

# Siege of the South: Hunga Tonga-Hunga Ha'apai Water Vapor Excluded from 2022 Antarctic Stratospheric Polar Vortex

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## Key Points:

- MLS trace gas data show that the Hunga Tonga-Hunga Ha'apai H<sub>2</sub>O plume was effectively excluded from the 2022 Antarctic polar vortex
- Antarctic lower stratospheric vortex strength, size, and longevity were among the largest on record, but within the range of previous years
- Antarctic chemical ozone loss in 2022 was unexceptional, with MLS ozone and related trace gases observed to be near average

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## Abstract

We use Aura Microwave Limb Sounder (MLS) trace gas measurements to investigate whether water vapor ( $\text{H}_2\text{O}$ ) injected into the stratosphere by the Hunga Tonga-Hunga Ha’apai (HTHH) eruption affected the 2022 Antarctic stratospheric vortex. Other MLS-measured long-lived species are used to distinguish high HTHH  $\text{H}_2\text{O}$  from that descending in the vortex from the upper-stratospheric  $\text{H}_2\text{O}$  peak. HTHH  $\text{H}_2\text{O}$  reached high southern latitudes in June–July but was effectively excluded from the vortex by the strong transport barrier at its edge. MLS  $\text{H}_2\text{O}$ , nitric acid, chlorine species, and ozone within the 2022 Antarctic polar vortex were near average; the vortex was large, strong, and long-lived, but not exceptionally so. There is thus no clear evidence of HTHH influence on the 2022 Antarctic vortex or its composition. Substantial impacts on the stratospheric polar vortices are expected in succeeding years since the  $\text{H}_2\text{O}$  injected by HTHH has spread globally.

## Plain Language Summary

The 2022 Hunga Tonga-Hunga Ha’apai eruption injected vast amounts of water vapor into the stratosphere. There has been much speculation that this large increase in water vapor could impact the Antarctic stratospheric polar vortex and Antarctic ozone hole: Water vapor plays an important role in polar vortex ozone depletion by providing the necessary conditions for the formation of polar stratospheric clouds. These clouds provide surfaces on which ozone-depleting chemical reactions can occur. The excess water vapor could also change the vortex evolution via water vapor’s effects on temperature, which could in turn affect the strong band of winds demarcating the polar vortex edge. We use satellite measurements of water vapor and other gasses to show that by the time the water vapor from the Hunga Tonga volcanic eruption reached the south polar regions in June–July 2022, the polar vortex was too strong for it to penetrate. Measurements of water vapor, ozone, and chemicals involved in destroying ozone all showed near-average amounts and evolution within the vortex. In future years, larger effects on the polar vortex and chemical processing are expected because water vapor from Hunga Tonga that has spread globally will be entrained into the polar vortex.

## 1 Introduction

The 15 January 2022 eruption of the underwater volcano Hunga Tonga-Hunga Ha’apai (HTHH) injected an unprecedented amount of water vapor ( $\text{H}_2\text{O}$ ) directly into the stratosphere, increasing the stratospheric  $\text{H}_2\text{O}$  burden by approximately 10% (e.g., Millán et al., 2022; Vömel et al., 2022). It also resulted in substantial, though not unprecedented, enhancements in volcanic aerosol loading (Khaykin et al., 2022; Sellitto et al., 2022; Taha et al., 2022). Numerous studies have already explored aspects of the stratospheric impacts of HTHH enhancements in aerosol and  $\text{H}_2\text{O}$ ; of particular relevance here are suggestions that  $\text{H}_2\text{O}$  and aerosol from HTHH injected into the Southern Hemisphere (SH) stratosphere took many months to reach high latitudes and did not extend poleward of about  $60^\circ\text{S}$  (e.g., Legras et al., 2022; Khaykin et al., 2022; Schoeberl et al., 2022; Zhu et al., 2022). In the lowermost stratosphere (at and below approximately the 380 K isentropic surface), a few studies suggest that some  $\text{H}_2\text{O}$  and aerosol were transported to high SH latitudes within days to weeks via the shallow branch of the Brewer-Dobson circulation (e.g. Taha et al., 2022; Schoeberl et al., 2022; Khaykin et al., 2022). Radiative cooling from HTHH  $\text{H}_2\text{O}$  led to unprecedented cold in SH mid/low latitudes, with associated circulation and transport anomalies (Coy et al., 2022; Schoeberl et al., 2022; Sellitto et al., 2022).

It was suggested that transport of HTHH aerosol and  $\text{H}_2\text{O}$  into high SH latitudes might impact the composition of the 2022 SH stratospheric polar vortex, and that circulation changes associated with the HTHH  $\text{H}_2\text{O}$  plume might affect the strength, size, and / or longevity of that vortex (e.g., Taha et al., 2022; Zhu et al., 2022). Here we use Aura Microwave Limb Sounder (MLS) data to analyze the evolution of the SH stratospheric polar vortex in 2022, transport of the HTHH  $\text{H}_2\text{O}$  plume in relation to it, and chemical processing within it. We find no evidence of substantial impacts of HTHH on the 2022 SH polar vortex or the chemical processing and ozone loss within it. We use temperature,  $\text{H}_2\text{O}$ ,  $\text{N}_2\text{O}$ , CO, HCl, ClO, and  $\text{O}_3$  from v5 MLS “level 3” (L3)

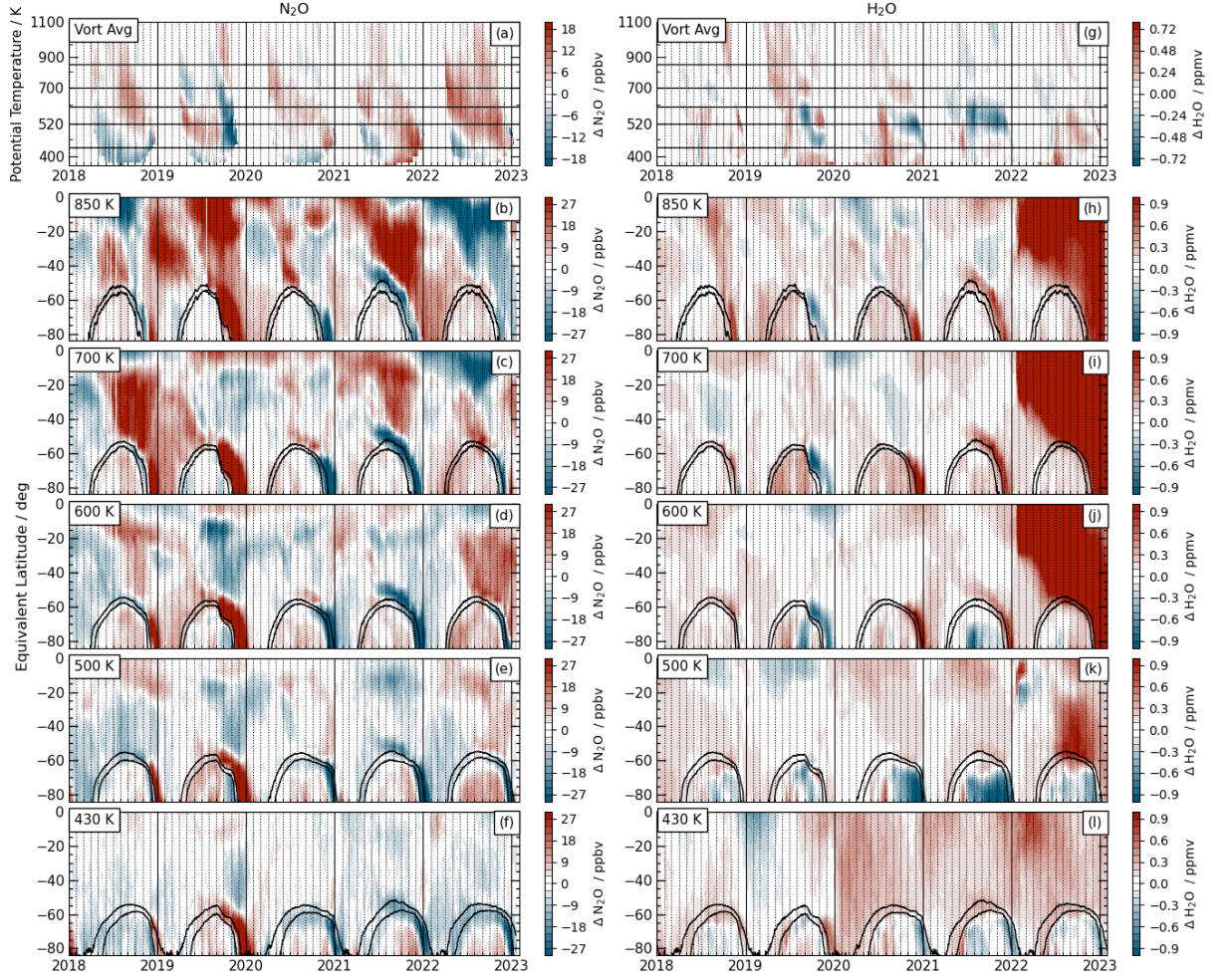
data (Livesey et al., 2020), along with meteorological fields from NASA’s Modern Era Retrospective-analysis for Research and Applications Version 2 (MERRA-2) dataset (Gelaro et al., 2017; Global Modeling and Assimilation Office (GMAO), 2015).

Immediately following the eruption, standard MLS v5 quality screening (Livesey et al., 2020) flagged many of the profiles most affected by HTHH as suspect retrievals (Millán et al., 2022); thus the  $\text{H}_2\text{O}$ ,  $\text{N}_2\text{O}$ , and  $\text{HNO}_3$  anomalies shown here may be artificially small for up to three weeks after the eruption. Since our focus is on the subsequent transport and relationship to the SH polar vortex, our results are unaffected.

## 2 Transport of HTHH Stratospheric $\text{H}_2\text{O}$

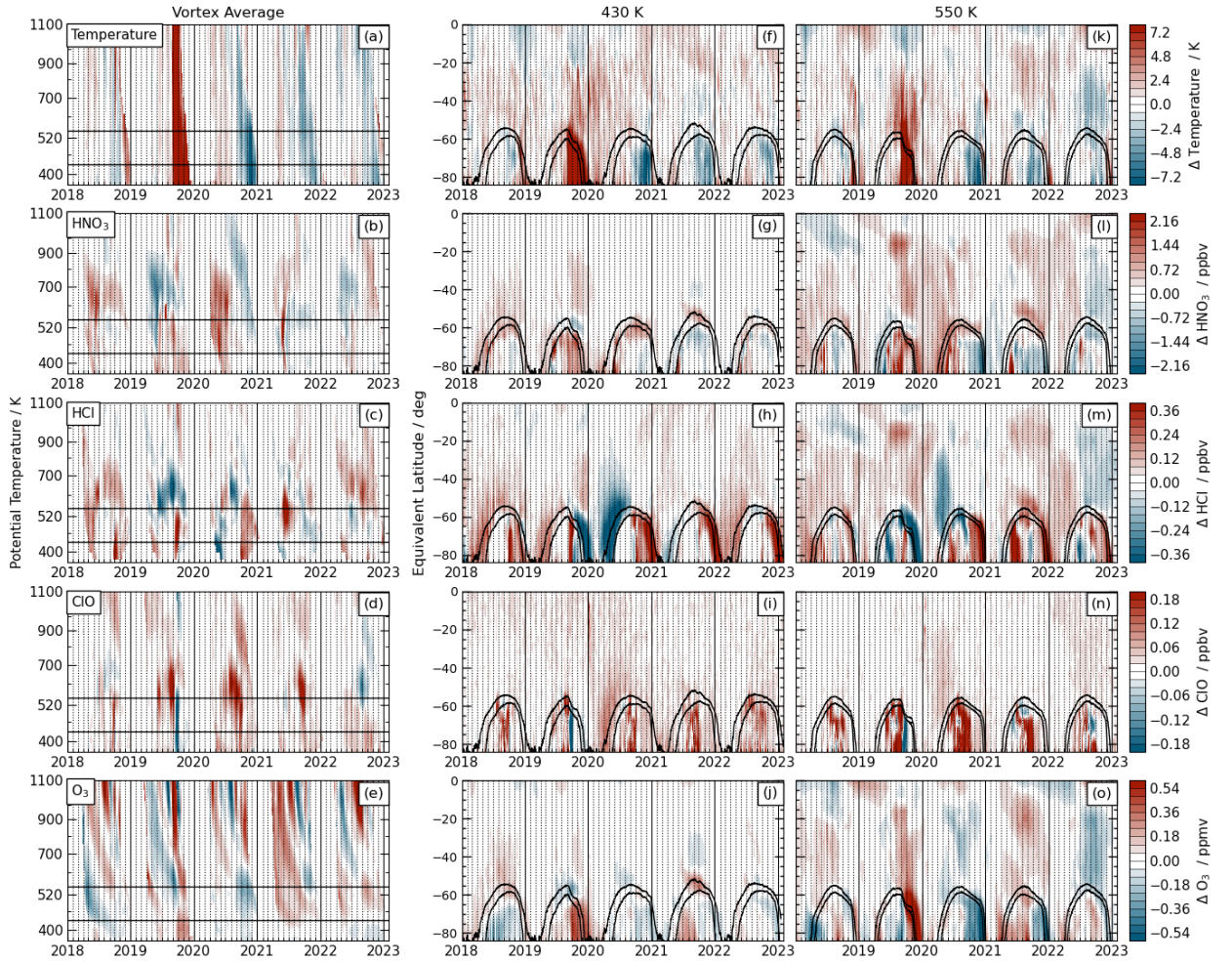
Figure 1 shows the evolution of  $\text{N}_2\text{O}$  and  $\text{H}_2\text{O}$  (both generally long-lived tracers of transport in the stratosphere) anomalies in the SH lower through middle stratosphere, in vortex averages as a function of height (expressed as potential temperature,  $\theta$ ) and as a function of equivalent latitude (EqL, the latitude enclosing the same area between it and the pole as a given potential vorticity, PV, contour, Butchart & Remsberg, 1986) on several isentropic ( $\theta$ ) surfaces. The past five years include seasons with exceptionally warm / short-lived (2019) and cold / long-lived (2020 and 2021) springtime polar vortices, as well as a year (2018) with more typical vortex characteristics (WMO, 2023). (Figs. S1–S2 in the Supporting Information, SI, show the full-mission and include MLS temperature.) The evolution of vortex-average  $\text{N}_2\text{O}$  (Fig. 1a) in 2022 is unexceptional, showing positive anomalies except at the lowest levels; such a vertical dipole pattern of  $\text{N}_2\text{O}$  anomalies is common, with primarily higher values in 2020, 2021, and 2022 consistent with lower vortex temperatures (see below and Figs. S1–3) and accompanying weaker diabatic descent (Fig. S4).  $\text{N}_2\text{O}$  EqL/time evolution (Fig. 1b–f) is also fairly typical; recurring changes above 430 K from high to low anomalies extending from low latitudes show quasi-biennial oscillation (QBO) related transport (e.g., Baldwin et al., 2001; Diallo et al., 2019). Low  $\text{N}_2\text{O}$  anomalies in austral spring 2020 and 2021 are related to the delayed vortex breakup in those years, with low  $\text{N}_2\text{O}$  values remaining confined longer in a more persistent vortex. Spring 2022 shows similar, but weaker, anomalies, suggesting a long-lived vortex. In contrast, high anomalies in 2019 result from a rare SH sudden stratospheric warming that led to a small, warm, and short-lived vortex (e.g., Wargan et al., 2020).

$\text{H}_2\text{O}$  anomalies (Fig. 1g–l) in the SH lower stratospheric vortex are typically dominated by interannual variations in polar stratospheric cloud extent; strong low  $\text{H}_2\text{O}$  anomalies in spring 2020 and 2021 at 500 K and surrounding levels arose from persistent cold anomalies in unusually long-lasting vortices. Outside the vortex (Fig. 1h–l Fig. 1g–l), high  $\text{H}_2\text{O}$  anomalies often accompany low  $\text{N}_2\text{O}$  anomalies because  $\text{H}_2\text{O}$  and  $\text{N}_2\text{O}$  have opposite vertical and horizontal gradients in the lower to middle stratosphere. For example, low (high) springtime  $\text{H}_2\text{O}$  ( $\text{N}_2\text{O}$ ) anomalies just outside the vortex edge in 2019, and opposite patterns in 2020 and 2021 at 600–850 K; similar patterns are seen in mid-EqLs in earlier years (Fig. S1). (Note that typical  $\text{H}_2\text{O}$  anomalies prior to 2022 are washed out by the large colorbar range needed to portray the HTHH  $\text{H}_2\text{O}$ .) Above 500 K, typical signatures of extra-vortex transport of  $\text{H}_2\text{O}$  are overwhelmed by the arrival of HTHH  $\text{H}_2\text{O}$  (Fig. 1h–j, Fig. S1). HTHH  $\text{H}_2\text{O}$  reached the vortex edge in early June 2022, after the vortex was fully developed except in the lowermost stratosphere. Above 500 K, extremely strong gradients along the vortex edge suggest that the HTHH plume could not penetrate the vortex edge. Pervasive high  $\text{H}_2\text{O}$  anomalies since early 2020 below about 500 K may reflect lingering enhancements from the 2020 Australian New Years fires (e.g., Santee et al., 2022). While small positive anomalies encroach into the vortex region in late winter 2022 at 500 K (near the lowest altitude of large HTHH enhancement) and 430 K, similar features are common (e.g., in 2018 and 2021), so it is unclear whether they are related to the HTHH plume. At all levels examined (including the lowermost stratosphere, e.g., Fig. S3),  $\text{H}_2\text{O}$  anomalies inside the vortex are within the typical range.



**Figure 1.** Evolution of MLS-observed SH anomalies from the baseline 2005–2021 climatology of  $\text{N}_2\text{O}$  (a–f) and  $\text{H}_2\text{O}$  (g–l) from January 2018 through January 2023: (a,g) vortex-averaged values; (b–f, h–l) evolution as a function of EqL at levels in the middle through lower stratosphere (horizontal lines in a,g). Black contours in b–f and h–l are sPV values indicating the vortex edge region.





**Figure 2.** As in Fig. 1, but for MLS temperature,  $\text{HNO}_3$ ,  $\text{HCl}$ ,  $\text{ClO}$ , and  $\text{O}_3$ ; (a–e) vortex averages, (f–j) 430 K, and (k–o) 550 K EqL timeseries, for January 2018 through January 2023. Black contours (f–o) are sPV values demarking the vortex edge region.

### 3 Polar Vortex Composition and Chemical Processing

Figure 2 shows a similar view of MLS measurements of temperature and species involved in polar chemical processing (Figs. S1–3 show 550 K, 430 K, and 380 K for the full mission). The Antarctic vortex was unusually cold and persistent in spring 2022, but less so than in 2020 and 2021. Vortex  $\text{HNO}_3$  values were near average throughout the season. Vortex  $\text{HCl}$  and  $\text{ClO}$  commonly oscillate between high and low anomalies, and thus they are also generally unexceptional within the 2022 vortex; the high  $\text{HCl}$  anomalies in spring are related primarily to longer-than-usual confinement of the very high values that ensue from chlorine deactivation. Consistent with near-average vortex values of chlorine species,  $\text{O}_3$  anomalies in 2022 were also relatively small. Both 2020 and 2021 showed lower  $\text{O}_3$ , consistent with larger cold anomalies and even longer-lived (see below) vortices in those years than in 2022. Outside the vortex, temperature anomalies (arising from radiative effects of HTHH  $\text{H}_2\text{O}$ , e.g., Coy et al., 2022; Schoeberl et al., 2022) and associated mid-latitude transport anomalies (Coy et al., 2022) appear consistent with the extravortex high  $\text{N}_2\text{O}$  anomalies seen near 500–600 K (Fig. 1), and suggest that accompanying extravortex  $\text{HCl}$ ,  $\text{HNO}_3$ , and  $\text{O}_3$  anomalies are at least partially transport-driven.

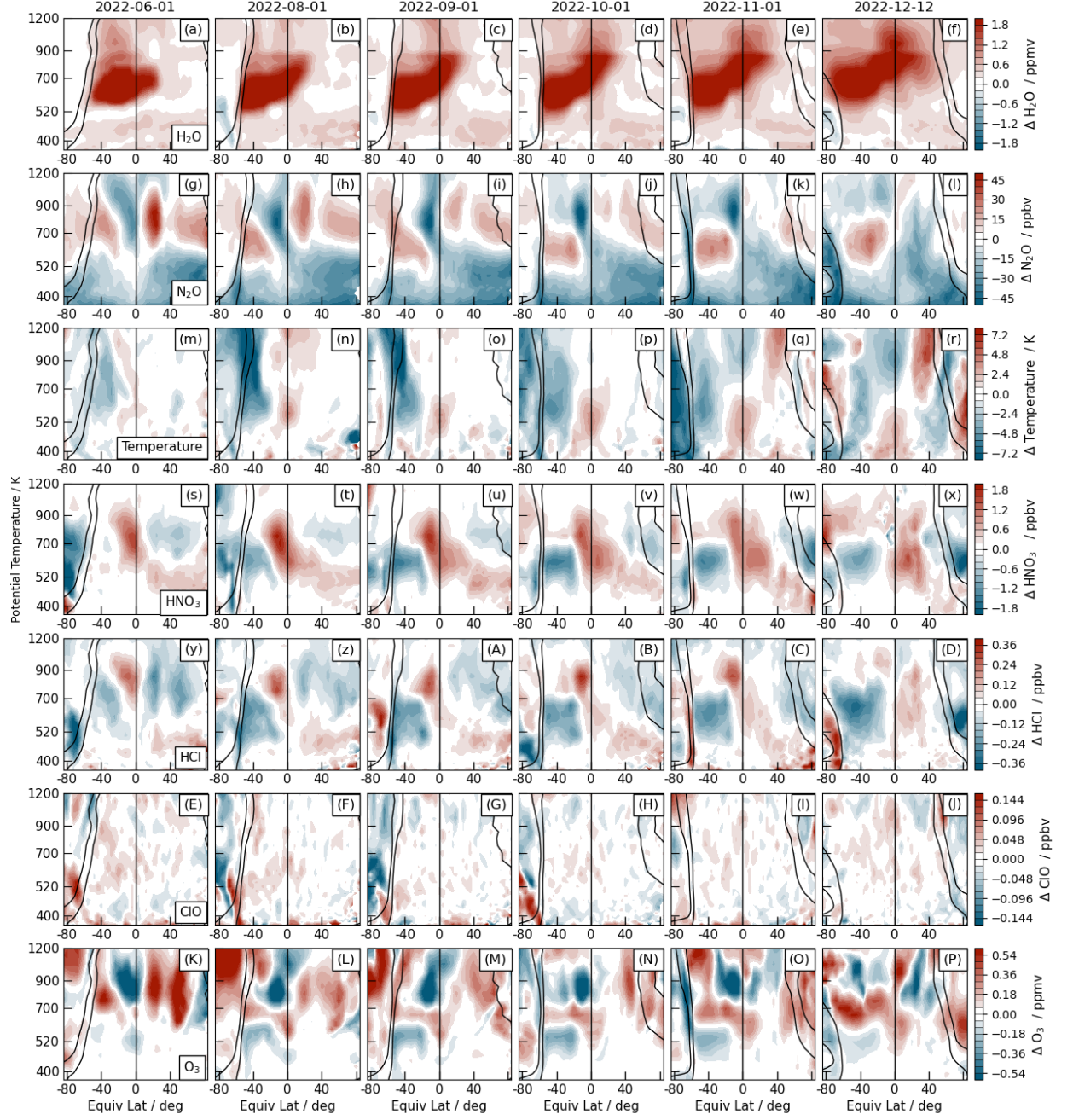
Figure 3 provides a closer look at the  $\text{EqL}/\theta$  evolution of MLS trace gases in 2022, showing snapshots of anomalies from climatology (similar anomaly plots in 2018, 2020, and 2021 are shown in Figs. S5–S7). The  $\text{H}_2\text{O}$  plume first approached the SH polar vortex edge in early to mid-June. Subsequently, extremely strong  $\text{H}_2\text{O}$  gradients developed along the vortex edge over 520–800 K and persisted through October (into December below about 700 K; see also Fig 1). By mid-December, only a weak remnant of the vortex remained below about 520 K, and the  $\text{H}_2\text{O}$  enhancement extended into high latitudes above that level. MLS data show no indication of air from the HTHH  $\text{H}_2\text{O}$  plume penetrating substantially into the SH vortex before its breakup.  $\text{N}_2\text{O}$  anomalies within the vortex were generally small until austral spring; below about 700 K, these anomalies were near zero from August through October. Low  $\text{N}_2\text{O}$  anomalies along the vortex edge beginning in early November are consistent with confinement in an unusually persistent vortex. Mid-latitude cold anomalies throughout the middle stratosphere (e.g., Coy et al., 2022; Schoeberl et al., 2022) are apparent from June through mid-December. Vortex temperatures were below average through much of the season, with largest cold anomalies in October and November (also see Fig. 2). High extra-vortex  $\text{N}_2\text{O}$  anomalies through this period are consistent in extent and location with the circulation anomalies reported by Coy et al. (2022). The co-location of  $\text{N}_2\text{O}$  anomalies with those in  $\text{HNO}_3$ ,  $\text{HCl}$ , and  $\text{O}_3$  suggests that transport plays a role in all of them; work is in progress analyzing the relative effects of dynamical and chemical processes.

Within the vortex,  $\text{HNO}_3$  is slightly lower than usual, consistent with a colder-than-average vortex.  $\text{HCl}$  ( $\text{ClO}$ ) shows low (high) anomalies during much (but not all, e.g., Fig. 3A,G) of the winter. As noted above, high  $\text{HCl}$  anomalies appear along the vortex edge in November and in the vortex remnant in mid-December, consistent with high values resulting from deactivation into  $\text{HCl}$  (as is typical in the SH, e.g., Santee et al., 2008) followed by unusually enduring confinement in the persistent vortex. Lower stratospheric  $\text{O}_3$  anomalies in the early winter (before extensive chemical loss) are slightly positive and remain so through October (e.g., Fig. 3O). Taken together, the results in Figs. 2 and 3 suggest that the modest low anomalies in  $\text{O}_3$  seen in austral spring 2022 (e.g., Fig. 3P) result primarily (if not entirely) from the unusual persistence of the vortex.

#### 4 Vortex Evolution and Trace Gas Confinement

Figure 4 summarizes the evolution of the 2022 SH vortex in the context of the 43-year MERRA-2 record and the evolution of trace gases in the context of the 18-year MLS record, both in relation to the previous three SH winters. Figure S8 shows profiles of additional MERRA-2 diagnostics of vortex strength and longevity. Consistent with the indications in the trace gases of its unusual persistence, the 2022 SH late winter and spring vortex was among the largest on record at levels up to about 650 K, approximately matching the maximum size and persistence seen prior to 2020 (Fig. 4a–d; Fig. S8b,d). In spring, the 2021 vortex area was slightly larger, and the 2020 vortex area substantially larger than that in 2022 from about 460 K to 650 K, with 2020 setting the record for lower-stratospheric vortex persistence (Fig. 4a–c, S8b–d). Maximum PV gradients, indicating vortex strength (that is, robustness as a transport barrier), show unusually strong springtime vortices in 2020 through 2022 below about 500 K, but only the 2020 vortex was stronger than average above about 600 K (Fig. 4e–h; Fig. S8a). Below about 520 K, the area with temperatures below the nitric acid trihydrate (NAT) and ice polar stratospheric cloud (PSC) thresholds was larger than usual (Fig 4m,n,q,r) and PSCs persisted later than usual (Fig. 4m–t, Fig. S8e,f) in spring 2020, 2021, and 2022, but only exceeded previous springtime records in 2020; above about 600 K PSC area and duration were near average.

The unexceptional MLS trace gas evolution in the 2022 Antarctic vortex is highlighted in Fig. 4A–P (Fig. S9 shows the vertical structure). Interannual variability in SH polar chemical processing is relatively small, but, with few exceptions, all of the trace gases show 2022 evolution that is well within the previously observed range. Over  $\sim 450$ –600 K, persistently low  $\text{H}_2\text{O}$  after October in 2022, and to an even greater extent in 2020 and 2021, is consistent with confinement of dehydrated air in long-lived vortices. Chlorine evolution (seen in  $\text{HCl}$  and  $\text{ClO}$ , Fig. 4E–



**Figure 3.** Snapshots on selected days in 2022 of anomalies from the baseline 2005–2021 climatology of MLS (a–f)  $\text{H}_2\text{O}$ , (g–l)  $\text{N}_2\text{O}$ , (m–r) temperature, (s–x)  $\text{HNO}_3$ , (y–D)  $\text{HCl}$ , (E–J)  $\text{ClO}$ , and (K–P)  $\text{O}_3$ . Black contours show sPV values demarcating the vortex edge region.



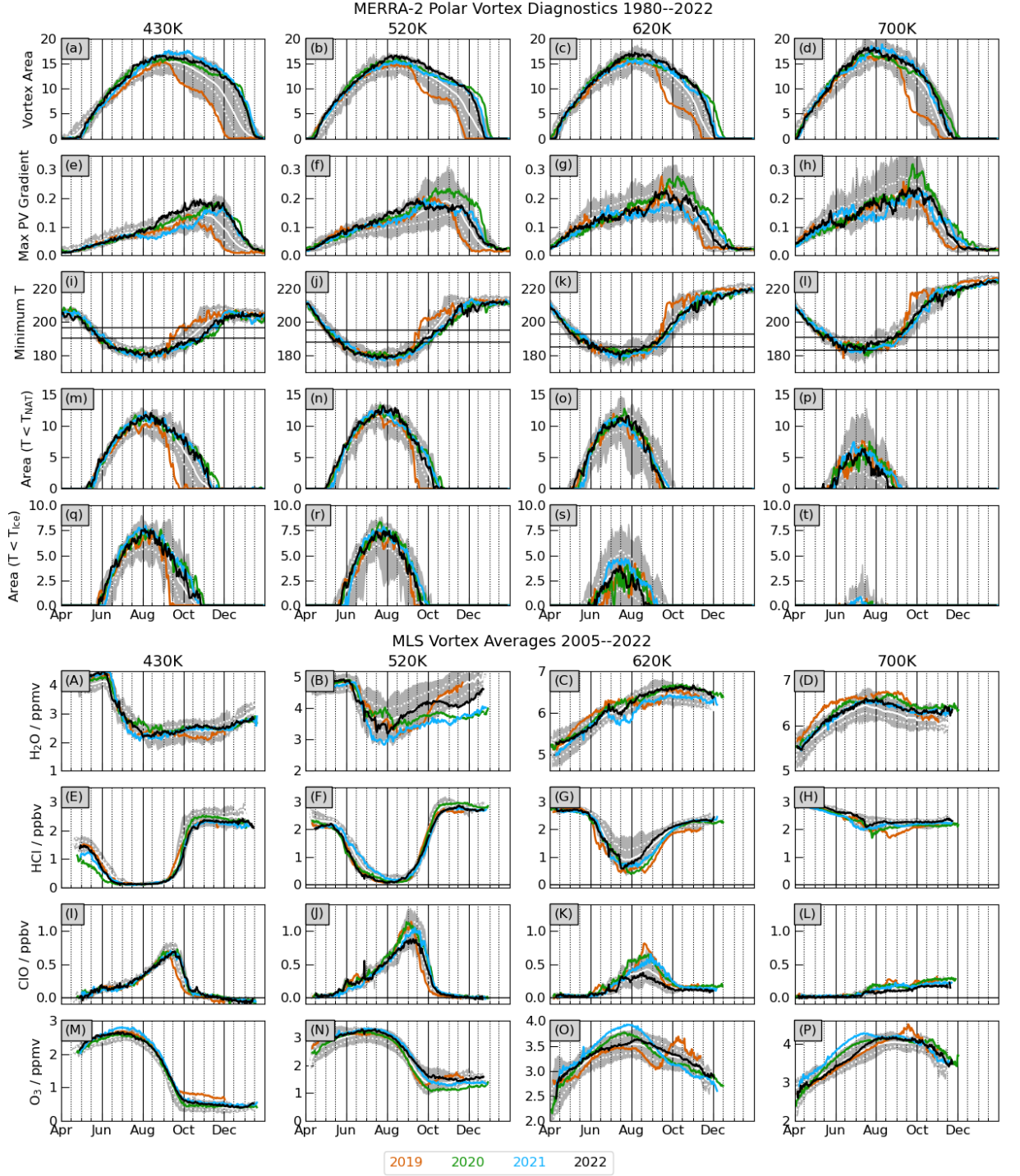
L; Fig. S9q–x) was fairly typical throughout the season. Observed O<sub>3</sub> evolution in 2022 was remarkably near average throughout the season (Fig. 4M–P; Fig. S9y–B).

The above results provide visual evidence that the vortex edge presented an effective transport barrier, preventing substantial penetration of the HTHH H<sub>2</sub>O plume. To look more closely at the robustness of the vortex edge transport barrier, Fig. 5 shows scatter and density plots of H<sub>2</sub>O versus N<sub>2</sub>O and sPV for representative days in 2022 compared with the evolution in all prior years in the MLS record. Low N<sub>2</sub>O (relative to the range of values at a given level) and high-magnitude sPV identify vortex air parcels. In the lower stratosphere (exemplified by 550 K), increasingly low vortex H<sub>2</sub>O through the season results from dehydration and is very similar to that previously observed by MLS (density plots, right columns, emphasize the similarity of the main distributions in 2022 to those in earlier years). Extravortex H<sub>2</sub>O at 550 K does not stand out from the previous record before July, but after that the HTHH enhancement manifests as a distinct cluster of high H<sub>2</sub>O with N<sub>2</sub>O near 200 ppbv and sPV magnitudes  $< 1 \times 10^{-4} \text{ s}^{-1}$  (both values that are unambiguously extravortex) that is unique to 2022 (compare yellow-orange / purple H<sub>2</sub>O / sPV values with grey dots; orange with grey contours). In the middle stratosphere (exemplified by 700 K), vortex H<sub>2</sub>O values first increase via descent of the upper stratospheric peak, then decrease as continuing descent brings low mesospheric H<sub>2</sub>O into the stratospheric vortex (e.g., Ray et al., 2002; Lee et al., 2011); both the high (e.g., Fig. 5a–d) and the low (e.g., Fig. 5e–l) H<sub>2</sub>O values that descend through the vortex (low N<sub>2</sub>O, high-magnitude sPV end of the x-axis) at 700 K are distinct from the extravortex population of high H<sub>2</sub>O from HTHH, and that is in turn distinguished from extravortex air in previous years by higher H<sub>2</sub>O values at extravortex N<sub>2</sub>O (~150–200 ppbv) and sPV (magnitude  $< \sim 1 \times 10^{-4} \text{ s}^{-1}$ ). These correlations of H<sub>2</sub>O with N<sub>2</sub>O and sPV (especially the density plots versus sPV) show clearly that the air with enhanced H<sub>2</sub>O from HTHH remained well separated from that within the vortex until vortex breakup at each level (as suggested in Figs. 1 and 3). MLS H<sub>2</sub>O / CO correlations show a similar picture in the middle (Fig. S10) and upper stratosphere, with HTHH H<sub>2</sub>O associated with low CO values characteristic of extravortex air. Further, because the seawater from HTHH has a higher ratio of HDO to H<sub>2</sub>O than background water vapor in the extravortex stratosphere (e.g., Randel et al., 2012; Khaykin et al., 2022), an unprecedented increase in that ratio in SH midlatitudes also marks the HTHH air as separate from (and excluded from) that in the vortex (Figs. S11–12).

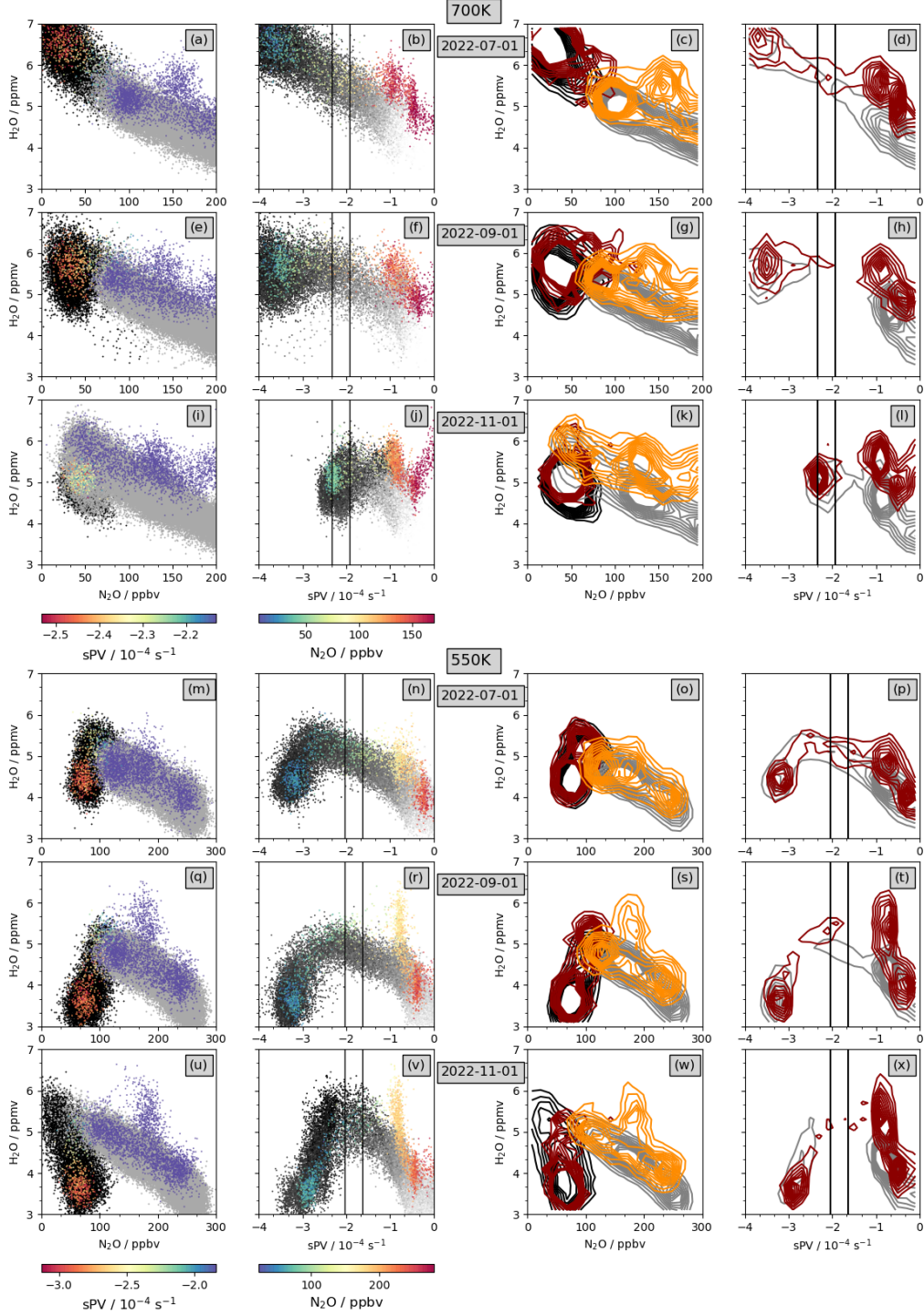
## 5 Summary

The unprecedented water vapor injection into the stratosphere by HTHH is tracked using MLS and reanalysis data. The H<sub>2</sub>O plume [for The enhanced H<sub>2</sub>O]] is shown to have been effectively excluded from the 2022 Antarctic polar vortex until the vortex breakdown. In contrast to speculation that HTHH stratospheric H<sub>2</sub>O and aerosol injections would lead to substantial anomalies in the Antarctic polar vortex and lower stratospheric polar processing and ozone loss within it (e.g., Taha et al., 2022; Zhu et al., 2022), our analysis suggests that HTHH did not cause substantial changes in polar processing and ozone loss within the vortex: MLS observations of HNO<sub>3</sub>, HCl, ClO, and O<sub>3</sub> inside the vortex through the depth of the lower stratosphere all show evolution well within the range of previous years during the MLS mission, with near-average O<sub>3</sub> loss. Evidence for possible dynamical impacts on the vortex is likewise not unequivocal: The vortex was among the larger, stronger, and longer-lived in the SH lower stratosphere, but these conditions were matched or exceeded by those in 2020, 2021, and several previous years in the MERRA-2 record since 1980; vortex cold anomalies were even less exceptional. Thus, despite large radiative, dynamical, and composition perturbations in midlatitudes, the observational evidence shows that chemical processing within the 2022 Antarctic stratospheric polar vortex was fairly typical, and does not show clear evidence of substantial dynamical vortex perturbations. The dispersal of HTHH H<sub>2</sub>O following the Antarctic vortex breakup (e.g., Fig. 1) led to unprecedented high H<sub>2</sub>O anomalies throughout the SH, which are expected to linger for at least several years (e.g., Millán et al., 2022; Khaykin et al., 2022), raising the expectation of large perturbations to Antarctic polar vortex chemistry and the ozone hole in 2023 and beyond. HTHH H<sub>2</sub>O has also been transported into the Northern Hemisphere (e.g., Schoeberl et al., 2023), but reached the Arctic vor-





**Figure 4.** (a–t) Time series at four levels in the lower to middle stratosphere of vortex area, maximum PV gradients, high latitude (poleward of  $30^\circ$ ) minimum temperature, and area below NAT and ice PSC thresholds, comparing 2019 (orange), 2020 (green), 2021 (cyan), and 2022 (black) with the range (shading), mean (solid white line), and one standard deviation envelope (dotted white lines) over 1980–2018. (A–P) Vortex-averaged  $\text{H}_2\text{O}$ ,  $\text{HCl}$ ,  $\text{ClO}$ , and  $\text{O}_3$  in same format as for the dynamical fields, with the range over 2005–2018.



**Figure 5.** Scatter (left two columns) and density (right two columns) plots of MLS H<sub>2</sub>O (y-axis) versus N<sub>2</sub>O (first and third columns) and sPV (second and last columns). Grey and black dots (contours) show values from 2005–2021 in the scatter (density) plots; for those years, black (grey) indicates x-axis values of N<sub>2</sub>O or sPV characteristic of inside (outside) the vortex. For 2022, colored (purple) dots or dark red (orange) contours show sPV values inside (outside) the vortex. 2022 N<sub>2</sub>O (second column) is colored such that blue/blue-green shows typical vortex values. Black vertical lines on the plots versus sPV indicate the vortex edge region.

tex edge after the vortex was well-developed and was only dispersed through the NH after a strong sudden stratospheric warming starting in mid-February (paper in preparation). Thus large effects on Arctic polar vortex chemistry are also expected to manifest starting in the 2023/2024 cool season.

## 6 Open Research

The data used herein are publicly available as follows:

- MERRA-2: (Global Modeling and Assimilation Office (GMAO), 2015)  
<https://disc.sci.gsfc.nasa.gov/uui/datasets?keywords=%22MERRA-2%22>
- Aura MLS Level-2 and Level-3 data: (Lambert, Read, & Livesey, 2020; Lambert, Livesey, & Read, 2020; Lambert et al., 2021b, 2021a; Schwartz, Pumphrey, et al., 2020; Schwartz, Froidevaux, et al., 2020; Schwartz, Pumphrey, et al., 2021; Schwartz, Froidevaux, et al., 2021)  
<https://disc.gsfc.nasa.gov/datasets?page=1&keywords=AURA%20MLS>
- ACE-FTS v4.1/4.2 data: <http://www.ace.uwaterloo.ca> (registration required)
- ACE-FTS v4.1/4.2 error flags: <https://dataverse.scholarsportal.info/api/access/dataset/:persistentId/versions/:latest?persistentId=doi:10.5683/SP2/BC4ATC>
- MLS & ACE-FTS derived meteorological products: <https://mls.jpl.nasa.gov/eos-aura-mls/dmp> (registration required).

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