expressed not only by high field values, but also by  
 345 directional deviations from a dipole field (Osete et al., 2020; Shaar et al., 2018; Shaar et al., 2016).  
 346 Taken altogether, there are low chances that the scattered and sparse paleointensity database  
 347 spanning the geological record can reveal short-lived spikes. Moreover, from the comparison with  
 348 the global paleointensity database we can conclude that spikes represent the highest value the  
 349 geomagnetic field can reach and can serve as a robust upper boundary for the maximum strength  
 350 of the geomagnetic field at a given location,



351

352 *Figure 8: Maximum rate of change of the geomagnetic field intensity between 1840-2020. a-b) Intensity*  
 353 *and VADM rate of change in today's field calculated from IGRF13 model (Alken et al., 2021) epochs*  
 354 *2020,2015. c-d) Maximum intensity and VADM rate of change between 1840-2020 calculated from IGRF*  
 355 *models (Alken et al., 2021) between 1950-2020 and gufm1 model (Jackson et al., 2000) between 1840-*  
 356 *1950. e) Field intensity and rate of change at the location with the maximum B change (40°S, 45°E). f)*  
 357 *VADM and rate of change at the location with the maximum VADM change (20°N, 65°W).*

358

## 4.2 What is the maximum rate of geomagnetic field changes?

359 We inspect now the rate of change of the geomagnetic field. The rates associated with the spikes  
360 range between  $\sim 0.35 - 0.55 \mu\text{T}/\text{year}$ , or  $0.7 - 1.1 \text{ Z Am}^2/\text{year}$  in VADM values. To place these  
361 values within the context of the global geomagnetic field behavior, we calculate in Fig. 8 the  
362 maximum rate in today's field by looking at the difference between IGRF models epochs 2015  
363 and 2020 (Alken et al., 2021). On most of Earth surface the rates do not exceed  $0.1 \mu\text{T}/\text{year}$  ( $\sim 0.2$   
364  $\text{Z Am}^2/\text{year}$ ) and a maximum rate of  $0.12 \mu\text{T}/\text{year}$  occurs only in limited areas. VADM  
365 transformation that accounts for the latitudinal dependency of the field yields a maximum rate of  
366  $0.25 \text{ Z Am}^2/\text{year}$ . We expand the calculation back to 1840 using the *gufm1* model (Jackson et al.,  
367 2000) for years 1840-1950 and IGRF models for years 1950-2020 (Alken et al., 2021), by looking  
368 at the maximum difference in field intensity every 5 years at any point on Earth surface. The  
369 maximum rate of change during the past 190 years is not significantly different than today:  $0.18$   
370  $\mu\text{T}/\text{year}$  or  $0.33 \text{ Z Am}^2/\text{year}$ . These rates are thus significantly lower than the rates observed during  
371 the spikes time interval. Hence, LAC.v.1.0 also places new robust constraints for how fast the field  
372 can change.

373 We note that the rates calculated using LAC.v.1.0 appear more moderate than first considered in  
374 previous studies (Ben-Yosef et al., 2009; Shaar et al., 2011), which had raised questions regarding  
375 the dynamo processes behind them (Livermore et al., 2014; Troyano et al., 2020). Although the  
376 spikes are now within the range of variations permitted by our current understanding of the  
377 geodynamo (Davies & Constable, 2017; Davies & Constable, 2018; Livermore et al., 2014), the  
378 fluctuations associated with spikes remain unprecedented in their amplitude and rate.

379

## 4.3 Long term evolution of the Levantine Iron Age Anomaly

380 The archeo-magnetostratigraphy of Tel Megiddo reveals the Holocene greatest amplitude change  
381 at multi-century scales between 1750 BCE to 1000 BCE. The increase began with a minimum of  
382  $73 \text{ Z Am}^2$  in the 18<sup>th</sup> century BCE - a period characterized by the lowest intensities in the Near  
383 East over the last five millennia, comparable to the low values at the beginning of the 3<sup>rd</sup>  
384 millennium BCE (Fig. 6b). The 700 years-long increase is nearly continuous, punctuated by a  
385 century-scale peak around 1500 BCE. The spike period may be, therefore, a climax of a long-term  
386 evolution in geomagnetic field intensity in the Near East, which could result from the occurrence  
387 of intense and rapidly evolving flux bundles at the core-mantle boundary. It also contrasts with the  
388 period spanning from the third millennium BCE to the 18<sup>th</sup> century BCE, which shows intensity  
389 peaks of moderate amplitude associated with lower rates of change of less than  $0.2 \mu\text{T}/\text{year}$  (Figs.  
390 6b,c; see also Gallet et al. (2020)). The significant increase in intensity would then result in spikes

391 instead of intensity peaks, with the first events also being shorter and apparently more frequent,  
392 which could reflect a remarkable change in core dynamics.

## 393 **5 Conclusions**

394 We report here archaeomagnetic data from Tel Megiddo, which can be considered as the largest  
395 archaeomagnetic dataset from a single site, with 23 groups of indicative pottery collected from 18  
396 consecutive radiocarbon-dated archaeological strata.

397 We assemble a new archaeomagnetic compilation of the Levant and Western Upper Mesopotamia  
398 between 3000 BCE to 550 BCE with 142 different groups of samples. The interval 1700 BCE to  
399 550 BCE is based mostly on contexts that are directly dated using radiocarbon and/or clear  
400 association with historical dating, providing an unprecedented sub-century resolution over long  
401 millennial timescale. We use this compilation to calculate the Levantine Archaeomagnetic Curve  
402 (LAC.v.1.0) – a Bayesian regional curve for high-precision archaeomagnetic dating.

403 The LAC depicts four geomagnetic spikes between 1050 BCE to 600 BCE, defining new upper  
404 limits for both the maximum local field values and rate of change. Considering the overall  
405 uncertainty, we suggest 155  $ZAm^2$  and 0.5  $\mu T/year$  (1  $ZAm^2/year$  in VADM values) as  
406 conservative upper boundaries for these quantities.

407 As a concluding remark we highlight the challenge in constructing a robust continuous  
408 geomagnetic intensity curve at a sub-centennial resolution over a large time interval. This is made  
409 possible in this study by exploiting the advantage of the Near East that allows acquisition of large  
410 multiple archaeomagnetic datasets, with precise dating, stratigraphic constraints, and cross-  
411 correlation between sites. In this respect, Tel Megiddo is an exemplary case study demonstrating  
412 a strong link between archaeology, radiocarbon and geomagnetism.

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420

## 421 **Open Research**

422 All measurements data are available in the MagIC database ([earthref.org/MagIC/19395](http://earthref.org/MagIC/19395))

423

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