

<sup>1</sup> **A Wall-like Sharp Downward Branch of the Walker**  
<sup>2</sup> **Circulation above the Western Indian Ocean**

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3 Key points:

4 • Climatology and interannual variability of the zonally-thin downward branch above the  
5 Indian Ocean are discussed.

6 • Model experiments confirm that the sharp downward branch is sustained by East African  
7 topography, in addition to radiative cooling.

8 • Without mountains in East Africa, the eastern Horn of Africa would exhibit wetter and  
9 more convective climatology.

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10 **Abstract.** In the zonal direction, the downward branch of the Walker cir-  
11 culation above the Indian Ocean is only 20 degrees thick, whereas the Pa-  
12 cific counterpart is 90 degrees thick. This zonal sharpness is remarkable be-  
13 cause atmospheric disturbances smaller than the planetary scale, such as the  
14 Asian Summer Monsoon, can interact with the planetary-scale Walker cir-  
15 culation through this branch. As a moist circulation, this zonal sharpness  
16 is imprinted on a unique zonal discontinuity of the Intertropical Convergence  
17 Zone, which has direct implications for the dry climate in the Northeast Africa.  
18 Therefore, in this study, we refer to this zonally-thin downward branch as  
19 the “Wall”, investigate its climatology and interannual variability, and aim  
20 at determining its reason for existence.

21 The strongest season of the lower tropospheric Wall in boreal summer is  
22 sustained by horizontal cold advection associated with the Asian Summer  
23 Monsoon. Two weak phases of the Wall correspond to two rainy seasons at  
24 the Eastern Horn of Africa, which are not reproduced well by state-of-the-  
25 art global climate models. As to interannual variability, a mass-weighted ver-  
26 tical mean of vertical motion at the Wall exhibits a tight linkage to the trop-  
27 ical Pacific, though total variance is explained more by local sea surface tem-  
28 perature.

29 Model experiments using a convection-permitting atmospheric general cir-  
30 culation model show that mountains in East Africa are necessary for the ex-  
31 istence of the Wall. Vertical mixings forced by mountain waves play a fun-  
32 damental role in strengthening the Wall. After flattening the East African

<sup>33</sup> topography, zonal discontinuity of the Intertropical Convergence Zone dis-  
<sup>34</sup> appears.

## 1. Introduction

35 The Walker circulation is the most prominent planetary-scale tropical atmospheric cir-  
36 culation in the zonal direction [e.g., *Walker*, 1923, 1924; *Bjerknes*, 1969]. It has been  
37 understood that, to first order, the vertical motion associated with the Walker circula-  
38 tion consists of upward branches over relatively warm surface (e.g., the warm pool in the  
39 western Pacific) and downward branches over relatively cool surface (e.g., the cold tongue  
40 in the eastern Pacific) [e.g., *Lau and Yang*, 2003]. In the context of climate variability,  
41 the Pacific branches of the Walker circulation have received particular attention, because  
42 its interannual fluctuation serves as the atmospheric component of the El Niño Southern  
43 Oscillation, the most dominant interannual climate mode on Earth [e.g., *Bjerknes*, 1969].

44 As a mean state, however, a downward branch of the Walker circulation above the  
45 western Indian Ocean also exhibits a strong subsidence, which is at least comparable to  
46 the Pacific downward branch. Figure 1a shows the annual-mean equatorial vertical motion  
47 calculated by taking the meridional mean over the equatorial region ( $10^{\circ}\text{S}$ - $10^{\circ}\text{N}$ ). The  
48 strong and sharp downward branch stands at the western edge of the Indian Ocean ( $40^{\circ}\text{E}$ -  
49  $60^{\circ}\text{E}$ ), whereas the weak and wide downward branch lies over the eastern Pacific ( $90^{\circ}\text{W}$ -  
50  $150^{\circ}\text{W}$ ). Considering the size of the two oceanic basins, one might find this interbasin  
51 contrast counterintuitive.

52 In fact, the interbasin difference in apparent strength of the downward motion emanates  
53 from a minor cause. An important caveat of this meridional-mean view shown in Fig. 1a  
54 is that the strength of the downward motion depends upon latitudinal choices of the  
55 equatorial belt. Figure 2a shows the annual-mean meridional-mean equatorial vertical

56 motion over the equatorial belt of  $10^{\circ}\text{S}$ - $10^{\circ}\text{N}$ ,  $5^{\circ}\text{S}$ - $5^{\circ}\text{N}$ , and  $2^{\circ}\text{S}$ - $2^{\circ}\text{N}$ . While the downward  
57 branch above the Indian Ocean relatively keeps its strength notwithstanding the latitudi-  
58 nal choices, that of the eastern Pacific becomes stronger when narrower equatorial belts  
59 are chosen. A key to understand this difference is the zonal-mean crosssection (Fig. 2b).  
60 The outstanding difference between the Indian and the Pacific downward branches is the  
61 degree of meridional asymmetry. The apparent weak downward motion above the Pacific,  
62 shown in Fig. 1a, originates from an offset of a strong upward branch over the northern  
63 off-equatorial region ( $4^{\circ}\text{N}$ - $10^{\circ}\text{N}$ ) against a strong equatorial downward motion over  $10^{\circ}\text{S}$ -  
64  $2^{\circ}\text{N}$ . On the other hand, the downward branch above the Indian Ocean shown in Fig. 1a  
65 consists of a strong downward branch over  $4^{\circ}\text{S}$ - $10^{\circ}\text{N}$  with a hint of weak upward branch  
66 located to the south. Though it could be misleading in some context, we will carefully  
67 keep using the equatorial belt of  $10^{\circ}\text{S}$ - $10^{\circ}\text{N}$  in this study, because we are interested in the  
68 meridionally-broad equatorial downward branch above the Indian Ocean.

69 A more fundamental difference between the two downward branches lies in the zonal  
70 thickness. That is, the downward branch above the Indian Ocean is only 20 degrees thick  
71 in the zonal direction, whereas the Pacific counterpart is about 90 degrees thick. As we  
72 shall see in later sections, this zonal sharpness of the Indian Ocean branch is remarkable  
73 in the sense that phenomena smaller than the planetary scale, such as the Asian Sum-  
74 mer Monsoon and possibly mountain waves, can easily interact with the planetary-scale  
75 Walker circulation via this branch. This zonal sharpness is particularly notable in the  
76 lower tropospheric layer, where the downward branch subsides only onto the coastal west-  
77 ern Indian Ocean and not onto the African continent (Fig. 2c), which is consistent with  
78 *Yang et al.* [2015]. In addition, from the viewpoint of moist circulations, the zonally-thin

79 downward branch is imprinted on a unique zonal discontinuity of the Intertropical Conver-  
80 gence Zone (ITCZ). Figures 1b and 1c show the annual mean outgoing longwave radiation  
81 (OLR) and precipitation, respectively, over the tropics. ITCZ is typically characterized  
82 by the narrow convective band that circles the Earth along the equatorial region, but if  
83 we carefully look at ITCZ, a discontinuity of ITCZ is found at the western edge of the  
84 Indian Ocean.

85 One of major implications of this ITCZ discontinuity, and thereby, of the zonally-thin  
86 downward branch, is the relatively dry climate at the so-called “Eastern Horn of Africa”,  
87 whose mean state, annual cycle, variability, and change have long been investigated in  
88 many previous studies [e.g., *Camberlin, 1995; Schreck III and Semazzi, 2004; Liebmann*  
89 *et al., 2014; Lyon, 2014; Tierney et al., 2015; Liebmann et al., 2017*]. By investigating  
90 the zonally-thin downward branch, we expect a better understanding of the climatology  
91 at the Eastern Horn of Africa, whose annual cycle of precipitation is poorly reproduced  
92 by state-of-the-art global climate models [*Tierney et al., 2015*].

93 Therefore, in this study, we refer to this meridionally-broad, zonally-thin sharp down-  
94 ward branch above the Indian Ocean as the “Wall” of the Walker Circulation, and will  
95 investigate its climatology and interannual variability in the hope that its reason for exis-  
96 tence will be determined. Data and methods are described in the next section. In section  
97 3, we describe the seasonality of the Wall, and highlight a role of horizontal cold advection  
98 to support its strongest phase. Next, in section 4, we define the Wall index to describe  
99 the interannual variability of the Wall and point out that both remote and local SST  
100 explain the interannual variance. In section 5, we then perform model experiments to  
101 identify the East African topography as a necessary condition for the existence of the

102 Wall, and discuss implications for the climate at the Eastern Horn of Africa. Conclusions  
103 are presented in section 6.

## 2. Data and Model

### 2.1. Data

104 Observed vertical motion, wind, and temperature data are from the European Center for  
105 Medium range Weather Forecasting (ECMWF) ERA-Interim reanalysis data [Dee *et al.*,  
106 2011], and the time span used in this study is from 1979 through 2017. Observed OLR  
107 data is from the National Oceanic and Atmospheric Administration (NOAA) interpolated  
108 OLR [Liebmann and Smith, 1996], whose time span used in this study is from June 1974  
109 through December 2018. Observed precipitation data is from the Global Precipitation  
110 Climatology Project (GPCP) [Adler *et al.*, 2003], and the time span used in this study is  
111 from January 1979 through January 2020. The horizontal resolutions used in this study  
112 are  $3^\circ$  for vertical motion, wind, and temperature, and  $2.5^\circ$  for OLR and precipitation.

### 2.2. Atmospheric General Circulation Model (AGCM) experiments

113 We use the Nonhydrostatic Icosahedral Atmospheric Model (NICAM) [Tomita and  
114 Satoh, 2004; Satoh *et al.*, 2008, 2014], the version of which used for our experiments  
115 is the latest stable version, NICAM16-S [Kodama *et al.*, 2020]. The condensation pro-  
116 cesses are explicitly calculated using the single moment water 6 microphysics scheme  
117 [Tomita, 2008a]. Sub-grid scale turbulence is calculated by a modified version of the  
118 Mellor-Yamada scheme [Mellor and Yamada, 1982; Nakanishi and Niino, 2004; Noda  
119 *et al.*, 2010]. The radiation model with two stream radiative transfer scheme employs  
120 a correlated  $k$ -distribution method (mstrnX) [Sekiguchi and Nakajima, 2008]. Surface

121 fluxes are calculated with a modified version of the Louis scheme [*Louis, 1979; Uno et al.,*  
122 1995]. For the land processes, the minimal advanced treatments of surface interaction  
123 and runoff (MATSIRO) land model [*Takata et al., 2003*] is used. Orographic gravity wave  
124 drag is considered to be sufficiently resolved in our simulations and to be opted out of  
125 employing the parameterization for the sub-grid scale orographic gravity wave drag.

126 The horizontal resolution is approximately 14 km on an icosahedral hexagonal-  
127 pentagonal mesh [*Tomita, 2008b*]. A terrain following vertical grid coordinate is employed  
128 with the model top of approximately 40 km and 38 vertical layers, whose thickness in-  
129 creases with height. The model time step is 60 seconds. Our simulations are initialized on  
130 00 UTC 28 June 2013, 2015, and 2016, and are integrated for 93 days for each year. Initial  
131 conditions of the atmosphere and the ocean are derived from the National Centers for En-  
132 vironmental Prediction (NCEP) Final Operational Model Global Tropospheric Analysis  
133 (NCEP-FNL) [*NCEP, 2015*]. Time evolution of the sea surface temperature is prescribed  
134 externally from the interpolation of the NCEP-FNL data at 00 UTC on each day. To mit-  
135 igate the effect of the model bias over land, the initial conditions of the land surface are  
136 taken from the monthly climatology derived from the last 5 years of a 10-year simulation  
137 of NICAM at 220 km horizontal resolution following *Kodama et al. [2015, 2020]*.

138 Because it takes approximately 45 days for the values of vertical motions to converge to  
139 realistic climatological values, the first 63 days of the integrations are taken as a spin-up  
140 period, and the last 30 days of the integrations starting from 1 September are analyzed  
141 in this study. For comparison, we have also performed the same simulation but initialized  
142 on 00 UTC 28 April 2016 to capture the strongest month of the Wall, i.e., July. Clima-  
143 tology of the Walker circulation, however, is not reproduced well, presumably because the

144 observed downward branches are not established until the end of May, during which the  
145 integration cannot be used as a spin-up period to capture the target circulation. Though  
146 the AGCM used in this study realistically simulates the mean field over long time peri-  
147 ods, the reproducibility of quick variations within relatively short time scales, such as a  
148 transition of seasons, is insufficient, to which future improvement is needed.

### 3. Climatology of the Wall

149 In this section, we first overview the seasonality of the Wall and the consistency with  
150 the local rainy seasons. Then, from the energetic viewpoint, we show that the strongest  
151 phase of the Wall is supported by horizontal cold advection associated with the Asian  
152 Summer Monsoon.

#### 3.1. Two-peak seasonality of the Wall

153 The Wall exhibits two-peak seasonal variability in its strength of the subsidence. The  
154 left panels of Fig. 3 shows the monthly climatological-mean equatorial vertical motion  
155 averaged over the base period of 1979-2017. The Wall exhibits moderate subsidence from  
156 January through March, almost disappears from April through May, reaches its strongest  
157 phase from June through September, and becomes weak from October through December.

158 The phase of this two-peak seasonality corresponds well to the annual precipitation  
159 cycle of the Eastern Horn of Africa, where two rainy seasons are known to exist. In this  
160 region, the term “Long Rains” denotes the longest and wettest rainy season that lasts  
161 from March through May, and the term “Short Rains” denotes the shorter and drier rainy  
162 season that peaks in October. Presumably, the lack of this seasonality in state-of-the-

163 art GCMs [*Tierney et al.*, 2015] is inseparable from the reproducibility of the seasonal  
164 variability of the Wall.

### 3.2. Role of horizontal cold advection for building the rigid Wall in its strongest phase

165 One of the essential features of the strongest phase of the Wall is that the subsidence  
166 reaches the surface, which is not the case in weaker phases (Fig. 3). In the Wall, horizontal  
167 temperature advection plays a key role for lower tropospheric atmospheric subsidence to  
168 extend to the surface. The right panels of Fig. 3 shows the mean horizontal temperature  
169 advection, which is defined as the inner product of climatological horizontal wind and the  
170 horizontal gradient of climatological temperature. Our definition of the mean horizontal  
171 advection does not take eddy heat transport into account.

172 The strongest subsidence observed in the lower troposphere from June through Septem-  
173 ber is supported by the mean horizontal advection. In general, adiabatic heating of large-  
174 scale downward motions is needed to balance radiative cooling in the tropics, and this  
175 energy budget is mostly true for the Walker circulation as well [*Veiga et al.*, 2011]. As  
176 described in the previous paragraph, however, it is not the case for the Wall. The contribu-  
177 tion from the horizontal cold advection makes this “subsidence extension” be a distinctive  
178 feature that can be observed particularly in the strong phases of the Wall.

179 The horizontal cold advection is tightly connected to the Asian Summer Monsoon. The  
180 top panels of Fig. 4 shows vertical-mean views of the horizontal temperature advection  
181 decomposed into zonal and meridional components. The horizontal advection cools the  
182 Wall region in boreal summer, when the Wall reaches its maximum phase (Fig. 4, top  
183 left). This horizontal advection is realized by the meridional advection (Fig. 4, top

184 right), rather than the zonal component (Fig. 4, top middle). In the bottom panels of  
185 Fig. 4, these components are further decomposed into zonal wind, zonal temperature  
186 gradient, meridional wind, and meridional temperature gradient. Based on these panels,  
187 the maximum horizontal advection in boreal summer originates from the southerly winds  
188 associated with Asian Summer Monsoon, which blows toward the upgradient direction of  
189 the temperature field in this season.

190 Presumably, the aforementioned cooling effect drags down the lower tropospheric Walker  
191 Circulation to the surface, which is capable of strengthening the downward flow of the  
192 Wall further. This notion is consistent with the disappearance of the Wall from April  
193 through May, because this season is the period when the Asian summer monsoon are  
194 weakened to switch its direction before the onset of the strong Somali jet in early June  
195 [e.g., *Findlater*, 1969]. From an energetic viewpoint, the relevance of the Somali jet is also  
196 consistent with *Heaviside and Czaja* [2013], who showed that the Somali jet dominantly  
197 accomplishes the atmospheric cross-equatorial heat transport.

#### 4. Interannual variability of the Wall

198 In this section, we define the Wall index to point out that both remote and local sea  
199 surface temperature (SST) variability explains the interannual variations of the Wall.

##### 4.1. Definition of the Wall index

200 To highlight the interannual variability of the Wall, we define the Wall index as the  
201 average over vertical motions at 250, 550, and 850 hPa. This index should be physically  
202 interpreted as the mass-weighted average of vertical motions for the upper (100-400 hPa),  
203 middle (400-700 hPa), and lower (700-1000 hPa) tropospheric layers. Figure 5a shows that,

204 to first order, the downward motion for the interannual scale is vertically constant, which  
205 justifies the definition of the Wall index for the purpose of representing the downward  
206 motion throughout the troposphere.

#### 4.2. Relationship with ENSO and local SST variability

207 The interannual variability of the Wall is explained by the El Niño Southern Oscillation  
208 (ENSO), which reminds us of the notion that the Wall is a part of the Walker circulation.  
209 Figure 5b shows the 5-month running-meaned timeseries of the Wall index. Also shown is  
210 the 5-month running-meaned Niño 3.4 index, which is defined as SST anomalies averaged  
211 over the Niño 3.4 region ( $5^{\circ}\text{S}$ - $5^{\circ}\text{N}$ ,  $170^{\circ}\text{W}$ - $120^{\circ}\text{W}$ ), but the sign is reversed. These two  
212 indices exhibits a correlation of 0.54 (1982-2017), which is significant at the 95% confidence  
213 level. During an El Niño event, the Wall region exhibits a weakly downward motion, and  
214 vice versa for a La Niña event. In particular, the strong negative peaks of the Wall index  
215 in 1982, 1997, and 2015 are well-explained by big El Niño events, whereas the strong  
216 positive peaks of the Wall index in 1998, 2010, and 2016 are well-explained by big La  
217 Niña events.

218 The association with ENSO is also verified by SST spatial patterns. Figure 5c shows the  
219 regression map of SST anomalies on the Wall index. This map clearly shows that the Wall  
220 variability is projected onto the equatorial Pacific SST variability. Because interannual  
221 variance of the tropical SST variability in the Pacific is dominated by ENSO, it is virtually  
222 certain that ENSO is one of the key factors to understand the Wall climate variability.

223 That being said, the clear regression pattern in the Pacific does not necessarily mean  
224 that the Wall variance is predominantly explained by ENSO. Figure 5d shows the same  
225 map as Fig. 5c but for correlation coefficients. Based on this correlation map, though

226 ENSO still retains its importance, local SST variability in the western equatorial Indian  
227 Ocean explains more variance of the Wall. Also notable is the high positive correlations  
228 over the maritime continent, presumably because the strength of the upward motions  
229 allowed in this region is also inseparable from the amount of the downward motions  
230 realized in the whole tropics, following the continuity equation.

## 5. Role of the East African topography for sustaining the Wall

231 Though we have concluded in section 3 that strong subsidence in the lower troposphere  
232 is associated with horizontal temperature advection, it remains unexplored what makes  
233 the subsidence in the Wall so strong that the Wall penetrates the entire troposphere in the  
234 vertical direction. In particular, cooling mechanisms of the upper troposphere has been  
235 largely unexplored. Therefore, in this section, we perform model experiments to highlight  
236 the role of topography for sustaining the Wall. Some implications for the climate of the  
237 Eastern Horn of Africa are also discussed.

### 5.1. Model experiments with flat East African topography

238 Our experiments are inspired by *Naiman et al.* [2017], who showed, in an interesting way,  
239 that topography can play major roles in determining the tropical circulation. Using the  
240 Geophysical Fluid Dynamics Laboratory (GFDL)-Earth System Model (ESM) 2M, they  
241 performed an experiment called “Pancake”, in which they removed all the topography  
242 on Earth and simulated the air-sea coupled system with flat lands. Because the Wall  
243 disappears in their “Pancake” run, we have hypothesized that, by flattening topography  
244 regionally rather than globally, it is possible to pinpoint the location of mountains that  
245 directly contribute to the realization of the Wall. In this regard, *Ogwang et al.* [2014]

246 also investigated a precipitation response to regionally flattened African topography and  
247 demonstrated that the mean rainfall significantly reduces at the west of the Wall region.  
248 Based on their results, the hydrology at the center of the Wall may be also modulated.

249 In this study, in addition to a control run, an AGCM experiment is arranged where  
250 the East African topography is regionally flattened. The experiment is named “Flat East  
251 Africa (FEA)”, in which the elevations are set to be 1 meter over the entire East African  
252 region ( $30^{\circ}\text{S}$ - $30^{\circ}\text{N}$ ,  $30^{\circ}\text{E}$ - $50^{\circ}\text{E}$ ) (Fig. 6a). The top and middle panels of Fig. 6b shows the  
253 monthly-mean equatorial vertical motion in September 2013 (ENSO neutral) and 2016 (La  
254 Niña) based on observations and the control experiment, respectively. The control runs in  
255 both years simulates the observed vertical motion associated with the Walker circulation  
256 well, so it is justified to investigate the Wall using this AGCM. We have also performed  
257 the same experimental sets but for 2015, an El Niño year, but the control run does not  
258 reproduce the observed features of the Walker circulation. This poor reproducibility of  
259 the Wall during an El Niño year appears to be because the observed Wall is weaker than  
260 those of other years (Fig. 5b). In our model, the water vapor supply from the anomalously  
261 warm eastern equatorial Pacific is biased to be excessive. The Walker Circulation over the  
262 tropics is distorted by this bias, to which the zonally-thin downward branch is sensitive.  
263 Therefore, for the rest of this article, only the ENSO neutral and La Niña years will be  
264 discussed.

265 The FEA experiment reveals that, without the East African topography, the Wall dis-  
266 appears almost entirely through the troposphere (Fig. 6b, bottom). By comparing the  
267 control and FEA runs, at least in both 2013 and 2016, the East African topography is a  
268 necessary condition for the existence of the Wall.

A promising hypothesis is that the lack of turbulence generated by mountain waves suppresses vertical mixings. Because the lower troposphere generally has lower potential temperature than the upper troposphere, the reduction of vertical heat exchange weakens the subsidence of upper tropospheric air. Figure 7a shows the difference of equatorial vertical motions between the control and FEA runs, which should be interpreted as the downward motion forced by the East African topography. In this model, the downward branch is shifted west compared to observations, so 30°E-45°E is drawn. Also shown in the left panel of Fig. 7b is the energetic tendency contributions by the sum of eddy heat transport (EHT) and longwave radiation. Here, the EHT contribution is calculated as follows.

$$\text{EHT} = -\frac{\overline{\partial u' T'}}{\partial x} - \frac{\overline{\partial v' T'}}{\partial y} - \frac{\overline{\partial w' T'}}{\partial z} \quad (1)$$

where  $x$ ,  $y$ , and  $z$  denotes zonal, meridional, and vertical coordinates, respectively;  $u$ ,  $v$ , and  $w$  denotes zonal wind, meridional wind, and vertical motion, respectively;  $T$  denotes temperature. The overlines denote the mean over September simulated in the model, and the primes denote deviations from the mean.

EHT and longwave radiation explain how East African topography works for realizing the vertical motion. In particular, the downward motion at higher levels than 10 km is predominantly explained by the eddy heat transport (Fig. 7b, middle). The phase of heat and momentum transport is shifted, which suggests a hint of stationary gravity waves forced by mountains. Because the mountain waves suppress the upper tropospheric cloudiness, radiative cooling is enhanced in the middle tropospheric layer (Fig. 7b, right), which in turn strengthens the downward motion further.

280 A caveat of this heat budget analysis is that the cooling effects of eddy transport and  
281 longwave radiation is quantitatively insufficient to explain the total downward motion.  
282 By assuming the dry adiabatic lapse rate, the vertical motion of 500 m/day requires  
283 approximately 5 °C/day of cooling, but the aforementioned effects only explains 1 °C/day  
284 of cooling. Though more dominant effect may exist, it is still hard to close perfectly  
285 the heat budget based on an analysis of the 6-hourly snapshots. Nevertheless, by using  
286 the convection permitting model without artificial gravity wave drags, our experiments at  
287 least confirm that decent amount of the upper tropospheric downward motion is sustained  
288 by eddy heat transport forced by East African topography, rather than radiative cooling.  
289 Radiative cooling, enhanced by the clearer condition, is only capable of contributing the  
290 downward motion in lower altitudes where specific humidity is higher.

291 This vertical mixing effect serves as a good example where interscale interaction plays  
292 a fundamental role in downward branches, in addition to convective upward branches, to  
293 realize the large-scale atmospheric circulation in the current tropical climate. Specifically,  
294 the narrowly localized downward branch above the western Indian Ocean is realized with  
295 a help of interactions between large-scale motions and disturbances in smaller horizontal  
296 scales than the weak temperature gradient approximation [*Sobel et al.*, 2001].

## 5.2. Implications for the climate of the Eastern Horn of Africa

297 Without the East African topography, the relatively dry climate at the Eastern Horn of  
298 Africa would become wetter than in the real world. Figure 8 shows the monthly-mean OLR  
299 and precipitation near the Wall in 2013 and 2016. The control run of the high-resolution  
300 convection-permitting model reproduces both OLR and precipitation well, particularly  
301 the discontinuity of ITCZ. In the FEA run, as the East African topography is flattened,

302 the discontinuity of ITCZ disappears, which makes the Eastern Horn of Africa be covered  
303 by ITCZ. Our result is consistent with a model experiment without topography performed  
304 by *Chou and Neelin* [2003], which does not exhibit the ITCZ discontinuity.

305 Both local processes and remote forcings can contribute to the “closing” of the ITCZ  
306 discontinuity in the FEA run. Locally, the removal of the Wall enhances convection  
307 above the Eastern Horn of Africa. This enhancement is due to the reduction of large-  
308 scale atmospheric subsidence as discussed in the last subsection. In addition to this local  
309 instability effect, clouds and moist air, which are advected remotely by zonal winds, are  
310 also allowed to enter the Eastern Horn of Africa from the interior of the African continent,  
311 because topographic obstacles do not exist.

## 6. Summary and Discussions

312 We have reconsidered the climatology and the interannual variability of the Walker  
313 circulation by focusing on its sharp downward branch, which we refer to as the Wall,  
314 observed at the western edge of the Indian Ocean (Figs. 1 and 2). A distinctive feature  
315 of the Wall is the two-peak seasonality (Fig. 3). The two weak phases of the Wall, one in  
316 boreal spring and the other in boreal fall, correspond well to the two rainy seasons at the  
317 Eastern Horn of Africa, which is not reproduced well by state-of-the-art GCMs. Another  
318 distinctive feature is that the subsidence of the Wall in its strongest phase reaches the  
319 surface (Fig. 1a). This “subsidence extension” appears to be sustained by horizontal cold  
320 advection associated with the Asian Summer Monsoon (Figs. 3 and 4). The interannual  
321 variability of the Walker circulation is in no doubt associated with ENSO, but more  
322 variance is explained by SSTs in western equatorial Indian Ocean and over the maritime  
323 continent (Fig. 5).

324 AGCM experiments show that the East African topography determines the strength  
325 of the Wall (Fig. 6). In the FEA experiment, where the East African mountains are  
326 broadly flattened, the Wall almost vanishes throughout the entire tropospheric layer.  
327 This result leads to a conclusion that the East African topography is necessary for the  
328 existence of the Wall. We hypothesize that the role of topography is to generate mountain  
329 waves in response to large-scale circulation. The stationary vertical mixings enhance  
330 vertical heat exchange to cool the upper troposphere, which makes the Wall rigid (Fig. 7).  
331 Assuming this mechanism, climate variability of the Wall could also be understood based  
332 on interscale interactions between macroscopic large-scale circulation and microscopic  
333 mountain waves. In addition, based on *Chakraborty et al.* [2009], we could also hypothesize  
334 that the modulation of the Somali jet strength by the African topography could control  
335 the lower tropospheric downward motion. Our simulation, however, does not reproduce  
336 the lower tropospheric downward motion realistically enough to make a case for the role  
337 of horizontal advection. Further process studies are needed to improve the robustness of  
338 these physical processes.

339 An implication of our conclusion is that the dry and clear climate at the Eastern Horn  
340 of Africa is sustained by the East African topography (Fig. 8). As a local effect, the large-  
341 scale subsidence associated with the Wall suppresses the local convection by drying the  
342 environment and by suppressing upward motion. At the same time, the high mountains  
343 in the East Africa serves as obstacles that prevents clouds and moist air from being  
344 conveyed from the interior of the African continent. Because both of these local and  
345 remote processes are consistent with the essentiality of the East African topography, it

346 remains to be an open question which process serves as the dominant cause of the ITCZ  
347 discontinuity.

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351 [gridded/data.interp\\_OLR.html](https://psl.noaa.gov/data/gridded/data.interp_OLR.html), and the GPCP dataset available online at <https://psl.noaa.gov/data/gridded/data.gpcp.html>. The first author is supported by the  
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355 cal computations using NICAM is performed on a super computer, Oakforest-PACS.

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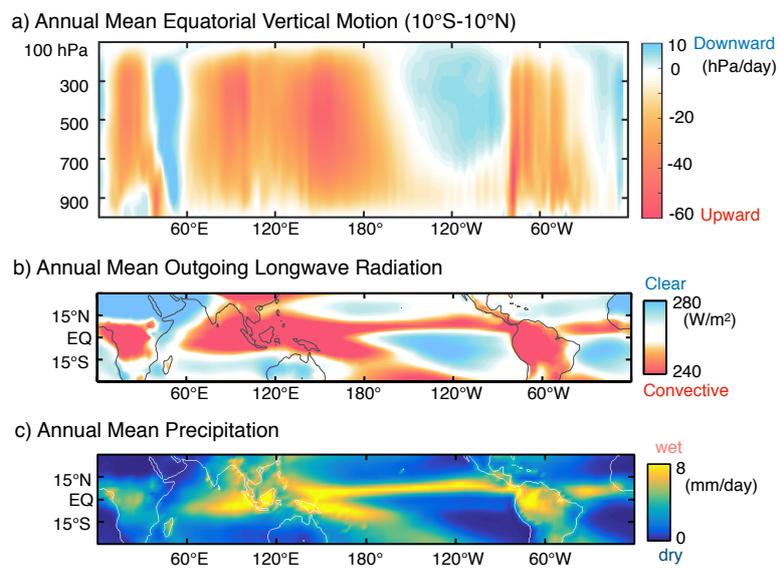
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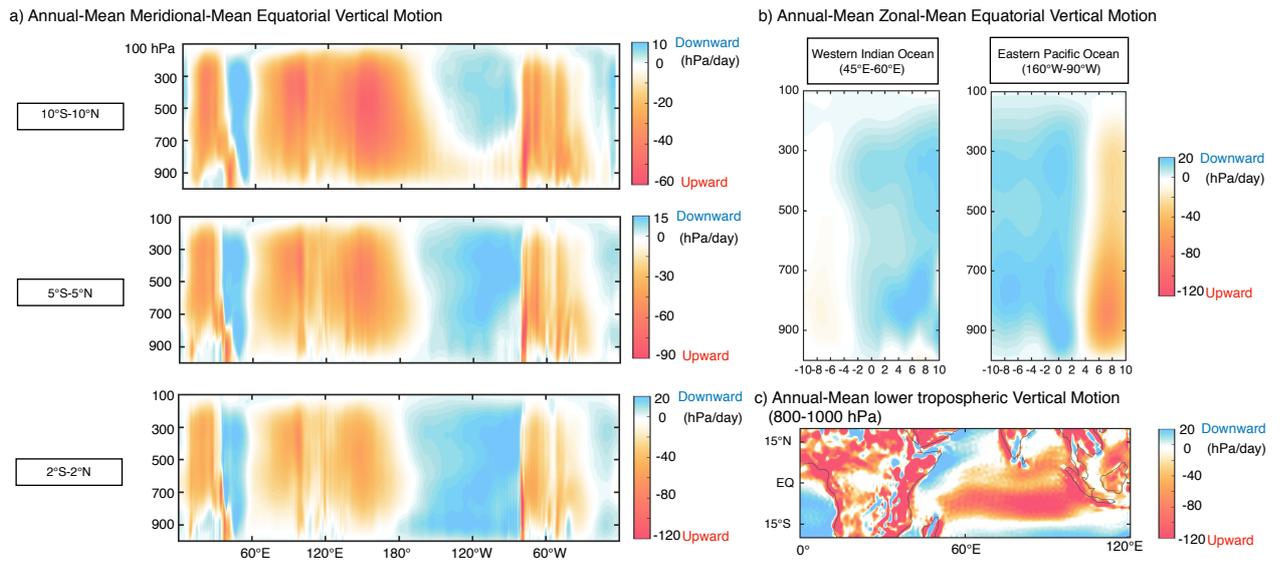
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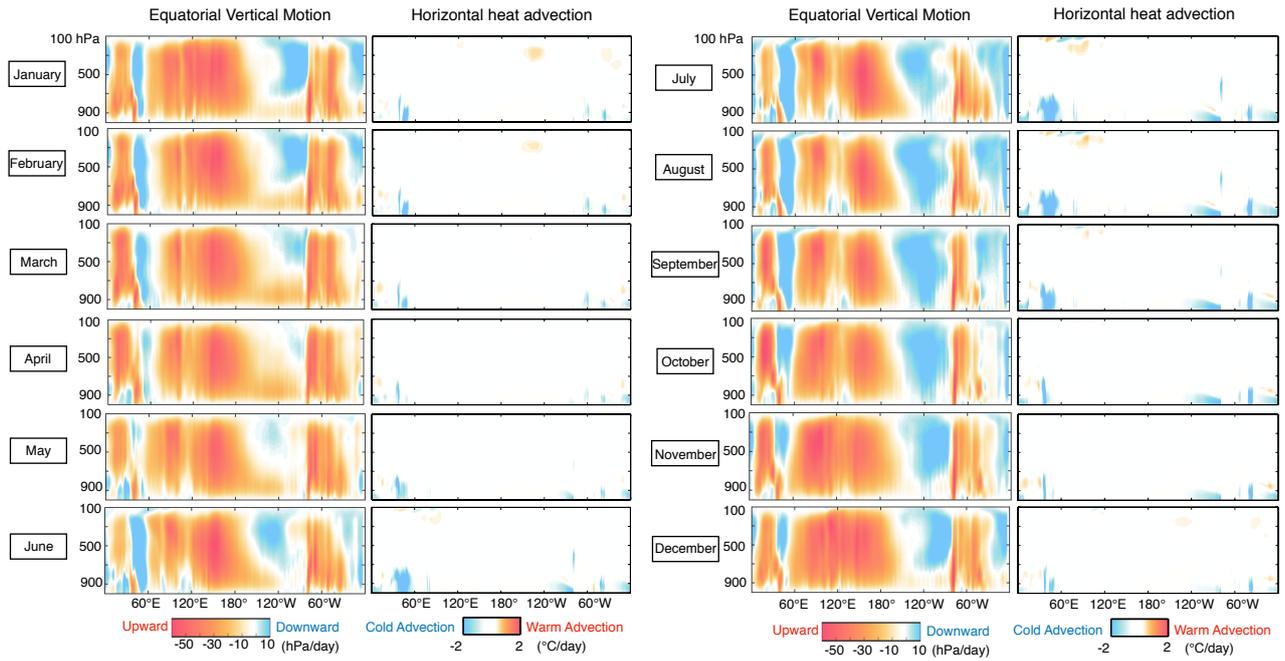
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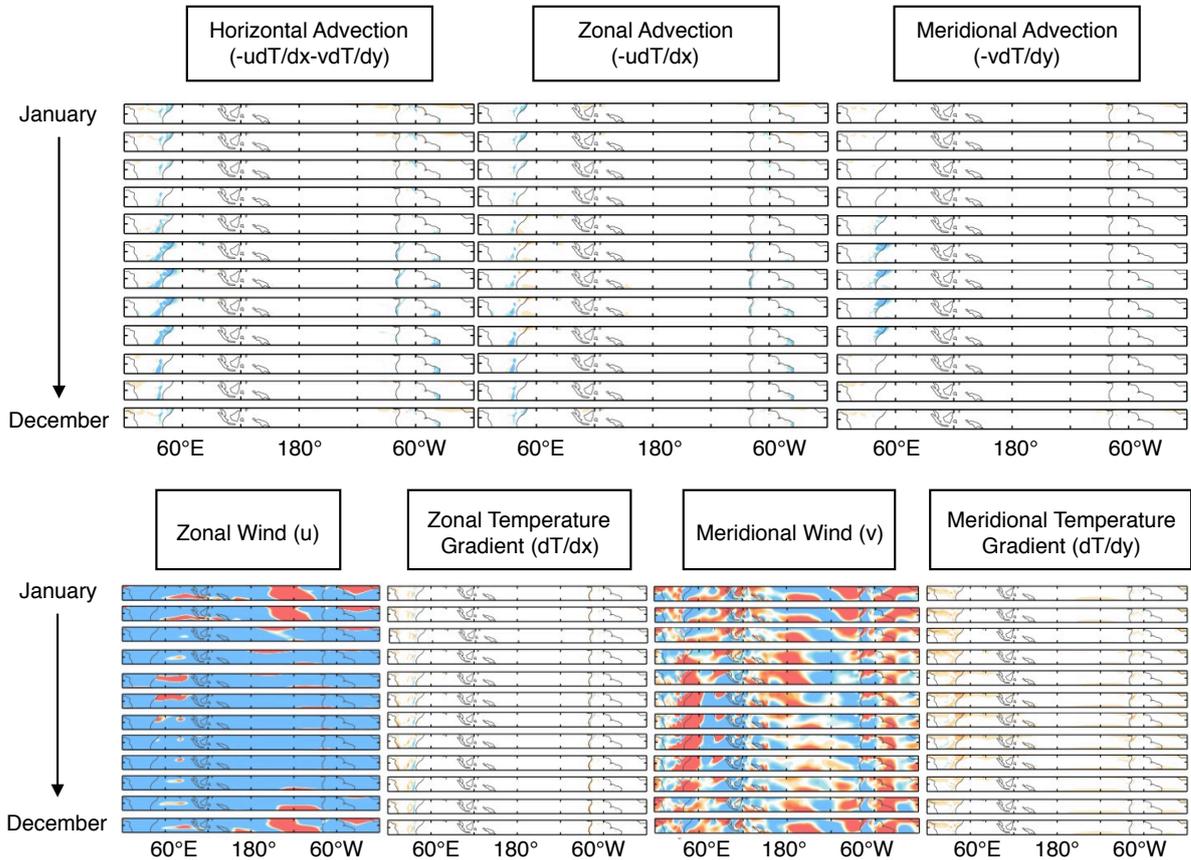
**Figure 1.** (a): Observed annual-mean vertical motion averaged meridionally over the equatorial region (10°S-10°N). (b): Observed annual-mean outgoing longwave radiation (OLR). (c): Observed annual-mean precipitation.



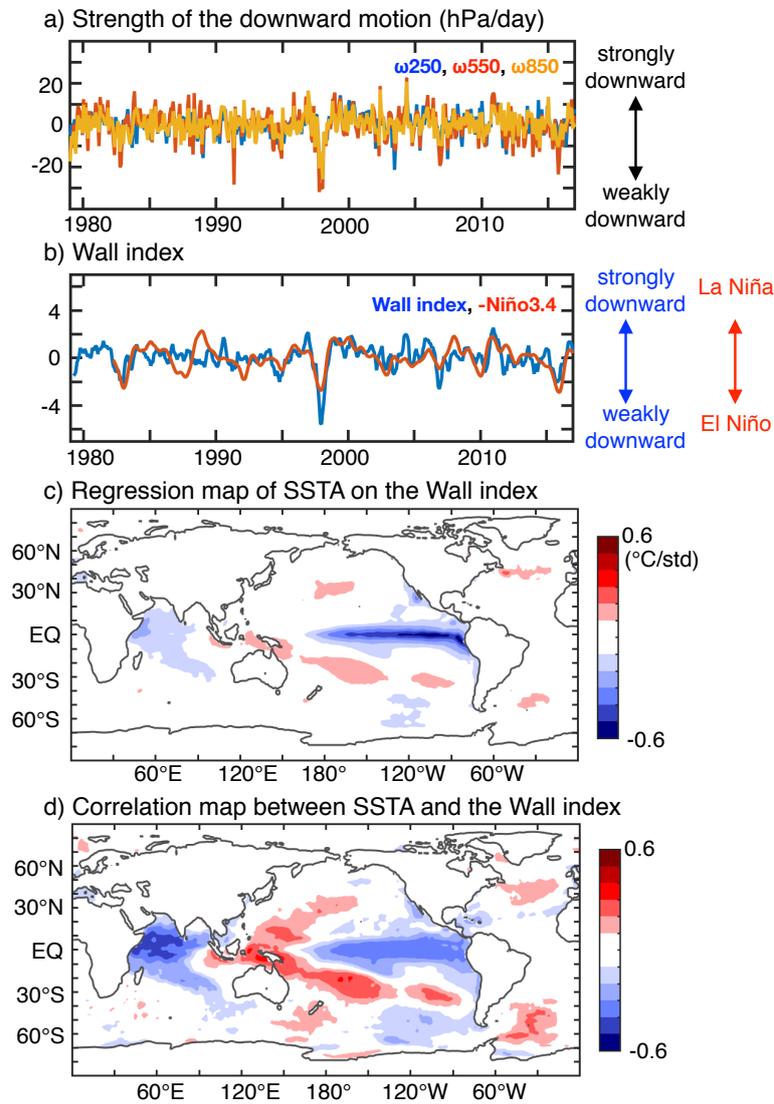
**Figure 2.** (a): As in Fig. 1a, but over the equatorial belt of 10°S-10°N (top) , 5°S-5°N (middle), and 2°S-2°N (bottom). (b): As in Fig. 1a, but averaged zonally over the Western Indian Ocean (45°E-60°E) and the Eastern Pacific Ocean (160°W-90°W). (c): As in Fig. 1a, but averaged vertically over the lower troposphere (800-1000hPa).



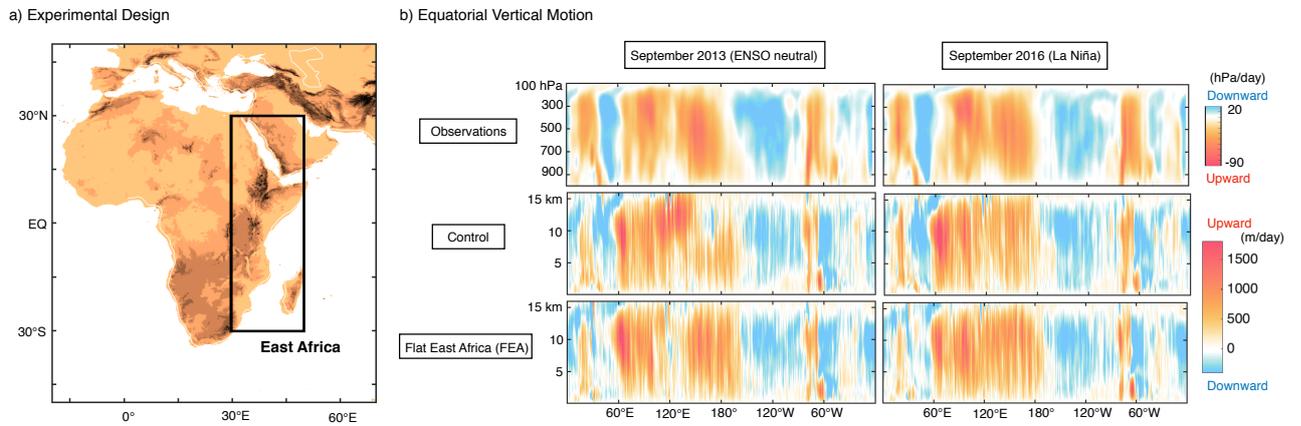
**Figure 3.** Left columns, As in Fig. 1a, but monthly mean values for each month averaged over 1979-2017. Right columns, As in right, but for mean horizontal advection defined as the inner product of mean horizontal wind and the horizontal gradient of mean temperature.



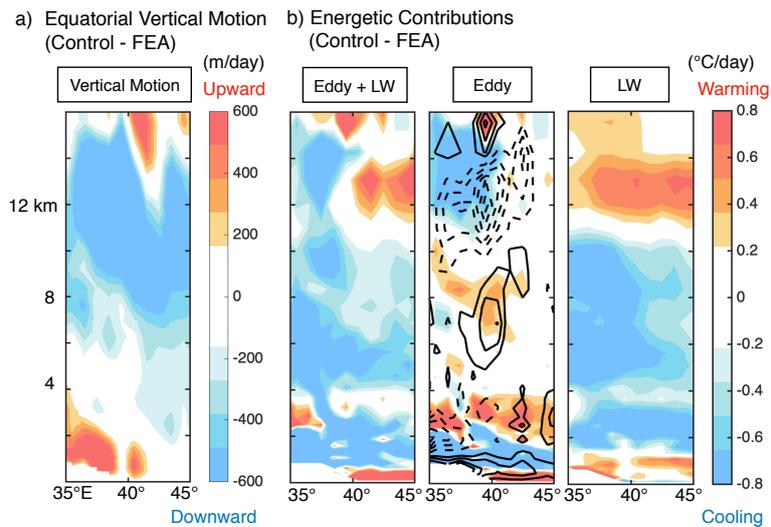
**Figure 4.** Top, As in the right columns of Fig. 3, but for vertical-mean tropospheric horizontal advection (left), zonal advection (middle), and meridional advection (right). Contribution of eddy transport is not considered. Bottom, As in top, but for zonal wind (left), zonal temperature gradient (second from the left), meridional wind (second from the right), and meridional temperature gradient (right).



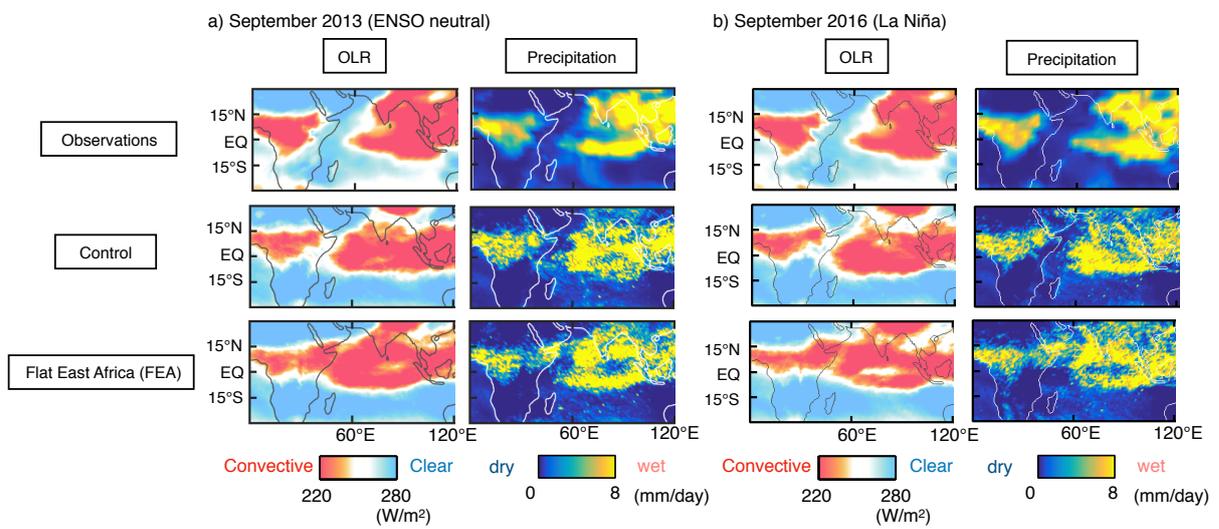
**Figure 5.** (a): Time series of the observed monthly-mean vertical motion at 250 hPa (blue), 550 hPa (red), and 850 hPa (yellow), averaged over the Wall region (10°S-10°N, 40°E-60°E). (b): Blue, Time series of the Wall index defined as the mean of the three time series shown in (a) standardized by its own standard deviation. Red, Niño 3.4 index defined as the regional-mean sea surface temperature anomalies (SSTA) over the Niño 3.4 region (5°S-5°N, 170°W-120°W) standardized by its own standard deviation and the sign is reversed. (c): Regression map of SSTA on the Wall index. (d) Correlation map between SSTA and the Wall index.



**Figure 6.** (a): Topography of the entire African continent. Black box shows the East African region (30°S-30°N, 30°E-50°E). (b): As in Fig. 1a, but for one-month mean values calculated for September 2013 (left) and 2016 (right) based on observations, the control, and the Flat East Africa (FEA) experiment in this order from the top panel. In the FEA experiment, the topography in the East African region, shown as the black box in (a), is flattened and set to be 1 meter.



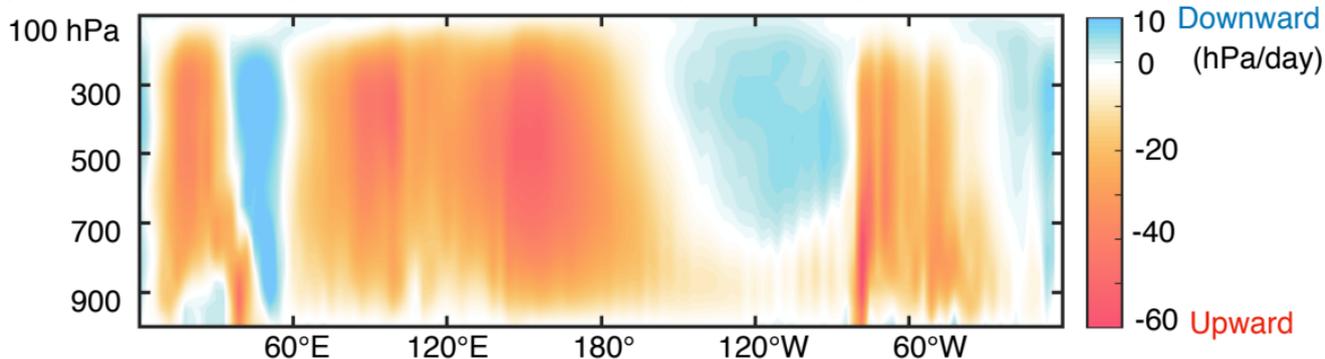
**Figure 7.** (a): As in the right panel of Fig. 6b, but the difference between the control and FEA runs in 35°E-45°E. (b): As in (a), but the energetic tendency contributions by the sum of eddy heat transport and longwave radiation (left), eddy heat transport (middle), and longwave radiation (right). Also shown as contours in the middle panel is eddy vertical momentum transport. Solid (dashed) curves denote upward (downward) vertical momentum transport. Contour interval is 0.05 (m/s)/day.



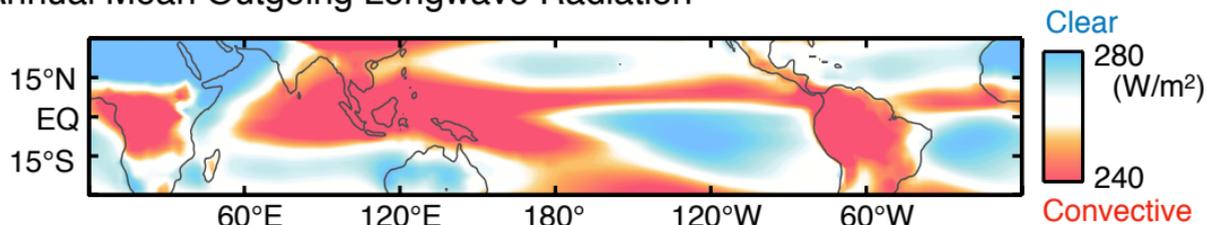
**Figure 8.** (a): As in Fig. 6b, but for OLR (left) and precipitation (right) for 2013. (b): As in (a), but for 2016.

Figure 1.

a) Annual Mean Equatorial Vertical Motion (10°S-10°N)



b) Annual Mean Outgoing Longwave Radiation



c) Annual Mean Precipitation

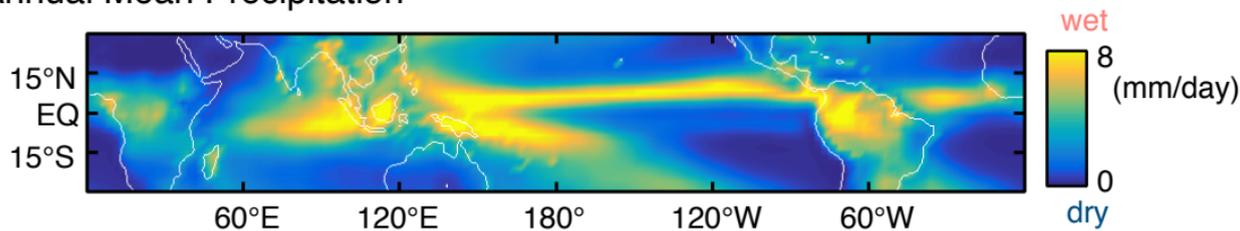
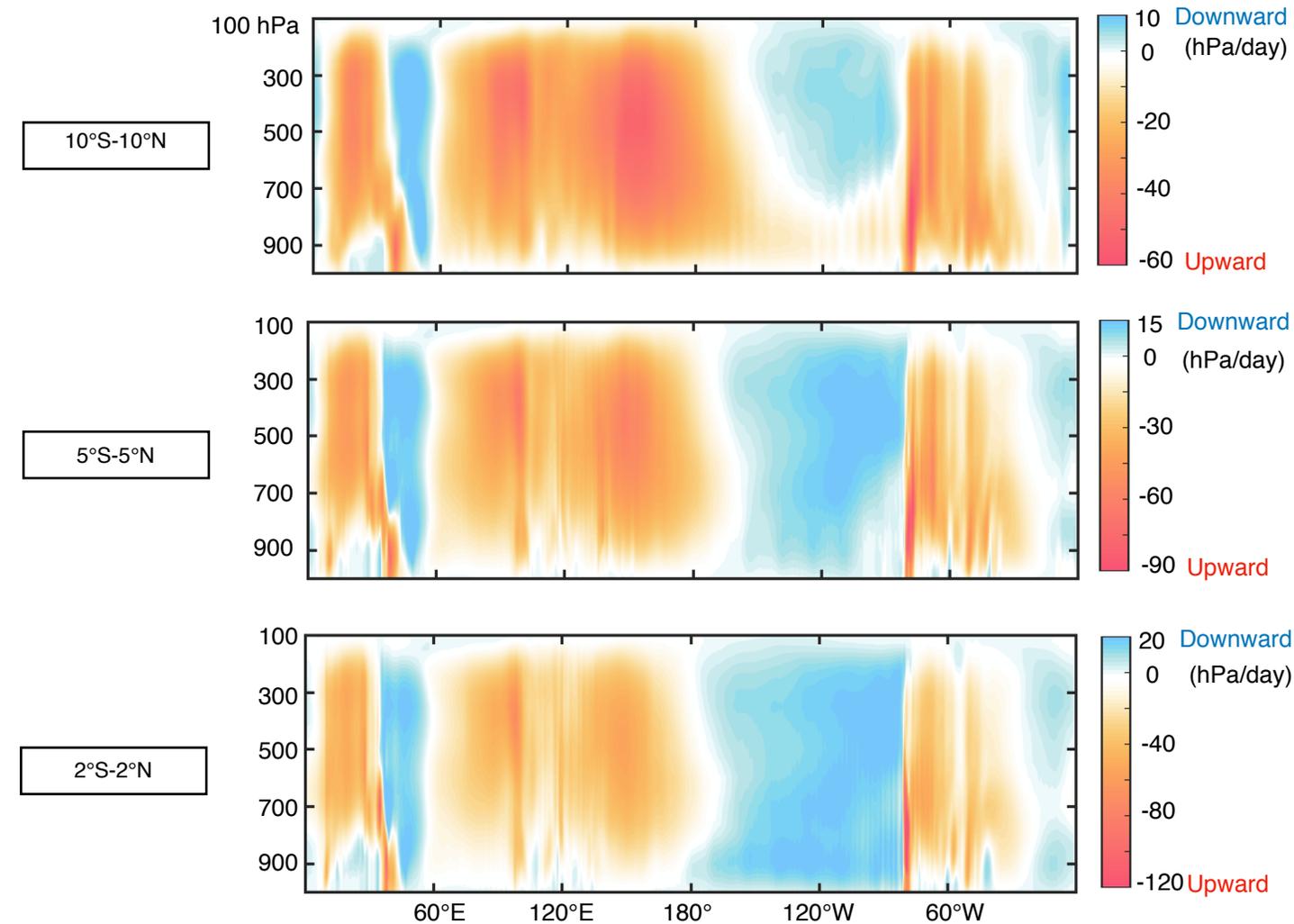
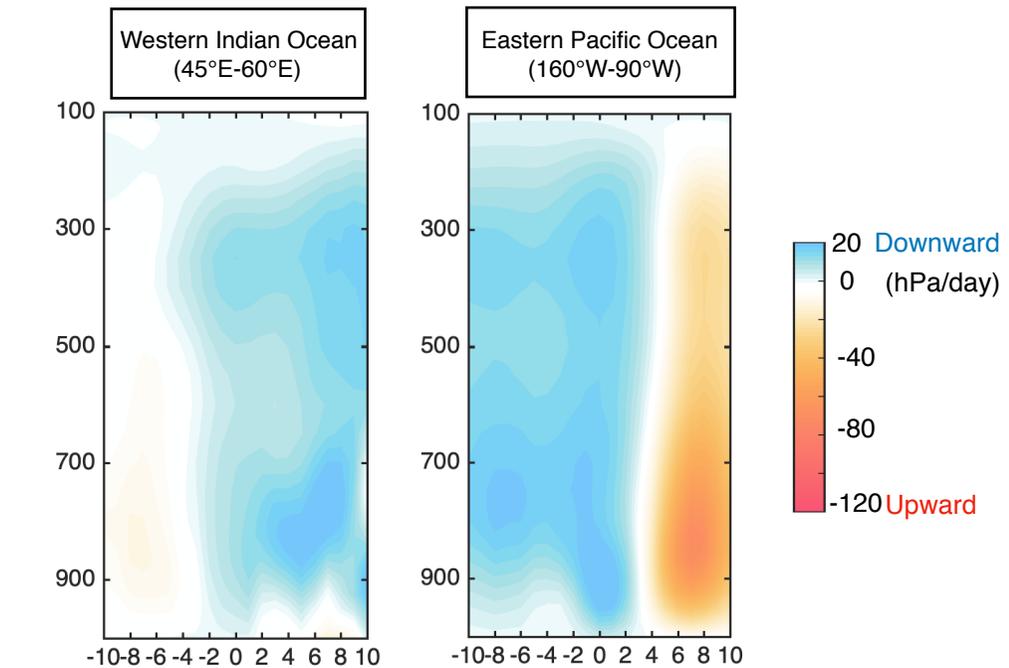


Figure 2.

a) Annual-Mean Meridional-Mean Equatorial Vertical Motion



b) Annual-Mean Zonal-Mean Equatorial Vertical Motion



c) Annual-Mean lower tropospheric Vertical Motion (800-1000 hPa)

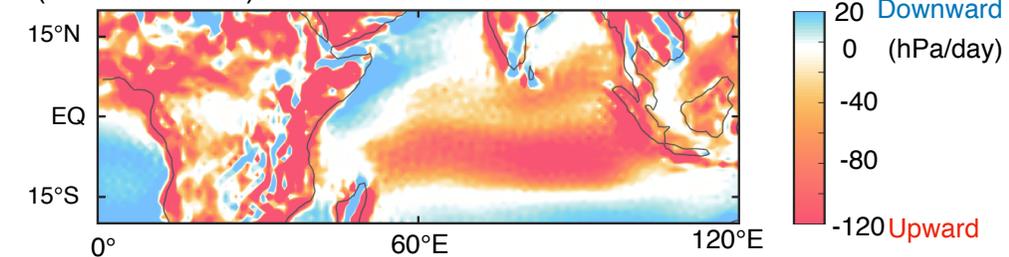
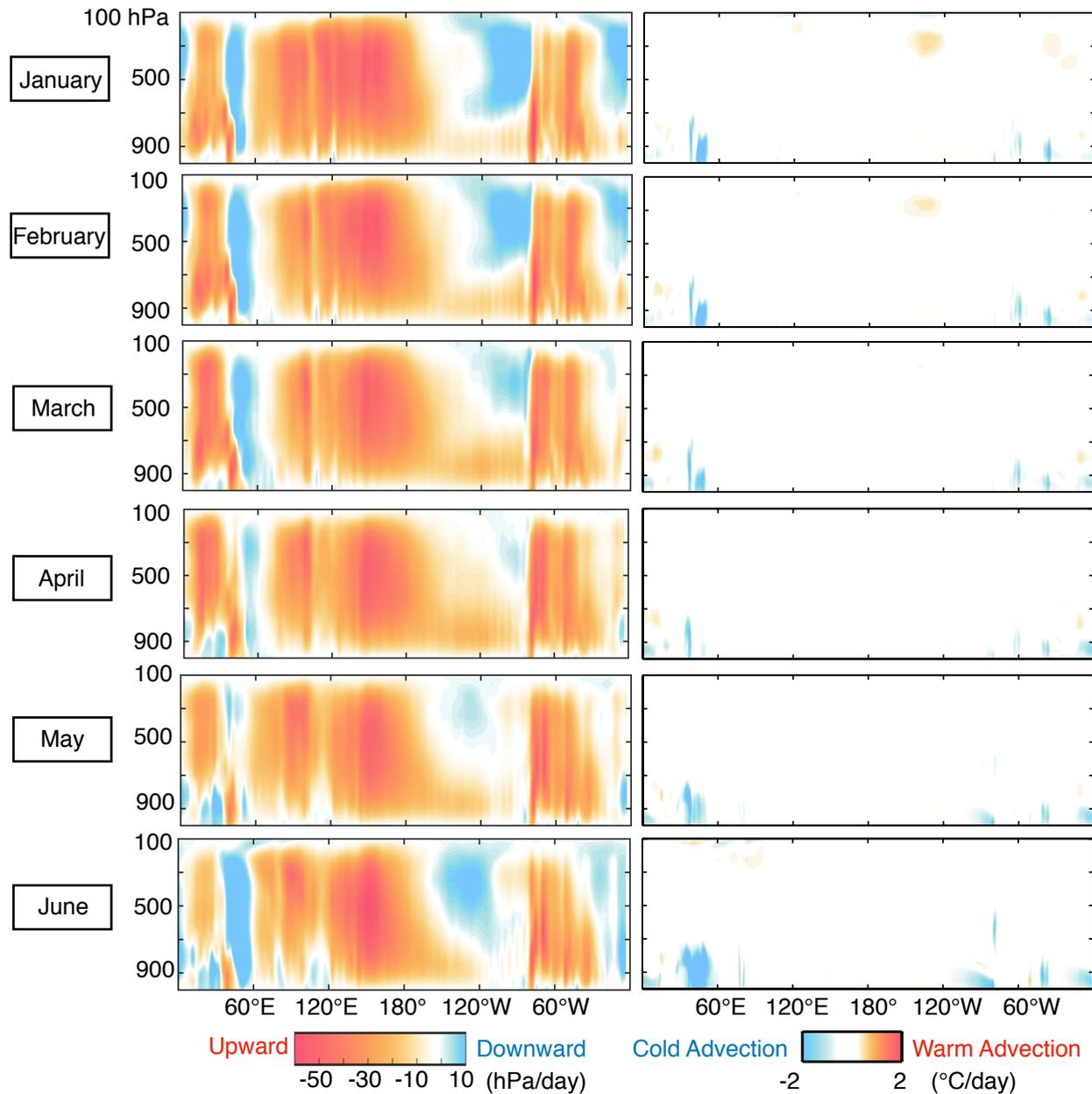


Figure 3.

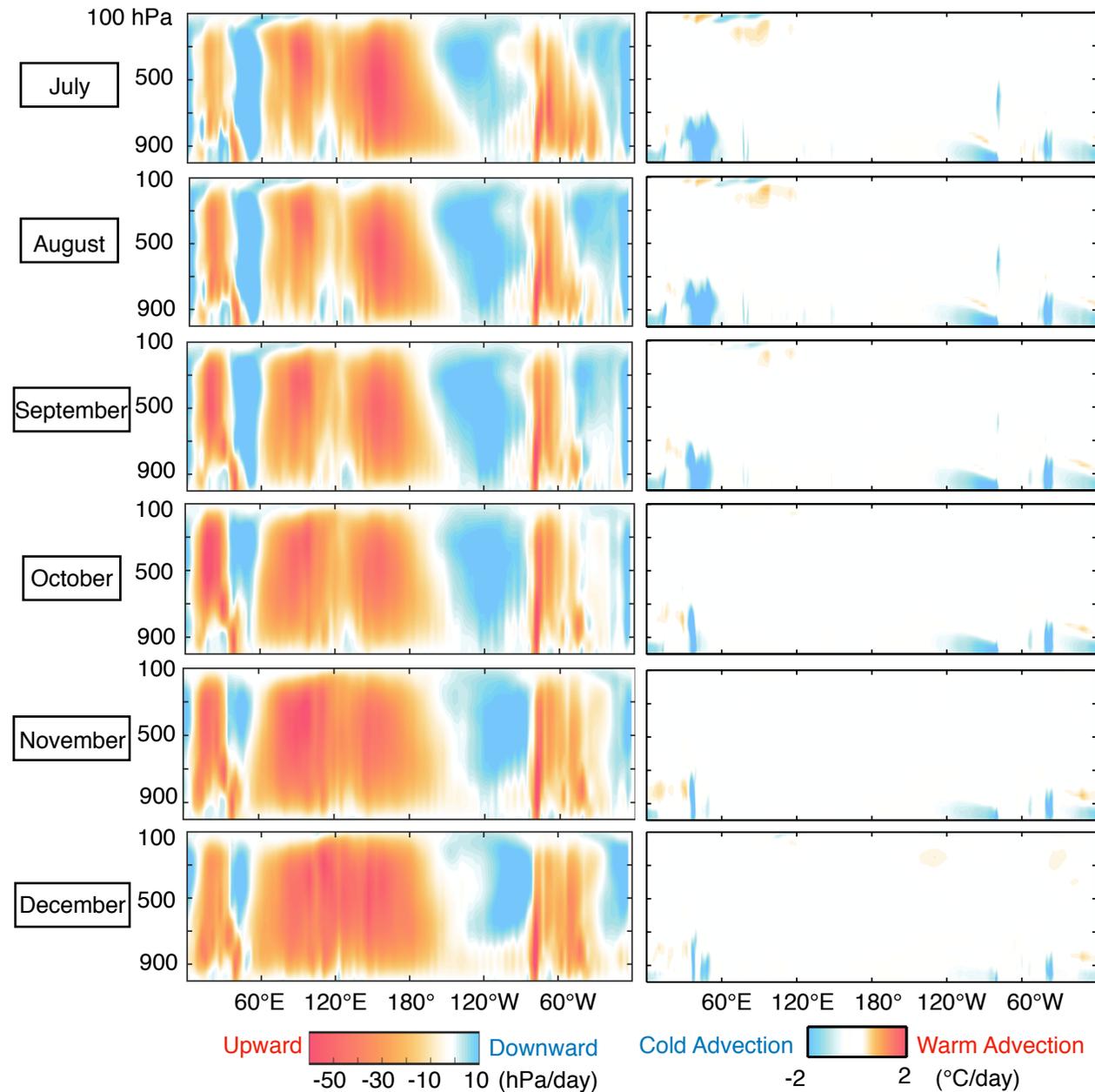
Equatorial Vertical Motion

Horizontal heat advection



Equatorial Vertical Motion

Horizontal heat advection



**Figure 4.**

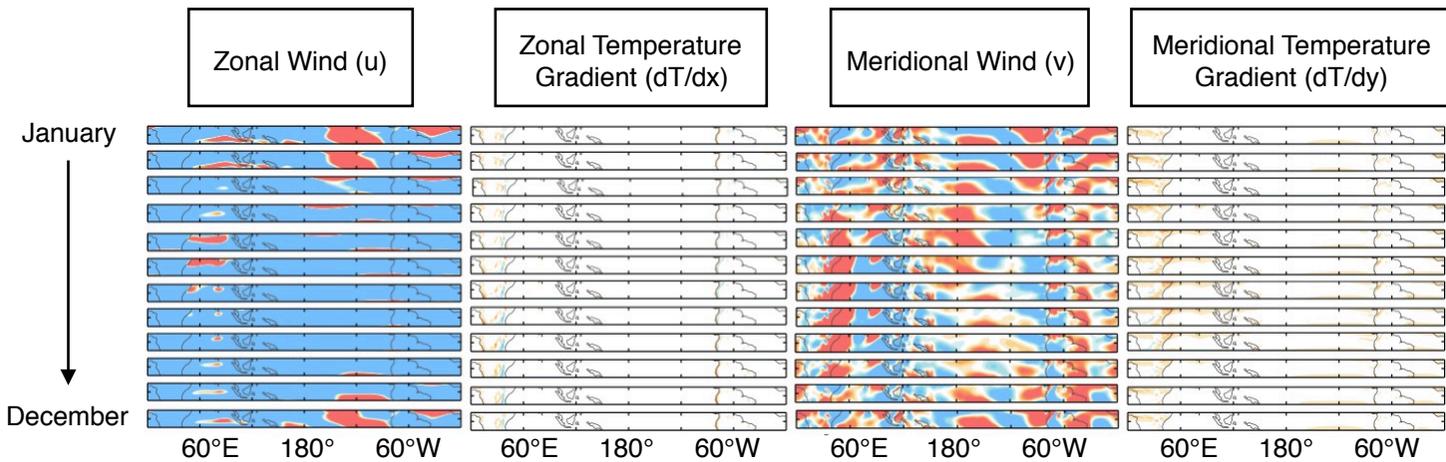
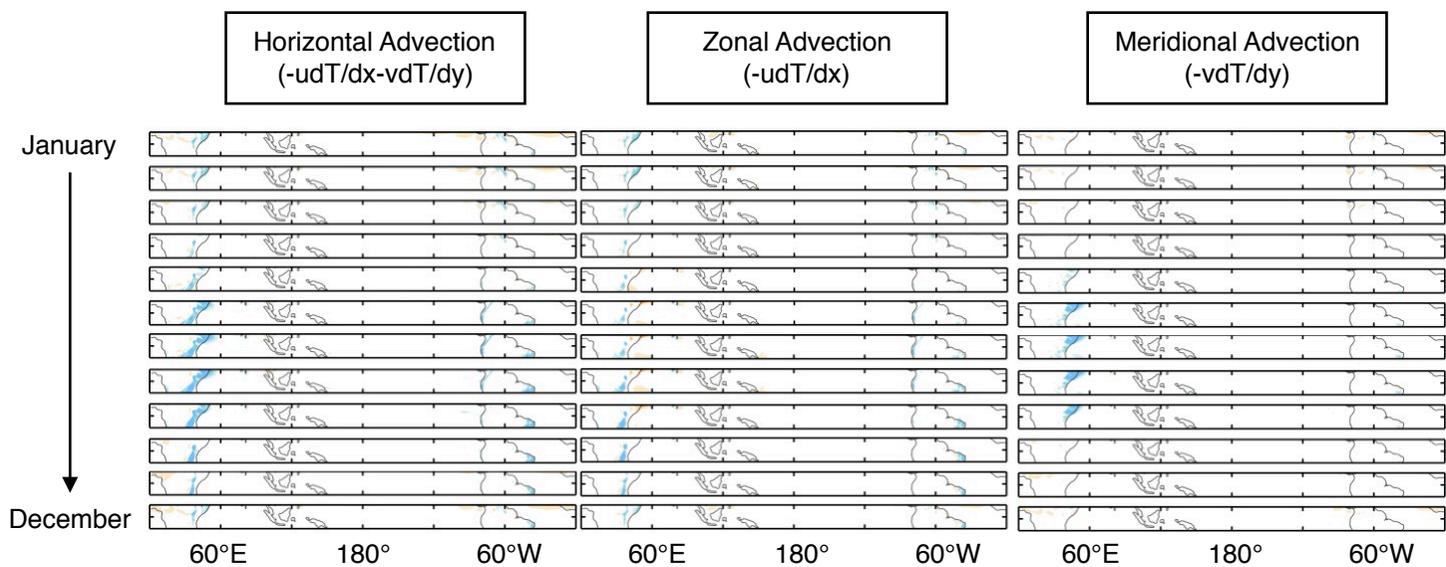
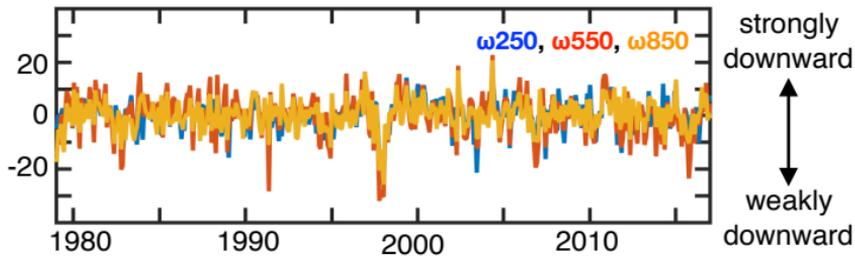
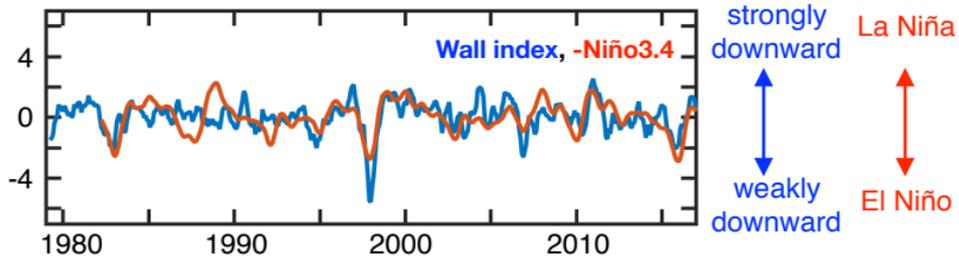


Figure 5.

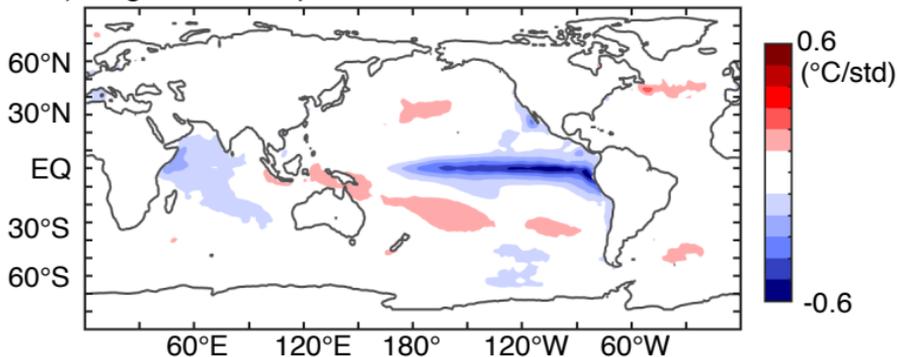
a) Strength of the downward motion (hPa/day)



b) Wall index



c) Regression map of SSTA on the Wall index



d) Correlation map between SSTA and the Wall index

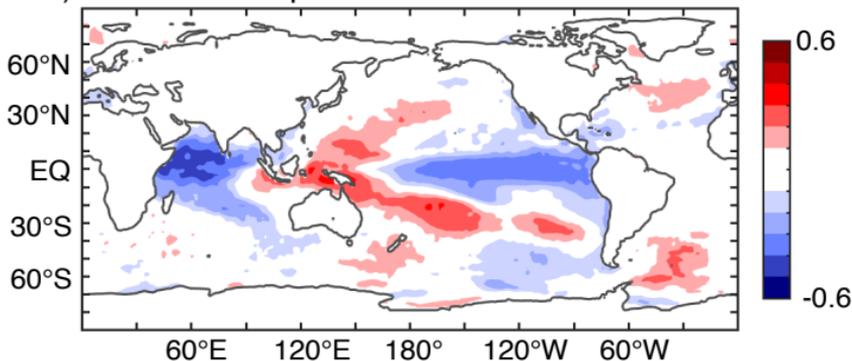
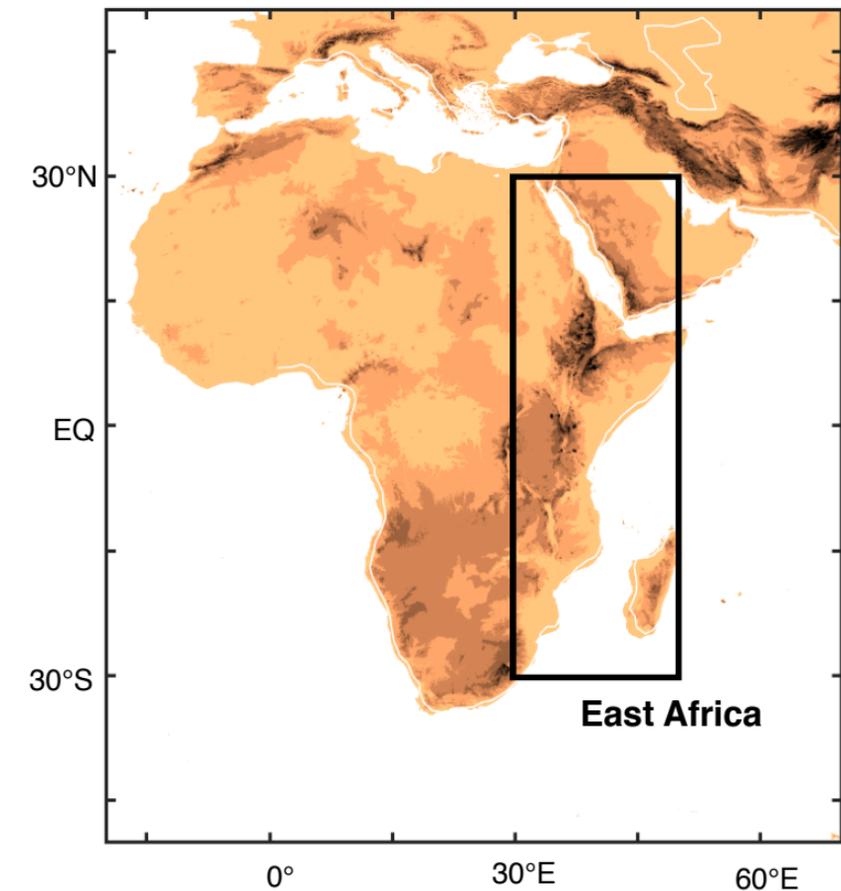


Figure 6.

a) Experimental Design



b) Equatorial Vertical Motion

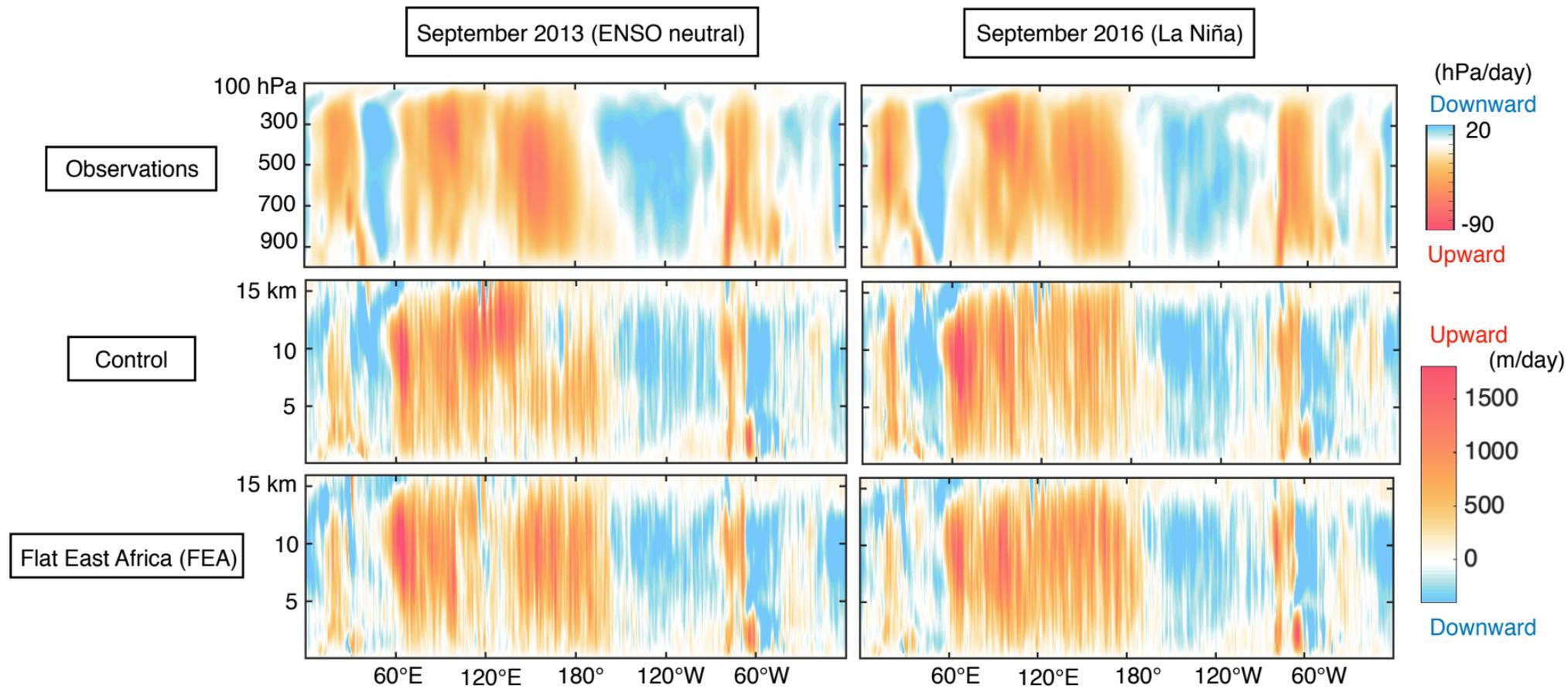
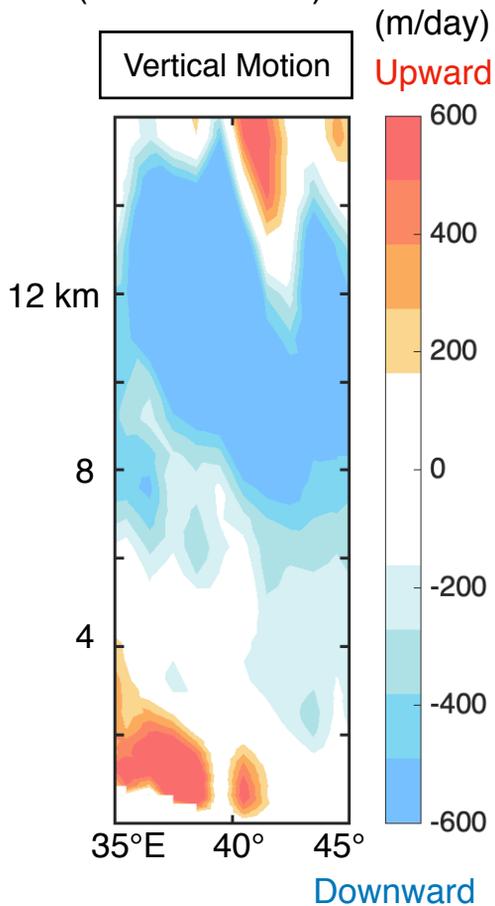


Figure 7.

a) Equatorial Vertical Motion  
(Control - FEA)



b) Energetic Contributions  
(Control - FEA)

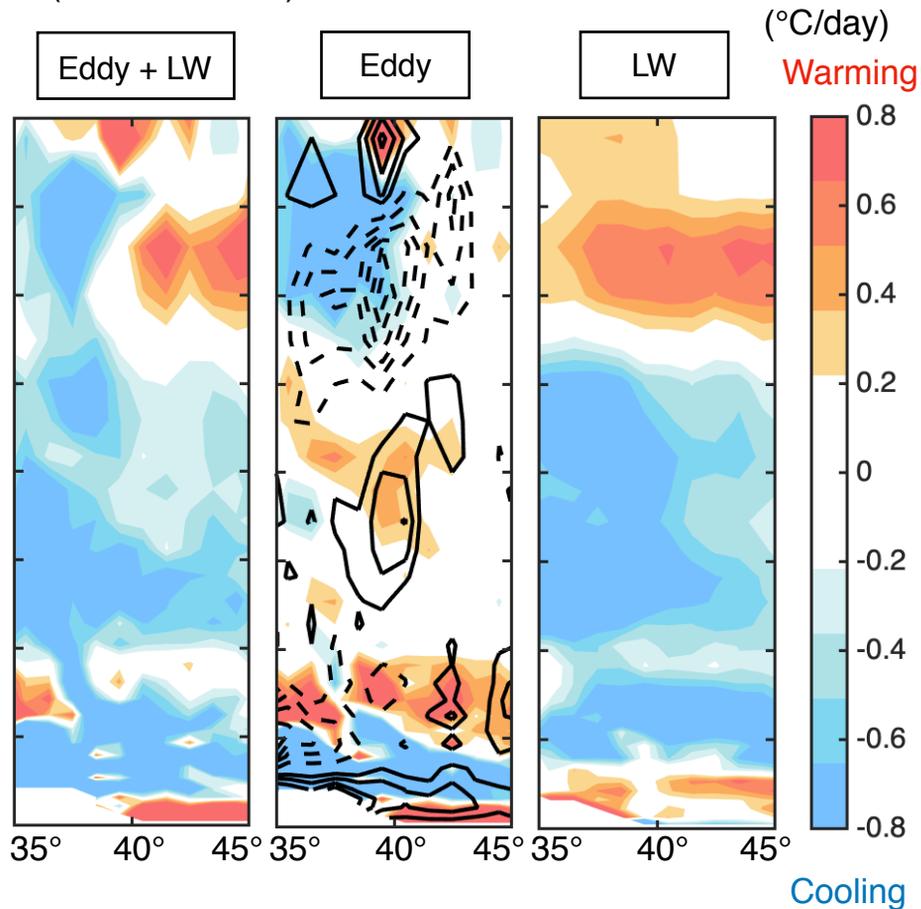


Figure 8.

a) September 2013 (ENSO neutral)

b) September 2016 (La Niña)

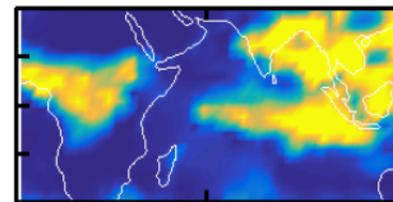
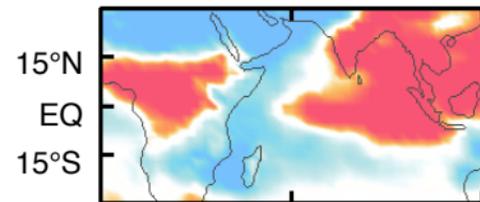
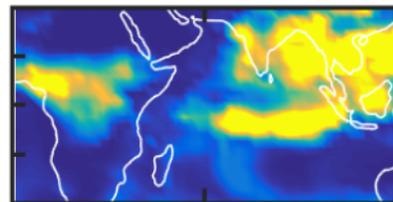
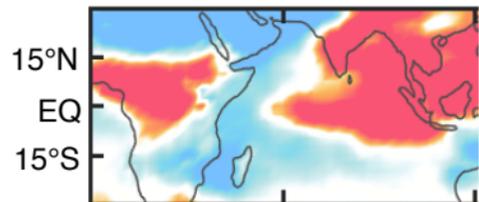
OLR

Precipitation

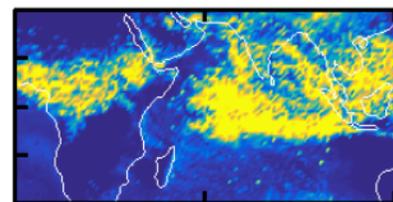
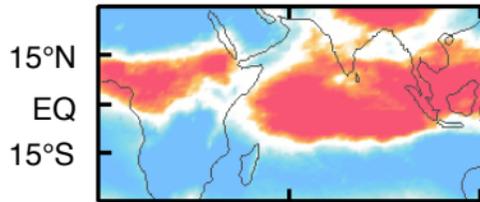
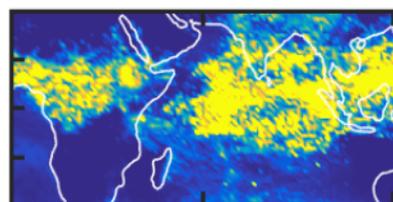
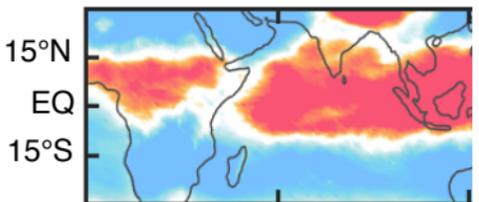
OLR

Precipitation

Observations



Control



Flat East Africa (FEA)

