

1 **MJO Initiation Triggered by Amplification of**
2 **Upper-tropospheric Dry Mixed Rossby–gravity Waves**

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9 **Key Points:**

- 10 • Amplification of upper-level MRGs above the western Indian Ocean (WIO) can
11 lead to MJO initiation via low-level MRG-wave packet formation
- 12 • Upper-level MRGs propagating into the WIO are amplified by wave accumulation
13 through the mean Walker circulation and then dispersed downward
- 14 • Upper-level circumnavigating Kelvin waves can assist MRG-induced MJO initi-
15 ation by promoting MRG wave accumulation in the WIO

Abstract

A possibly important dynamical process for the Madden–Julian oscillation (MJO) convective initiation is proposed. An MJO event during the “CINDY2011” field campaign is triggered by eastward-moving lower-tropospheric mixed Rossby-gravity (MRG) wave packets, and its leading precursor is predominance of upper-tropospheric MRGs in the Indian Ocean (IO). Simple three-dimensional model experiments reveal that the upper-tropospheric MRGs in the IO are amplified particularly in the western IO (WIO) by their westward advection and wave accumulation due to the upper-level convergence in mean easterlies of the Walker circulation. The model also predicts downward dispersion of the amplified upper-tropospheric MRGs and resultant lower-tropospheric MRG wave packet formation. This MRG evolution consistently explains the MJO initiation process during CINDY2011, which is further verified by ray tracing for MRGs. Upper-tropospheric circumnavigating Kelvin waves assist the proposed mechanism by promoting MRG-wave accumulation (advection) in their westerly (easterly) phases via enhanced zonal convergence and weakened easterlies (enhanced easterlies).

Plain Language Summary

In the tropics, there exists a huge cluster of clouds and rainfall systems moving from the western Indian Ocean (WIO) to the western Pacific, called the Madden–Julian Oscillation (MJO). It is of great interest how and when MJO clouds are formed in the WIO because of their large impacts on global weather patterns, but we have not fully understood it yet. Using a simplified computer simulation and observational data, we find that the formation of MJO tall clouds can start from the “upper-sky” (~ 10 km altitude) short-period wind variations. During the no-cloud period of MJO, energy of upper-sky wind variations can be input above the central Indian Ocean. Then, it is carried to the WIO and amplified there by upper-sky easterly winds and their convergence associated with the seasonal atmospheric circulation. Because the amplified upper-sky wind energy above the WIO is further dispersed downward, near-surface wind variations become active, which triggers MJO clouds. This new mechanism is theoretically plausible but has been confirmed only in a limited case. It thus should be evaluated for more observations.

1 Introduction

The Madden–Julian oscillation (MJO) is the most prominent intraseasonal variability in the tropics (Madden & Julian, 1972), observed as an eastward-propagating large-scale organized convective system over the Indo-Pacific region. The whole picture of the MJO cannot be explained by classical equatorial wave theories (Matsuno, 1966; Takayabu, 1994; Wheeler & Kiladis, 1999). Also, the MJO has extensive impacts on global weather patterns (Zhang, 2013). It thus has been of great interest to scrutinize the MJO mechanics and to improve the MJO prediction capability over the past decades. In particular, revealing the physics underlying MJO initiation is one of challenging tasks, as inferred from the fact that many pathways to MJO onset in the Indian Ocean (IO) have been proposed (Jiang et al., 2020, and references therein).

MJO initiation processes can be largely divided into two stages; S_1) establishment of large-scale environments favorable for MJO initiation, and S_2) MJO convective outbreaks under the S_1 . The S_1 is often explained as “preconditioning” via MJO-scale anomalous horizontal moisture advection (e.g., Kiranmayi & Maloney, 2011; Zhao et al., 2013) and gradual shallow-to-deep pre-moistening called the “discharge-recharge mechanism” (e.g., Bladé & Hartmann, 1993; Benedict & Randall, 2007). Presumably, these processes commonly help MJO convective organization by promoting moisture accumulation.

It is non-trivial when MJO convection is triggered during the preconditioning, however. For instance, Xu and Rutledge (2016) showed that the transition into deep con-

vection is sometimes more rapid than the prediction from the discharge-recharge mechanism. To fill this deficiency in *thermodynamic*-driven processes, we should scrutinize *dynamic* variations as external forcing at the S_2 . Specifically, equatorially circumnavigating Kelvin waves (e.g., Seo & Kim, 2003; Powell & Houze, 2015; Chen & Zhang, 2019) and extratropical disturbances (e.g., Hsu et al., 1990; Ray & Zhang, 2010; Gahtan & Roundy, 2019) can trigger MJO convection by inducing upward motions directly, although it is still debated how robust and plausible the proposed processes are.

This paper focuses on the S_2 , particularly dynamic roles of mixed Rossby–gravity waves (MRGs) in triggering of MJO convection, which is motivated by several observational studies (Straub & Kiladis, 2003; Yasunaga et al., 2010; D. Yang & Ingersoll, 2011; Takasuka et al., 2019; Takasuka & Satoh, 2020). Field observations clearly detected MRGs enhanced in the mid-to-upper troposphere during the MJO-suppressed phase (Yasunaga et al., 2010; Takasuka et al., 2019), which supports a notion that those MRGs determine the timing of MJO convective outbreaks in the IO through rapid moistening and/or the development of low-level convergence. Moreover, some previous studies showed that eastward group velocity of MRGs assists the start of MJO propagation (D. Yang & Ingersoll, 2011; Takasuka et al., 2019; Takasuka & Satoh, 2020).

The aforementioned findings imply that mid-to-upper-tropospheric MRGs may be sometimes influential precursors of MJO initiation. A question here is how upper-tropospheric MRGs finally initiate MJO convection, which is more likely to be affected by lower-tropospheric moisture fields. Takasuka and Satoh (2020) statistically suggested that upper-tropospheric MRG energy input by more diabatic heating associated with MRG–convection coupling results in downward dispersion of MRGs and the formation of low-level MRG wave packets leading to MJO initiation. However, because upper-tropospheric diabatic heating is rooted in lower-tropospheric moisture/wind variations, diabatic processes may not be a primary trigger for low-level MRGs stemming from the upper troposphere. Hence, it is worth examining whether, as an intrinsic mechanism for MJO initiation in which amplification of upper-tropospheric MRGs is involved, there exists a process more in line with upper-tropospheric dynamics.

In this regard, we shed light on the dry interaction between upper-tropospheric MRGs and a wall-like sharp downward branch (SDB) of the Walker circulation (WC) above the western IO (WIO) (Kohyama et al., 2021). Because SDB climatologically forces upper-tropospheric zonal convergence in easterlies over the IO, MRGs approaching there may be amplified by wave accumulation (Hoskins & Yang, 2016) and then be dispersed downward and eastward. Motivated by this insight, we aim to verify the possibility that dry MRG dynamics can play an essential role in MJO initiation, based on simple model simulations and observational data analyses. A possible role of circumnavigating Kelvin waves in the MJO–MRG relationship is also discussed.

2 Data and Model Descriptions

2.1 Observational Data

To provide observational evidence for our hypothesis, we analyze “MJO2” event initiated in mid-November 2011 during a field campaign CINDY2011 (Yoneyama et al., 2013). We use 3-hourly radiosonde observations at Gan Island (0.7°S , 73.2°E), 6-hourly atmospheric fields from ERA-Interim (Dee et al., 2011) with 27 vertical layers spanning 1000–100 hPa, and 6-hourly rainfall data from the Global Satellite Mapping of Precipitation (GSMaP; Okamoto et al., 2005). A horizontal grid interval of the ERA-Interim (GSMaP) is 0.5° (0.1°). The ERA-Interim data and others covered the entire period of October/November and November 2011, respectively. Note that the boreal-winter (November to March) climatology used in section 4 is derived from the period of 1979–2012.

114 Anomalies are calculated by subtracting the mean during the data period. To cap-
 115 ture MRG variations, we filter 6-hourly anomalies for westward-propagating wavenum-
 116 bers and periods of 3.5–8 days (cf. section 3), using fast Fourier transforms in space and
 117 a 101-point Lanczos filter in time (Duchon, 1979).

118 2.2 Simple Dry Model

119 Based on Stechmann et al. (2008), a simple dry model with the barotropic and the
 120 first and second baroclinic modes for the vertical depth of $H = 16$ km is constructed
 121 on the equatorial β -plane. The model equations are

$$122 \quad \frac{\partial \zeta_0}{\partial t} + v_0 = -\nabla \times \mathcal{D}_{\mathbf{u}_0}(\mathbf{u}_0, \mathbf{u}_1, \mathbf{u}_2) - \frac{\zeta_0}{\tau_{\mathbf{u}}} + K_{\mathbf{u}} \nabla^4 \zeta_0 \quad (1)$$

$$123 \quad \frac{\partial \mathbf{u}_j}{\partial t} + y \mathbf{u}_j^\perp - \nabla \theta_j = -\mathcal{D}_{\mathbf{u}_j}(\mathbf{u}_0, \mathbf{u}_1, \mathbf{u}_2) - \frac{\mathbf{u}_j}{\tau_{\mathbf{u}}} + K_{\mathbf{u}} \nabla^4 \mathbf{u}_j \quad (\text{for } j = 1, 2) \quad (2)$$

$$124 \quad \frac{\partial \theta_1}{\partial t} - \nabla \cdot \mathbf{u}_1 = -\mathcal{D}_{\theta_1}(\mathbf{u}_0, \mathbf{u}_1, \mathbf{u}_2, \theta_1, \theta_2) - \frac{\theta_1}{\tau_{\theta}} + S_{\theta_1} + K_{\theta} \nabla^4 \theta_1 \quad (3)$$

$$125 \quad \frac{\partial \theta_2}{\partial t} - \frac{1}{4} \nabla \cdot \mathbf{u}_2 = -\mathcal{D}_{\theta_2}(\mathbf{u}_0, \mathbf{u}_1, \theta_1, \theta_2) - \frac{\theta_2}{\tau_{\theta}} + S_{\theta_2} + K_{\theta} \nabla^4 \theta_2 \quad (4)$$

126 where $\mathbf{u} = (u, v)^T$ is the horizontal wind vector; $\mathbf{u}^\perp = (-v, u)^T$; θ is potential tem-
 127 perature; ζ is relative vorticity; $\tau_{\mathbf{u}}$ (τ_{θ}) is the time scale of damping (cooling) for \mathbf{u} (θ);
 128 and S_{θ} is the heat source. Subscripts j for prognostic variables represent the barotropic
 129 ($j = 0$) and first and second baroclinic modes ($j = 1, 2$), and the nonlinear advection
 130 terms $\mathcal{D}_{\mathbf{u}_j, \theta_j}$ are given by

$$131 \quad \mathcal{D}_{\mathbf{u}_0} = \sum_{j=0}^2 \mathbf{u}_j \cdot \nabla \mathbf{u}_j + \sum_{j=1}^2 (\nabla \cdot \mathbf{u}_j) \mathbf{u}_j$$

$$132 \quad \mathcal{D}_{\mathbf{u}_1} = \mathbf{u}_0 \cdot \nabla \mathbf{u}_1 + \mathbf{u}_1 \cdot \nabla \mathbf{u}_0 + \frac{1}{\sqrt{2}} \left[\mathbf{u}_1 \cdot \nabla \mathbf{u}_2 + \mathbf{u}_2 \cdot \nabla \mathbf{u}_1 + 2(\nabla \cdot \mathbf{u}_1) \mathbf{u}_2 + \frac{1}{2}(\nabla \cdot \mathbf{u}_2) \mathbf{u}_1 \right]$$

$$133 \quad \mathcal{D}_{\mathbf{u}_2} = \mathbf{u}_0 \cdot \nabla \mathbf{u}_2 + \mathbf{u}_2 \cdot \nabla \mathbf{u}_0 + \frac{1}{\sqrt{2}} [\mathbf{u}_1 \cdot \nabla \mathbf{u}_1 - (\nabla \cdot \mathbf{u}_1) \mathbf{u}_1]$$

$$134 \quad \mathcal{D}_{\theta_1} = \mathbf{u}_0 \cdot \nabla \theta_1 + \frac{1}{\sqrt{2}} \left[2\mathbf{u}_1 \cdot \nabla \theta_2 - \mathbf{u}_2 \cdot \nabla \theta_1 + 4(\nabla \cdot \mathbf{u}_1) \theta_2 - \frac{1}{2}(\nabla \cdot \mathbf{u}_2) \theta_1 \right]$$

$$135 \quad \mathcal{D}_{\theta_2} = \mathbf{u}_0 \cdot \nabla \theta_2 + \frac{1}{2\sqrt{2}} [\mathbf{u}_1 \cdot \nabla \theta_1 - (\nabla \cdot \mathbf{u}_1) \theta_1]$$

136 Equations (1)–(4) start with the three-dimensional Boussinesq system (Majda, 2003),
 137 and they have been nondimensionalized by the scaling used in Stechmann et al. (2008).
 138 The derivation of (1)–(4) is provided in the supporting information (Text S1).

139 Solutions to the present model are numerically obtained for specific $S_{\theta_{1,2}}$ distribu-
 140 tions and initial conditions given to examine the interaction between WC and MRGs (see
 141 section 4.1 for details). We assume a zonally-periodic meridionally-bounded channel of
 142 which the zonal and meridional extent is 40,000 km (nearly the circumference along the
 143 equator) and 8,000 km, respectively. In all simulations, a grid spacing of 100 km on the
 144 Arakawa C-grid and a time step of 15 min for the third-order Runge-Kutta scheme are
 145 used. For the fourth-order horizontal diffusion (the damping/cooling) term in (1)–(4),
 146 we adopt $K_{\mathbf{u}} = K_{\theta} = 1.6 \times 10^{14} \text{ m}^4 \text{ s}^{-1}$ ($\tau_{\mathbf{u}} = \tau_{\theta} = 20$ days).

147 3 Observational Evidence of MRG Variations Leading to MJO initi- 148 ation

149 “MJO2” event during CINDY2011, initiated in the WIO around 17 November (see
 150 Figs. 1c–e for 10°N – 10°S rainfall variations in the time–longitude sections), stems from

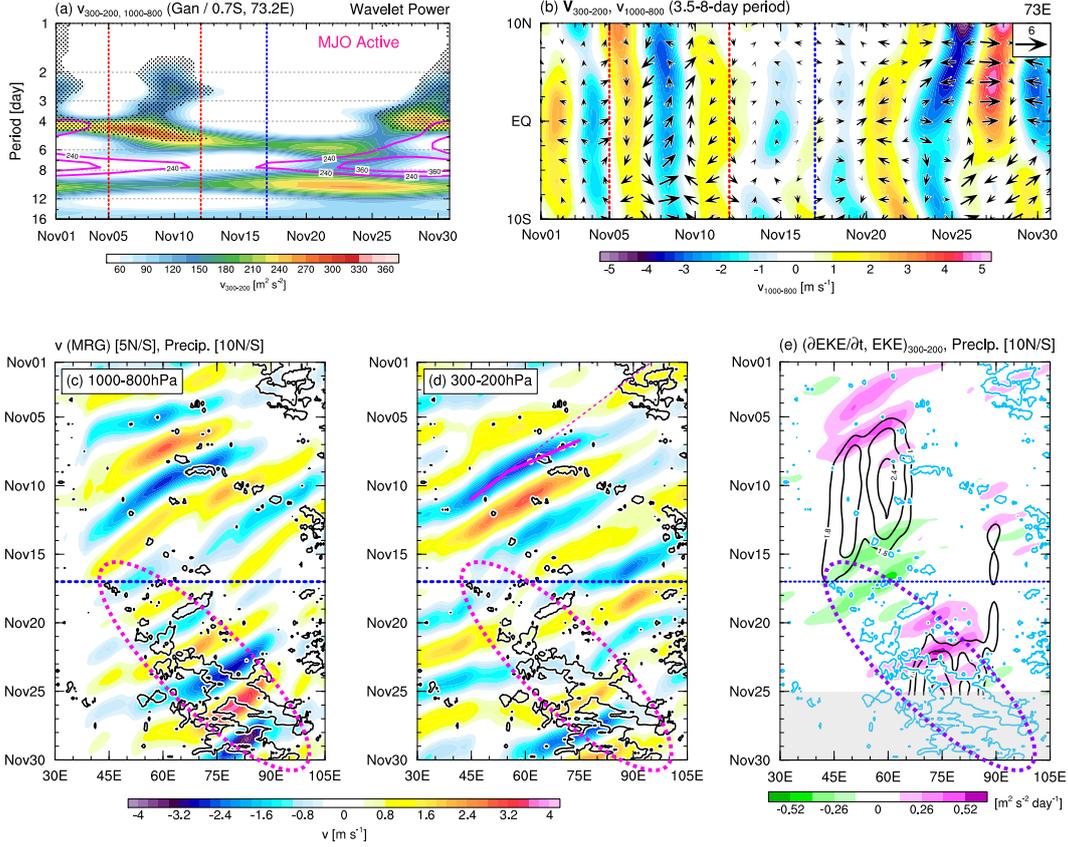


Figure 1. (a) Wavelet power of radiosonde-derived 300–200-hPa and 1000–800-hPa meridional winds (shading and contours) at Gan. Stippling for shading denotes statistical significance at the 95% level. (b) Time-latitude diagram at 73°E of 3.5–8-day bandpass-filtered horizontal winds at 300–200 hPa (vectors) and meridional winds at 1000–800 hPa (shading). (c,d) Time-longitude diagrams of 5°S–5°N averaged MRG-filtered meridional wind anomalies at (c) 1000–800 hPa and (d) 300–200 hPa (shading) and 10°S–10°N averaged precipitation with 0.8 mm/hr (contours). Ellipses indicate MJO2. (e) As in (c,d), but for 10°S–10°N averaged MRG-related EKE at 300–200 hPa (black contours) and its tendency (shading).

151 amplification of upper-tropospheric MRGs. In Fig. 1a, the wavelet analysis (Torrence
 152 & Compo, 1998) for radiosonde-derived meridional winds at Gan highlights significant
 153 4–5.5-day period variations at 300–200 hPa during 5–12 November (shading), after en-
 154 hanced lower-tropospheric variations in the 6–8-day cycle (contours). These wind vari-
 155 ations, detected from a 3.5–8-day-filtered data, are associated with cross-equatorial cir-
 156 culations with equatorially symmetric meridional wind signals (Fig. 1b), indicating the
 157 robust MRG structure. This amplification of upper-tropospheric MRGs is followed by
 158 re-intensification of lower-tropospheric MRGs in the end of November during the MJO-
 159 active phase (Figs. 1a,b).

160 The aforementioned fact is reinforced by the time–longitude diagrams of equatorial
 161 MRG-filtered meridional wind anomalies in the upper/lower troposphere and non-
 162 filtered precipitation field (Figs. 1c,d). The eastward propagation of MJO2 precipita-
 163 tion in the IO appears to collocate with the eastward formation of lower-tropospheric
 164 MRG wave packets beginning with northerlies in 45°–60°E (Fig. 1c). In fact, low-level
 165 MRG convergence successively triggers MJO convection from the WIO (Fig. S1), con-

166 sistent with the view that MRGs can actively contribute to MJO convective initiation
 167 (Takasuka et al., 2019; Takasuka & Satoh, 2020). Before this situation, around 10 Novem-
 168 ber, upper-tropospheric MRG variations begin to strengthen over the WIO in conjunc-
 169 tion with the slowdown of their westward propagation (Fig. 1d; magenta lines), which
 170 slightly precedes the development of the lower-tropospheric MRG wave packets in 45°–
 171 60°E. This evolution is also reconfirmed from the MRG-related eddy kinetic energy (EKE)
 172 field, defined by $K' = (u'^2 + v'^2)/2$ where primes denote MRG-filtered values; the pos-
 173 itive tendency and subsequent accumulation of upper-level EKE is evidently observed
 174 over the WIO before MJO2 initiation (Fig. 1e).

175 4 Mechanism

176 Based on the analyses in section 3, we raise two questions: Why are upper-tropospheric
 177 MRGs amplified in the WIO?; How are low-level MRG wave packets leading to MJO ini-
 178 tiation formed? As for the former question, the amplifying upper-tropospheric MRGs
 179 with the slowdown of their phase propagation are reminiscent of the interaction with zon-
 180 ally varying background flows (Hoskins & Yang, 2016). Inspired by this idea, we deduc-
 181 tively examine the above questions with simple dry model experiments by focusing on
 182 a role of WC, which has the wall-like SDB above the WIO (Kohyama et al., 2021). In
 183 parallel, we show that the presented mechanism is applicable to the MJO2 event.

184 4.1 Relationship Between the Walker Circulation and MRGs

185 First, WC in the model is obtained as the steady-state response to the time-invariant
 186 heat source S_{θ_1, θ_2} . Here, S_{θ_1, θ_2} are set so that $\hat{S}_{\theta_j} = \sqrt{2}S_{\theta_j} \sin(jz\pi/H)$ follows the struc-
 187 ture of boreal-winter mean apparent heating \bar{Q}_1 (Yanai et al., 1973) computed from the
 188 ERA-Interim; the formulation for S_{θ_1, θ_2} is provided in Text S2. For example, Fig. 2a com-
 189 pares the equatorial zonal variations of \hat{S}_{θ_1} at $z = 8$ km and \bar{Q}_1 at 400 hPa, where the
 190 first baroclinic components is dominant, subtracted from their zonal mean. \hat{S}_{θ_1} captures
 191 both amplitudes and zonal distributions of \bar{Q}_1 . Similarly, S_{θ_2} is given to match S_{θ_1} vari-
 192 ations except for its amplitudes with reference to Q. Yang et al. (2019).

193 These S_{θ_1, θ_2} produce the realistic WC after the 200 day from the state of rest, as
 194 recognized by a comparison of WC for the model and observed boreal-winter mean (Figs.
 195 2b,c); the wall-like SDB and associated upper-tropospheric zonal convergence over the
 196 WIO are reproduced. As expected, the same features as climatology are also realized in
 197 the 11-day running mean zonal-vertical circulations before MJO2 initiation (during 5–
 198 15 November; Fig. 2d), except for stronger zonal convergence than for the climatology
 199 (or the model), which will be discussed later.

200 Under the simulated WC, we examine how upper-tropospheric MRGs as observed
 201 before MJO initiation evolve. Referring to observations (Figs. 1d and S4), we set the ini-
 202 tial MRG structure as the zonal wavenumber-8 mode confined in $7500 \leq x \leq 9000$
 203 km (i.e., the eastern side of SDB) with maximum amplitudes at the model top. The hor-
 204 izontal structure of MRGs is derived following Aiyyer and Molinari (2003), and its de-
 205 tails are provided in Text S2 and Fig. S2. From the initial condition prepared by super-
 206 imposing the derived MRG field onto the steady state obtained from a 200-day spin-up
 207 integration, we run the model for 30 days.

208 Figure 3a shows the time-longitude diagram of equatorial upper-/lower-tropospheric
 209 meridional wind anomalies for the model. The initial MRG given in a limited area im-
 210 mediately excites westward-propagating MRGs in the upper troposphere. These upper-
 211 tropospheric MRGs experience wave contraction and deceleration of phase propagation,
 212 which occurs in the upper-level background zonal convergence area with easterlies (Fig.
 213 3c). Along with this contraction, upper-tropospheric MRGs are gradually amplified in

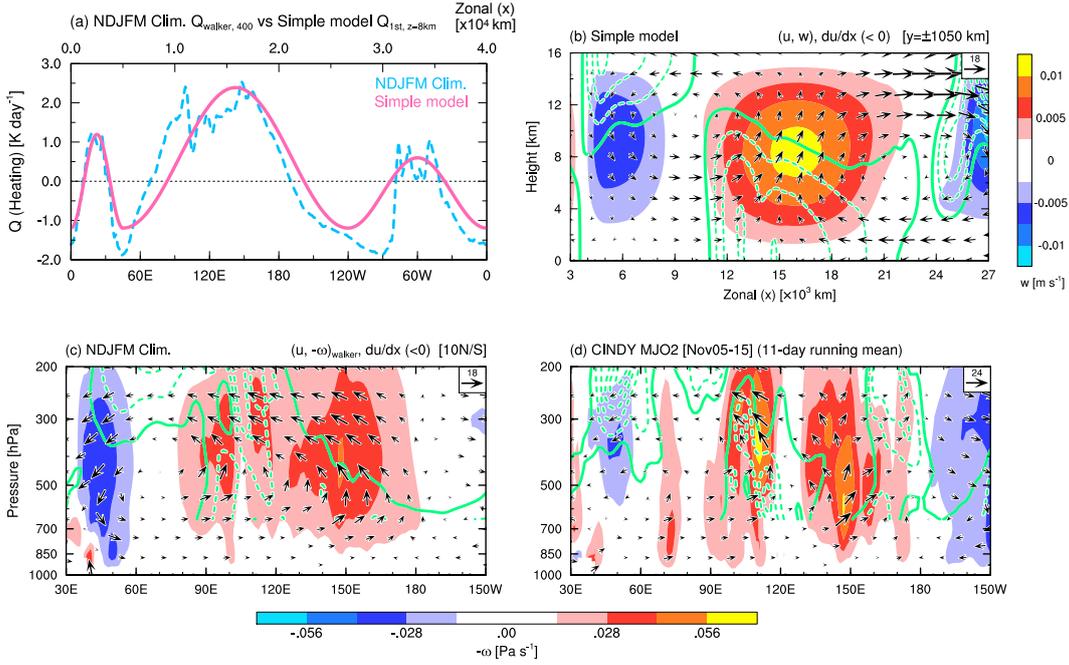


Figure 2. (a) Zonal distributions of \hat{S}_{θ_1} at $z = 8 \text{ km}$ (pink) and $\overline{Q_1}$ at 400 hPa (blue). (b–d) Zonal-height sections of vertical velocity (shading), zonal-vertical winds (vectors), and zonal convergence (contours) for (b) the spin-up model simulation at day 200, (c) boreal-winter mean, and (d) 11-day running mean fields during 5–15 November. All fields are subtracted from their zonal mean, and averaged over $y = \pm 1050 \text{ km}$ (10°S – 10°N) range for the model (observation). Contour interval in (b,d) [(c)] is 1.0 [0.5] $\times 10^{-6} / \text{s}$, with zero contours bolded. Contours below 750 hPa in (c,d) are masked for visibility, and vertical velocity for vectors in (b) and (c,d) is multiplied by 1000 and 400, respectively.

214 SDB ($4500 \leq x \leq 5500 \text{ km}$) until around day 15 when they begin to exhibit eastward
 215 group velocity, and then lower-tropospheric MRG wave packets are radiated eastward.

216 In Figs. 3b,d, which are the same as Figs. 3a,c but for the observed MJO2, the pro-
 217 cesses predicted by the model are similarly detected. After 5 November, upper-tropospheric
 218 MRGs propagating westward with small positive group velocity are decelerated and am-
 219 plified in 45° – 60°E , where the zonal convergence associated with SDB is realized. Then,
 220 lower-tropospheric MRG wave packets moving eastward are established, which charac-
 221 terizes MJO2 initiation.

222 Despite much consistency between the model and MJO2, there are some notewor-
 223 thy differences. One is faster group velocity of the lower-tropospheric MRGs in the model
 224 (Figs. 3a,b). This is attributed to the doppler shift by stronger background low-level west-
 225 erlies (Figs. 3c,d) and deeper equivalent depth in the dry model. The latter reflects the
 226 limitation that dry dynamics cannot represent wave–convection coupling effects that are
 227 important after MJO initiation.

228 Another difference is the stronger upper-tropospheric background zonal convergence
 229 around SDB before MJO2 initiation (Figs. 3c,d). This is because the observed background
 230 WC for MRGs are contributed by not only the climatology but also large-scale circum-
 231 navigating Kelvin waves with their evolution slower than MRGs. In fact, upper-tropospheric
 232 westerlies associated with circumnavigating Kelvin waves intrude into the WIO (Fig. 3e),

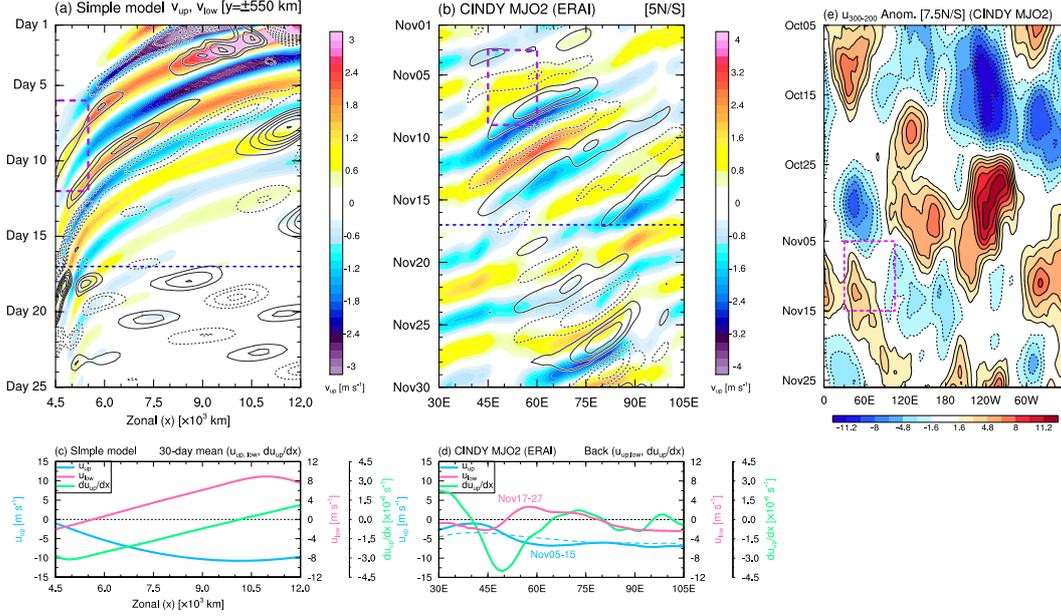


Figure 3. (a,b) Time-longitude diagrams of 24-hr running mean (MRG-filtered) meridional wind anomalies in the upper/lower troposphere (shading/contours) for the model (MJO2). Contours denote $\pm 0.2, \pm 0.3, \dots$ ($\pm 1.0, \pm 2.0, \dots$) m/s with negative values dashed. (c,d) Zonal distributions of upper-/lower-tropospheric zonal winds (blue/pink) and upper-tropospheric zonal convergence (green) for the simulation-period mean in the model (11-day running mean for MJO2). Upper-tropospheric (Lower-tropospheric) fields for MJO2 are computed during 5–15 (17–27) November. Broken lines in (d) denote the boreal-winter mean. (e) Time-longitude diagram of 5-day running mean upper-tropospheric zonal wind anomalies for MJO2. All fields in (a,b) and (c–e) are averaged over $y = \pm 550$ km (5° S– 5° N) and $y = \pm 850$ km (7.5° S– 7.5° N) range for the model (MJO2), respectively. The upper and lower troposphere for the model (MJO2) are defined as the 13.2–16 km and 0–2.8 km (300–200 hPa and 1000–800 hPa) layer, respectively.

233 which is implied by the stronger background westerlies to the west of 50° E than the boreal-
 234 winter mean (Fig. 3d). This process, which is not incorporated in the model, promotes
 235 convergence with climatological upper-level easterlies. Considering that zonal conver-
 236 gence can amplify MRGs (see section 4.2), upper-tropospheric circumnavigating Kelvin
 237 waves could serve as a catalyst of MRG-induced MJO initiation.

238 4.2 Amplification of Upper-tropospheric MRGs and Its Impacts on the 239 Lower Troposphere

240 To reveal why upper-tropospheric MRGs are amplified around SDB and then lower-
 241 tropospheric MRG wave packets are formed there, we conduct the EKE budget analy-
 242 sis. The budget equation for the model is

$$243 \frac{\partial \overline{K'}}{\partial t} = - \underbrace{\overline{\mathbf{v}' \cdot \nabla \mathbf{u}}}_{K_m K_e} - \underbrace{\overline{\mathbf{v} \cdot \nabla K'}}_{A_m K_e} - \underbrace{\overline{\mathbf{v}' \cdot \nabla K'}}_{A_e K_e} + \underbrace{\overline{w' \theta'}}_{P_e K_e} + \underbrace{\overline{\nabla \cdot (\mathbf{v}' \theta')}}_{G K_e} + (Res.) \quad (5)$$

244 where \mathbf{v} is the three-dimensional wind vector; w is vertical velocity; and overbars (primes)
 245 denote 11-day running mean (deviations from the mean of the 30-day simulation). For
 246 the ERA-Interim, primes denote MRG-filtered values, and $P_e K_e$ and $G K_e$ terms are re-
 247 placed with $-(R/p)\overline{\omega' T'}$ and $-\overline{\nabla \cdot (\mathbf{v}' \Phi')}$, respectively, where ω is vertical p -velocity; T

248 is temperature; Φ is geopotential; and R is the gas constant. Note that real sources/sinks
 249 of EKE are $K_m K_e$ and $P_e K_e$.

250 Figures 4a and 4d compare upper-tropospheric EKE budget terms averaged in the
 251 time-area domain where MRG amplification occurs for the model and MJO2 (see broken-
 252 line squares in Figs. 3a,b). This comparison shows physical consistency with each other;
 253 upper-tropospheric MRGs are amplified by EKE advection by background flows ($A_m K_e$)
 254 and the barotropic conversion from the background ($K_m K_e$). In Figs. 4b,e, the decom-
 255 position of these terms,

$$256 \quad (A_m K_e) = -\overline{u \frac{\partial K'}{\partial x}} - \overline{v \frac{\partial K'}{\partial y}} - \overline{w \frac{\partial K'}{\partial z}} \quad (6)$$

$$257 \quad (K_m K_e) = -\overline{u'^2 \frac{\partial \bar{u}}{\partial x}} - \overline{u'v' \frac{\partial \bar{u}}{\partial y}} - \overline{u'w' \frac{\partial \bar{u}}{\partial z}} - \overline{v'u' \frac{\partial \bar{v}}{\partial x}} - \overline{v'^2 \frac{\partial \bar{v}}{\partial y}} - \overline{v'w' \frac{\partial \bar{v}}{\partial z}} \quad (7)$$

258 reveals that $A_m K_e$ and $K_m K_e$ processes are dominantly contributed by $-\overline{u(\partial K'/\partial x)}$ and
 259 $-\overline{u'^2(\partial \bar{u}/\partial x)}$, respectively. This result ensures the following interpretation for upper-tropospheric
 260 MRG amplification: upper-level easterlies of WC into SDB efficiently advects MRG en-
 261 ergy from the east of SDB, and advected energy is further amplified by wave accumu-
 262 lation due to zonal convergence arising from SDB. Because westward-propagating MRGs
 263 (with typical group velocity ~ 5 m/s) are accumulated for their positive ground group
 264 velocity in zonal convergence (Hoskins & Yang, 2016), the region near SDB with $\bar{u} >$
 265 -5 m/s is indeed appropriate for MRG accumulation (cf. Fig. 3).

266 Also in the lower-troposphere, the EKE tendency is positive for both the model and
 267 MJO2 (Figs. 4c,f), corresponding to the formation of low-level MRG wave packets. For
 268 the model (Fig. 4c), this positive tendency almost originates from the EKE redistribu-
 269 tion via potential eddy flux convergence (GK_e). Because EKE source here is only upper-
 270 tropospheric $K_m K_e$ (Figs. 4a,c), the lower-tropospheric EKE is brought by the energy
 271 dispersion from the upper troposphere. Positive lower-tropospheric GK_e with positive
 272 upper-tropospheric $K_m K_e$ is also observed for MJO2 (Figs. 4d,f), supporting a notion
 273 that process found in the model operates before MJO2 initiation, despite the difference
 274 in lower-tropospheric $K_m K_e$ contributions.

275 The downward impacts of amplification of upper-tropospheric MRGs are qualita-
 276 tively inferred from the vertically eastward-tilted MRG structure (Fig. S3) and equa-
 277 torial zonal-height sections of $-\overline{u'^2(\partial \bar{u}/\partial x)}$ and GK_e (Figs. 4g,h). For the model (Fig.
 278 4g), MRG-related EKE is accumulated especially in the inner SDB ($4500 \leq x \leq 5500$
 279 km) in the upper troposphere, and as indicated by positive GK_e below it, the accumu-
 280 lated EKE is redistributed to the mid-to-lower troposphere. This situation reasonably
 281 holds true for MJO2 (Fig. 4h), although more EKE redistribution by GK_e is realized
 282 to the west of 45°E and around 60°E .

283 To make the above view more compelling for observation, we conduct 10-day ray
 284 tracing of MRGs from around 49°E , 300 hPa ($z \sim 9680$ m), and 10 November, where
 285 and when MRG amplification is clearly observed (Figs. S3b and S4). The initial zonal
 286 and vertical wavelength (λ_x and λ_z) for ray tracing is roughly estimated as $\lambda_x \sim 47^\circ$
 287 and $\lambda_z \sim 20$ km from the vertical structure (Fig. S3b). For those parameters and $\bar{u} =$
 288 -5.5 m/s, the MRG dispersion relation predicts ground zonal phase speed $c_{px} \sim -17$
 289 m/s, consistent with MRGs propagating into the WIO (Fig. S4). Practically, λ_z is dif-
 290 ficult to be identified from the vertically-coarse data, so initial λ_z is determined by the
 291 MRG dispersion relation with initial λ_x and $c_{px} = -17$ m/s.

292 This ray tracing reconfirms the downward-eastward energy dispersion of amplified
 293 upper-tropospheric MRGs. In Fig. 4h, the rays for 45 initial conditions considering their
 294 estimation uncertainties (see Text S3 for method details) indicate that a fraction of rays
 295 reach the mid-to-lower troposphere in $50^\circ\text{--}70^\circ\text{E}$ after “reflection” in SDB, although oth-
 296 ers go through SDB westward (as indicated by GK_e distributions).

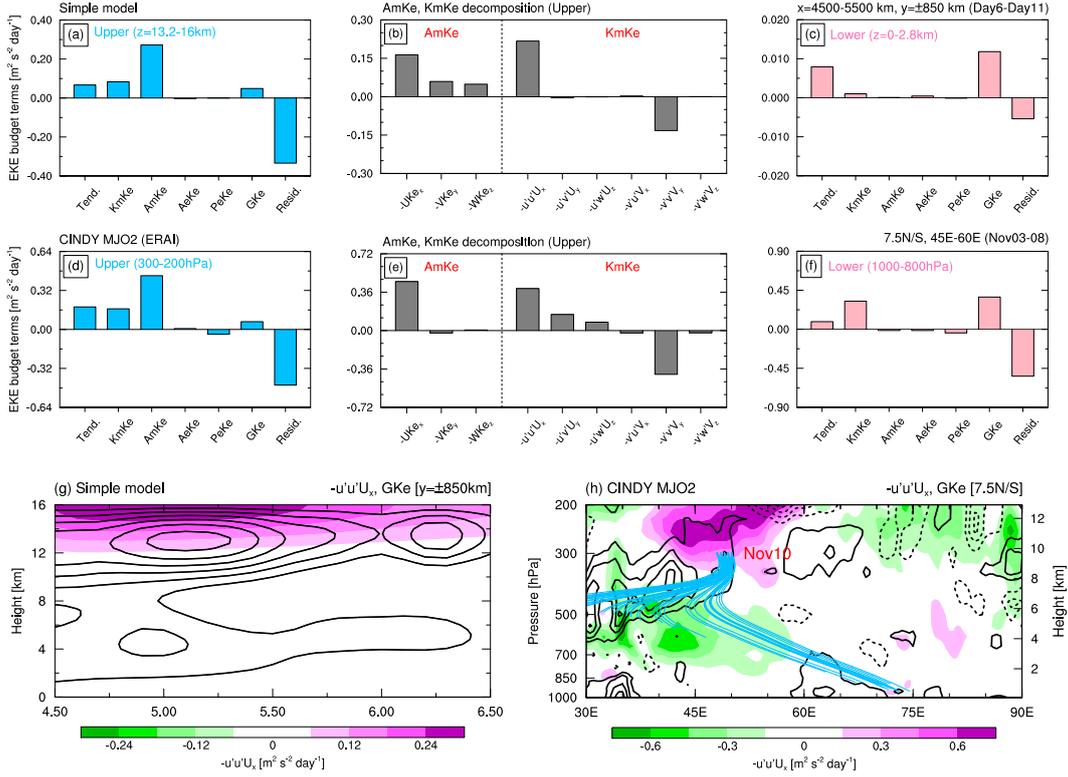


Figure 4. (a,d) All upper-tropospheric EKE budget terms, (b,e) decomposition of upper-tropospheric $A_m K_e$ and $K_m K_e$, and (c,f) all lower-tropospheric EKE budget terms for the model (top) and MJO2 (bottom). All values are averaged over the time-longitude domain indicated by broken-line squares in Fig. 3a (Fig. 3b) within $y = \pm 850$ km (7.5°S – 7.5°N) meridional bands for the model (MJO2). (g,h) Longitude-height sections of $-u'^2(\partial\bar{u}/\partial x)$ (shading) and GK_e (contours) for the model (MJO2). Contour interval is 0.015 (0.7) $\text{m}^2 \text{s}^{-2} \text{day}^{-1}$ for the model (MJO2), with negative (zero) values dashed (omitted). Blue lines in (h) denote MRG rays calculated for 45 different initial conditions.

5 Summary and Discussion

In this study, we have presented a new pathway to MJO initiation that stems from dry upper-tropospheric westward-propagating MRGs above the IO. This is inspired by initiation processes of the “MJO2” event during CINDY2011, in which upper-tropospheric MRG amplification in the WIO is followed by MJO2 initiation (Fig. 1). Here we hypothesize that the interaction between MRGs and the Walker circulation (WC) is the key.

To verify our hypothesis, we perform numerical simulations using a simple dry model with three vertical modes, comparing the model output with observations for MJO2. The model captures the essence of the boreal-winter mean WC above the IO: upper-level zonal convergence in mean easterlies blowing into the WIO, where the sharp downward branch (SDB) of WC exists (Figs. 2b,c). In the model with this idealized WC, upper-tropospheric MRGs propagating into SDB are amplified in the inner region of SDB. Then, lower-tropospheric MRG wave packets start to propagate eastward (Figs. 3a,c), resembling the processes of MJO2 initiation triggered by low-level MRG wave packets with eastward group velocity (Figs. 3b,d).

312 The energetics for this MRG evolution is discussed by both the model experiment
 313 and observations (Fig. 4). The initial amplification of upper-tropospheric MRGs in SDB
 314 results from MRG energy advection to SDB and wave accumulation due to upper-level
 315 easterlies of WC and their zonal convergence arising from SDB. Subsequently, the eastward-
 316 downward dispersion of the amplified upper-level MRG energy is activated, which forms
 317 lower-tropospheric MRG wave packets leading to MJO initiation.

318 A difference of WC between the model and MJO2 (Figs. 2b–d) has implication that
 319 upper-tropospheric circumnavigating Kelvin waves make the presented mechanism more
 320 efficient by modulating background WC additionally. For MJO2, upper-level zonal con-
 321 vergence in SDB are enhanced by cooperation between the westerly phase of Kelvin waves
 322 propagating into the WIO and climatological easterlies of WC above the IO (Fig. 3e),
 323 which promotes MRG-wave accumulation. In addition, upper-tropospheric Kelvin-wave
 324 westerly anomalies help the realization of positive ground group velocity of MRGs by
 325 weakening upper-tropospheric mean easterlies, which is advantageous to triggering the
 326 wave accumulation (Hoskins & Yang, 2016). Furthermore, the easterly phase of Kelvin
 327 waves before the westerly phase can enhance westward advection of MRG energy into
 328 the WIO. For these reasons, equatorial circumnavigation of Kelvin waves assists MRG-
 329 induced MJO initiation cooperatively with the climatological WC.

330 The idea proposed in this study for MJO initiation does not require moist processes
 331 at all, which provides several debatable topics. First, we may reconsider roles of diabatic
 332 processes in the similar MRG-related mechanism suggested by Takasuka et al. (2019) and
 333 Takasuka and Satoh (2020). A possible interpretation for this is that dry dynamics are
 334 sufficient for an initial trigger of amplification of upper-tropospheric MRGs, although
 335 diabatic heating can accelerate and/or maintain MRG amplification in a later stage when
 336 MRG–convection coupling becomes evident. Secondly, our idea does not necessarily con-
 337 tradict with the preexisting hypotheses that put emphasis on moisture variations (e.g.,
 338 Benedict & Randall, 2007; Zhao et al., 2013), because we have tackled MJO initiation
 339 in terms of convective triggering by gravity wave dynamics (e.g. Tulich & Mapes, 2008),
 340 assuming a favorable environment for organized convection regulated by moisture fields.
 341 Nevertheless, if dry MRG dynamics by itself can determine the timing of MJO initia-
 342 tion, it would be misleading to emphasize only the moisture variations for understand-
 343 ing MJO initiation. Because a simple dynamical model theoretically predicts the dry in-
 344 teraction between upper-tropospheric MRGs and WC as observed for a single MJO event,
 345 the next step is to examine its robustness and relationship with moist processes statis-
 346 tically for multiple cases.

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 354 GSMaP product is distributed online ([https://sharaku.eorc.jaxa.jp/GSMaP/index
 355 .htm](https://sharaku.eorc.jaxa.jp/GSMaP/index.htm)); and radiosonde-derived data during CINDY2011 are archived at [https://data
 356 .eol.ucar.edu/dataset/347.008](https://data.eol.ucar.edu/dataset/347.008).

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