

1                   **AVA Inversion of PP Reflection in a VTI Medium using**  
2                   **Proximal Splitting Algorithm**

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4                   **Key Points:**

- 5                   • To alleviate the ill-posedness of the seismic amplitude inversion, and obtain a unique and  
6                   stable solution.

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## 7 Abstract

8 Amplitude variation with offset (AVO) inversion, particularly for more than two model parameters,  
 9 is a highly ill-posed problem and, hence, regularization is indispensable. Here, we propose a regularized  
 10 inverse problem to mitigate the ill-posedness of the amplitude inversion. The regularization is  
 11 added to measure the difference in information between the a priori probability density function  
 12 and the predicted probability density of the inverted parameters. Information theory provides a  
 13 collection of contrast functions which quantify the divergence from one probability distribution  
 14 to another, such as the relative entropy. The a priori density is approximated by a Gaussian mixture  
 15 model, obtained from well logs and rock physics model. The mixture model is a density estimator,  
 16 providing the statistical properties of the model parameters of interest. The likelihood of the data  
 17 and the divergence are combined in an augmented Lagrangian scheme, the alternating direction  
 18 method of multipliers (ADMM), to obtain a unique solution that best generate the recorded seismic  
 19 data and satisfy the geological constraints conveyed by the a priori probability density function.  
 20 The proposed inversion scheme is then applied to the anisotropy AVO inversion, for estimating  
 21 the elastic and seismic anisotropy parameters of shale formations. Compared to the unconstrained  
 22 minimization, the P- and S-wave velocity, and  $\varepsilon$  are better recovered, moreover, density and Thomsen's  
 23  $\delta$  are well-constrained.

## 24 Plain Language Summary

25 Geophysical inverse problems are highly ill-posed, due to variable sensitivities of different  
 26 model parameters, measurement errors, among other reason. To alleviate ill-posedness of the seismic  
 27 amplitude inversion, a regularization function that measures the distance to a priori mixture model  
 28 is proposed. The a priori mixture model is first obtained from well log and rock physics data, independently  
 29 of the seismic data. The a priori model conveys the statistical properties of the parameters of interest.  
 30 Then, the functional and regularization are combined via a proximal splitting scheme, to obtain  
 31 a unique solution that is close to the a priori mixture model and best generate the observed seismic  
 32 data.

## 33 1 Introduction

34 Seismic amplitude inversion has been successfully applied to predict the elastic properties  
 35 of the subsurface (Hampson et al., 2005; Russell & Hampson, 2006); jointly with rock physics  
 36 inversion, models for reservoir lithology and fluid properties are built (Bosch et al., 2010; Grana  
 37 & Della Rossa, 2010). In a strong vertical transverse isotropy (VTI) medium, i.e., shale, isotropic  
 38 approximations of the Zoeppritz equations fail to predict the correct AVO responses, which leads  
 39 to incorrect lithology and fluid predictions, and fail to accurately describe the geomechanical behavior  
 40 of the shale rocks (Sayers, 2013b). The anisotropy of shale is attributed to preferential orientation  
 41 of clay minerals, micro-cracks, organic matter, amongst others, hence, anisotropy exists at all scale  
 42 lengths (Sayers, 2013a). Since there is no direct method to measure anisotropy in vertical wells,  
 43 rock physics modeling (RPM) is used to estimate the low-to-medium frequency anisotropy parameters  
 44 necessary to initialize the seismic amplitude inversion for the lithology and fluid predictions. Bandyopadhyay  
 45 (2009) showed that, increasing the volume fraction of kerogen increases the rock stiffnesses, and  
 46 subsequently the elastic anisotropy. Consequently, the estimated seismic anisotropy parameters  
 47 are of great relevance for the organic shale exploration and production.

48 Daley & Hron (1977) gave the exact solution for reflection and transmission coefficients  
 49 in a transversely isotropic medium. Many linear approximations have been subsequently introduced  
 50 assuming weak elastic contrasts (Thomsen et al., 1993; Rüger, 1997). Despite a vertical transverse  
 51 medium (VTI) is described by the Thomsen's parameters  $\varepsilon$ ,  $\delta$ , and  $\gamma$  (Thomsen, 1986), the  
 52 P-P reflection coefficients in a VTI medium depends only on  $\varepsilon$  and  $\delta$  (Rüger, 1997). The effect  
 53 of the NMO anisotropy  $\delta$  appears at the small-to-middle angles of incidence (Banik, 1987), while  
 54 P-wave anisotropy  $\varepsilon$  influences larger angles of incidence greater than  $30^\circ$  (Kim et al., 1993). Plessix  
 55 & Bork (2001) studied the AVO response in a VTI medium based on the least-squares function;

56 showing that, it is difficult to obtain all five parameters for small-to-medium angles of incidence  
 57 due to the existence of local minimum solutions. F. Zhang et al. (2019) used implicit constraints  
 58 to derive an approximation of the P-P reflection coefficients in a VTI medium, consisting of only  
 59 three parameters. In addition, joint P-P and P-S amplitude inversion of a VTI medium has been  
 60 proposed to better constrain the density and seismic anisotropy parameters (Nadri & Hartley, 2007;  
 61 Lu et al., 2018; Luo et al., 2020; Zhou et al., 2020).

62 Seismic amplitude inversion is highly ill-posed due to variable-sensitivities among model  
 63 parameters; moreover, a small amount of noise in the observed seismic data results in a very large  
 64 change in the estimates (Tarantola, 2005). Furthermore, the density and Thomsen's anisotropy  
 65 inversions are very challenging due to less information in the near-angle traces, whereas the far-angle  
 66 traces are typically distorted (non-flatten) (L. Liu et al., 2013). To quote Treitel & Lines (2001),  
 67 "A good match between the observed and theoretical geophysical responses provides us with a  
 68 necessary, but by no means sufficient condition for the calculation to converge to the *ground truth*  
 69 below us.", a solution to ill-posed problems is to add regularization that enhances the stability  
 70 of the inversion, such as  $L_2$ -norm promoting smoothness in the solution (Velis, 2008), and  $L_1$ -  
 71 norm to promote sparsity in the solution (Y. Wang, 2010) and as an anti-noise functional (C. Liu  
 72 et al., 2015). Total variation ( $TV$ -norm) has been successfully used to promote sparsity of the  
 73 reflection coefficients (F. Zhang et al., 2014) and acoustic impedance (Wu, He, et al., 2019). Zand  
 74 et al. (2020) used the  $TV$  - norm regularization to obtain a stable solution of the least-square  
 75 reverse-time migration (LSRTM). Despite the norm regularization functions to provide a desired  
 76 shape of the solution, it can not infer the statistical information about the model parameters of  
 77 interest. Different geoscience data follow different distributions, such as the P- and S-wave velocity  
 78 follow a Gaussian distribution (Hernlund & Houser, 2008) and a Lévy-stable distribution (Painter  
 79 & Paterson, 1994), the P- and S-wave reflection coefficients follow a Lévy-stable distribution (Painter  
 80 et al., 1995) and a Gaussian mixture distribution (Mukerji et al., 2009), and the porosity follows  
 81 a log-normal distribution (Berezin, 1982). Accordingly, the Bayesian inference uses a prior probability  
 82 distribution that infer the statistical properties of the model parameters to obtain the posterior distribution  
 83 of the unknown model parameters, using deterministic approaches (Downton & Lines, 2001; Buland  
 84 & Omre, 2003; Downton, 2005; Grana, 2020), or stochastic methods (D. Zhu & Gibson, 2018;  
 85 Wu, Li, et al., 2019; K. Li et al., 2020).

86 Zidan (2022) proposed a generalized framework for Bayesian inversion using a lower-bound  
 87 estimate of the statistical information of the model parameters of interest, along with a Bregman  
 88 divergence to regularize the amplitude inversion by measuring the distance to the a priori model.  
 89 However, different rock facies and fluids saturations result in different elastic and petrophysical  
 90 properties, and subsequently lead to multimodal behavior of the marginal and joint distributions  
 91 of the model parameters of interest. Examples of such problems are petrophysical inversion of  
 92 seismic data (Connolly & Hughes, 2016; de Figueiredo et al., 2017, 2018a; Kolbjørnsen et al.,  
 93 2020), and anisotropy AVO inversion (Lu et al., 2018; Zhou et al., 2020), where more parameters  
 94 must be estimated. To this end, a prior mixture model is necessary to infer the correct statistical  
 95 information about the parameters of interest. Additionally, the mixture models provide good fit  
 96 to heavy-tailed data (Mukerji et al., 2009; Grana & Bronston, 2015). Grana & Della Rossa (2010)  
 97 proposed a Gaussian mixture model for the litho-fluid classes. Grana et al. (2017) obtained an  
 98 analytical solution of the posterior distribution using a prior Gaussian mixture model. Fjeldstad  
 99 & Grana (2018) used Markov chain Monte Carlo (McMC) to sample from a Gaussian mixture  
 100 posterior density for predicting lithology-fluid classes. de Figueiredo et al. (2019) jointly inverted  
 101 for facies and elastic properties, using a Gaussian mixture prior density. Consequently, a single-component  
 102 density function is incompetent to represent the model space, and subsequently infer the wrong  
 103 statistical information about the model parameters.

104 To quote Goodfellow et al. (2016), "The basic intuition behind information theory is that  
 105 learning that an unlikely event has occurred is more informative than learning that a likely event  
 106 has occurred.", we can calculate the amount of information conveyed by the inverted parameters  
 107 using the relative entropy (Kullback & Leibler, 1951). Here, we propose a regularized inverse  
 108 problem that holds for mixture probability density functions. The regularization is based on  $f$ -divergences,

109 which measure the difference between two probability distributions, i.e., the a priori distribution  
 110 and predicted distribution. Firstly, an inverse rock physics (IRPM) problem is solved using well  
 111 log and seismic data for estimating the P-wave anisotropy  $\varepsilon$  and normal-moveout (NMO) anisotropy  
 112  $\delta$  parameters, in order to fully cover the low-to-medium frequency gap of the limited-bandwidth  
 113 seismic data (Zidan et al., 2021). Then, a priori joint distribution is obtained using all five parameters  
 114  $[V_p, V_s, \rho, \delta, \varepsilon]$ . The mixture model contains the statistical information about the model parameters  
 115 of different litho-facies. Next, the anisotropy AVO modeling based on Rüger's P-P reflection coefficients  
 116 in a VTI medium is performed. A regularized inverse problem that measures the residual between  
 117 the amplitudes of observed and synthetic seismic angle gathers, and the divergence of the predicted  
 118 probability density from the a priori probability density is constructed. Subsequently, the alternating  
 119 direction of multipliers method (ADMM) is used to minimize the functional and regularization  
 120 alternately. The performance of the proposed approach is subsequently tested on synthetic and  
 121 real seismic angle gathers. The regularized optimization successfully constrains the low-sensitivity  
 122 density and  $\delta$  parameters, and a better recovery of the elastic velocities and  $\varepsilon$ .

## 123 Methodology

### 124 1.1 Anisotropic AVO modeling

Reflection amplitudes are parameterized by the angles of incidence and elastic properties  
 of rock layers across the interface. Assuming incident plane waves at an interface between two  
 VTI media, the energy partitioning is described by Daley & Hron (1977). Various weak-contrast  
 approximations have been introduced, including Ursin & Haugen (1996); Vavryčuk (1999); Shaw  
 & Sen (2004). Thomsen et al. (1993) derived a linear approximation of the PP reflection coefficients  
 for weak anisotropy VTI media:

$$\begin{aligned}
 R_{PP}^{VTI}(\theta) &= \frac{1}{2} \left[ \frac{\Delta Z_0}{\bar{Z}_0} \right] \\
 &+ \frac{1}{2} \left[ \frac{\Delta V_{p0}}{\bar{V}_{p0}} - \left( \frac{2\bar{V}_{s0}}{\bar{V}_{p0}} \right) \frac{\Delta G_0}{\bar{G}_0} + (\delta_2 - \delta_1) \right] \sin^2 \theta \\
 &+ \frac{1}{2} \left[ \frac{\Delta V_{p0}}{\bar{V}_{p0}} - (\delta_2 - \delta_1 - \varepsilon_2 + \varepsilon_1) \right] \tan^2 \theta \sin^2 \theta,
 \end{aligned} \tag{1}$$

where  $Z_0 = \rho V_{p0}$  and  $G_0 = \rho V_{s0}^2$  are the vertical P-wave impedance and vertical S-wave modulus  
 (Castagna & Backus, 1993). The  $\Delta\delta$  appears on both  $\sin^2 \theta$  and  $\tan^2 \theta \sin^2 \theta$  terms, resulting  
 in inaccurate reflection coefficients for large angles of incidence ( $> 20^\circ$ ). Moreover, for angles  
 larger than  $45^\circ$ , the  $\tan^2 \theta \sin^2 \theta$  dominates the AVO-gradient term, and equation 1 breaks down  
 (Rüger, 2002). Rüger (1997) decomposed the reflection coefficients into isotropic and anisotropic  
 terms:

$$R_{PP}^{VTI}(\theta) = R_{PP}^{iso}(\theta) + R_{PP}^{ani}(\theta), \tag{2}$$

where,

$$\begin{aligned}
 R_{PP}^{iso}(\theta) &= \frac{1}{2} \left[ \frac{\Delta Z_0}{\bar{Z}_0} \right] \\
 &+ \frac{1}{2} \left[ \frac{\Delta V_{p0}}{\bar{V}_{p0}} - \left( \frac{2\bar{V}_{s0}}{\bar{V}_{p0}} \right) \frac{\Delta G_0}{\bar{G}_0} \right] \sin^2 \theta \\
 &+ \frac{1}{2} \left[ \frac{\Delta V_{p0}}{\bar{V}_{p0}} \right] \tan^2 \theta \sin^2 \theta,
 \end{aligned} \tag{3}$$

$$R_{PP}^{ani}(\theta) = \frac{1}{2} \Delta\delta \sin^2 \theta + \frac{1}{2} \Delta\varepsilon \tan^2 \theta \sin^2 \theta. \tag{4}$$

In the above equation,  $\Delta\delta$  controls the small-angle reflection coefficients through the  $\sin^2 \theta$  term, whereas  $\Delta\varepsilon$  controls the large-angle reflection coefficients with the  $\tan^2 \theta \sin^2 \theta$  term. Consequently, equation 2 is more stable for large angles of incidence. Figure 1a shows comparison of the exact isotropic and exact anisotropic solutions in a VTI medium, the difference between the two solutions particularly for the mid-to-far angle of incidence may lead to the wrong lithology and fluid content. Equation 2 is extended to a time-continuous reflectivity function reads per Stolt & Weglein (1985):

$$R_{PP}^{VTI}(\theta) = a_{V_p}(\theta) \frac{\partial \ln V_p}{\partial t} + a_{V_s}(\theta) \frac{\partial \ln V_s}{\partial t} + a_\rho(\theta) \frac{\partial \ln \rho}{\partial t} + a_\delta(\theta) \frac{\partial \delta}{\partial t} + a_\varepsilon(\theta) \frac{\partial \varepsilon}{\partial t}, \quad (5)$$

where,

$$a_{V_p}(\theta) = \frac{1}{2}(1 + \sin^2 \theta + \sin^2 \theta \tan^2 \theta), \quad (6)$$

$$a_{V_s}(\theta) = -\left(\frac{2\bar{V}_s(t)}{\bar{V}_p(t)}\right)^2 \sin^2 \theta, \quad (7)$$

$$a_\rho(\theta) = \frac{1}{2}\left(1 - \left(\frac{2\bar{V}_s(t)}{\bar{V}_p(t)}\right)^2 \sin^2 \theta\right), \quad (8)$$

$$a_\delta(\theta) = \frac{1}{2} \sin^2 \theta, \quad (9)$$

$$a_\varepsilon(\theta) = \frac{1}{2} \sin^2 \theta \tan^2 \theta. \quad (10)$$

125 In the above equation, accurate estimation of the background  $\bar{V}_s/\bar{V}_p$  is important, particularly for  
 126 relatively strong elastic contrasts (Alemie, 2010); in practice, it is usually taken as a constant calculated  
 127 from well logs. Figure 1b shows comparison between the Daley & Hron (1977) and the R uger  
 128 (1997) solutions in a VTI medium. In spite of the good fit between the two solutions, equation 5  
 129 is highly ill-posed and rarely inverted directly.

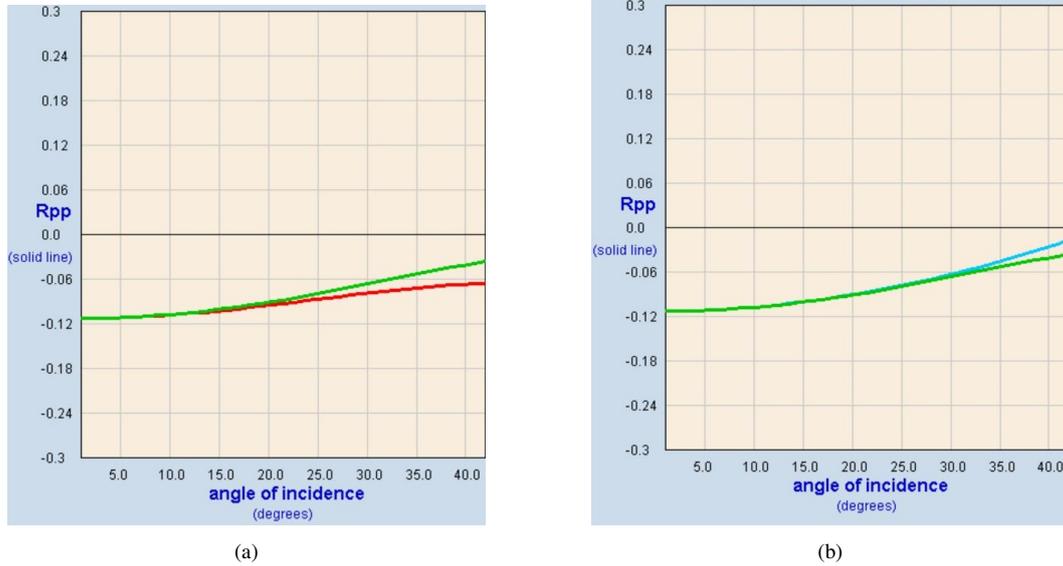


Figure 1: (a) Comparison between the exact isotropic (red) and exact anisotropic (green) solutions in a VTI medium. (b) Comparison between the exact anisotropic (green) and R uger anisotropic (cyan) solutions in a VTI medium. We have used the CREWES VTI Explorer 1.1 to perform the above simulation.

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## 1.2 Alternating direction method of multipliers (ADMM)

In constrained optimization, the goal is to solve:

$$\hat{x} = \underset{x}{\operatorname{argmin}} f(x) + g(x), \quad (11)$$

where  $f(x)$  and  $g(x)$  are convex functions. The two functions can be separate using variable splitting:

$$\hat{x} = \underset{x,z}{\operatorname{argmin}} f(x) + g(z), \quad (12)$$

$$\text{s.t. } x = z.$$

The consensus constraint ( $x = z$ ) can be presented as a penalty term using the augmented Lagrangian (Parikh & Boyd, 2014):

$$L_\rho(x, z, y) = f(x) + g(z) + y^T (x - z) + \frac{\rho}{2} \|x - z\|_2^2, \quad (13)$$

where  $y$  is the Lagrange multiplier vector, and  $\rho > 0$ . Equation 13 can be solved iteratively using the scaled dual variable  $u^k = (1/\rho) y^k$  at  $k^{\text{th}}$  iteration and  $\lambda = (1/\rho)$  (Boyd et al., 2011; Parikh & Boyd, 2014):

$$x^{k+1} = \mathbf{prox}_{\lambda f} (z^k - u^k), \quad (14)$$

$$z^{k+1} = \mathbf{prox}_{\lambda g} (x^{k+1} + u^k), \quad (15)$$

$$u^{k+1} = u^k + x^{k+1} - z^{k+1}. \quad (16)$$

The above equation is the proximal form of the ADMM, where  $\mathbf{prox}_{\lambda f}$ ,  $\mathbf{prox}_{\lambda g}$  are the proximity operators for  $f$  and  $g$  with penalty  $\lambda$ , respectively. The proximity operator of a function, e.g.,  $g$ , is defined in equation 17, where we try to minimize  $g(x)$ , yet stay close to a given value  $y$ .

$$\mathbf{prox}_{\lambda g}(y) = \underset{x}{\operatorname{argmin}} g(x) + \frac{1}{2\lambda} \|x - y\|_2^2. \quad (17)$$

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For  $f$  and  $g$  convex, and  $k^{\text{th}}$  iteration, the ADMM has  $\mathcal{O}(1/k)$  rate of convergence (He & Yuan, 2012; Hong et al., 2016). Many variants of the ADMM have been developed, such as linearized (Goldfarb et al., 2013), online (H. Wang & Banerjee, 2013), stochastic (Ouyang et al., 2013; Huang et al., 2019), and accelerated (J. Zhang et al., 2019).

The function  $f(x)$  can be set as the mean-square error (MSE) of the amplitude differences between synthetic and observed seismic data:

$$MSE(\hat{y}_\theta, y_\theta) = \frac{1}{n} \sum_i (\hat{y}_{i,\theta} - y_{i,\theta})^2, \quad (18)$$

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where  $\hat{y}_\theta$ ,  $y_\theta$  are the synthetic and observed amplitudes at angle of incidence  $\theta$ , respectively. The  $L_2$  loss function is strictly convex, with a positive-definite Hessian. The  $L_2$  loss function is more sensitive to noise, as it is related to the short-tailed Gaussian density (Guitton & Symes, 2003; Tarantola, 2005). In general, seismic data contains different types of noises, and therefore noise-insensitive loss functions are favorable, such as the  $L_1$  and Huber misfit functions. However, the mean-squares and Huber errors are not normally distributed, and therefore can lead to erroneous results (Kosheleva & Kreinovich, 2017) and, hence, a regularization  $g(z)$  may help in stabilizing the inversion.

### 1.3 $f$ -divergences

$f$ -divergences (a.k.a. Csiszár  $f$ -divergences) are contrast functions that measure the dissimilarity between two probability distributions (Csiszár, 2008). Let  $p$  a prior probability density defined from the available well logs and rock physics model, and  $q$  is the predicted probability distribution from seismic amplitude inversion, then for a convex function  $f$  with  $f(1) = 0$ , the  $f$ -divergence of  $p$  from  $q$  is:

$$D_f(p, q) = \mathbb{E}_q [f(u(x))], \quad (19)$$

where  $u$  is the density ratio  $u(x) = p(x)/q(x)$ . The divergence measure depends on the choice of the function  $f$ , i.e., the KL divergence is a special case of the  $f$ -divergences for  $f(x) = x \log x$  (Polyanskiy & Wu, 2014). Examples of probability metrics  $f$  include (Liese & Vajda, 2006):

$$\text{(Forward KL-div): } D_{KL}(p\|q) = u(x) \log u(x) - (u(x) - 1), \quad (20)$$

$$\text{(Reverse KL-div): } D_{KL}(q\|p) = -\log u(x) + (u(x) - 1), \quad (21)$$

$$\text{(Jensen-Shannon): } D_{JS}(p\|q) = \frac{1}{2} \left[ (u(x) + 1) \log \left( \frac{2}{u(x) + 1} \right) + u(x) \log u(x) \right]. \quad (22)$$

The  $f$ -divergences are non-negative, and convex:

$$D_f[\lambda p_1 + (1 - \lambda)p_2 \| \lambda q_1 + (1 - \lambda)q_2] \leq \lambda D_f(p_1 \| q_1) + (1 - \lambda)D_f(p_2 \| q_2). \quad (23)$$

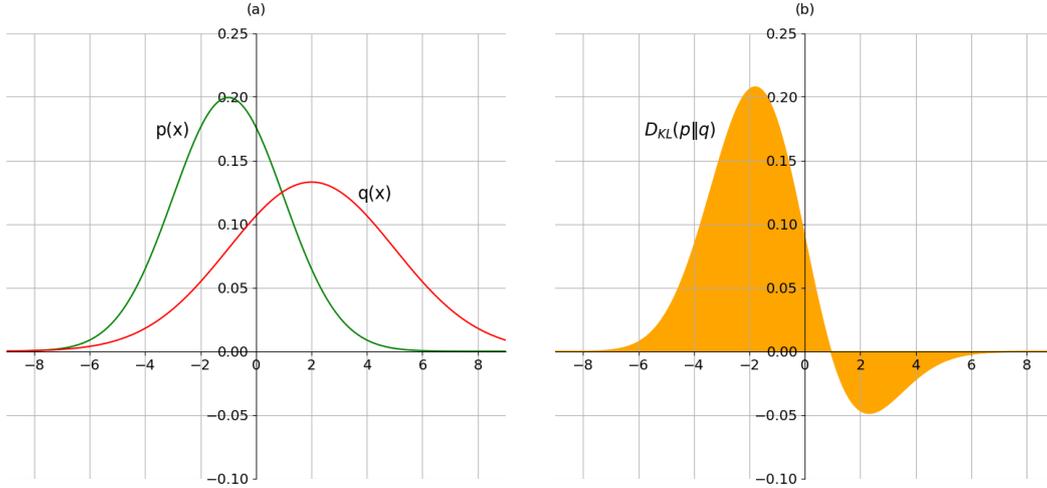


Figure 2: (a) Illustration of the difference in information between the true distribution  $p(x)$  and observed distribution  $q(x)$ . (b) The Kullback–Leibler metric is calculated as the area under the curve  $D_{KL}(p\|q) = \int p(x) \ln \frac{p(x)}{q(x)}$ .

143  $f$ -divergences appear in many machine learning applications and related fields (Nowozin  
 144 et al., 2016; Ke et al., 2020; Ghasemipour et al., 2020; Yu et al., 2020; Gimenez & Zou, 2019).  
 145 Figure 2 shows how to estimate the similarity between two Gaussian probability distributions via  
 146 information. The Kullback–Leibler metric (area under the curve) increases much faster with increasing  
 147 the difference in the mean value as compared to the difference in the variance value between the  
 148 two distributions.

To solve equation 19, *i.i.d.* (independent and identically distributed) samples are generated  $\{x_i\}_{i=1}^M$  from  $q$ , and then the Monte Carlo estimation is used as following (Mnih & Rezende, 2016):

$$D_f(p, q) = \frac{1}{M} \sum_{i=1}^M f(u(x_q^{(i)})), \quad x_q^{(i)} \sim q(x). \quad (24)$$

149 In figure 3a, despite starting with a normal Gaussian distribution  $q^0(x)$ , we iteratively fit  
 150 a mixture model of same number of components as  $p(x)$ ; nonetheless, we can always choose to  
 151 fit a simpler model as shown in figure 3b. In equation 19,  $q(x)$  is calculated using the inverted  
 152 model parameters  $\{x\}$ , and thereby would always have a significant mass that leads to zero-forcing  
 153 of  $q$ , i.e.,  $p(x) = 0 \Rightarrow q(x) = 0$ , hence, minimization of the reverse KL divergence is prompted  
 154 as it steers clear of regions where  $q(x)$  is high and  $p(x)$  is small (a.k.a. mode-seeking). On the  
 155 other hand, the forward KL divergence approximates  $p$  distribution across all its modes, by seeking  
 156 the mean (a.k.a. mean-seeking). For our proposal, due to variable-sensitivities among model parameters,  
 157 noises in the observed data, and bad starting model, the inverted parameters are incorrect leading  
 158 to a variational distribution  $q^0(x)$  (orange) that is far away from the a priori distribution (red) and,  
 159 hence, minimizing the divergence between the two distributions can mitigate ill-posedness of the  
 160 inverse problem.

161 Intuitively, we use  $f$ -divergence to measure how far is the predicted parameters from the  
 162 a priori probability density function, and consequently rejects local minima that have large divergence  
 163 values. Figure 4 illustrates how a solution with larger divergence value from the a priori density  
 164 is identified as a local minimum solution and, hence, inversion is proceeded until reaching a solution  
 165 with a small divergence value, and best generate the observed seismic data. In a proximal splitting  
 166 framework, both the likelihood of data and regularization are minimized, alternately. For  $f$ -divergence,  
 167 minimization is done with respect to the model parameters  $\{x\}$  that used to approximate the variational  
 168 distribution  $q$ , and without needing to sample from  $q$ . The seismic amplitude inversion is an underdetermined,  
 169 hence, a direct method might not be accurate for the  $x$ -minimization step. Moreover, there is no  
 170 closed form solution of the  $f$ -divergence proximity operator. Assuming  $f$  and  $g$  are smooth functions,  
 171 we use the L-BFGS-B iterative solver to carry out the two primal variables' minimization, the  
 172  $x$ -minimization and  $z$ -minimization (Boyd et al., 2011). The L-BFGS-B algorithm is useful to  
 173 handle bound constraints on the variables, where the Hessian is updated at each iteration (Byrd  
 174 et al., 1995; C. Zhu et al., 1997).

#### 175 **1.4 The a priori model**

176 Because seismic anisotropy can not be measured directly in vertical wells, a rock physics  
 177 model is built to predict the elastic anisotropy parameters  $\delta$  and  $\varepsilon$ . A constrained inverse rock  
 178 physics problem is constructed, in which the Hudson-Cheng crack model is combined with the  
 179 anisotropy AVO convolution model to predict the Thomsen's anisotropy parameters (Zidan et al.,  
 180 2021).

A prior probability density function of the elastic and seismic anisotropy parameters is then defined, using a Gaussian mixture model and a multivariate Gaussian model. The a priori distribution provides the statistical information that helps in stabilizing the amplitude inversion. A Gaussian mixture model is defined as a weighted sum of various Gaussian probability density functions (Reynolds, 2009):

$$p(x|\lambda) = \sum_{i=1}^M w_i g(x|\mu_i, \Sigma_i), \quad (25)$$

with,

$$\lambda = [w_i, \mu_i, \Sigma_i], \quad (26)$$

$$\sum_{i=1}^M w_i = 1, \quad (27)$$

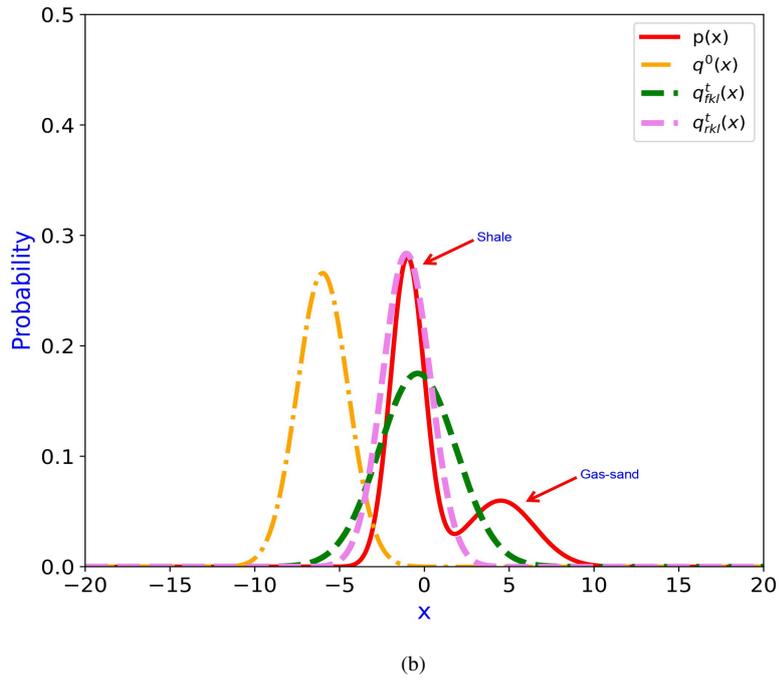
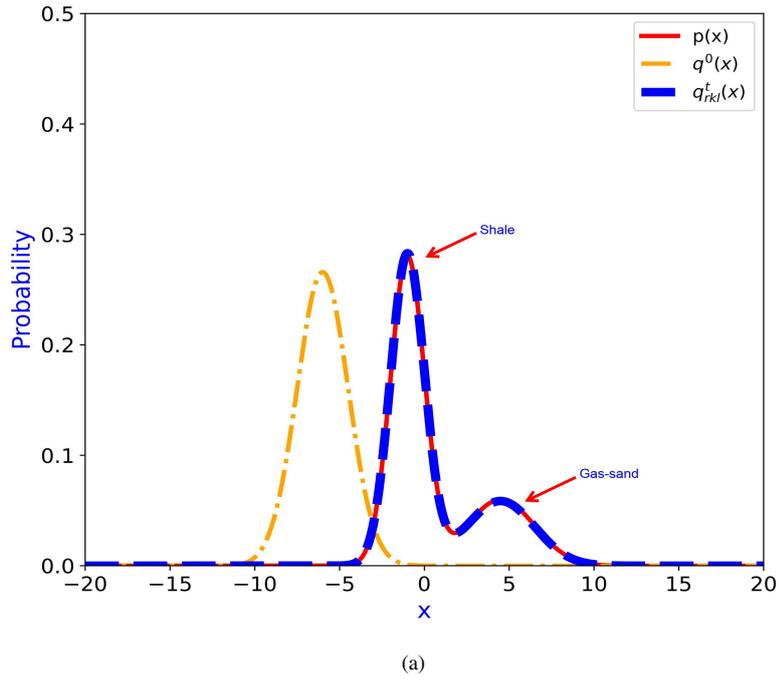


Figure 3: Minimization of  $f$ -divergence between (a) two mixture models using the reverse KL divergence, and (b) a single and mixture models using the forward and reverse KL divergences.

where  $x$  is an  $N$ -dimensional data vector;  $g(x|\mu_i, \Sigma_i)$  are the Gaussian densities components, each defined by a mean vector  $\mu_i$ , and a covariance matrix  $\Sigma_i$ .  $w_i$  is a weight assigned to each component:

$$g(x|\mu_i, \Sigma_i) = \frac{1}{(2\pi)^{\frac{N}{2}} |\Sigma_i|^{\frac{1}{2}}} \exp\left\{-\frac{1}{2} (x - \mu_i)' \Sigma_i^{-1} (x - \mu_i)\right\}. \quad (28)$$

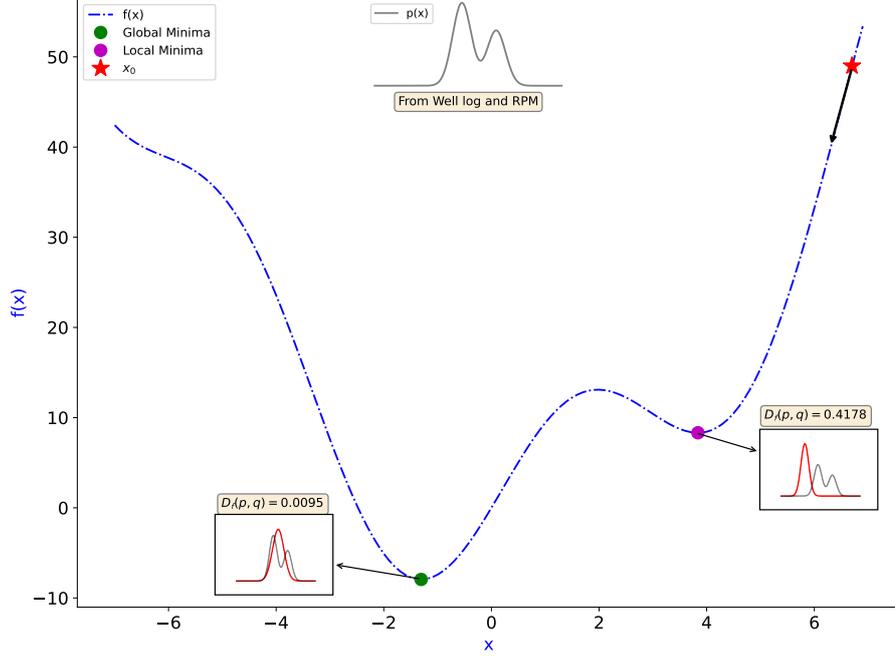


Figure 4: Illustration of the proposed constraint optimization, which aims to skip local minima solutions with large  $f$ -divergence values from the a priori model. The probability density in (red) obtained from the amplitude inversion, is compared with the probability density obtained from well logs and rock physics data (gray).

Here, the Gaussian mixture model is emphasized as a density estimator that best describe the input parameters and fully represents the model space. To achieve the best density estimation of the GMM, a *full*-type covariance matrix is used. In addition, the optimal number of components is estimated by adjusting the model likelihood, via minimizing the Bayesian information criterion (BIC) and the Akaike information criterion (AIC) (VanderPlas, 2016). Smooth Gaussian components, each with a mean vector and a covariance matrix, are fitted using Expectation–Maximization approach (E-M) (Reynolds, 2009). The E-M algorithm is an iterative approach, where the mixture parameters  $\lambda$  are updated to increase the likelihood of the model (VanderPlas, 2016):

$$\hat{w}_i = \frac{1}{T} \sum_{t=1}^T Pr(i|x_t, \lambda), \quad (29)$$

$$\hat{\mu}_i = \frac{\sum_{t=1}^T Pr(i|x_t, \lambda) x_t}{\sum_{t=1}^T Pr(i|x_t, \lambda)}, \quad (30)$$

$$\hat{\Sigma}_i = \frac{\sum_{t=1}^T Pr(i|x_t, \lambda) x_t^2}{\sum_{t=1}^T Pr(i|x_t, \lambda)} - \hat{\mu}_i^2, \quad (31)$$

with,

$$Pr(i|x_t, \lambda) = \frac{w_i g(x_t|\mu_i, \Sigma_i)}{\sum_{k=1}^M w_k g(x_t|\mu_k, \Sigma_k)}, \quad (32)$$

181 where  $\lambda = [w_i, \mu_i, \Sigma_i]$  are the initial mixture parameters.  $\hat{w}_i, \hat{\mu}_i, \hat{\Sigma}_i$  are the updated mixture  
 182 parameters at  $i^{th}$  iteration.  $Pr(i|x_t, \lambda)$  is the posterior probability for  $i^{th}$  component. Figure 5  
 183 shows the Bayesian information criterion and Akaike information criterion as a function of the  
 184 number of the GMM components. The number of Gaussian components that enhances density  
 185 estimation of the GMM is around 25 – 30 for the AIC, whereas BIC suggest a simpler model  
 186 of 9 components; the simple model of 9 components is selected. The inversion results would therefore  
 187 depend on the choice of the a priori density function, i.e., number of the mixture model components.  
 188 The Gaussian mixture and multivariate Gaussian models are shown in figures 6a and 6b, respectively.  
 189 All the 2D projections of the negative log-likelihood of the a priori distributions and the data points  
 190 are plotted. The GMM preferably addresses the multimodal behavior of the model parameters.  
 191 Nevertheless, there would be a covariance matrix for each component, and subsequently the Mahalanobis  
 192 distance is not applicable (Zidan, 2022). Furthermore, the analytical solution of the posterior distribution  
 193 for a GMM is not tractable and, hence, requires a stochastic sampling algorithm to explore the  
 194 mixture prior density function, e.g., Markov chain Monte Carlo. However, the stochastic sampling  
 195 algorithms are computationally expensive, and converges slowly in high-dimensional models.  
 196 Moreover, the AVO models depend on the boundary properties rather than layer properties, which  
 197 necessitates a longer MCMC chain. Alternatively, we propose a deterministic approach to regularize  
 198 the seismic amplitude inversion using a divergence measure from the a priori probability density  
 199 function.

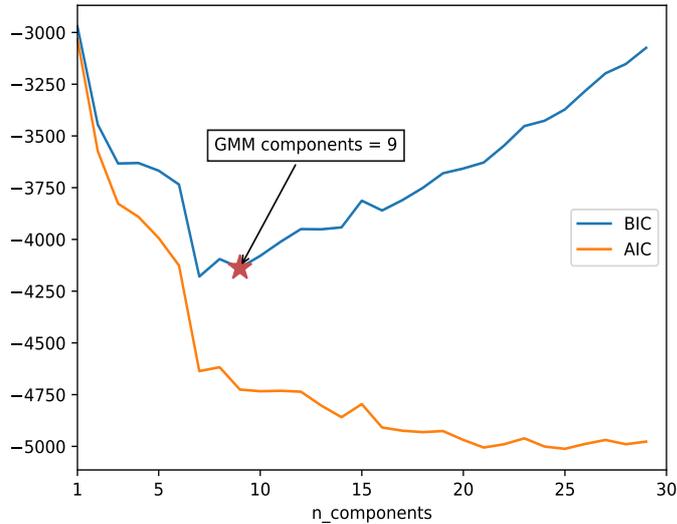


Figure 5: The optimal number of Gaussian components based on the AIC and BIC criterion.

## 200 Results

201 Figure 7 shows cross-plot of the AVO intercept-and-gradient attributes, calculated from the  
 202 pre-stack seismic angle gathers using the two term Aki-Richard's equation. The AVO response  
 203 of the organic-rich shale is of class IV (negative intercept and positive gradient), where the absolute  
 204 values of the intercept and gradient increase with the kerogen content and porosity (Y. Li et al.,

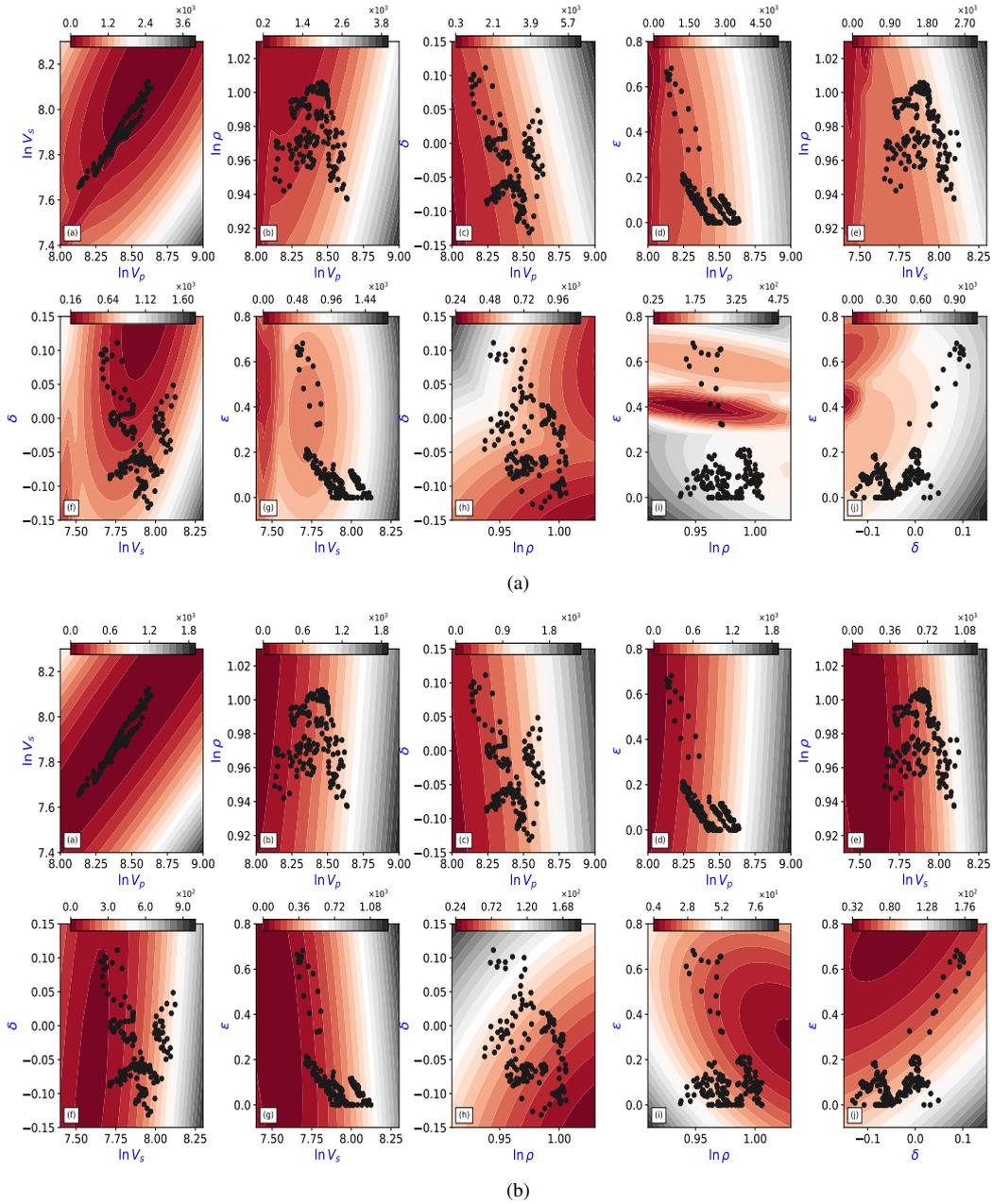


Figure 6: The negative log-likelihood of the a priori probability distribution using a (a) Gaussian mixture model, and (b) multivariate Gaussian model. The GMM components is set to 9. Black dots are the input data samples from well logs and rock physics modelling.

205 2015). As shown, it is hard to distinct class IV AVO response of the organic-rich shales from the  
 206 surrounding rocks. Despite conventional AVO attributes provide information about interfaces (top  
 207 and base of the shale reservoir), they cannot infer the effective moduli of the shale reservoirs. Consequently,  
 208 it is necessary to estimate supplementary parameters that are sensitive to the organic matters and  
 209 correlated with the gas content of the sweetest intervals. The feasibility of the proposed regularization  
 210 is demonstrated using synthetic seismic angle gathers. The model parameters consist of the P-  
 211 and S-wave velocity, density, and Thomsen's parameters  $\delta$  and  $\epsilon$ . The synthetic angle gathers are

212 obtained by convolving the P-P reflection coefficients calculated using equation 5 with a statistical  
 213 wavelet. Figure 8 shows the synthetic angle gather consisting of 41 traces of range  $1 - 41^\circ$ .

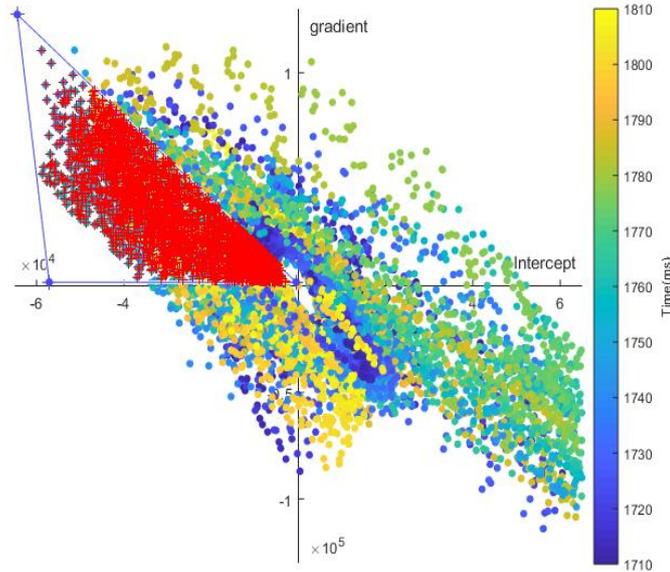


Figure 7: Cross-plot of the AVO intercept and gradient attributes calculated from pre-stack seismic angle gathers, using the two term Aki-Richard's equation. The red dots represent the class IV AVO responses of the organic-rich shales.

214 The information contained in the a priori mixture density and the difference in information  
 215 of the a priori distribution from the variational (predicted) distribution is measured via  $f$ -divergence.  
 216 Subsequently, the amplitude inversion can be regularized by measuring how much information  
 217 lost when substituting the a priori probability density  $p$  with the variational density  $q$ . In an augmented  
 218 Lagrangian scheme, i.e., ADMM, the data residual between synthetic and real angle gathers is  
 219 measured, next regularization is imposed by measuring the information lost in the model space,  
 220 and then the functional and regularization are coordinated globally via constraints, i.e., the dual  
 221 variable for the ADMM.

222 To assess the proposed approach, results are compared with the unconstrained minimization  
 223 of the L-BFGS-B algorithm. We use the same low-frequency model and boundaries as for the  
 224 proposed approach. The initial model is necessary to fully cover the low-to-medium frequency  
 225 gap of the seismic data. The proposed method is first tested on synthetic data with high signal-to-noise  
 226 ratio of  $S/N = \infty$ . For this test, the mean-square error misfit and the Kullback-Leibler divergence  
 227 regularization are used. Figure 9 shows the inversion results of the unconstrained L-BFGS-B and  
 228 constrained ADMM methods. Despite the high signal-to-noise ratio, the unconstrained L-BFGS-B  
 229 algorithm converges to a local minimum and couldn't update all the five parameters properly due  
 230 to the variable-sensitivities among the model parameters and missing information in the observed  
 231 data (limited aperture and S-wave), particularly the density and Thomsen's  $\delta$  as shown in figure 9a.  
 232 Using ADMM with the KL divergence regularization of the GMM a priori model, the model parameters  
 233  $V_p$ ,  $V_s$  and  $\varepsilon$  have been recovered successfully, and  $\rho$  and  $\delta$  are preferably constrained, as shown  
 234 in figure 9b. The constrained minimization is next run using the multivariate Gaussian a priori  
 235 model, as shown in figure 9c. Similarly, all five parameters have been recovered fairly well, yet  
 236 the a priori density of the multivariate Gaussian could not properly model particular litho-facies,

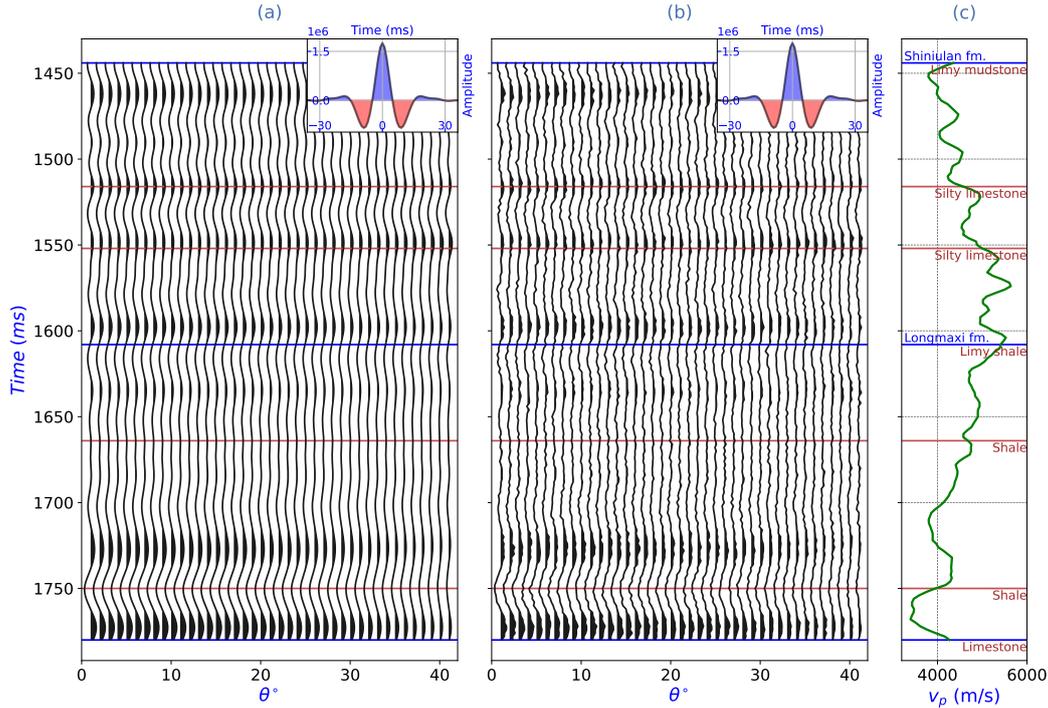


Figure 8: Synthetic pre-stack angle gathers with (a)  $S/N = \infty$ , (b)  $S/N = 9$ , along with (c) the P-wave velocity from well logs, and the statistical wavelet.

237 e.g., the shale gas zone as compared to the GMM. Notably, the log interval is consisting of alternating  
 238 shale and mudstone, and thereby there are no significant variations in the lithofacies.

239 Additionally, a simpler (single-component) variational distributions of a multivariate Gaussian  
 240 and a multivariate student's t- are used to model the inverted parameters, and to approximate the  
 241 a priori Gaussian mixture model. Figures 10a and 10b show the inversion results of the multivariate  
 242 Gaussian and multivariate student's t-distributions. Despite using simpler models of the same (Gaussian)  
 243 and different (student's t-) density functions to fit the 9-components GMM, the  $f$ -divergence minimization  
 244 successfully regularizes the parameters' updates corresponding to the a priori Gaussian mixture  
 245 model.

246 Next, the proposed regularization is tested against noise by adding a random noise to the  
 247 pre-stack angle gather corresponding to signal-to-noise ratio of  $S/N = 9$  as shown in figure 8b.  
 248 The Huber misfit and KL divergence are used. The unconstrained optimization fails to obtain the  
 249 correct updates of model parameters, and the Huber misfit function fails to handle the low signal-to-noise  
 250 ratio of the seismic data, as shown in figure 11a. In the constrained optimization, all five model  
 251 parameters are fairly recovered, by constraining the model updates using the statistical properties  
 252 of the unknown parameters. When the information lost from the a priori probability distribution  
 253 is large, the dual variable is updated and proceed to next iteration until a solution of lower divergence  
 254 and residual values is reached, as shown in figure 11b.

255 Then, the regularization is tested with a poor starting model, represented by a constant mean  
 256 value for each parameter. Because the initial model is far from the true solution, the unconstrained  
 257 optimization fails to converge to the true solution, as illustrated in figure 12a. On the contrary,  
 258 minimization step of the KL divergence imposes regularization that prompts solutions close to  
 259 the a priori information, hence, fairly recover all the five model parameters as shown in figure 12b.  
 260 Additionally, a single-component multivariate Gaussian and student's t-distributions are used as

261 variational distributions to fit the a priori model of 9-components GMM. Figure 13 shows that  
 262 the regularization functions, likewise, properly constrain the model parameters.

263 Finally, the proposed constrained optimization is applied to pre-stack seismic data acquired  
 264 at the Zhaotong national shale gas demonstration area, Sichuan basin. The marine shales of the  
 265 Silurian Longmaxi formation occur along synclinal belt, the sweetest intervals have an average  
 266 thickness of  $35.5m$  and rich in organic matter ( $TOC \approx 3.5\%$ ). The estimated anisotropy parameters  
 267 are therefore used to accurately map the occurrence of the sweetest intervals across the seismic  
 268 array. The maximum angle of incidence at the reservoir interval is estimated as  $42^\circ$ , and all the  
 269 ( $1-42^\circ$ ) are used in the inversion process. The seismic convolution model of the Rüger equation,  
 270 and the statistical wavelets estimated from the seismic data are used to synthesize angle gathers,  
 271 which subsequently compared with the real angle gathers at each CDP location. The low-frequency  
 272 models for the elastic and Thomsen's anisotropy parameters are built based on the well log data  
 273 and the structure and stratigraphy interpretation. A priori Gaussian mixture model is built, and  
 274 a single-component Gaussian density function is used as the variational distribution. The sweetest  
 275 shale intervals are well-identified with a lower P- and S-wave velocity, and density, and higher  
 276  $\delta$  and  $\varepsilon$  along the synclinal structure at about  $1750ms$  as shown in figures 14 and 15. The near-angle  
 277 seismic traces have less information about density, while most information is in the far-angle traces,  
 278 hence, the density result is less stable as it is more sensitive to the noise level in the data, particularly  
 279 misalignment of reflectors at far angles of incidence.

## 280 Discussion

281 To alleviate ill-posedness of the amplitude inversion, the objective function is split into a  
 282 loss, and regularization that addresses the statistical properties conveyed by a priori density function  
 283 of the unknown parameters. The regularization is based on measuring the distance between two  
 284 probability distributions. A priori mixture model, and a variational distribution of the same number  
 285 of components or simpler are used as regularization. The proposed approach works as following,  
 286 first the data residual between the observed and synthetic data is minimized, next the information  
 287 loss in the model space is minimized, and then both the functional and regularization are coordinated  
 288 globally via the dual variable. Despite the proposed regularization provides better constrains on  
 289 the density and Thomsen's parameter  $\delta$ , P-S data and larger angles of incidence are required to  
 290 fully constrain the inversion results. The proposed approach is very useful in solving geophysical  
 291 inverse problems that involve different moduli and elastic attributes, such as the Young's modulus,  
 292 Poisson's ratio, incompressibility, rigidity, Lamé parameters, and density (Goodway et al., 1997;  
 293 Xu & Bancroft, 1998; Gray et al., 1999; Golalzadeh et al., 2008; Zong et al., 2013; Yin et al., 2015);  
 294 moreover, the joint elastic and petrophysical models, such as the P- and S-wave velocity, density,  
 295 effective porosity, clay volume, and water saturation (Bosch, 2004; Z. Li et al., 2016; M. Liu &  
 296 Grana, 2018; de Figueiredo et al., 2018b; Guo et al., 2021; K. Li et al., 2021). The latter AVO  
 297 models necessitate robust a priori information to better addresses the joint distributions of the different  
 298 model parameters.

299 The performance of the proposed regularization is demonstrated using a high signal-to-noise  
 300 seismic, low signal-to-noise seismic, and bad starting model. In the high signal-to-noise seismic,  
 301 the proposed regularization successfully eases the ill-posedness due to variable-sensitivities among  
 302 model parameters. With bad starting model, the regularization fairly mitigate the problem of frequency  
 303 gap between the seismic data and prior model, by updating the dual variable of the functional and  
 304 regularization. The regularization is also capable of mitigating the low signal-to-noise observed  
 305 seismic data. Furthermore, the proposed regularization is suited for many probability distributions  
 306 that can best describe the statistics of the well logs data, such as parametric distributions (e.g.,  
 307 log-normal distribution), and non-parametric distributions (e.g., kernel density estimation). However,  
 308 the resulting solution from ADMM is not exact because of the partial updates for the dual variable.  
 309 Using a Bregman divergence as a regularization function (Zidan, 2022), a stationary statistical  
 310 correlation information matrix is used, which can be extended in time using a temporal correlation  
 311 function (Buland & Omre, 2003), and in space using a spatial correlation model (de Figueiredo  
 312 et al., 2018b). Nonetheless,  $f$ -divergence is considered non-stationary and can be used for each

313 interface, without a temporal nor a spatial correlation functions (for sufficient well-control), hence,  
 314 can be used for the entire well log interval and across the seismic array. Furthermore, the ADMM  
 315 provides an adaptable framework to incorporate additional regularization functions, such as noise-reduction  
 316 and sparsity-promoting functions. The ADMM can be extended for parallel implementation, to  
 317 solve  $N$  sub-problems in parallel at each iteration (Boyd et al., 2011; Deng et al., 2017).

318 Nevertheless, the divergence measure heavily depends on the choice of the probability metric;  
 319 moreover, it is required to approximate a variational distribution from the inverted model parameters.  
 320 Furthermore, optimization is done over the same data points (inverted model) that are used to approximate  
 321 the variational distribution. This deterministic update might lead to biased estimates of the descent  
 322 directions on the  $f$ -divergence. To enhance the estimates of the functional value and gradients,  
 323 the Auxiliary  $f$ -divergence or the Fenchel-conjugate  $f$ -divergence can be used, which yield the  
 324 upper and lower bounds on the divergence measure, respectively (M. Zhang et al., 2019). Furthermore,  
 325 the tail-adaptive  $f$ -divergence can be used to achieve mass-covering of the target distribution (D. Wang  
 326 et al., 2018). Other divergence measures such as the Alpha-, Beta- and Gamma-divergences can  
 327 also be adapted for better convergence (Cichocki & Amari, 2010). Having said that, it is necessary  
 328 to estimate the descent direction without the need to approximate a variational distribution. Zidan  
 329 (2022) used the Stein's method to estimate the optimal descent direction that maximally decreases  
 330 the KL divergence without needing to approximate a variational distribution. The estimated descent  
 331 direction depends only on the a priori probability density through a score function.

## 332 Conclusion

333 We proposed a constrained optimization scheme that can be used with a priori mixture models.  
 334 We then applied the proposed regularization to anisotropy amplitude inversion in a VTI medium.  
 335 We first combined the well logs and rock physics data to build a joint mixture prior probability  
 336 distribution that conveys the statistical properties of the parameters of interest. We then set up  
 337 a regularized inverse problem using the alternating direction of multipliers method (ADMM),  
 338 in which the functional and regularization are solved separately. The goal is to obtain a single  
 339 stable solution that minimizes the data residual, yet stay close to the a priori mixture probability  
 340 distribution. Such constraints are necessary when the starting model is far away from the true solution.  
 341 In comparison with the unconstrained optimization, the vertical velocities are better recovered,  
 342 and the density and anisotropy parameters are well-constrained.

## 2 ACKNOWLEDGMENTS

The authors acknowledge BGP for the well log data and seismic data. The work is supported by the Singapore Economic Development Board Petroleum Engineering Professorship program. A. Zidan is also supported by SINGA scholarship. The data that support the finding of this study are available online at <https://doi.org/10.5281/zenodo.4304555>.

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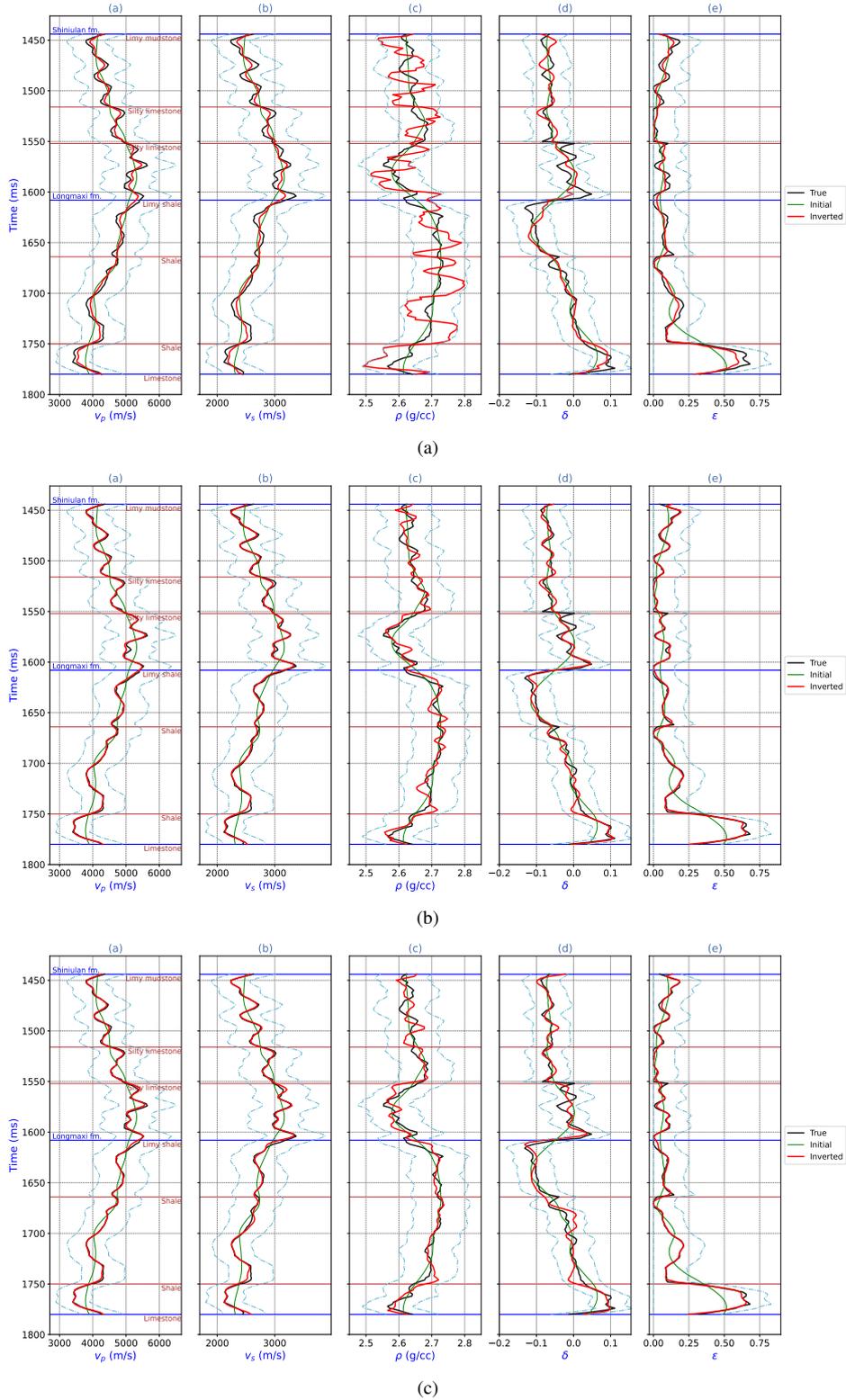


Figure 9: Inversion results of the (a) unconstrained L-BFGS-B, (b) ADMM of the 9-components-GMM, and (c) ADMM of the multivariate Gaussian. The signal-to-noise ratio is set as  $S/N = \infty$ . Cyan lines represent the upper and lower boundary constraints.

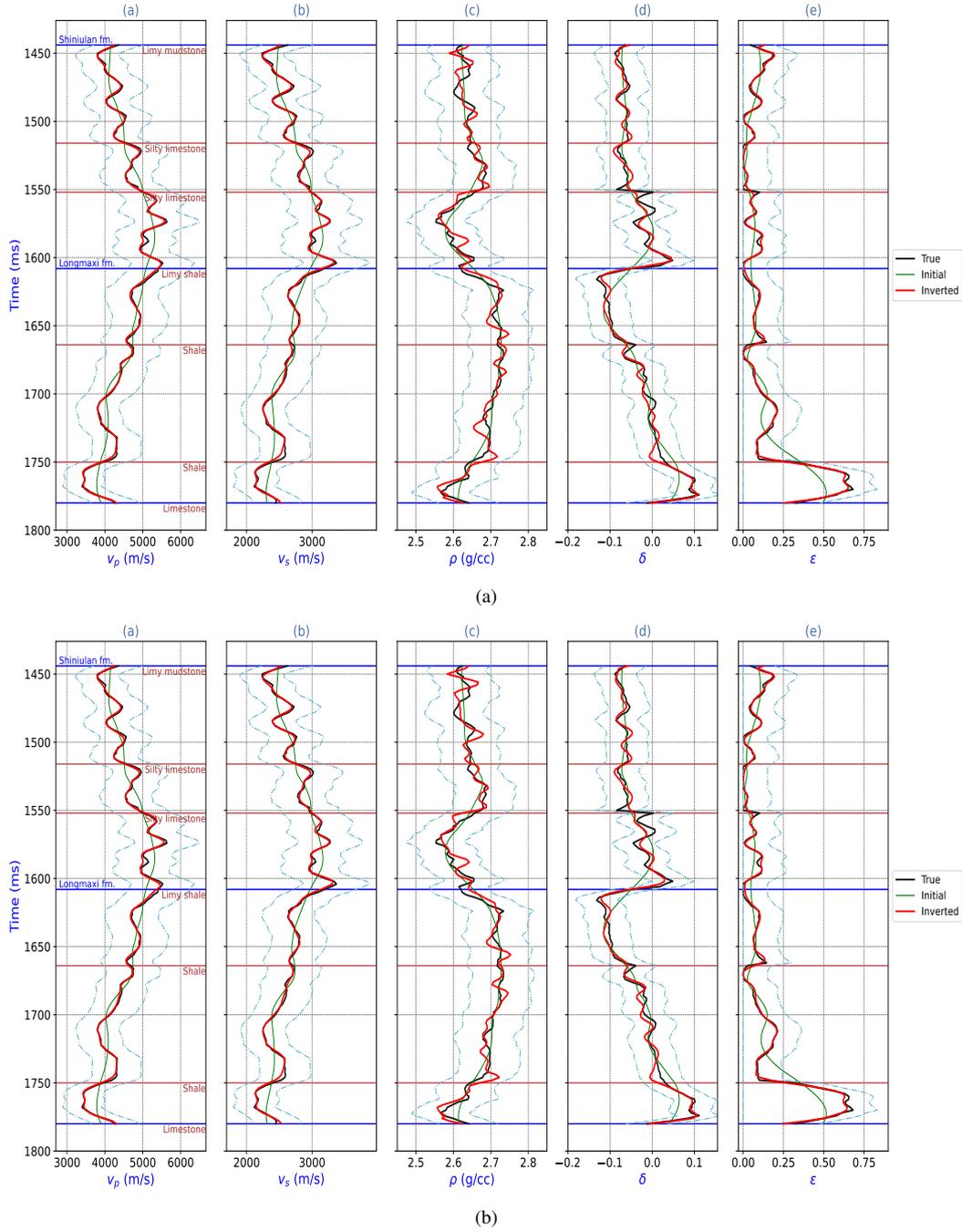


Figure 10: Inversion results of using a (a) single-component multivariate Gaussian, and (b) single-component multivariate student's t-distribution as the variational distribution to fit the 9-components Gaussian mixture a priori model. The signal-to-noise ratio is set as  $S/N = \infty$ . Cyan lines represent the upper and lower boundary constraints.

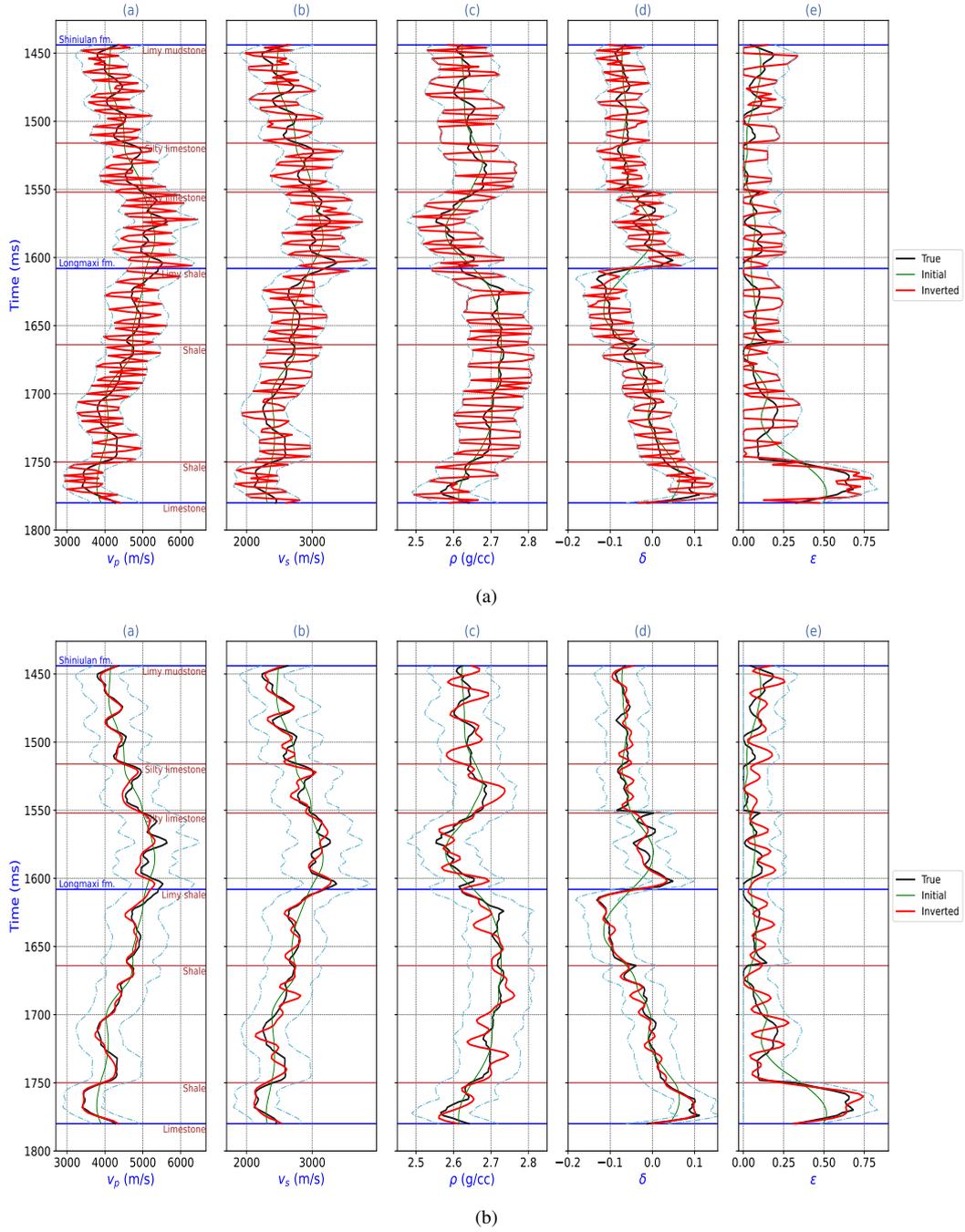


Figure 11: Inversion results of the (a) unconstrained optimization of the  $L=BFGS-B$ , and (b) ADMM of 9-components GMM. The signal-to-noise ratio is set as  $S/N = 9$ . Cyan lines represent the upper and lower boundary constraints.

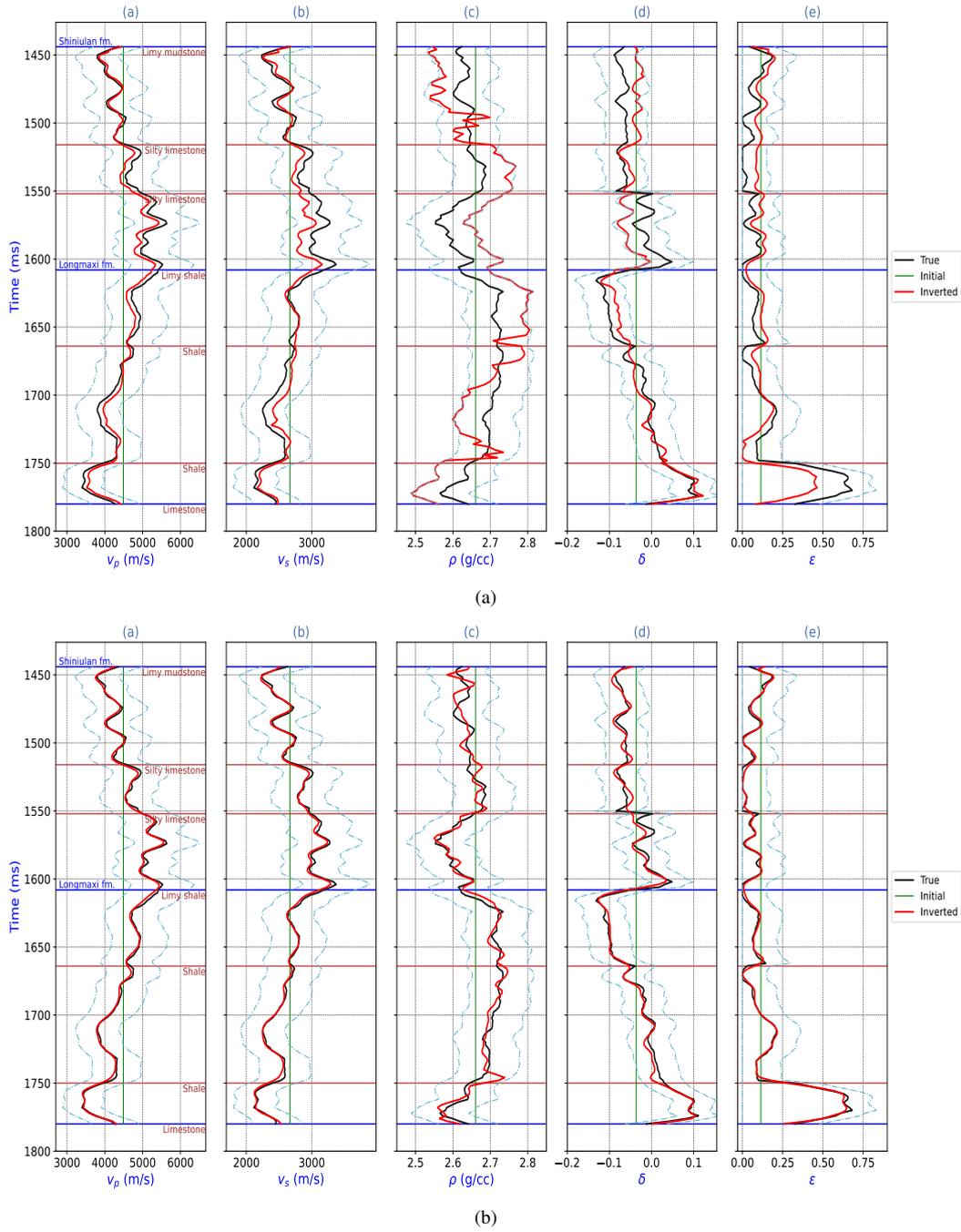


Figure 12: Inversion results of the (a) unconstrained L-BFGS-B, and (b) ADMM of 9-components GMM. The signal-to-noise ratio is set as  $S/N = \infty$ , and with a bad initial model (the mean value for each parameter along the time axis). Cyan lines represent the upper and lower boundary constraints.

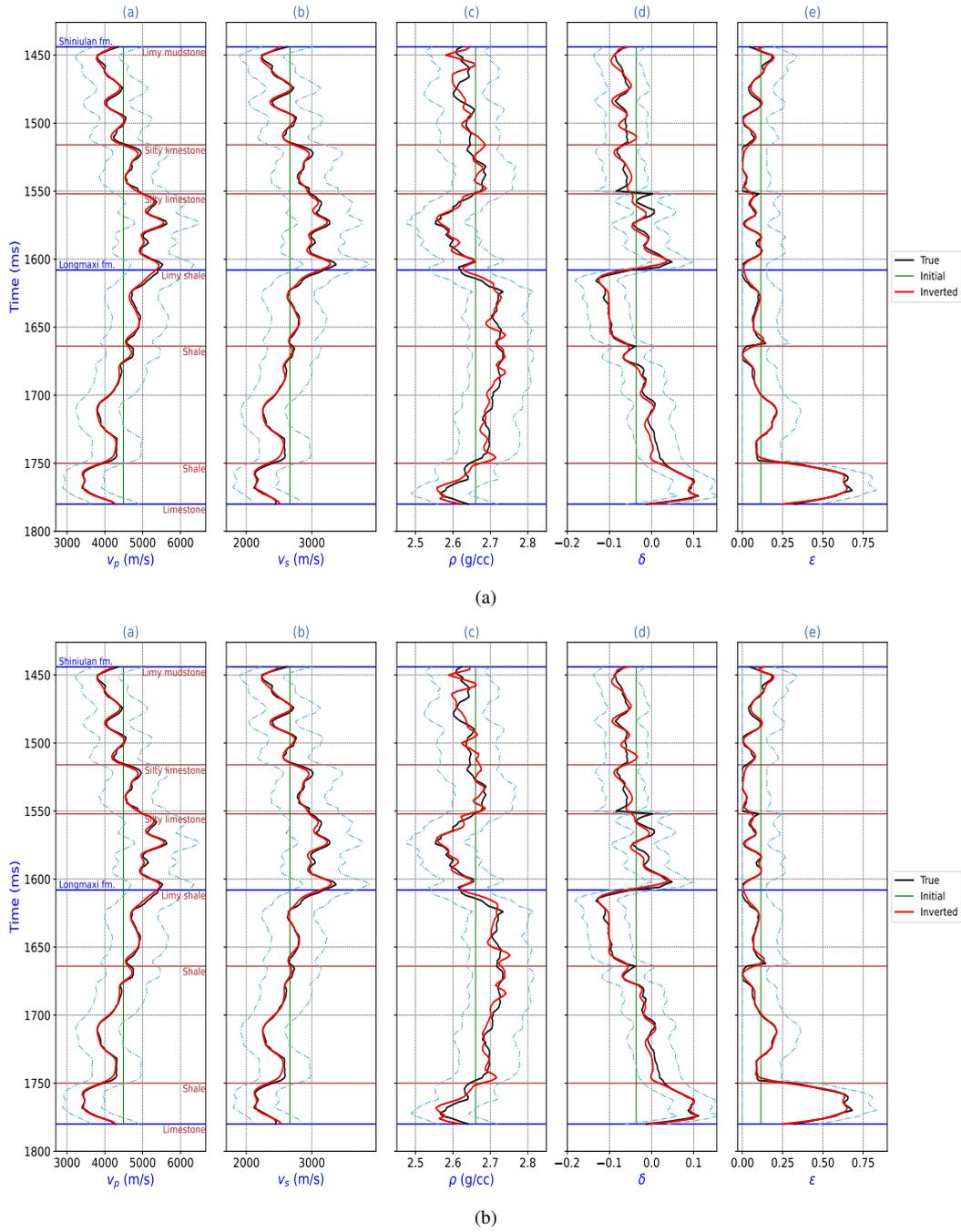


Figure 13: Inversion results of using a (a) single-component multivariate Gaussian, and (b) single-component multivariate student's t-distribution as the variational distribution to fit the 9-components Gaussian mixture a priori model. The signal-to-noise ratio is set as  $S/N = \infty$ , and with a bad initial model (the mean value for each parameter along the time axis). Cyan lines represent the upper and lower boundary constraints.

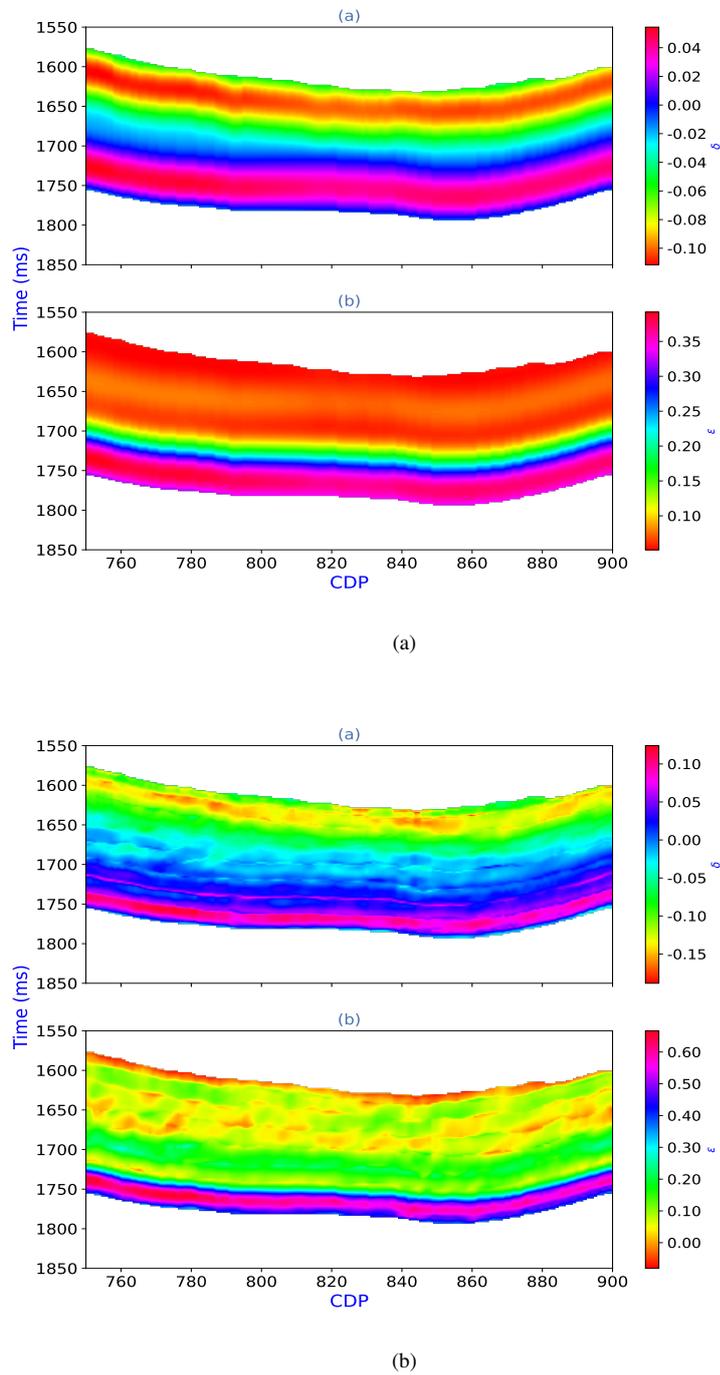
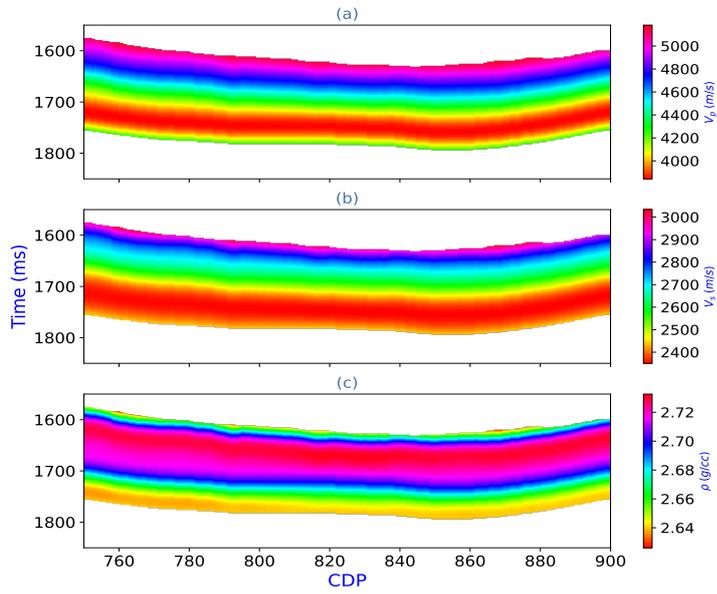
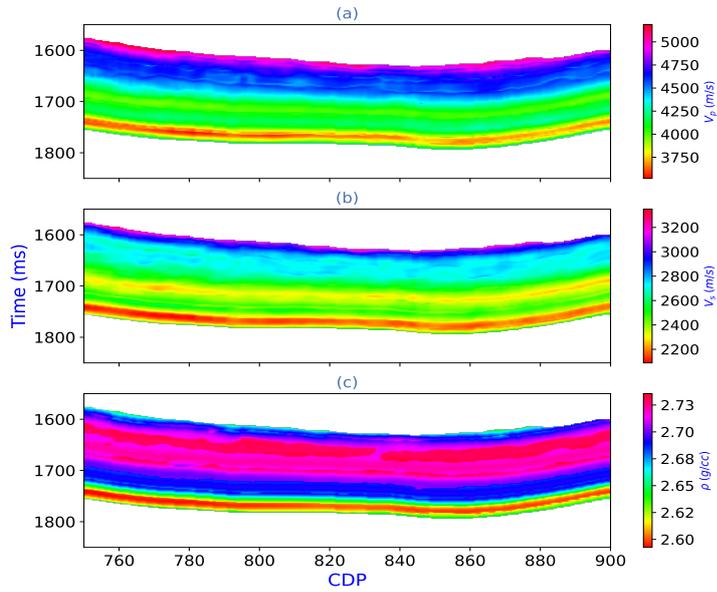


Figure 14: (a) Low-frequency anisotropy models of the (a) near-vertical anisotropy ( $\delta$ ), and (b) P-wave anisotropy ( $\epsilon$ ), which used to initialize (b) the amplitude-versus-offset inversion using the proposed constrained optimization.



(a)



(b)

Figure 15: (a) Low-frequency elastic models of the (a) P-wave velocity ( $V_p$ ), (b) S-wave velocity ( $V_s$ ), and (c) density ( $\rho$ ), which used to initialize (b) the amplitude-versus-offset inversion using the proposed constrained optimization.