

1 **Permian magnetostratigraphy and end of the Kiaman Reverse Polarity**
2 **Superchron from the southeast Karoo basin, South Africa.**

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9
10 **Key points**

- 11 • The end of the Kiaman Reverse Polarity Superchron is identified in the
12 southeast Karoo basin allowing regional correlation
13 • The effect of Karoo large igneous province dolerite intrusions is limited, and
14 primary Permian remanence define magnetostratigraphy
15 • The diachronous boundary between the Ecca and Beaufort groups is calibrated

16
17 **Abstract**

18 Paleomagnetic results and a ~2353 m-thick magnetostratigraphic section for
19 undeformed late to early Permian rocks in the south-eastern part of the Karoo basin of
20 South Africa are reported. Pseudo-single domain or single domain titanomagnetite as
21 well as pyrrhotite were identified as possible remanence carriers of a dual polarity
22 magnetization interpreted as the record of the Permian geomagnetic field during the
23 Kiaman Reverse Polarity Superchron and subsequent Illawara mixed polarity zone.
24 The timing of remanence acquisition is further constrained by the effect of Jurassic-
25 aged dolerite intrusions, which either partially or wholly overprint the Permian
26 remanence in their immediate vicinity. A paleopole at 53.2°S; 46.9°E and $dp/dm =$
27 $5.9^\circ/6.3^\circ$ is calculated from the bedding-corrected primary remanence corrected for the
28 effects of inclination shallowing using a correction factor of $f = 0.6$. This is comparable
29 to known Permian paleopoles from the southwestern Karoo basin and overlaps with
30 the 290-230 Ma segment of the Gondwana apparent polar wander path. The end of
31 the Kiaman Reverse Polarity Superchron can be correlated between the extremes of
32 the southern Karoo basin and reveal a diachronous boundary between the Ecca and
33 the Beaufort groups that is calibrated for the first time.

34

35 **Plain Language Summary**

36

37 Magnetostratigraphy is used as a dating tool for calibrating the correlation of middle to
38 late Permian sedimentary rock units. During the Permian, the Kiaman Reverse Polarity
39 Superchron is a prominent ca. 318 Ma to 265 Ma magnetostratigraphic interval of
40 reverse polarity described from east Australia and is followed by the Illawara interval
41 of mixed polarity. The end of the Kiaman Reverse Polarity Superchron is a useful
42 chronostratigraphic marker horizon. We identify the end of the Kiaman Reverse
43 Polarity Superchron within drill core from the southeast Karoo basin and correlate this
44 to its record within the southwest of that basin at localities that are 900 km apart.
45 Another prominent marker horizon in the Karoo basin (i.e., the boundary between the
46 Ecca and the Beaufort groups) is characterized by a change in depositional
47 environment. That is a change from a marine environment to deposition in a terrestrial
48 environment. Although this important lithostratigraphic boundary is known to be
49 diachronous across the Karoo basin, we here illustrate and calibrate the degree of
50 time transgression of this boundary in the southern part of the basin.

51

52 **Index terms**

53 1520 Magnetostratigraphy; 1535 Reversals: process,
54 timescale, magnetostratigraphy; 1540 Rock and mineral magnetism

55

56 **Keywords**

57 Magnetostratigraphy; Permian; Kiaman Reversed Polarity Superchron; Karoo large
58 igneous province; Karoo basin; Gondwana

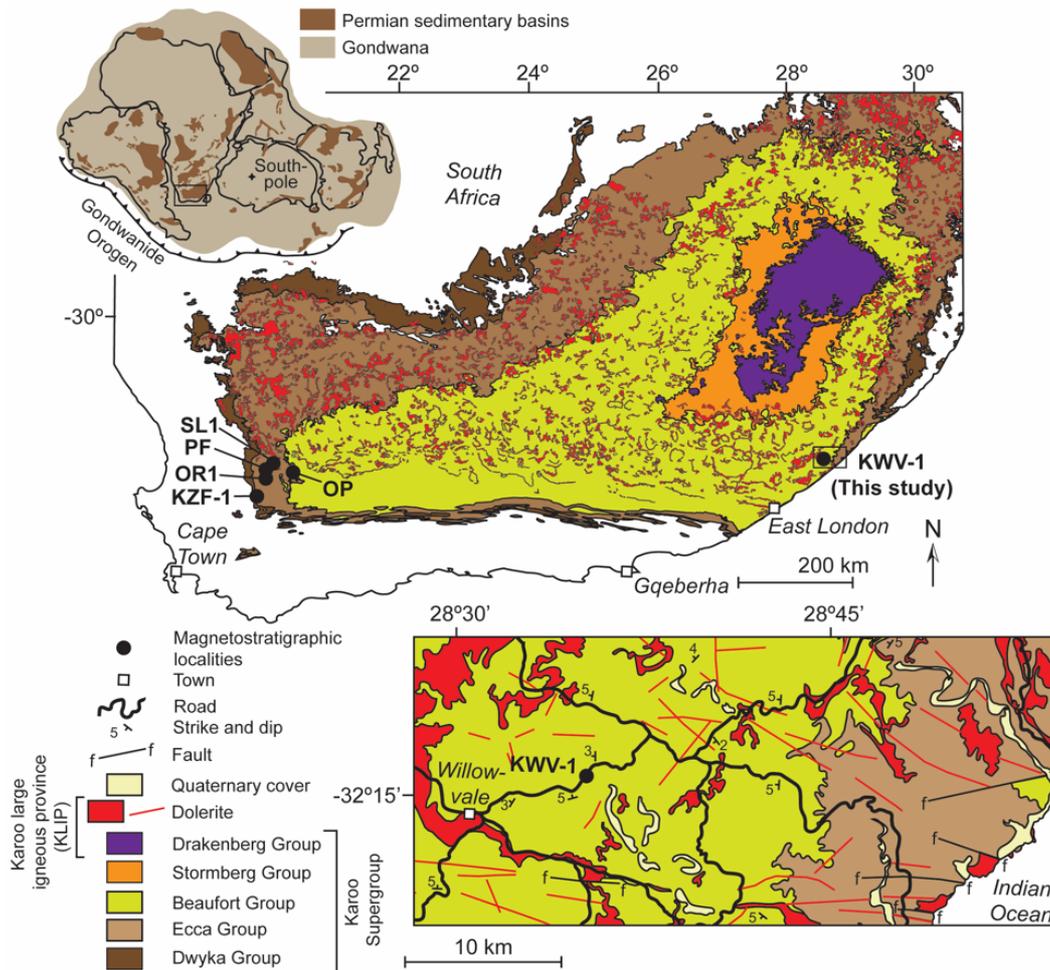
59 **1. Introduction**

60
61 Within the southern part of the main Karoo basin, the late Carboniferous to mid-
62 Jurassic Karoo Supergroup of South Africa (Figure 1) represents a 12 km-thick,
63 laterally continuous, clastic sedimentary succession that was deposited in a retro-arc
64 foreland setting in response to subduction and accretionary tectonics along the
65 southwestern margin of Gondwana (Johnson et al., 1996; Johnson et al., 2006).
66 During the Permian an interplay between regional tectonics, an icehouse to
67 greenhouse climatic transition, and a marine incursion resulted in deepening of the
68 southern main Karoo basin and deposition of the Ecca Group (Catuneanu et al., 2002;
69 Cole, 1992).

70 The Ecca Group is well documented in the southwest section of the basin in terms
71 of lithostratigraphy (Cole, 1992; Wickens, 1994), geochronology (Bangert et al., 1999;
72 Belica et al., 2017a; Fildani et al., 2007; Fildani et al., 2009; Griffis et al., 2019; McKay
73 et al., 2016; McKay et al., 2015; Turner, 1999) and magnetostratigraphy (Abubakre
74 and De Kock, 2021; Belica et al., 2017a; Lanci et al., 2013; Tohver et al., 2015). It,
75 however, remains underexplored in the southeast where outcrops are fewer, the
76 lithofacies very different (Johnson, 2009), and the stratigraphy interrupted by dolerite
77 sills of the ca. 183 Ma Karoo large igneous province (KLIP). Magnetostratigraphy can
78 be used as a dating tool for calibrating the lithostratigraphic correlation of the Ecca
79 Group across basin between the dated southwest and the undated southeast sections.
80 For this the Kiaman Reverse Polarity Superchron is a prominent Paleozoic
81 magnetostratigraphic interval. It is an interval of reverse polarity described from eastern
82 Australia that extend from ca. 318 Ma to 267-265 Ma, followed by the Illawara interval
83 of mixed polarity (Belica et al., 2017b; Cortrell et al., 2008; Irving, 1963; Irving and
84 Parry, 1963). The end of the Kiaman Reverse Polarity Superchron is particularly useful
85 as a chronological marker (Hounslow et al., 2016; Kirschvink et al., 2015). It has been
86 assigned an approximate age of 267 Ma (Steiner, 2006) but recently it was constrained
87 by Belica et al. (2017b) to coincide with the Wordian-Capitanian boundary at ~265 Ma
88 (Figure 2). In the southwest Karoo basin, Lanci et al. (2013) identified normal polarity
89 intervals within the Waterford Formation of the Ecca Group and the Abrahamskraal
90 Formation of the Beaufort Group, thus placing these rocks in the Illawara mixed polarity
91 interval with an mid-Wordian age, based magnetostratigraphic age assignments of
92 Steiner (2006). This was incorporated in the latest Permian magnetostratigraphy of

93 Hounslow and Balabanov (2018) (Figure 2), but the age assignments can now be
 94 revised to approximate the Wordian-Capitanian boundary based on the work of Belica
 95 et al. (2017b).

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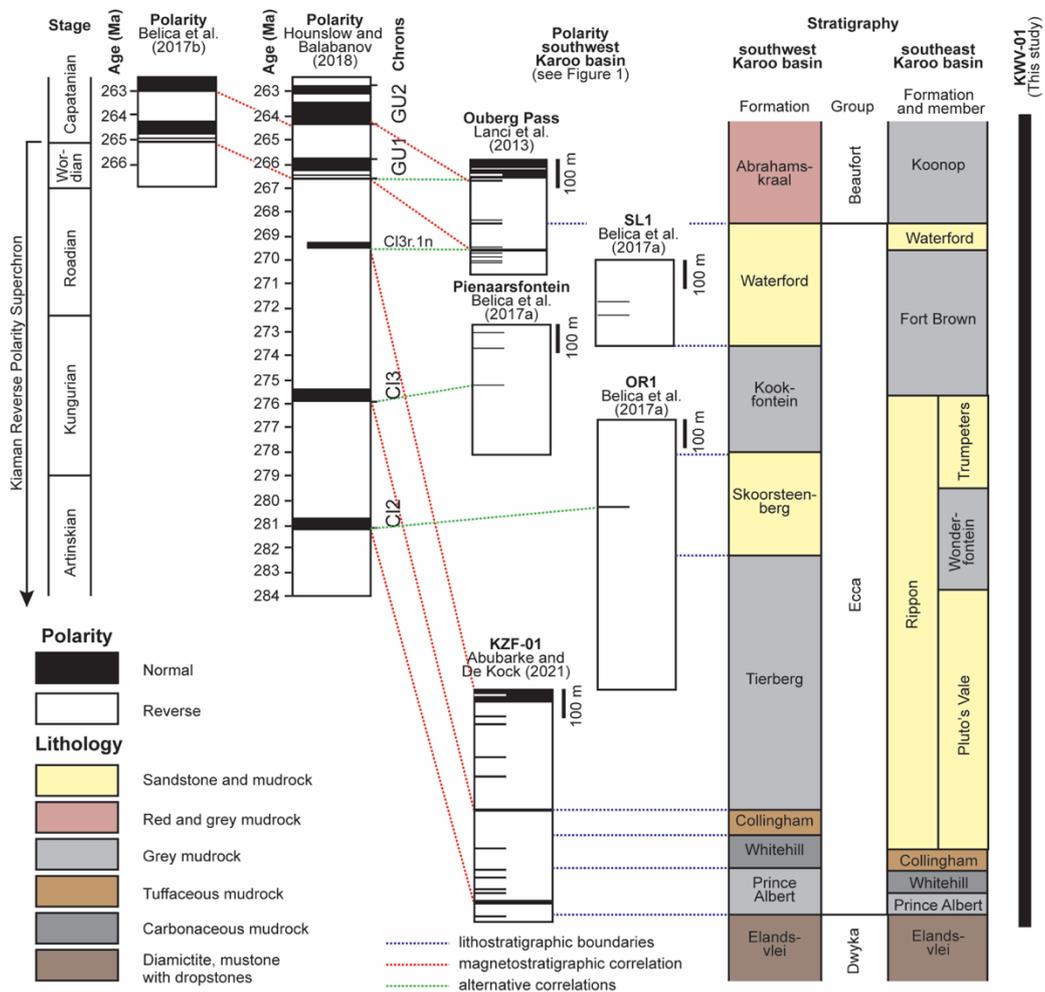


97
 98 **Figure 1.** Simplified geology of the main Karoo basin of South Africa and locality of
 99 borehole KVV-1 (i.e., enlarged inset from the 1:250 000 3228 Kei Mouth map sheet)
 100 shown in relation to published magnetostratigraphic localities (modified from the 1:1
 101 000 000 geological map of South Africa, Council for Geoscience, 1997). OP = Ouberg
 102 Pass, PF = Pienaarsfontein. Top left inset illustrates an Early Permian reconstruction
 103 of Gondwana with sedimentary basins (modified from Sirevaag et al., 2018).

104

105 Here, an integrated rock magnetic and paleomagnetic approach is used to study a
 106 2353 m-long intersection (KVV-1) of the Ecca Group and dolerite intrusions from the
 107 southeastern main Karoo basin. The magnetic and thermal effect of the KLIP intrusions
 108 on surrounding sedimentary rocks are characterized and magnetostratigraphic

109 constraints are presented for the Eccca Group and lower Beaufort Group to assist
 110 lithostratigraphic correlation between extremes of the southern Karoo basin. The
 111 effects of KLIP intrusions on the Eccca Group are important to evaluate because
 112 carbonaceous mudstone units of the Eccca Group have been identified as a potential
 113 shale-gas resource (Decker, 2014; Decker and Marot, 2012; Kuuskraa et al., 2013).
 114 The scale of this resource, however, is unconstrained and depends amongst other
 115 factors on the lithostratigraphy of the Eccca Group throughout the basin and thermal
 116 effects associated with KLIP intrusions (Cole, 2014; De Kock et al., 2017; Mowzer and
 117 Adams, 2015).



118
 119 **Figure 2.** Mid and late Permian composite magnetostratigraphy of Hounslow and
 120 Balabanov (2018) and corrected chronology of the end of the Kiaman Reversed
 121 Polarity Superchron by Belica et al. (2017b) compared to magnetostratigraphic
 122 constraints from the southwestern Karoo basin (see Figure 1 for localities) and
 123 stratigraphy of the Eccca Group (after Johnson, 2009). The stratigraphic position of
 124 borehole KWV-01 is indicated for comparison.

125 **2. Geological setting**

126

127 The Karoo basin fill commences with the glaciogenic Dwyka Group followed by the
128 marine Ecca Group and the fluvial-deltaic Beaufort Group (Johnson et al., 2006). The
129 subsequent clastic sedimentary sequence of the Stormberg Group reflects
130 increasingly arid conditions (Bordy et al., 2004; Johnson et al., 1996) and ends
131 abruptly with the rapid KLIP sill emplacement (Hastie et al., 2014; Jourdan et al., 2005)
132 and eruption of the Drakensberg Group basalts (Figure 1).

133

134 KVV-1 (S 32°14'43.10" E 28° 35'08.10") was drilled near Willowvale, within a borrow
135 pit for road aggregate, in the Amathole district in the Eastern Cape Province of South
136 Africa in the southeast extreme of the main Karoo basin and within a network of multiple
137 dolerite sills (Figure 1). Bedding in the core is variable, but generally near horizontal
138 with an average dip to the northwest based on electrical dip meter readings (dip
139 direction = 302°, dip = 2°). KVV-1 provides a continuous 2352.39 m-thick succession
140 from the uppermost layers of the Dwyka Group (i.e., the Elandsvlei Formation, below
141 2331.7 m depth), the entire Ecca Group (subdivided into Prince Albert [2302.4-2331.7
142 m], Whitehill [2269.2-2302.4 m], Collingham [2002.4-2269.2 m], Ripon [742.8-2002.4
143 m], Fort Brown [262-742.8 m] and Waterford [241.8-262 m] formations) and the
144 lowermost part of the Beaufort Group (i.e., Koonap Formation [0-241.8 m], Figure 3).
145 The lithostratigraphy of the Ecca Group here stands in contrast to the southwestern
146 part of the basin where the Tierberg Formation is developed above the Collingham
147 Formation (Figure 2). The Ripon Formation, which occurs east of 22° consists of fine-
148 grained greywacke turbidite beds of varying thicknesses and is subdivided into the
149 Pluto's Vale Member sandstone (1388-2002.4 m), the Wonderfontein Member shale
150 (948.1-1388 m), and the Trumpeter's Member sandstone (742.8-948.1 m; Figure 3).
151 Another marked difference compared to the southwestern part of the basin, is the
152 presence of numerous dolerite intrusions that intersect the clastic sedimentary rock
153 succession. A 18.82 m-thick fine-grained dolerite intrudes the top of the Whitehill
154 Formation with a sharp contact in KVV-1. The top of this dolerite intrudes sharply into
155 the base of the 712.14 m-thick arenaceous Pluto's Vale Member of the Rippon
156 Formation. The Pluto's Vale Member is further intersected six times by dolerite
157 intrusions with a total cumulative thickness of 234.86 m. The argillaceous
158 Wonderfontein Member is intruded by a 46.07 m-thick dolerite intrusion between

159 1249.79 and 1203.72 m. The Trumpeter's Member envelopes a 9.6 m-thick dolerite
160 intrusion between 965.12 and 955.48 m. In the upper parts of the core, the Fort Brown
161 Formation is cut by eight thin dolerite intrusions with a total cumulative thickness of
162 71.75 m, and the Koonap Formation is intruded by two dolerite sills (i.e., 34.87 m and
163 1.87m thick) between 127.11 and 92.24 m-depth, and between 16.57 and 14.7 m-
164 depth.

165

166 **3. Method**

167

168 As the KVV-1 core was retrieved, pieces of core was fitted for continuity before being
169 marked with a reference line and indicating the upward direction with an arrow head
170 by the managing geologists at the drill site (e.g., Figure 3D). Where successive pieces
171 of core could not be fitted for continuity due to excessive breaking or core loss the core
172 was marked with two short parallel lines perpendicular to the reference line, a new
173 reference line was then selected and drawn along the next section of continuous
174 retrieved core. The core is thus made up of several sections of continuous rock
175 intersection, with pieces of core from a section being oriented relative to each other.
176 Such sections ranged from less than a meter to over 100 meters. Orientation of the
177 reference line on any section was determined relative to the bedding dip azimuths
178 intersected in the core and dip meter readings down the hole. Generally the orientation
179 line coincided with bedding dip azimuth. Electronic dip meter logging was not possible
180 beyond 2000 m due to hole collapse after core retrieval. Orientation of the reference
181 line from here to the end of hole was attempted by using the average bedding in the
182 core (i.e., $302^{\circ}/2^{\circ}$). Core logging and sampling was done at the Council for
183 Geosciences National Core Library at Donkerhoek, South Africa. Plug samples for
184 paleomagnetic analyses were taken at ~3m intervals, but this spacing was reduced
185 near dolerite contacts. A piece of core earmarked for sampling was first split along the
186 reference line, before a paleomagnetic sample was plugged perpendicular to the core
187 axis and from the right-hand half of the split core using a benchtop drill press. Plugs
188 were labelled with their depth and their azimuth was marked along their upper margins.
189 Plug azimuths are equal to the orientation of the reference line minus 90° . Samples
190 were oriented and prepared into specimens for step-wise demagnetization at the
191 paleomagnetism laboratories at the University of Johannesburg (UJ) and California

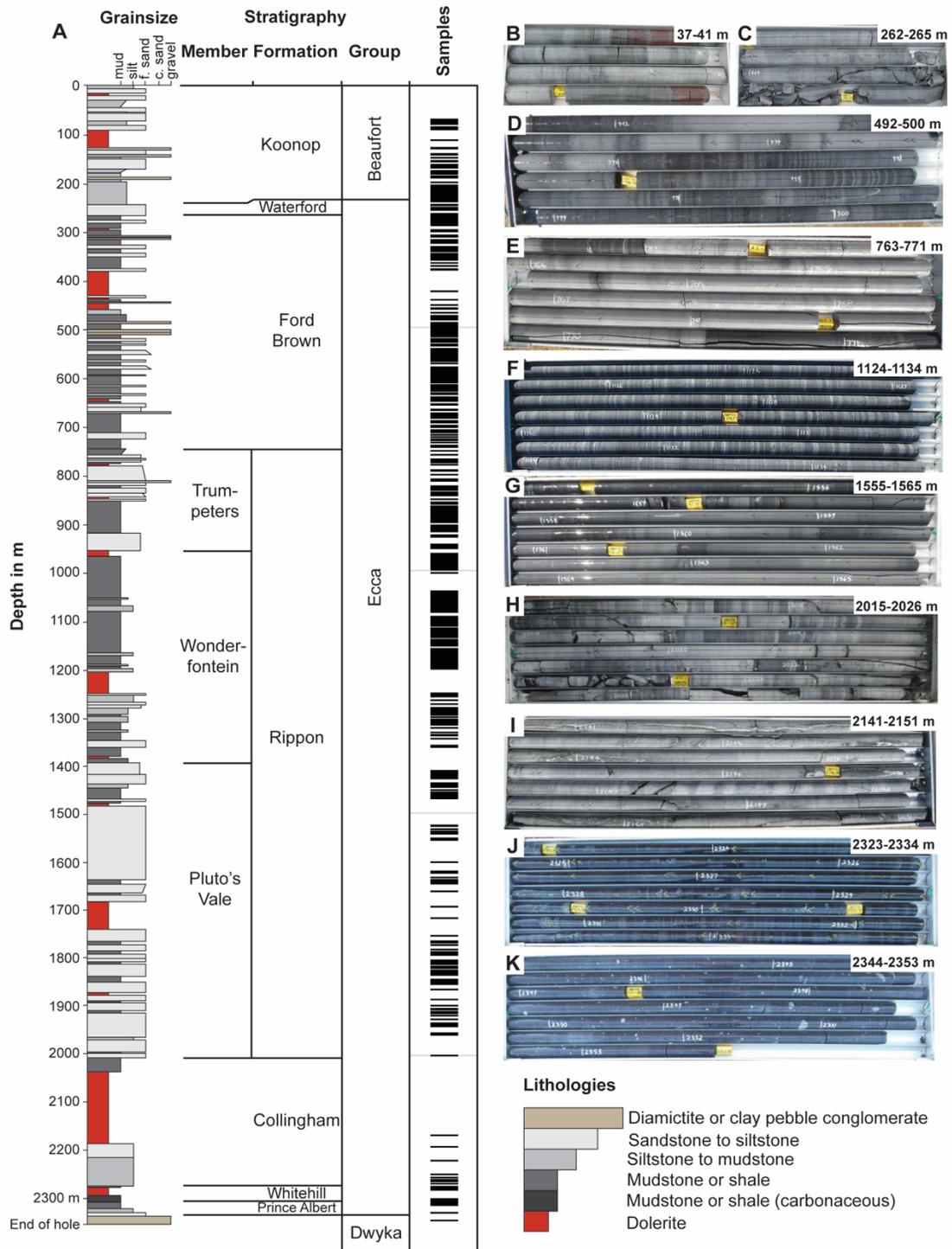
192 Institute of Technology (Caltech). Demagnetization consisted of low-intensity
193 progressive alternating field (AF) demagnetization in three to four 25 mT steps up to
194 75 mT or 100 mT. Thermal demagnetization was achieved through stepwise heating
195 in 20-25°C steps from 100°C to 600°C. All measurements of remanence were made
196 using 2G Enterprises 755 or DC-4K superconducting rock magnetometers.

197 Remanence components were identified and quantified via principal component
198 analysis (Kirschvink, 1980) using PaleoMag 3.2 (Jones, 2002) with individual
199 components calculated from three or more points for line fits and four or more points
200 for plane fits. Component fits were used in subsequent analysis if fits mean angular
201 deviation (MAD) values of $\leq 15^\circ$ for line fits and $\leq 20^\circ$ for plane fits. Component means
202 were calculated using PaleoMag 3.2, which employs the methodology of McFadden
203 and McElhinny (1988) for the combination of line and plane fits. Component means
204 were used to calculate virtual geomagnetic poles (VGPs) as per Butler (1992). The
205 VGPs were visualized using GPlates 2.2 (Williams et al., 2012). Remanence
206 components interpreted as being primary were used to develop a magnetostratigraphic
207 record by considering the stratigraphic variation of declination, inclination and the
208 corresponding VGP latitude of specimens.

209 Fourteen representative samples were subjected to a suite of rock magnetic
210 experiments as described by Kirschvink et al. (2008). This includes progressive
211 isothermal remanent magnetization (IRM) acquisition up to 1000 mT followed by
212 progressive AF demagnetization, progressive anhysteretic remanent magnetization
213 (ARM) acquisition with a variable direct current biasing field of 0 to 1 mT, an ARM
214 version of the Lowrie-Fuller test (e.g., Johnson 1975), and a comparison of natural
215 remanent magnetization (NRM) and ARM stability with IRM stability after Fuller et al.
216 (2002). IRM acquisition curves were unmixed using the MAX Unmix software
217 (Maxbauer et al., 2016), which employs a skewed-normal distribution (Wuertz and
218 Chalabi, 2015) to represent individual coercivity components.

219

220



221

222 **Figure 3.** Simplified stratigraphic log of KVV-01 and sample positions (A).
 223 Representative photographs of core from the Koonop Formation (B), Waterford
 224 Formation (C), Fort Brown Formation (D), Trumpeters Member (E), Wonderfontein
 225 Member (F), Pluto's Vale Member (G), Collingham Formation (H), dolerite (I), Whitehill
 226 Formation (J), and the Dwyka Group (K). Note the reference line marking relative
 227 orientation and up direction on the core.

228 4. Results

229 4.1. Rock magnetic results

230 Nine of 14 samples acquire IRM linearly (on a log scale) from 20-30 mT to 300 mT,
231 but only reach or approach saturation above 700 mT (Figure 4A). AF demagnetization
232 of IRM define mean destructive fields of 61-69 mT, while H_{CR} values are typically 89-
233 100 mT with R-ratios of 0.35-0.40 (after Cisowski, 1980). IRM is acquired more steeply
234 from 20 mT to 100 mT in two samples (i.e., 475.14 and 2245.3), and saturation is
235 reached at 200-400 mT (Figure 4B). These samples have low mean destructive fields,
236 H_{CR} values and R-ratios. IRM is acquired non-linearly in three samples (i.e., 191.13,
237 916.45, and 996.51). The samples approach saturation at fields above 700 mT.
238 Steeper acquisition is seen in 191.13 and 996.51 before 100 mT, followed by less
239 steep acquisition that becomes steeper again after 300 mT. These samples have
240 mean destructive fields of 36-48 mT, H_{RC} values of 48-60 mT, and R-ratios of 0.40-
241 0.44 (Figure 4C). Sample 916.45 show shallower acquisition before 100 mT, after
242 which acquisition becomes steeper. It has a mean destructive field of 121 mT, H_{RC} of
243 139 mT, and a R-ratio of 0.46 (Figure 4D).

244 IRM acquisition curve decomposition identifies six possible components. When
245 fitting such models, a strategy was followed to keep components within a relatively
246 constant coercivity range throughout the sample set. Each component is described by
247 its median coercive field, dispersion parameter, relative proportion, and skewness
248 (Supporting Information S1). Any specific sample require two, but more commonly
249 three, and rarely four phases to match the IRM acquisition spectra. Generally, IRM
250 acquisition is dominated by a medium coercivity or MC1 component (mean coercivity
251 = 122 ± 27 mT; Figure 4E). Samples that reach saturation above 700 mT display small
252 contributions of both low and high coercivity phases (LC1 with a mean coercivity of 26
253 ± 6 mT; and HC1 with a mean coercivity of 488 ± 51 mT) in addition to the main
254 contribution of MC1, and in one case there is also a small contribution from a very high
255 coercivity phase (i.e., HC2 with a mean coercivity of 1284 ± 401 mT). Samples that
256 reach saturation at 200-400 mT display either low (i.e., LC2 with a mean coercivity of
257 51 ± 7 mT) or medium coercivity phases (i.e., MC1) as main contributors to the IRM
258 together with a small contribution of LC1, but without higher coercivity phases
259 contributing (Figure 4F). Non-linear IRM acquisition is characterized by either
260 dominant LC2 contributions (in 191.13 and 996.51; Figure 4G) or by a dominant
261 medium coercivity (i.e., MC2 with a mean coercivity of 274 ± 47 mT) in 916.45 (Figure

262 4H). There are additional small contributions of LC1 and MC2 phases in 191.13, and
263 by MC2 and HC2 phases in 996.51. Apart from a dominant MC2 phase, a small LC1
264 contribution is also recognized in the IRM acquisition by sample 916.45.

265 An ARM version of the Lowrie-Fuller test (Johnson et al., 1975) shows harder
266 AF demagnetization of ARM than that of the IRM for all specimens (Figure 5A), thus
267 indicating the dominance of interacting single-domain (SD) particles. The ARM
268 acquisition for all the specimens confirm high inter-particle interaction. Interaction is
269 highest for 231.42, 425.45, and 475.14, which resemble ARM acquisition by chiton
270 teeth. Curves otherwise resemble ARM acquisition by partially collapsed bacterial
271 magnetosomes (Figure 5B). The least inter-particle interactions are seen in 2245.3,
272 996.51, 916.45, and 191.13.

273 Plots of NRM and ARM demagnetization values against IRM demagnetization
274 for select specimens reveal a conspicuous soft NRM component in most specimens
275 that is removed by 10 mT (Figure 5C). The ratio of NRM to IRMs is generally in the
276 1:1000 range for demagnetization between 10 mT and 50 mT for most specimens, but
277 for four specimens (716.74, 916.45, 996.51, and 1764.7) the ratio was in the 1:100
278 range. At higher field levels the AF demagnetization of NRM proceed much slower
279 than the AF demagnetization of IRM, and the curve is characteristically concave
280 upwards for all our specimens. The ARM behaves similarly, albeit less pronounced.
281 ARM:IRM values are above 0.1 in all specimens and are equal to, or exceed 1 at
282 demagnetization levels above 70 mT (Figure 5C).

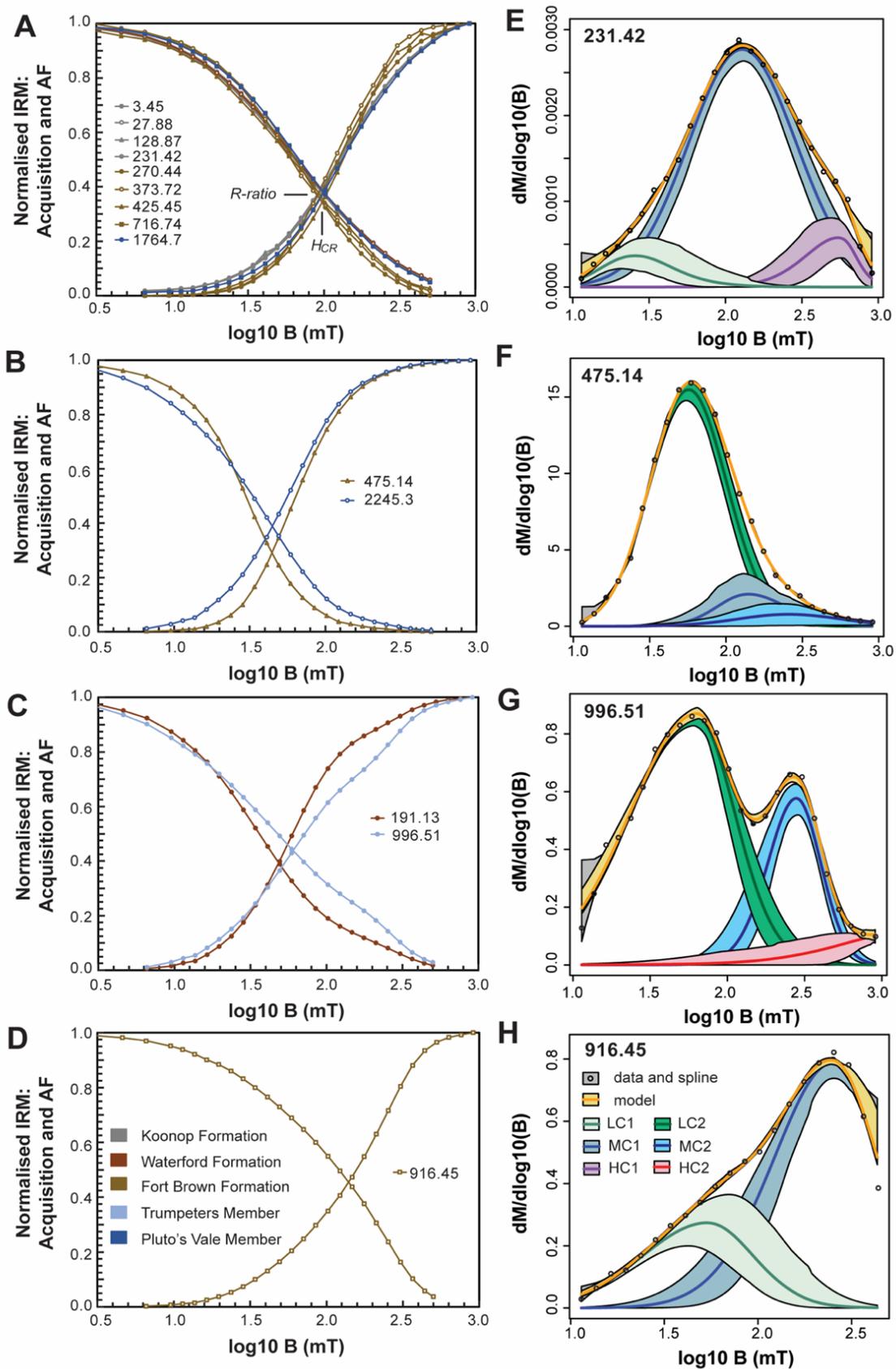
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284 **4.2. Demagnetization**

285 **4.2.1. Sedimentary rocks**

286 Four magnetic components were identified during demagnetization of sedimentary
287 rock specimens (Table 1). The demagnetization behaviour was variable as expected
288 for the range of different lithologies and grain sizes. The four components were a low
289 coercivity or soft magnetic component (SFT); a variable, but typically north to
290 northwest directed and moderately steep upward directed component (A); a northerly
291 and steep upward directed high stability magnetization (B-); and a southerly and steep
292 downward directed high stability magnetization (B+). Any specific specimen, however,
293 recorded a maximum of three components, and many were characterized by just two
294 components or a single magnetic component (Figure 6 and 7).

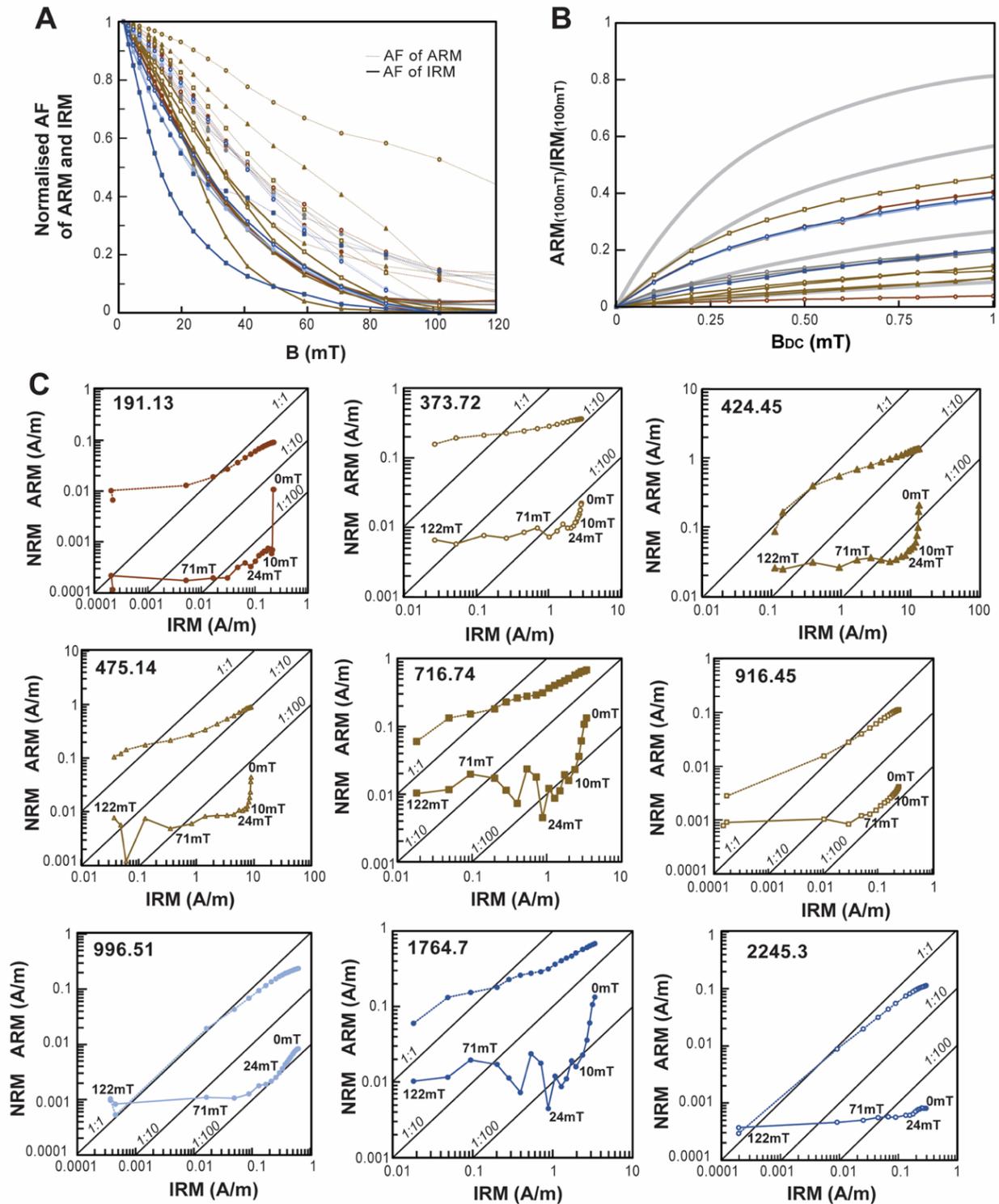
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297 **Figure 4.** Acquisition and AF demagnetization of IRM of selected samples (A-D) and

298 components of IRM identified through IRM acquisition curve decomposition (E-H).



299

300 **Figure 5.** ARM version of the Lowrie-Fuller test for selected samples (A); ARM
 301 acquisition (B) for selected samples compared against intact magnetotactic bacteria
 302 (top grey line), partially collapsed bacterial magnetosomes (middle grey line), and
 303 chiton teeth (lower grey line); and plots of NRM and ARM demagnetization values
 304 against IRM demagnetization for selected samples (C). The lines in the background
 305 mark different ratios of NRM or ARM to IRM (i.e., 1:1, 1:10 or 1:100).

306 The demagnetization behaviour of nine specimens proved too erratic for the
307 identification of any magnetic components (Figure 6A), and these are not further
308 considered.

309 A low coercivity SFT components was identified in 220 of 400 specimens during
310 AF demagnetization steps up to low temperature demagnetization steps (Figure 6, 7
311 and 8). The SFT component is poorly defined in most with MAD values $> 15^\circ$ for fits
312 from 56 of the specimens. These components are randomly directed and not
313 considered to be of any geological significance.

314 After the removal of SFT components, prominent remanence components (i.e.,
315 component A) unblocked as either linear trajectories towards the origin or as linear
316 trajectories that clearly miss the origin to reveal the presence of remanence
317 components of higher thermal stability (Figure 6 and 7). These A components were
318 identified in 311 specimens (plus 34 specimens in which it was identified, but which
319 were excluded due to their $> 15^\circ$ MAD values). A components are highly variable in
320 terms of their unblocking and direction. They are generally northerly and upward
321 directed with a component average declination of 350.2° , inclination of -75.5° and α_{95}
322 of 3.74° , but with a very low precision parameter ($k = 5.55$) and a correspondingly
323 large standard deviation (Figure 9A). The A components either unblock by 340°C or
324 by 460°C in general.

325 In specimens where A component trajectories do not demagnetize towards the
326 origin, higher stability components (B- and B+) are revealed as either stable end-points
327 of demagnetization, linear demagnetization trajectories towards the origin, or more
328 commonly as demagnetizations as planes away from the A component direction
329 towards B+ component directions seen on equal area plots (Figure 6C-F and Figure
330 7E-F). The B- and B+ components variably unblock between 250°C and 520°C , a
331 rather wide range of thermal demagnetization steps, but most commonly unblock
332 above 400°C . B- components were identified in 65 specimens (plus 2 specimens in
333 which it was identified, but which were excluded due to their $> 15^\circ$ MAD values). The
334 B- components are northerly and upwards directed and often steeper than A
335 components. The component average was calculated at declination = 319.3° ,
336 inclination = -73.6° , $\alpha_{95} = 5.86^\circ$, and $k = 10.02$ (Figure 9B). The distribution of B-
337 components is obviously elongated, but this is an artifact of imperfect orientation of
338 the drill core (see De Kock et al., 2009 for a description of this artifact). B+ components
339 were most commonly identified as demagnetization along planes away from the A

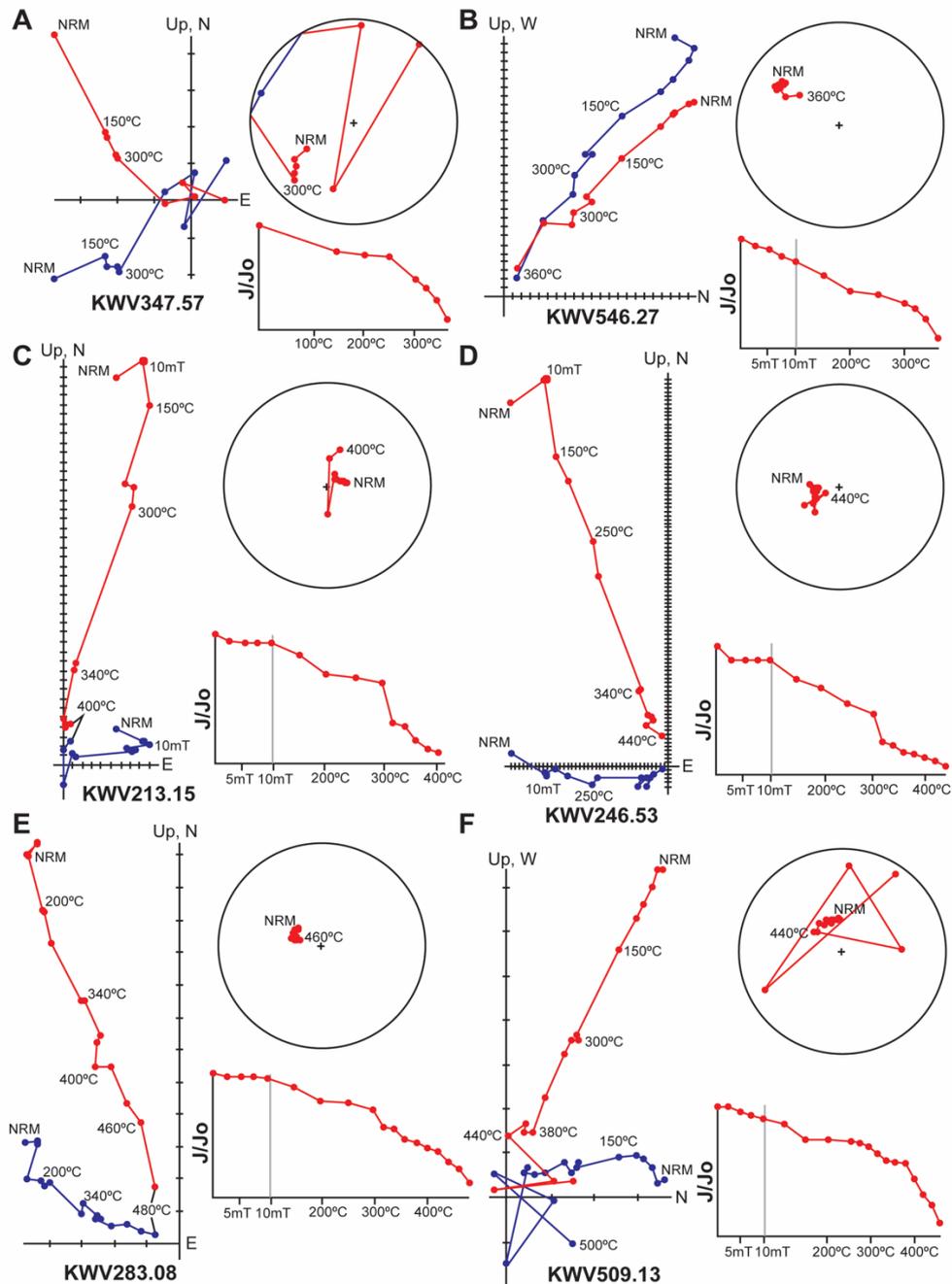
340 component direction (i.e., in 122 specimen after the exclusion of 10 specimen for
341 which planes had MAD values that were $> 20^\circ$), but also as components with linear
342 demagnetization trajectories towards the origin after the removal of SFT and A
343 components (i.e., in 37 specimens), and rarely (i.e., in 9 specimens) as single
344 remanence components (Figure 7A-F). Considering only line data, a B+ component
345 average is calculated at declination = 144.6° , inclination = 62.2° , $\alpha_{95} = 9.34^\circ$, and $k =$
346 6.05 (Figure 9C). Like the B- components, the B+ components also define an
347 elongated distribution. Combining fitted lines and planes yield a B+ component
348 average at declination = 153.5° , inclination = 65.3° , $\alpha_{95} = 3.88^\circ$, and $k = 8.74$ (Figure
349 9C). The B- and B+ components are near antipodal, but the line fit data fail to pass the
350 reversal tests of McFadden and McElhinny (1990) with in-situ remanence directions.
351 If the polarity of the B+ mean is reversed and compared to the B- mean it can be seen
352 that the angle that separates the two means is 11.56° , which is larger than the critical
353 angle of separation (10.44°) at which the two means would be statistically
354 indistinguishable (i.e., 180° apart). Repeating the same test with tilt-corrected
355 remanence directions, however, passes a C quality reversals test. The angle that
356 separates the two means is 11.75° , which is smaller than the critical angle of
357 separation of 14.36° . A combination of B- and B+ components (line fits only) yield an
358 overall tilt-corrected mean at declination = 324.4° , inclination = -70.5° , $\alpha_{95} = 5.25^\circ$, and
359 $k = 7.50$ for 111 specimens. Inclusion of plane fits yield an overall tilt-corrected mean
360 at declination = 333.3° , inclination = -69.0° , $\alpha_{95} = 3.39^\circ$, and $k = 8.31$ for 235
361 specimens.

362

363 **4.2.2. Dolerite**

364 A 34.8 m-thick dolerite that intrudes the Koonap Formation between 127.11 and 92.24
365 m-depth. The magnetization of two specimens (i.e., KVV92.38 and KVV113.7) are
366 characterized by small contributions of SFT components and dominated by northerly
367 up directed magnetizations that remain stable above 400°C (i.e., DOL components;
368 Figure 8). Sedimentary rock specimens above and below the dolerite are
369 characterized by demagnetization along planes away from prominent A components
370 (up) towards B+ components (down). The A components of specimens surrounding
371 the intrusion is comparable to the DOL direction, and it appears that the magnetic
372 effects of this intrusion on the surrounding sedimentary units was limited.

373



374

375 **Figure 6.** Representative demagnetization behaviour of selected specimens. A)

376 KWW347.75 display an A component only (150°C to 300°C) before becoming erratic.

377 B) KWW546.27 is dominated by an origin-seeking A component as a single magnetic

378 component. C) KWW213.15, D) KWW246.53, E) KWW283.08 and F) KWW509.13 all

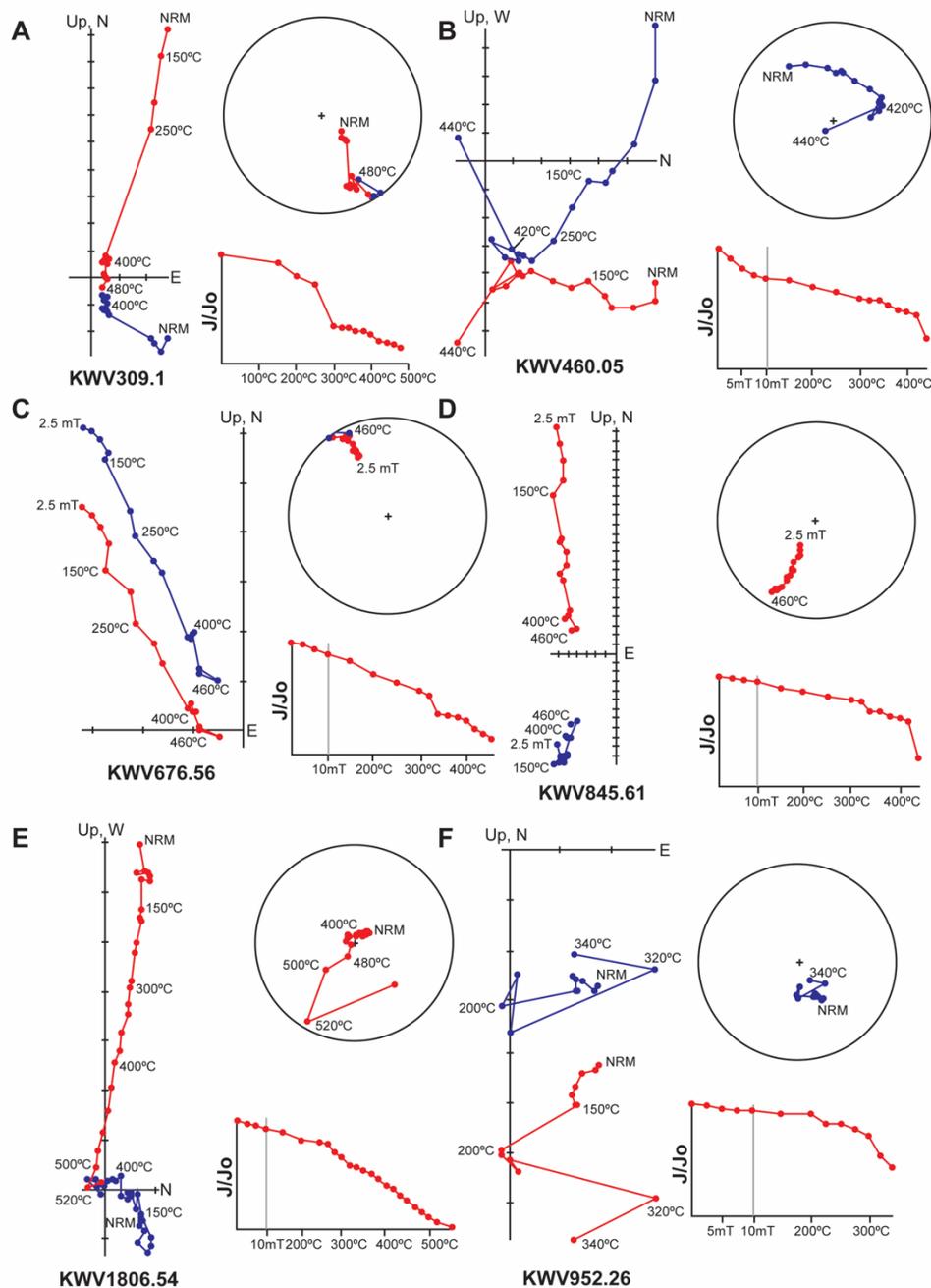
379 display A components that miss the origin up to ~350°C before steeper origin-seeking

380 trajectories define the B- components. KWW213.15 and KWW246.53 display small

381 contributions of SFT components to the remanence that are removed by AF

382 demagnetization steps. Zijderveld diagrams: red = inclination data, blue = declination

383 data. Equal-area nets: red = up, blue = down. J/Jo = normalized magnetization.



384

385 **Figure 7.** Representative demagnetization behaviour of selected specimens. A)

386 KVV309.1, B) KVV460.05, C) KVV676.56, D) KVV845.61, and E) KVV1806.56

387 display small contributions of SFT components after which an A component unblocks

388 as a linear trajectory that miss the origin. At high temperature demagnetization steps,

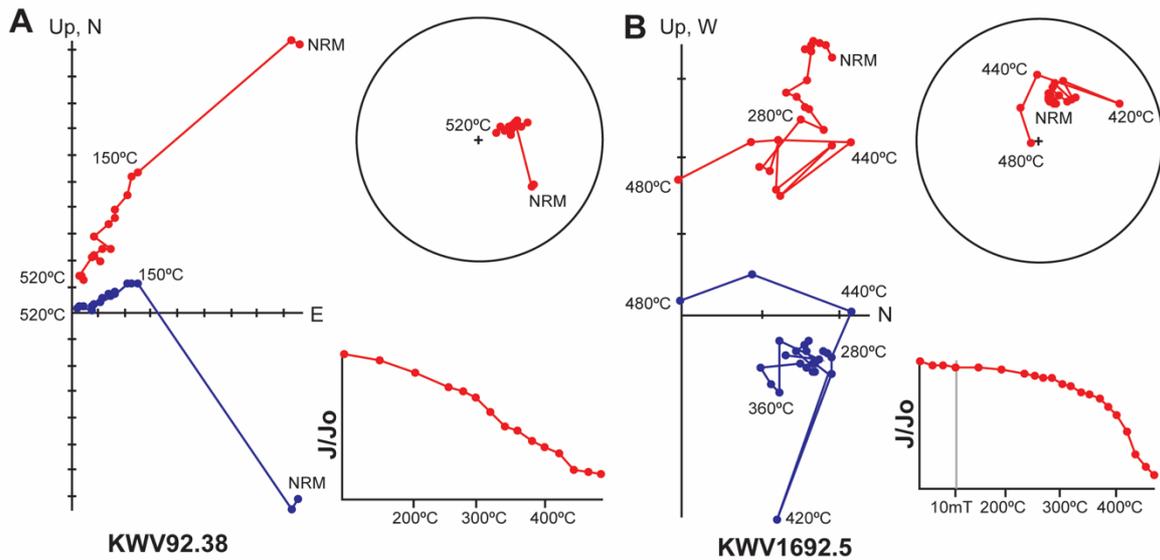
389 component B+ is revealed as planes of demagnetization away from the A component

390 direction to rarely reach stable end points of demagnetization (e.g., KVV460.05). F)

391 In a few cases like KVV952.26, component B+ is the only remanence component that

392 unblocks. Zijderveld diagrams: red = inclination data, blue = declination data. Equal-

393 area nets: red = up, blue = down. J/Jo = normalized magnetization.



394

395 **Figure 8.** Representative demagnetization behaviour of selected dolerite specimens.
 396 A) KVV92.38, B) KVV1692.5 are dominated by northeast up DOL components. Either
 397 as origin seeking-trajectories (e.g., KVV92.38) or as stable endpoints of
 398 magnetization (e.g., KVV1692.5). Zijderveld diagrams: red = inclination data, blue =
 399 declination data. Equal-area nets: red = up, blue = down. J/J_0 = normalized
 400 magnetization.

401

402 Several closely spaced dolerite intrusions with a cumulative thickness of 71.75 m
 403 intrude the upper part of the Fort Brown Formation between ~380 and ~460 m-depth.
 404 One specimen (i.e., KVV443.24) yielded a DOL component below 420°C, but another
 405 (i.e., KVV452.98) does not conform. It yielded a southwest and downward directed
 406 magnetization with a small SFT component contribution, but its demagnetization
 407 behaviour becomes increasingly erratic above 250°C. Sedimentary units up to 70m
 408 above the intrusion generally behaved erratically during demagnetization. One
 409 specimen from near the upper contact with the intrusion yielded a B+ component via
 410 a plane fit. Specimens from up to 50m below the intrusion yielded A components below
 411 ~300°C and B+ components constrained by line as well as plane fits above ~300°C .
 412 The contrast between these dolerite intrusions and the surrounding units is less clear.

413 Thin (< 5 m-thick) intrusions in the Trumpeters Member near 780 m-depth and
 414 850 m-depth were not sampled. Sedimentary rock specimens around the intrusion at
 415 780 m (i.e., KVV783.05KVV777.08) yield only north up (component A) with small SFT
 416 component contributions. Immediately above the intrusion at 850 m-depth specimen
 417 (KVV845.61) demagnetize away from a northerly up (component A) direction towards

418 a southerly down (component B+) direction. A specimen from immediately below the
419 intrusion (KVV853.18) record a single remanence component that is north-easterly
420 and up directed (component A).

421 A ~12 m-thick dolerite intrudes along with the contact between sandy units of the
422 Trumpeters Member and carbonaceous shale of the Wonderfontein Member around
423 960 m-depth. The dolerite was not sampled, but the sandy unit immediately above the
424 sill yielded well defined moderate to steep south-easterly down (B+) magnetizations
425 stable up to 460°C (i.e., KVV952.26 and KVV955.32). At lower thermal
426 demagnetization steps northerly up magnetizations (A components) were removed.
427 The Wonderfontein Member shale immediately below the sill also yielded well defined
428 moderate to steep south-easterly down (B+) magnetizations stable up to 460°C (i.e.,
429 KVV966.56) with a north up (A) component removed at lower thermal
430 demagnetization steps. The magnetic effect of this dolerite on the surrounding
431 sedimentary rocks seems to be minimal. Between 1200 to 1245 m-depth the
432 Wonderfontein Member is cut by ~45 m-thick coarse-grained dolerite. Two specimens
433 of dolerite near the top of this sill yielded northerly up magnetizations that were stable
434 up to ~300°C (DOL components), but with -17.9° the DOL component in specimen
435 KVV1203.66 was much lower inclined than the -44.4° of the DOL component in
436 KVV1200.86. Sedimentary specimen above the dolerite (KVV1197.87) yield
437 demagnetization trajectories towards a steep downward directed remanence (B+
438 component). Below the dolerite, sedimentary specimens also yield steep downward
439 directed directions. Neither sedimentary rocks associated with, nor a thin (< 5 m-thick)
440 dolerite near the base of the Wonderfontein Member around 1380 m-depth was
441 sampled.

442 A ~5 m-thick dolerite intrudes the upper Pluto's Vale Member around 1480 m-
443 depth. Neither this dolerite, nor sedimentary units in immediate association with it was
444 sampled. Multiple sills with a cumulative thickness of ~60 m intrude the upper Pluto's
445 Vale Member between 1680 and 1740 m-depth. Two dolerite specimens (KVV1692.5
446 and KVV1717.27) yielded north up DOL components that were stable up to 400°C
447 with minor contributions by SFT components. Sedimentary rock units immediately
448 above and below the intrusion displayed clear demagnetization trajectories along
449 planes away from the A components towards a B+ remanence.

450 A ~5m-thick dolerite intrusion between 1870 and 1880 m-depth was not sampled.
451 A very thick series of dolerite intrusions with a cumulative thickness of ~148 m is

452 present in the lower Pluto's Vale Member. One specimen (KVV2272.42) from these
453 intrusions reveal a northwest up DOL component that remains stable up to 300°C. A
454 second specimen (KVV2276.13) displayed erratic demagnetization behaviour and no
455 fits were attempted. Sedimentary rock units above and below the dolerite are mainly
456 characterized only by prominent A components. Only at 2251.11 m, some 20 m above
457 the intrusions is a B+ remanence recorded. Below the intrusions, the closest specimen
458 is at 2294.95 m, and it also record the B+ remanence.

459 A ~18 m-thick dolerite intrudes carbonaceous shale of the Whitehill Formation.
460 This intrusion was not sampled, but sedimentary rock specimens above and below the
461 dolerite (like the lowermost Pluto's Vale Member) at 2294.95 m and 2301.5 m record
462 the B+ remanence.

463 An average for the northerly up DOL components was calculated from six
464 specimens at declination = 57.3°, inclination = -68.8°, α_{95} = 24.04°, and k = 7.26
465 (Figure 9D).

466

467 **5. Discussion**

468 **5.1. Carriers of remanence**

469 NRM:IRM ratios in the 1:1000 range are consistent with the NRM largely being a post-
470 depositional or depositional remanent magnetization (pDRM or DRM), while those in
471 the 1:100 range are more characteristic of secondary magnetizations, possibly a
472 thermal remanent magnetization (TRM) due to the heating effects of dolerite
473 intrusions, or a chemical remanent magnetization (CRM) due to diagenetic alteration
474 in response to burial. The latter seems a less likely explanation, as there is no clear
475 stratigraphic relationship with NRM:IRM ratios. The pronounced concave upward
476 shape of the NRM:IRM curves suggest that magnetically soft phases (LC1 and LC2)
477 are not efficient carriers of weak field remanent magnetization. Samples dominated
478 by LC1 and LC2 phases reach saturation by 200 mT, suggesting these phases to be
479 multi-domain low-Ti magnetite. NRM carried in high-coercivity phases are effectively
480 immune to AF demagnetization and flattens out the NRM:IRM curves. The HC2
481 phase, which was only seen as a very minor contributor to IRM in three samples, is
482 likely very fine-grained hematite. This phase is not an important carrier of remanence,
483 because none of the specimens carried stable magnetizations after demagnetization
484 above 540°C. HC1 is present in most samples as a minor contributor to IRM and could
485 be pyrrhotite. Pyrrhotite has been identified in mudstones of the Whitehill and

486 Collingham formations in KVV by means of x-ray diffraction (Geel et al., 2021).
487 Pyrrhotite may be responsible for the drops in magnetization seen during heating of
488 specimens between 250°C and 300°C. The identities of the MC1 and MC2 phases,
489 the main carriers of the NRM are less clearly diagnosed but are likely pseudo-SD or
490 SD low-Ti magnetite phases, and together with the HC1 phase (i.e., pyrrhotite?)
491 carriers of the A, B+ and B- components.

492

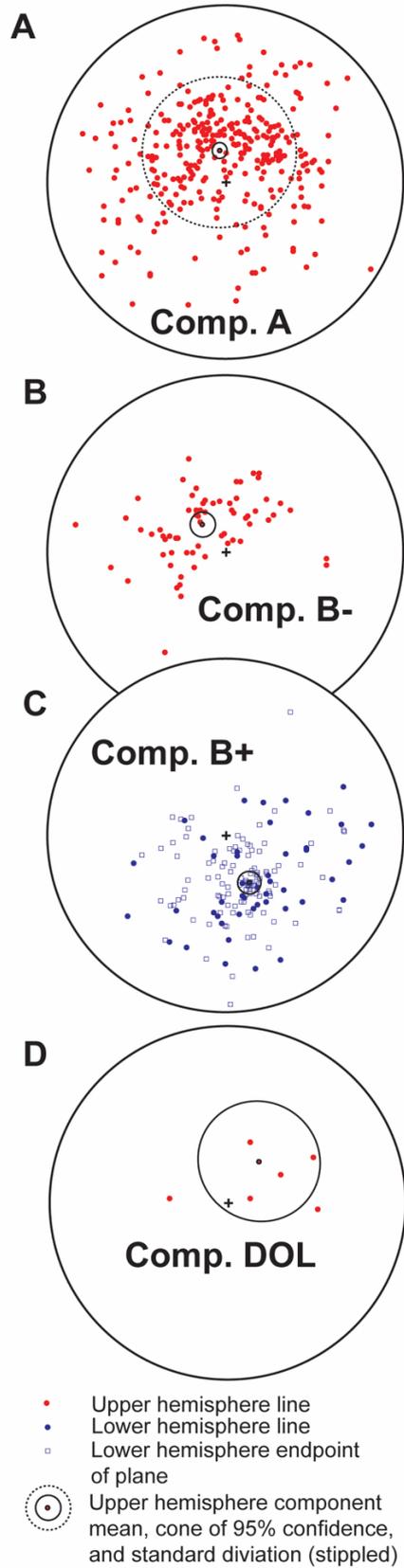
493 **5.2. Timing of remanence acquisition and paleopole calculation**

494

495 The DOL component seen in the KLIP intrusions is clearly different from the B+
496 magnetizations revealed during high temperature demagnetization steps in the
497 surrounding sedimentary units. The DOL direction is comparable to the B- directions,
498 although nowhere in direct association in the core. The DOL direction is also
499 comparable to the lower temperature A direction that is developed throughout most of
500 the core (Figure 6; Table 1). The DOL direction is, however, statistically distinguishable
501 from the B- component. The separation angle between the means is 28.3° and larger
502 than the critical angle of 20.7° at which the two means would be indistinguishable. A
503 direct comparison of the DOL and A directions shows that they are statistically the
504 same. The separation angle between the means is 20.1° and smaller than the critical
505 angle of 26.3° at which the two means would be indistinguishable. We therefore
506 interpret the A component as a younger remanence that was acquired during the
507 intrusion of the KLIP dolerites. The B+ and B- components are interpreted to be older
508 magnetizations, and likely primary Permian magnetizations recorded as pDRMs or
509 DRMs that pass a quality C reversals test for tilt-corrected remanence directions.
510 Paleopoles were calculated for the DOL and A components based on in-situ
511 remanence directions, and for the B+ and B- components based on remanence
512 directions after bedding was restored to paleohorizontal.

513

514



515

516 **Figure 9.** Identified magnetic components and their means. A) Component A, B)
 517 Component B-, C) Component B+, and D) Component DOL. All diagrams are of in-
 518 situ remanence directions.

519 B-, B+ and their combined paleopoles are comparable to the 240 Ma to 180 Ma
 520 segment of the Gondwana reference apparent polar wander path of Torsvik et al.
 521 (2012). We follow Torsvik et al. (2012) by applying a correction for inferred inclination
 522 shallowing for detrital sediments of $f = 0.6$, where:

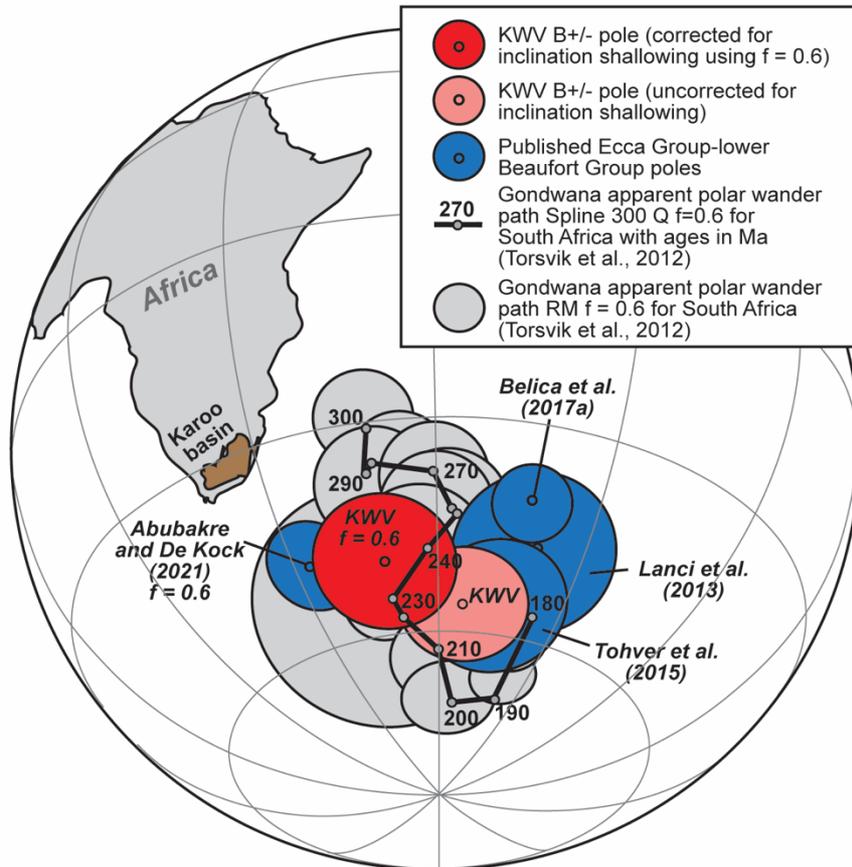
$$\tan(\text{observed inclination}) = f \tan(\text{corrected inclination})$$

523
 524
 525
 526 Such a correction yield B-, B+ and combined paleopoles that are comparable to the
 527 280 Ma to 210 Ma segment of the Gondwana reference apparent polar wander path
 528 (Figure 10). Our inclination corrected pole is also comparable to the Eccca Group pole
 529 presented for the KZF-01 core from the southwestern Karoo basin of Abubakre and
 530 De Kock (2021) as well as the lower Beaufort Group poles of Lanci et al. (2013) and
 531 Tohver et al. (2015) (Figure 10). However, our pole is distinct from the Eccca Group
 532 pole of Belica et al. (2017a), which plots further to the east from our pole as well as
 533 the 280-240 Ma segment of the Gondwana reference apparent polar wander path of
 534 Torsvik et al. (2012) (Figure 10).

535
 536 **Table 1: Summary of remanence components identified in KWV-01 and their**
 537 **corresponding paleopoles**

Com- ponent	n	In-situ remanence				Tilt-corrected remanence*				Paleopole		
		Decl. in °	Incl. in °	k	α_{95} in °	Decl. in °	Incl. in ° (f=0.6)	k	α_{95} in °	Plat in °N (f=0.6)	Plong in °E (f=0.6)	dp/dm in ° (f=0.6)
DOL	6	57.3	-68.8	7.26	24.04	58.6	-72.4	5.97	26.88	-44.6	342.2	34.5/40.7
A	311	350.2	-75.5	5.55	3.74	354.1	-76.2	5.26	3.86	-58.9	36.5	6.3/6.9
B-	65	319.3	-73.6	10.02	5.86	322.8	-75.2 (-81.0)	8.84	6.28	-51.9 (-45.4)	55.8 (43.7)	10.5/11.5 (11.7/12.1)
B+ (lines only)	46	144.6	62.2	6.05	9.34	145.7	63.5 (73.4)	6.49	8.96	-60.6 (-54.7)	82.7 (58.5)	11.2/14.2 (14.4/16.1)
B+ (lines and planes)	170	153.5	65.3	8.74	3.88	155.8	66.5 (75.4)	8.32	3.99	-65.3 (-56.1)	68.7 (48.4)	5.4/6.6 (6.7/7.3)
B+/- (lines only)	111	322.1	-69.1	7.70	5.18	324.4	-70.5 (-78.0)	7.50	5.25	-56.4 (-49.5)	66.0 (49.1)	7.9/9.1 (9.3/9.9)
B+/- (lines and planes)	235	150.6	67.7	8.89	3.27	153.3	69.0 (77.0)	8.31	3.39	-62.0 (-53.2)	64.3 (46.9)	4.9/5.8 (5.9/6.3)

538 *Structural correction applied to restore bedding as per electrical dip meter log.
 539 Abbreviations: n = number of specimens, Decl. = declination, Incl. = inclination, k = precision parameter, α_{95} =
 540 radius of the cone of 95% confidence around the mean, Plat = VGP latitude, Plong = VGP longitude, dp/dm = semi-
 541 minor/semi-major axes of oval of 95% confidence around the calculated pole, "f = 0.6" indicates inclination
 542 corrected for shallowing and corresponding paleopoles calculated from these corrected values.
 543



544

545 **Figure 10.** The Ecca Group-lower Beaufort Group B+/- pole from KVV-01
 546 (uncorrected and corrected for inclination shallowing) compared to published poles
 547 and the Gondwana apparent polar wander path of Torsvik et al. (2012) in South African
 548 coordinates.

549

550 **5.3. Magnetostratigraphy and end of the Kiaman Reverse Polarity**

551 **Superchron in KVV-01**

552

553 The B- and B+ remanence directions define stratigraphically bound groupings and can
 554 be used to construct a magnetostratigraphy for KVV-01 (Figure 11). The core is
 555 dominated by reverse polarity zones, with three normal polarity zones are recorded in
 556 KVV-01 and are labelled KVV2n, KVV3n and KVV4n. Normal polarity zone KVV4n
 557 is recorded in the uppermost specimens of the core and within the Koonop Formation.
 558 The remainder of the Koonop Formation, Waterford Formation and the uppermost Fort
 559 Brown Formation define polarity chron KVV3 between 325 m and 20 m depth. It is
 560 defined by a prominent normal polarity zone (137 m to 328 m) and reverse polarity
 561 zone (20 m to 137 m) with possible normal polarity subzones at 82 m depth, and

562 several possible reverse polarity subzones within KVV3n. The presence of a normal
563 magnetized dolerite sill within the Koonop Formation does not affect resolution of the
564 KVV3r polarity zone. The Fort Brown Formation host one magnetic polarity chron (i.e.,
565 KVV2). The KVV2r reverse polarity zone is well defined between 328 m and 446 m,
566 but normal polarity zone KVV2n is perhaps partially obscured by the effects of a
567 possible reversely magnetized thin dolerite sill (Figure 11). KVV2n does appear to
568 extend to 522 m. The remainder of the core is dominated by a very thick reverse
569 polarity zone (KVV1r), but four well defined normal polarity subzones are recorded in
570 the lower Fort Brown Formation (i.e., KVV1r.2n, KVV1r.3n, and KVV1r.4n) and the
571 middle of the Trumpeters Member (i.e., KVV1r.1n). Possible, but poorly defined
572 normal polarity subzones were identified within the Fort Brown Formation between
573 KVV1r.3n and KVV1r.4n at 600 m and within the Trumpeters Member between
574 KVV1r.1n and KVV1r.2n at 781.5 m. Additional poorly defined normal polarity
575 subzones occur in the Wonderfontein Member (at 981 m and 1141 m), the Pluto's Vale
576 Member (at 1784 m), and at the base of the Collingham Formation (at 2259 m). Normal
577 magnetized dolerite intrusions within the Wonderfontein Member, the Pluto's Vale
578 Member, the Collingham Formation, and the Whitehill Formation appears to have little
579 effect on the resolution of the KVVr1 reverse polarity zone, except for the possible
580 normal polarity subzone at 2259 m. It cannot be ruled out that this specimen's
581 magnetization is overprinted by the intrusion within the Whitehill Formation. It should
582 also be noted that sample coverage between 1923 m and 2200 m is poor. Although
583 this interval is largely represented by dolerite, the polarity interpretation for
584 sedimentary units in the interval is regarded as uncertain (Figure 11).

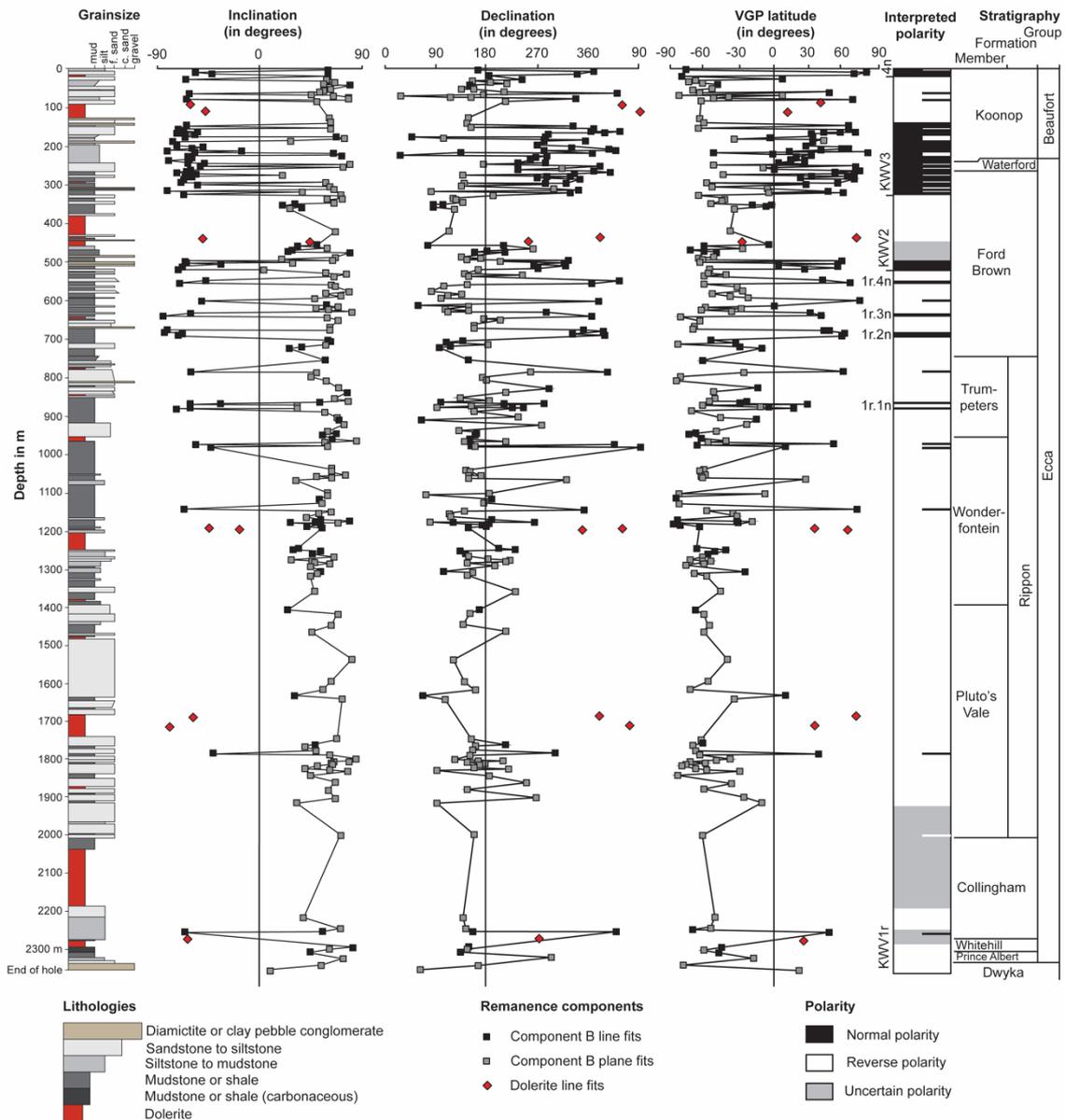
585

586 **5.4. Magnetostratigraphic correlation and geochronology of the Ecca Group** 587 **across the southern Karoo basin**

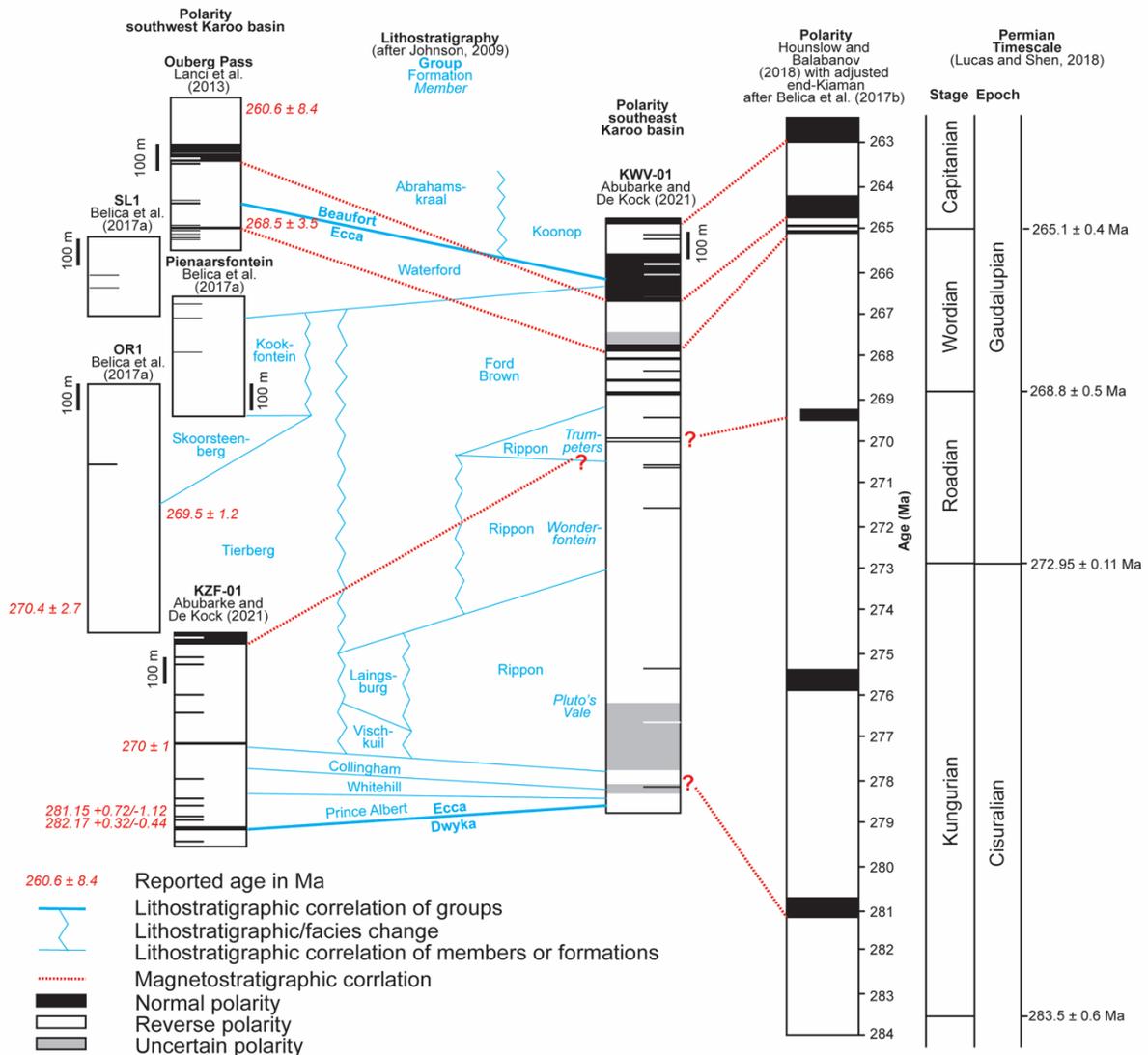
588

589 The magnetostratigraphy of KVV-01 can be correlated to that of southwestern Karoo
590 basin. In doing so, the KVV3n normal polarity zone in the upper Fort-Brown to lower
591 Koonop formations and the KVV4n normal polarity zone in the Koonop Formation can
592 be correlated to the N2 normal polarity zones identified by Lanci et al. (2013) within
593 the Abrahamskraal Formation at Ouberg Pass (Figure 12). The thin KVV2n normal
594 polarity zone in the middle of the Fort Brown Formation would then correlate to N3
595 identified by Lanci et al. (2013) in the Waterford Formation (Figure 12). Correlation of

596 normal polarity sub zones (i.e., KVV1r.4n, KVV1r, 3n, KVV1r.2n, and KVV1r.1n) are
 597 less obvious, but we suggest a correlation of the normal polarity zone within the
 598 Tierberg Formation (Abubakre and De Kock, 2021) at the top of borehole KZF-01 to
 599 KVV1r.1n (Figure 12). Two normal polarity zones at the top of the Collingham
 600 Formation and base of the Price Albert Formation in KZF-01 (Abubakre and De Kock,
 601 2021) are not clearly identified in KVV-01.



602
 603 **Figure 11.** Magnetic polarity interpretation for KVV-01 based on inclination,
 604 declination, and VGP latitude stratigraphic variation of the B- and B+ remanence
 605 directions (All indicated here as in-situ remanence directions). Magnetic components
 606 identified in dolerite intrusions are illustrated for comparison.



607

608 **Figure 12.** Magnetostratigraphic and lithostratigraphic correlation between the
 609 southwest and southeast Karoo basin and the Permian composite reference
 610 magnetostratigraphic record. See the text for details on the reported ages from the
 611 southwest Karoo basin.

612

613 The normal polarity zones at the top of KWV-01 are interpreted to mark the onset of
 614 the Illawara mixed polarity zone at the top of the Kiaman Normal Polarity Superchron.
 615 The end of the Kiaman Normal Polarity Superchron is placed at or near the base of
 616 the Capitanian, but the precise age was not well constrained (Hounslow and
 617 Balabanov, 2018; Menning et al., 2006; Steiner, 2006). Recently Belica et al. (2017b)
 618 constrained the end of the Kiaman Normal Polarity Superchron (i.e., the so-called
 619 Illawara reversal) to be between 265.05 ± 0.35 Ma (Ar-Ar on plagioclase) and a 263.51
 620 ± 0.05 Ma (U-Pb thermal ionization mass spectrometry or TIMS age by Metcalfe et

621 al., 2015). Lanci et al. (2013) reported U-Pb sensitive high-resolution ion microprobe
622 or SHRIMP zircon ages from tuffaceous units at Ouberg Pass within the Illawara
623 mixed polarity zone of between 264-268 Ma, while McKay et al. (2015) reported a U-
624 Pb SHRIMP zircon ages of 260.6 ± 8.9 Ma 264.7 ± 2.5 Ma from the Abrahamskraal
625 Formation. Lanci et al. (2013) deduced a ~ 269 Ma age for the end of the Kiaman
626 Normal Polarity Superchron. The relative imprecision of the SHRIMP ages, however,
627 does not exclude an interpretation that places the Illawara reversal at ~ 265 Ma. The
628 Tierberg, Kookfontein and much of the Waterford formations of the Eccca Group in the
629 southwestern part of the basin is characterized by dominantly reversed magnetic
630 polarity, with a normal polarity zone in the middle of the Tierberg Formation reported
631 from the top of KWZ-01 and several poorly constrained normal polarity subzones in
632 the Kookfontein, Skoorsteenberg and lower Tierberg formations (Abubakre and De
633 Kock, 2021; Belica et al., 2017a). Belica et al. (2017a) correlated poorly characterized
634 normal polarity subzones from OR1 and Pienaarsfontein to Kungurian normal polarity
635 zones of the Geomagnetic Polarity Time Scale. Such a correlation is, however, not
636 compatible with U-Pb SHRIMP zircon ages of 269.5 ± 1.2 Ma and 270.4 ± 2.7 Ma
637 reported by Belica et al. (2017a) and McKay et al. (2016) from the Tierberg Formation.
638 Here we suggest correlation of the normal polarity zone in the middle of the Tierberg
639 Formation of Abubakre and De Kock (2021) to the late Roadian normal polarity zone
640 in the reference composite of Hounslow and Balabanov (2018), which would account
641 for the reported U-Pb SHRIMP zircon ages. In KWV-01, this can possibly be correlated
642 to normal polarity subzone KWV1r.1n in the Trumpeters Member of the Rippon
643 Formation (Figure 12). The deposition of the Prince Albert, Whitehill and Collingham
644 formations are constrained between ~ 282 Ma to ~ 270 Ma by U-Pb SHRIMP zircon
645 ages from tuffaceous units in the southwestern part of the basin (Griffis et al., 2019;
646 Turner, 1999). Abubakre and De Kock (2021) thus correlated two short normal polarity
647 subzones in KZF-01 to Kungurian normal polarity zones in the reference composite of
648 Hounslow and Balabanov (2018). Normal polarity zones or subzones are not defined
649 in KWV-01, but a possible normal polarity subzone (represented only by one
650 specimen, and possibly affected by a nearby dolerite intrusion) in the Collingham
651 Formation could be Kungurian in age.

652

653 The top of KWV-01 in the Koonop Formation can be constrained to be near 263 Ma,
654 while the top of the Eccca Group in KWV-01 is well constrained as being deposited

655 during an early Capitanian normal polarity zone between 264 Ma and 265 Ma while
656 deposition of the Eccca Group likely commenced around ~281 Ma (Figure 12). As such
657 the KVV-01 span 18 million years of deposition of which the Eccca Group represents
658 ~16 million years. From the magnetostratigraphic correlation the lithostratigraphic
659 boundary between the Eccca Group and the Beaufort Group is shown to be time
660 transgressive. The boundary represents a change from marine to terrestrial facies and
661 has long been interpreted as diachronous and to young from the southwest to the
662 north based on sedimentology and biostratigraphy (Catuneanu et al., 2002; Rubidge,
663 2005; Rubidge et al., 1999). Here the position of the boundary is calibrated at near
664 265 Ma in the southwest at Ouberg Pass and closer to 264.5 Ma in the southeast in
665 KVV-01.

666

667 **6. Conclusions**

668 The Permian rocks of the southeast Karoo basin in KVV-01 revealed dual polarity
669 primary remanence components (components B- and B+) that pass a quality C
670 reversals test as well as baked contact tests, where the rocks are in contact with
671 younger dolerite intrusions of the KLIP. These intrusions have affected much KVV-01
672 magnetically, and likely caused widespread acquisition of a lower stability northerly up
673 remanence (component A), which is indistinguishable from magnetizations recorded
674 by dolerite specimens (i.e., component DOL). The combined B+/- paleomagnetic pole
675 compares favourably with previously published paleopoles from the Eccca Group and
676 lower Beaufort Group. Sedimentary rocks in KVV-01 recorded predominantly B+
677 magnetizations representing the Kiaman Reverse Polarity Superchron. B-
678 magnetizations in KVV-01 define three normal polarity zones in the lower Koonop,
679 Waterford, and Ford Brown formations representing the Illawara mixed polarity zone.
680 As such, the end of the Kiaman Reverse Polarity Superchron is recorded in KVV-01
681 and can be correlated to southwestern part of the basin, where it is recorded at Ouberg
682 Pass. Correlation to the Permian reference composite magnetostratigraphy suggest
683 an age of 264.5 Ma for the boundary between the Eccca and the Beaufort groups in
684 KVV-01, and ~265 Ma at Ouberg Pass, thus calibrating this diachronous boundary in
685 the southern Karoo basin for the first time.

686

687 **7. Acknowledgements**

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689 Innovation and National Research Foundation Centre of Excellence for Integrated
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694 magnetism laboratory at Caltech.

695

696 **8. Data Availability Statement**

697 The electrical dip meter log for KVV-01, details of core orientation, all demagnetization
698 data, least squares fits as well as rock magnetic data used to create this manuscript
699 are available at <http://dx.doi.org/10.17632/f7p753f2hp.1>. Least-squares analysis of
700 demagnetization data, orthogonal diagrams, equal area nets and J/Jo diagrams were
701 made using PaleoMag 3.2 (Jones, 2002), available at
702 <http://cires1.colorado.edu/people/jones.craig/PMag3.html>. Paleopoles were
703 visualized using GPlates 2.2 (Williams et al., 2012), available at
704 <https://www.gplates.org/>. IRM acquisition curves were unmixed using the MAX Unmix
705 software (Maxbauer et al., 2016) through the online application available at
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707

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