Interseismic Strain Accumulation across the Main Recent Fault, SW Iran, from Sentinel-1 InSAR Observations

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Key Points:

- We derive East and Vertical surface velocities from 5.5 years of Sentinel-1 synthetic aperture radar images over the western Zagros.
- We estimate an interseismic slip rate of 3.0 ± 1.0 mm/yr (2σ) for the Main Recent Fault, in agreement with previous GNSS studies.
- We estimate a geodetically-determined interseismic locking depth of 18–30 km, a first for this fault.

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Abstract

The Main Recent Fault is a major right-lateral strike-slip fault in the western Zagros mountains of Iran. Previous studies have estimated a wide range of slip rates from both sparse GNSS (1–6 mm/yr) and geological/geomorphological (1.6–17 mm/yr) methods. None of these studies have estimated the depth to the top of the locked seismogenic zone. Characterizing this "locking depth" for the Main Recent Fault, and more accurately constraining its interseismic slip rate, are both critical for estimating the seismic hazard posed by the fault, as well as for understanding how oblique convergence is accommodated and partitioned across the Zagros. To address this important knowledge gap for the MRF, here we use 200 Sentinel-1 SAR images from the past 5 years, spanning two ascending and two descending tracks, to estimate the first InSAR-derived slip rate and locking depth for a 300 km long section of the fault. We utilise two established processing systems, LiCSAR and LiCSBAS, to produce interferograms and perform time series analysis, respectively. We constrain north-south motion using GNSS observations, decompose our InSAR line-of-sight velocities into fault-parallel and vertical motion, and fit 1-D screw dislocation models to three fault-perpendicular profiles of fault-parallel velocity, following a Bayesian approach to estimate the posterior probability distribution on the fault parameters. We estimate an interseismic slip velocity of 3.0±1.0 mm/yr below a loosely constrained 18–30 km locking depth, the first such estimate for the fault, and discuss the challenges in constraining the locking depth for low magnitude interseismic signals.

Plain Language Summary

Convergence between the Arabian and Eurasian plates is causing deformation of the Earth's crust in Iran. Some of this motion is taken up by movement at depth on the Main Recent Fault, which is stuck by friction near the Earth's surface and is therefore accumulating strain which may then be released in an earthquake. We use five years of satellite radar images to measure the average velocity of the ground surface either side of the fault. By looking at the velocity difference across the fault, along with the gradient, we can estimate the rate at which the fault is accumulating strain and the depth below which this is occurring. Our estimated rate of 3.0 ± 1.0 mm/yr is in agreement with previous estimates from GPS studies, while our estimate of the locking depth, from 18–30 km, is the first such estimate for the fault. The broad range of possible values for the locking depth highlights the difficulties of studying tectonic signals when they are close in magnitude to the sensing limit of our satellite imagery method (1 mm/yr).

1 Introduction

The Main Recent Fault (MRF) is a 800 km long dextral stike-slip fault in the hinterlands of the Zagros mountains, Iran. The fault is one of the most seismically active in the northwestern Zagros, having experienced historical earthquakes up to M_s 7.4 (Ambraseys & Moinfar, 1973; Ghods et al., 2012; Karasözen et al., 2019), driven by convergence between the Arabian and Eurasian plates. During the interseismic period of the earthquake cycle, the MRF can be viewed as accumulating strain in the locked upper crust whilst slipping aseismically at depth, following that assumed for other strike-slip fault zones (Savage & Prescott, 1978; Thatcher, 1983; Savage, 2000; Wright et al., 2013). Estimates of interseismic slip rate and the depth-extent of the locked seismogenic zone, from here on referred to as the 'locking depth', are critical to our understanding of both the local seismic hazard (Smith-Konter & Sandwell, 2009), and the accommodation of oblique convergence across the Zagros. Despite the importance of the MRF for understanding Iranian tectonics and seismic hazard, its interseismic slip rate is still poorly constrained, and no estimates for the locking depth of the fault have been published. Previous studies of the MRF have used a range of geological markers, geomorphological offsets, cosmogenic isotope dating, and Global Navigation Satellite System (GNSS) measurements

to estimate a wide range (1–17 mm/yr) of possible slip rates (Table 1). The average slip rates determined from long-term geological/geomorphological offsets (1.6–17 mm/yr, Talebian & Jackson, 2002; Bachmanov et al., 2004; Copley & Jackson, 2006; Alipoor et al., 2012) and cosmogenic dating (3.5–12.5 mm/yr, Authemayou et al., 2009) cover a broader range with higher upper bounds than geodetic slip rates from decadal GNSS studies (1–7 mm/yr, Vernant et al., 2004; Hessami et al., 2006; Walpersdorf et al., 2006; Khorrami et al., 2019). The large variation in geological and geomorphological slip rates reflects differences in the time scales of the estimates (thousands vs. millions of years), uncertainties in the measured offsets, and uncertainties in the age of the MRF. Meanwhile, GNSS-derived estimates have suffered from the sparsity of instruments in Iran, especially in the south west and in northern Iraq.

Interferometric Synthetic Aperture Radar (InSAR) time series analysis is a well-established technique for measuring ground deformation linked to interseismic strain accumulation (Fialko, 2006; Jolivet et al., 2013; Tong et al., 2013; Hussain, Hooper, et al., 2016; Weiss et al., 2020), and has been shown in a couple of instances to be able to estimate interseismic fault slip rates down to a few millimeters per year (Bell et al., 2011; Mousavi et al., 2015). The European Space Agency's (ESA) Sentinel-1 C-band SAR satellites provide previously unprecedented temporal coverage of data suitable for interferometry, with all of Iran being imaged on average every six days. Current sensing limits for InSAR time series methods are 2–3 mm/yr for an average point velocity, given a large number (hundreds) of acquisitions over a long time period (several years) and in the presence of minimal noise (Morishita et al., 2020). The total time period covered has a greater influence on the sensing threshold than the number of images used (Morishita et al., 2020). Iran is generally a suitable target location for InSAR, given its relatively arid climate and sparse vegetation cover. InSAR time series methods have been applied to measure interseismic slip rates on a number of faults in the region, including the Ashkabad (5–12 mm/yr, Walters et al., 2013), Doruneh (2.5±0.3 mm/yr Mousavi et al., 2021), North Tabriz (6–10 mm/yr, Rizza et al., 2013; Karimzadeh et al., 2013; Su et al., 2016; Aghajany et al., 2017), Shahroud (4.75±0.8 mm/yr Mousavi et al., 2015), and the Minab-Zendan-Palami (10 mm/yr) and Sabzevaran-Kahnuj-Jiroft (7.4 mm/yr) fault systems (Peyret et al., 2009). However, InSAR has not previously been used to estimate the interseismic motion across the MRF, despite the potential of this technique to better constrain both the slip rate and locking depth of this important fault.

Using ESA's Sentinel-1 SAR satellites and 5.5 years of data covering a 400×200 km area centred on the MRF, we measure the relative horizontal motion in a velocity time series to the millimeter per year level. We mitigate atmospheric noise contamination and co- and post-seismic signals for the M_w 7.3 Sarpol-e Zahab earthquake in 2017, and from these corrected velocity data we investigate the fault kinematics of the MRF, namely the rate of interseismic fault slip and the depth above which the fault is considered locked, using a simple screw dislocation model (Savage & Burford, 1973) and following a Bayesian framework to assess the uncertainties based upon the data noise. We conclude with a discussion of the role of the MRF in accommodating convergence between the Arabian and Eurasian plates, the extent of slip localisation, and the limitations of measuring tectonic signals close to the current sensing limit of Sentinel-1 InSAR observations.

2 Tectonic Background

Iran constitutes one of the widest zones of continental convergence on a global scale (Allen et al., 2004, 2013). Present day convergence rates between Arabia and Eurasia are estimated at 15–25 mm/yr (McClusky et al., 2003; Khorrami et al., 2019) (Figure 1c). Convergence is roughly range-perpendicular in the southeastern Zagros, becoming increasingly oblique up to 45° in the northwest. Talebian and Jackson (2004) first suggested that deformation in the western Zagros is partitioned into range-perpendicular shortening and range-parallel strike-slip motion, later supported by Iranian GNSS ve-

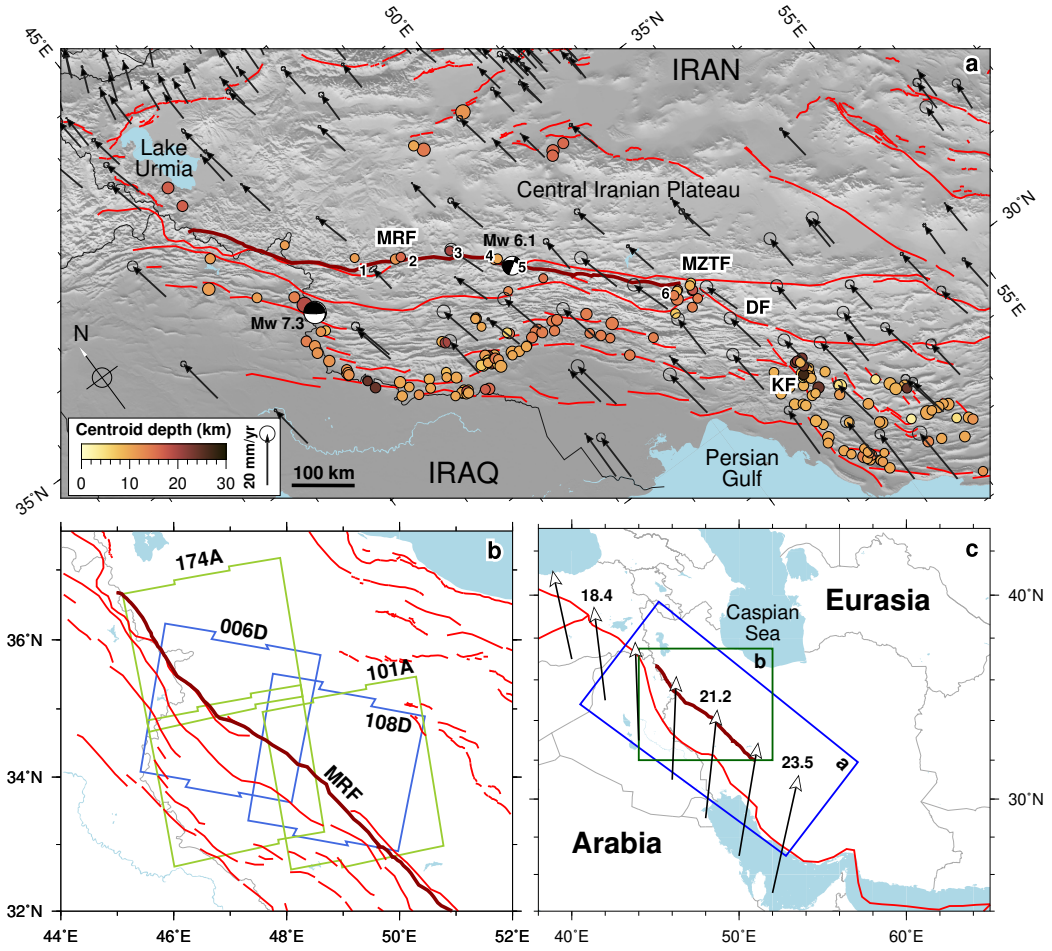

Figure 1. Overview of the study area in Western Iran. (a) Location of major faults (red lines) from Walker et al. (2010), with a 750 km section of the MRF highlighted in dark red. Vectors show GNSS velocities with 1σ uncertainties from Khorrami et al. (2019) with respect to a stable Eurasia. Circles show relocated seismicity from Karasözen et al. (2019) covering 1962–2017, scaled by magnitude and coloured by centroid depth. MRF = Main Recent Fault, MZTF = Main Zagros Thrust Fault, DF = Dena Fault, KF = Kazerun Fault. Fault segments are numbered as: 1 - Kamyanan, 2 - Sahneh, 3 - Nahavand, 4 - Borujerd, 5 - Dorud, 6 - Ardal. (b) LiCS defined frames for Sentinel-1 InSAR coverage. A = Ascending orbital track, D = Descending orbital track. Focal mechanisms are shown for the M_w 7.3 2017 Sarpol-e Zahab mainshock (Nissen et al., 2019) and the M_w 6.1 2006 Silakhour mainshock (Ghods et al., 2012). (c) Plate boundary from Bird (2003) shown in red, Arabian plate velocities (mm/yr) relative to stable Eurasia (Kreemer et al., 2014).

Table 1. Previously published long-term geologic and geodetic interseismic slip rate estimates for the Main Recent Fault (top) and adjacent faults (bottom).

Study	Fault	Method	Rate (mm/yr)
Talebian and Jackson (2002)	MRF	Geological/geomorphological features	10–17
Bachmanov et al. (2004)	MRF : Dorud & Nahavand segments	Geological/geomorphological features	10
Vernant et al. (2004)	MRF	GNSS (regional)	3 ± 2
Copley and Jackson (2006)	MRF	Geological/geomorphological features	2–5
Walpersdorf et al. (2006)	MRF	GNSS (regional)	4–6
Authemayou et al. (2009)	MRF	Cosmogenic ^{36}Cl dating	3.5–12.5
Alipoor et al. (2012)	MRF	Geological, geomorphological markers, pullapart basins, and drainage patterns	1.6–3.2
Khorrami et al. (2019)	MRF	GNSS (regional)	2.7–4.0
Hessami et al. (2006)	Kazerun	GNSS (campaign profiles)	4–5
Walpersdorf et al. (2006)	Kazerun	GNSS (regional)	3 ± 2
	Dena		3 ± 2
Tavakoli et al. (2008)	Dena	GNSS	3.7 ± 0.8
	Kazerun		3.6 ± 0.6
Authemayou et al. (2009)	Kazerun (Northern strand)	Cosmogenic ^{36}Cl dating	2.5–4
	Kazerun (Central strand)		1.5–3.5

Note. Adjacent fault are provided for comparison and labelled in Figure 1. Studies are ordered by year of publication.

locity fields (Walpersdorf et al., 2006) and crustal stress maps (Zarifi et al., 2014). However, this partitioning may be incomplete (Vernant & Chery, 2006; Nissen et al., 2019). At longitudes between 48°–55° E, 7–10 mm/yr of shortening is accommodated by range-perpendicular movement on thrust faults (Vernant et al., 2004; Walpersdorf et al., 2006; Khorrami et al., 2019). This decreases to 4–6 mm/yr moving westward to 42°–46° E.

The Zagros is also one of the most seismically active fold-and-thrust belts in the world (Talebian & Jackson, 2004; Hatzfeld & Molnar, 2010; Nissen et al., 2011) (Figure 1a). Focal depths generally range between 4–25 km, nucleating in both the basement and sedimentary cover in similar proportions, and with the majority failing to rupture to the surface (Karasözen et al., 2019). Seismicity accounts for around half of the geodetic shortening rate in the northwestern Zagros, and less than a third in the southeast, implying large amounts of folding, aseismic fault slip, and ductile shortening of the basement (Karasözen et al., 2019). Rates of seismicity drop off rapidly northwest of the MRF in the Central Iranian Plateau.

The Main Recent Fault trends NW-SE for over 800 km as a linear series of fault segments (Figure 1). These segments may be characterised by their strike, which changes from 330° northwest of Kamyanan, to 300° in the centre near Sahneh, and 315° southeast of Borujerd (Talebian & Jackson, 2002). The overall slip vector is believed to be parallel to the central section (47–50° E), between 300° and 310°, suggesting oblique motion on both northwestern and southeastern segments. The MRF cross-cuts the Main Zagros Thrust Fault (MZTF), having partially inherited its fault trace west of 49° E (Tchalenko & Braud, 1974). The MZTF traces the suture between the Arabian margin and the Iranian block and is currently thought to be inactive (Walpersdorf et al., 2006). Relocated seismic events also highlight slip on previously unmapped faults, suggesting continuing evolution of the geometry of the MRF (Ghods et al., 2012). The MRF contains multiple pull-apart basins (Talebian & Jackson, 2002; Authemayou et al., 2009; Sepahvand et al., 2012), related either to the change in strike (Talebian & Jackson, 2002) or the changing convergence direction (Copley & Jackson, 2006). To the southeast, the MRF terminates in a 'horse tail' structure formed by the Dena, Kazerun, Borazjan, Kareh Bas, and Sarvestan faults (Bachmanov et al., 2004; Authemayou et al., 2009; Khorrami et al., 2019). To the northwest, right-lateral motion continues into a complex dextral shear zone that runs along the Arabian-Eurasian suture north of 37° N (Niassarifard et al., 2021). This in turns joins two NNW-striking normal fault zones north of 37.5° N, which accommodate range-parallel motion through ESE-extension. Right-lateral strike-slip motion is re-

sumed on the North Tabriz fault in NW Iran (Aghajany et al., 2017) and the North Anatolian fault in Turkey (Hussain, Hooper, et al., 2016).

Calibrated earthquake relocations from Karasözen et al. (2019) show that the majority of earthquakes close to the MRF occur at depths shallower than 15 km. A minority of events are recorded at depths closer to 20 km, implying a likely locking depth of around 15–20 km. The 2006 Silakhour sequence, consisting of two foreshocks on the 30th March (M_b 4.8 and 5.2) and a M_w 6.1 mainshock on the 31st March, ruptured two patches of the MRF (Peyret et al., 2008; Ghods et al., 2012). These patches were on the western Nahavand-Borujerd and eastern Borujerd-Dorud segments, separated by 10 km of fault that did not rupture. The Nahavand-Borujerd patch was not associated with any known fault structure, suggesting ongoing development of the fault zone (Ghods et al., 2012). The largest historical earthquake believed to have occurred on the fault, a M_s 7.4 earthquake near Dorud in 1909, ruptured 45–65 km of the fault (Ambraseys & Moinfar, 1973).

Large earthquakes have also occurred on adjacent faults, including the M_w 7.3 Sarpol-e Zahab earthquake in 2017 (Barnhart et al., 2018; Nissen et al., 2019; K. Wang & Bürgmann, 2020) which ruptured a dextral-thrust fault beneath the Lurestan arc, 100–200 km south of the MRF. Ground surface deformation from the earthquake may be observable up to the MRF, which would constitute a source of error in our velocity estimates. The oblique slip direction highlights the incomplete partitioning of convergence onto reverse and strike-slip faults in the western Zagros. Similarly, the 2008 and 2012 Moosiyan earthquake sequences, which occurred on the Zagros foredeep fault, caused seismic and aseismic strike-slip motion on structures other than the MRF (Nippress et al., 2017). An accurate interseismic slip rate for the MRF will allow for an improved assessment of the degree of strain localisation and partitioning on the MRF. A measurable vertical velocity contrast across the MRF would suggest dip-slip motion and incomplete partitioning, as would a lower slip rate than needed to complete the plate-circuit-closure.

3 Methods

3.1 InSAR Processing

We process a total of 1038 Sentinel-1 Interferometric Wide Swath images across two ascending tracks (174 and 101) and two descending tracks (006 and 108), with an average time span of 5.6 years from late 2014 to mid 2020 (Table 2). From these images, we produce a total of 4634 interferograms. Interferograms are formed between each acquisition and the three previous epochs to form a redundant network with minimised temporal baselines (Figure S1). Additionally, we produced long temporal baseline interferograms to bridge periods of low coherence (e.g. winter). We manually remove a total of 52 interferograms due to unresolved processing errors.

We generate interferograms using the LiCSAR system, a set of high-level tools and algorithms that operate the GAMMA SAR and Interferometry software (Werner et al., 2000; Wegnüller et al., 2016). The reader is referred to Lazecky et al. (2020) for an in-depth description of the processing system. Images are processed in predefined frames that average 13 bursts across each of the three subswaths that were acquired in Terrain Observation with Progressive Scans (TOPS) mode. We remove topographic contributions to the phase return using the 1 arc-second SRTM DEM (Farr et al., 2007). We unwrap each interferogram in two dimensions using the statistical-cost network-flow algorithm (SNAPHU) version 2 (Chen & Zebker, 2000, 2001, 2002). Interferograms are multilooked by 20 in range and 4 in azimuth (46.6×55.9 m) during the processing, and then further downsampled to 100 × 100 m pixels outside of GAMMA.

Atmospheric noise is often the largest source of error in InSAR data and results from spatially-correlated radar path delays as waves are refracted through the troposphere

Table 2. Time extent and number of processed SAR data for each track used in this study.

Track	Start	End	Duration (yr)	no. of epochs	no. of ifgs generated	no. of ifgs used	mean ifg length (d)
006D	2014-10-06	2020-07-06	5.75	196	1022	997	10.8
108D	2014-10-25	2020-07-13	5.72	204	1024	987	10.3
174A-N	2014-11-10	2020-06-23	5.62	219	984	939	9.5
174A-S	2014-11-10	2020-06-23	5.62	206	884	878	10.0
101A	2015-02-09	2020-07-06	5.41	213	720	665	9.3

Note. A = ascending orbital tracks (acquisition time 14:43-14:52 UTC, 18:13-18:22 IRST, 19:13-19:22 IRDT), D = descending orbital tracks (acquisition time 02:45-02:52 UTC, 06:15-06:22 IRST, 07:15-07:22 IRDT). Typically about 200 SAR images are used to produce 700-1000 interferograms (ifgs) (no. of ifgs generated) spanning a 6 to 12 day period (10 days on average), which are then reduced in number by various quality checks (no. of ifgs used). 174A-N and 174A-S refer to the northern and southern frames shown in Figure 1b, respectively.
UTC = Coordinated Universal Time, IRST = Iranian Standard Time, IRDT = Iranian Daylight Time

(Zebker et al., 1997; Parker et al., 2015). We mitigate this error using the Generic Atmospheric Correction Online Service for InSAR (GACOS) (Yu et al., 2017; Yu, Li, & Penna, 2018, 2018), which provides tropospheric delay maps derived from the European Centre for Medium-Range Weather Forecasts (ECMWF) upscaled through use of a DEM. These maps include both hydrostatic and wet components. ECMWF models are provided every six hours and can be interpolated to any SAR acquisition time in between. For each interferogram, the respective tropospheric zenith delay maps are differenced, projected into the satellite line-of-sight (LOS), and then subtracted from the interferogram. Figure 2 shows the change in the standard deviation of each interferogram after the GACOS correction has been applied. On average, 81% of interferograms for each frame show a decrease in the standard deviation, associated with a reduction in atmospheric noise, following the application of the GACOS correction, assuming that no tectonic signal is visible in the short time span interferograms.

Next, we use LiCSBAS, a small-baseline time-series analysis package, to generate cumulative line-of-sight displacements and average velocities from our interferograms (Morishita et al., 2020; Morishita, 2021). We further downsample our 100×100 m interferograms to 1 km to reduce processing requirements while retaining sufficient resolution to capture short-wavelength tectonic signals. Interferograms with low average coherence (< 0.05) and low coverage (< 0.3) are identified and removed. We identify phase unwrapping errors by calculating the loop closure phase, Φ of every interconnected image triplet following Equation 1 (Biggs et al., 2007):

$$\Phi_{123} = \phi_{12} + \phi_{23} - \phi_{13} \tag{1}$$

where ϕ_{12}, ϕ_{23}, and ϕ_{13} are the interferograms formed from SAR images ϕ_1, ϕ_2, and ϕ_3. Near-zero values of Φ for a triplet indicate that the unwrapping is consistent between all three interferograms, while values near integer multiples of 2π indicate the presence of unwrapping errors in at least one interferogram. We calculate the root mean square (RMS) of the loop phase image for every triplet and remove a total of 116 interferograms where the RMS is greater than 1.5 rad for every loop that the interferogram is a part of. Cumulative LOS displacements are inverted for a linear velocity on a pixel-by-pixel basis following the NSBAS least-squares method (Doin et al., 2011), using the 4466 in-

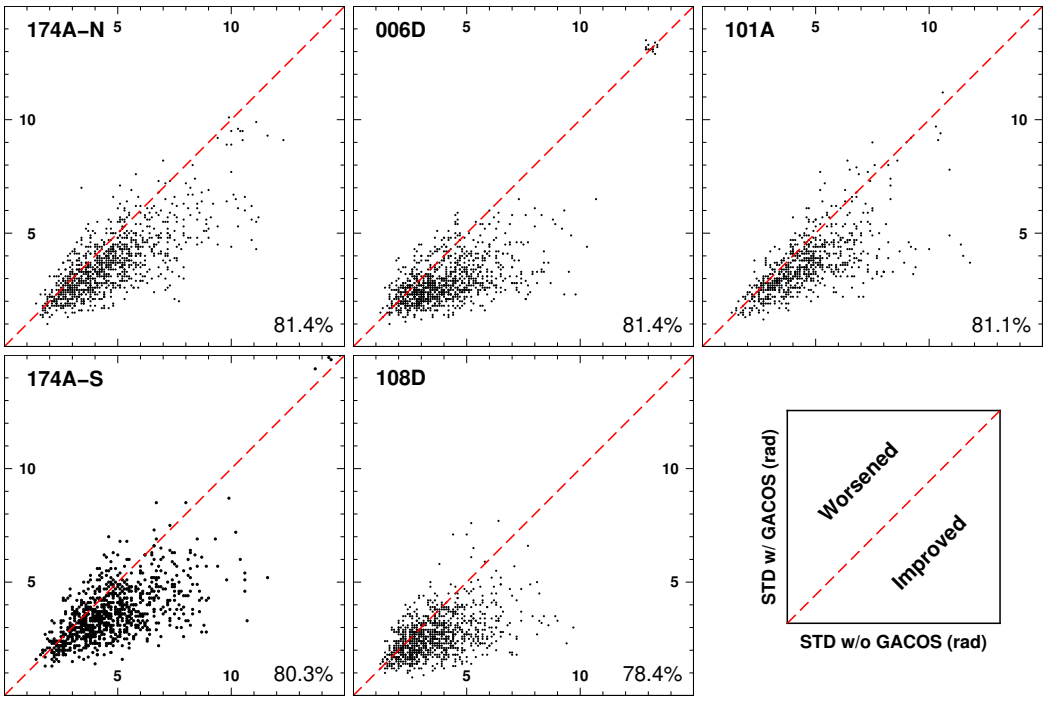

Figure 2. Change in the standard deviation (SD) of all pixels in each interferogram resulting from the application of the GACOS correction, with black dots below the red dashed line indicating an improvement in terms of interferogram noise from the application of the atmospheric model. The percentage of interferograms showing a reduction in SD for each frame is given in the bottom right of each subpanel. The distinct high SD clusters seen in 006D and 174A-S are associated with the M_w 7.3 earthquake.

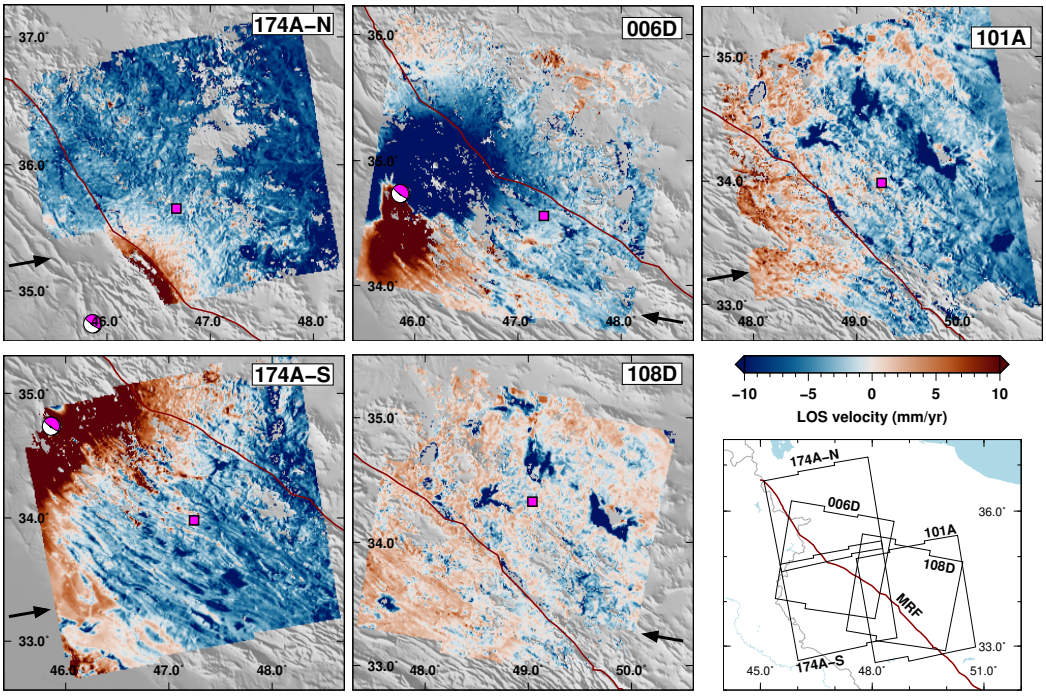

Figure 3. Average line-of-sight velocities for ascending and descending tracks at 1 km resolution. Black arrows show the horizontal projection of LOS vector. Red and blue indicate motion towards and away from the satellite, respectively, relative to a reference pixel (pink squares). The effect of the M_w 7.3 2017 Sarpol-e Zahab mainshock (pink focal mechanism) is clearly visible in frames 174A-N, 174A-S and 006D as the large positive and negative velocity areas saturated in this image. A number of subsiding basins are visible as regions with rates faster than -10 mm/yr.

terferograms that passed the quality checks. In the case of missing observations in the displacement time series (e.g. incoherence, missing acquisitions, masked pixels), LiCSBAS imposes a linear temporal constraint across the gap. These estimated displacements may be unreliable if the displacement series deviates significantly from a linear function. Therefore, we avoid network gaps where possible by generating additional interferograms. The average velocities are referenced to a stable pixel in each frame, chosen by calculating the RMS of all the loop closure phases for each pixel and selecting the lowest, with the requirement that the pixel must be unmasked in every interferogram. We estimate the uncertainty on the velocities by applying the percentile bootstrap method (Efron & Tibshirani, 1986) to the cumulative displacements for each pixel. Each displacement series is randomly resampled with data replacement 100 times and the velocity is re-calculated. The standard deviation of the final 100 velocities is taken as the uncertainty on the LOS velocity. A number of statistical quality checks are used to threshold a mask for the velocity map (Figures S2–S6). The LOS velocity maps for each frame are shown in Figure 3. Example displacement series for two pixels, one across-fault difference and one subsidence signal, are shown in Figures S7 and S8.

3.2 Evaluation of Co- and Post-seismic Signals

Tracks 174A and 006D span (both spatially and temporally) the 12 November 2017 M_w 7.3 Sarpol-e Zahab earthquake (Nissen et al., 2019) and subsequent aftershock sequences (Lv et al., 2020). The earthquake involved dextral-thrust slip on a 40×20 km basement fault in the Lurestan arc, potentially triggering aftershocks up to 80 km away.

Displacements of up to 90 cm in LOS were observed in Sentinel-1 interferograms spanning the event (Nissen et al., 2019). The first main aftershock sequence occurred on 11 January 2018, consisting of five events between M_w 5.1–5.5 (Lv et al., 2020). The second occurred on 25 November 2018 and consisted of M_w 6.3, 5.2, and 5.0 events. In the one year period following the mainshock, an additional 100 mm of LOS displacement was observed related to postseismic deformation (K. Wang & Bürgmann, 2020). For areas affected by the Sarpol-e Zahab earthquake, any estimate of the interseismic slip rate will be biased by the co- and post-seismic displacements (Figure 3). A full assessment of the magnitude and spatial extent of these signals in the InSAR velocity field is required so as to be able to robustly estimate the interseismic slip rate.

We first attempt to mitigate the coseismic signals by forwarding modelling the displacements and subtracting these from the cumulative displacements generated by LiCSBAS. For the mainshock, we use fault parameters estimated by Nissen et al. (2019) from Sentinel-1 interferograms and a uniform slip fault model (Table S1). We model both aftershock sequences as a single event using InSAR-derived fault parameters from Lv et al. (2020). The fault is modelled using a rectangular dislocation source (Okada, 1985) defined by nine parameters: x-position, y-position, strike, dip, rake, slip, fault length, top depth, and bottom depth. We chose InSAR-derived fault parameters, as opposed to those from seismology, as they may more accurately fit our own observed InSAR signals. Modelled surface displacements are projected into the satellite LOS for each frame and then removed from the cumulative displacement series.

Next, we calculate the change in the average LOS velocity following the 12 November 2017 M_w 7.3 mainshock. To do this, we split the cumulative displacement time series produced by LiCSBAS into two parts about 12 November 2017 and solve for the average velocity pre- and post-earthquake using least squares. We do not attempt to remove the post-seismic signal from the times series (K. Wang & Bürgmann, 2020), as the signal is difficult to separate from the noise and any errors could be of a similar magnitude to the interseismic signal. Pixels for which the average velocity changed significantly following the earthquake may have been affected by post-seismic deformation. The reduction in time series length to 2-3 years either side of the earthquake will increase the velocity uncertainties to 4-5 mm/yr (Morishita et al., 2020). Additionally, the time series before the earthquake contain fewer interferograms because Sentinel-1B was inactive for some of this time period (2014–2017). We calculate the velocity difference for the three affected frames and combine them by averaging overlapping pixels. We estimate the expected velocity difference as a result of noise and non-tectonic signals by calculating the velocity difference for 108D, which should be unaffected by the Sarpol-e Zahab earthquake sequence. We calculate a standard deviation of 11.4 mm/yr, and contour the merged velocity differences based on the 95% confidence interval. Figure 4 shows the merged and contoured velocity differences for frames 006D, 174A-N, and 174A-S. The primary post-seismic signal is highlighted by the pink square and covers an area of approximately 100×150 km. We observe a similar spatial extent to the cumulative post-seismic displacements observed by K. Wang and Bürgmann (2020, Figure 4). The true extent of the post-seismic deformation likely extends further than the highlighted area, given the uncertainty in our velocities and the threshold used, and so we avoid velocities west of 47° E when selecting profile lines.

3.3 Velocity Field Generation

Our initial LOS InSAR velocity fields are referenced relative to a stable pixel for each frame (pink squares in Figure 3). To better combine all four tracks, we shift the LOS velocities into a Eurasia-fixed reference frame (Hussain, Hooper, et al., 2016; Hussain et al., 2018; Weiss et al., 2020). Using horizontal GNSS velocities provided by Khorrami et al. (2019), we fit second-order polynomial surfaces to the East and North velocity components within one degree of our study area (Figure 5). The velocities are weighted us-

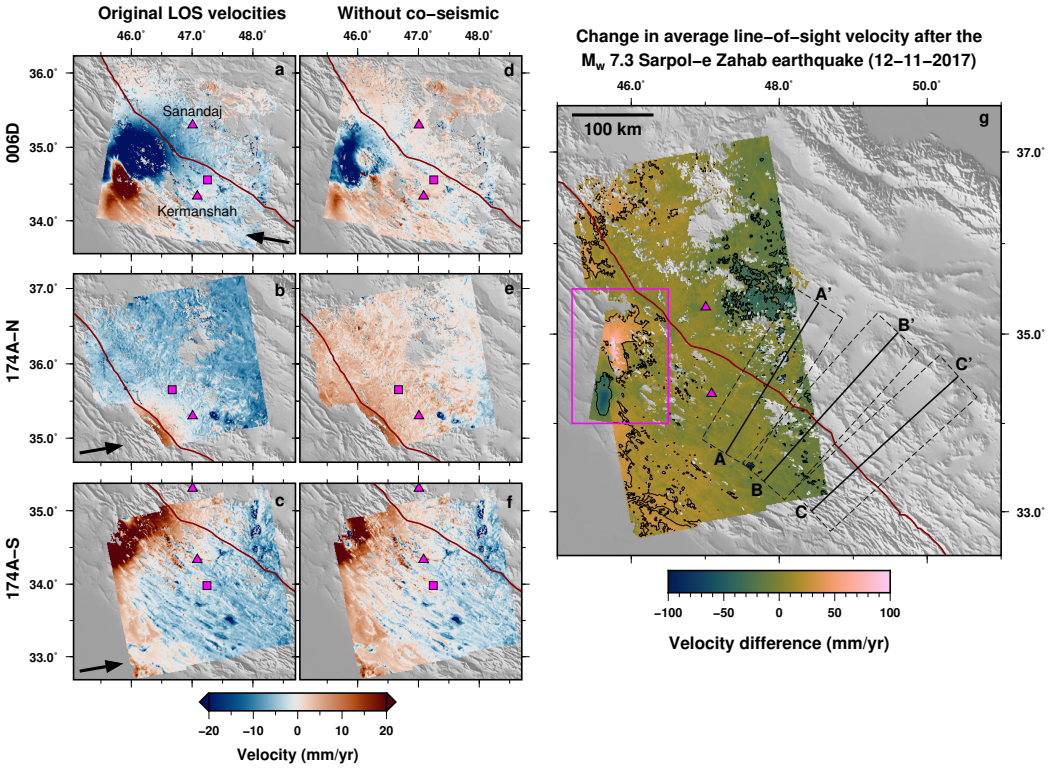

Figure 4. Analysis of co- and post-seismic signals following the M_w 7.3 2017 Sarpol-e Zahab earthquake sequence. (a-c) Original line-of-sight velocities for frames 006D, 174A-N, and 174A-S, relative to the reference pixel (pink square). Two cities, Sanandaj and Kermanshah, are marked as pink triangles for spatial reference, along with the trace of the MRF (red). Black arrows show the horizontal projection of the LOS vector. (d-f) LOS velocities after the forward models for the earthquakes shown in Table S1 have been removed from the displacement time series, relative to the reference pixel (pink square). (g) Difference in average line-of-sight velocities before and after the 12 November 2017 M_w7.3 mainshock, for frames 006D, 174A-N, and 174A-S. For overlapping pixels between frames, the values have been averaged. The values are contoured at ±23 mm/yr, based upon the 2σ value derived from frame 108D which is unaffected by the earthquake deformation. The pink rectangle highlights the region encompassing the main post-seismic signal. We choose profiles (A-A', B-B', and C-C') that avoid areas with significant velocity changes following the earthquake.

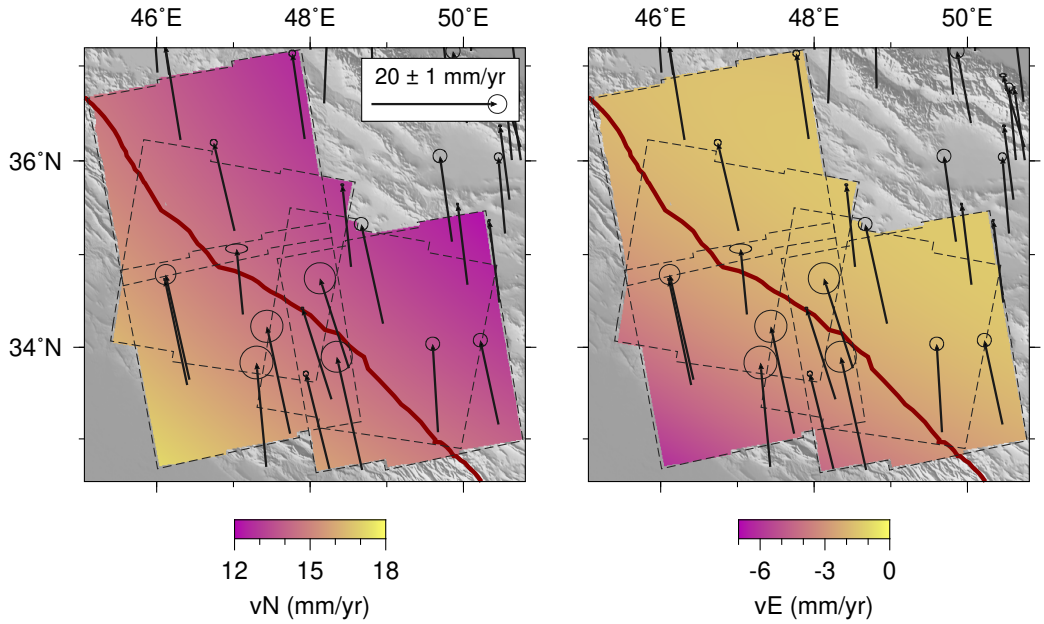

Figure 5. Spatially interpolated GNSS velocity fields generated by fitting a 2nd-order polynomial plane to North (vN) and East (vE) GNSS velocities from Khorrami et al. (2019) relative to a stable Eurasia, cropped to the area covered by the InSAR. GNSS velocities are located at the base of the arrows and given with 1σ uncertainties.

ing their respective bootstrapped uncertainties. We then project the GNSS velocity fields are projected into the satellite LOS for each frame. We calculate the residual between the projected GNSS velocities and the InSAR velocities, and fit a second-order polynomial surface to the result. Subtracting this function from the respective InSAR velocity field results in InSAR velocities in the same reference frame as the GNSS velocities.

To investigate interseismic slip along the MRF, we decompose our satellite LOS velocities into local geodetic coordinate velocities. The velocity for each pixel observed in the satellite LOS can be expressed as a linear combination of the East, North, and Up components:

$$V_{LOS} = \begin{bmatrix} sin(\theta)cos(\alpha) & -sin(\theta)sin(\alpha) & -cos(\theta) \end{bmatrix} \begin{bmatrix} V_E \\ V_N \\ V_U \end{bmatrix} \tag{2}$$

where θ is the radar incidence angle, measured from the vertical to the LOS, and α is the azimuth of the along-track satellite heading. The majority of our study area is covered by two overlapping tracks, one ascending and one descending. In this situation, we have two observations (V_{asc} and V_{desc}) and three unknowns (V_E, V_N, and V_U), making the inverse problem under-determined. In order to find a unique solution to the problem, we must add either further observations or a-priori constraints for one of the model parameters. Sentinel-1 InSAR observations are particularly insensitive to displacement in the north-south direction, as a result of the near-polar satellite orbit and sidewards look direction. We estimate the north contribution to the InSAR LOS velocities by projecting the interpolated north GNSS velocity field (Figure 5) into the respective satellite LOS for each track. This projected velocity is then subtracted from each frame, leaving LOS velocities that contain a negligible long-wavelength north-south component. For a point with observations from two look directions, the resulting simplified linear equa-

tion is given by Equation 3.

$$\begin{bmatrix} V_{asc} \\ V_{desc} \end{bmatrix} = \begin{bmatrix} -cos(\theta_{asc}) & sin(\theta_{asc})cos(\alpha_{asc}) \\ -cos(\theta_{desc}) & sin(\theta_{desc})cos(\alpha_{desc}) \end{bmatrix} \begin{bmatrix} V_U \\ V_E \end{bmatrix} \tag{3}$$

In the presence of three or more overlapping frames, the 2-by-2 matrix in Equation 3 can be expanded to a n-by-2 matrix, where n is the number of overlapping frames. We solve Equation 3 using weighted least squares and the data variance-covariance matrix (VCM). The data VCM is a n-by-n matrix containing the bootstrapped variance values generated by LiCSBAS for a given point, and assuming no covariance between the same point in different frames. We also calculate the model VCM following Equation 4:

$$Q_m = (G'WG)^{-1} \tag{4}$$

where Q_m is the 2-by-2 model VCM, G is the design matrix (the n-by-2 matrix in Equation 3), and W is the inverse of the data VCM. Uncertainties for the decomposed velocities (Figure S9) are typically in the range of 0.2–1.0 mm/yr in the East component and 0.5–1.5 mm/yr in the Up component.

3.4 Interseismic Fault Slip Modelling

We model profiles of the InSAR-derived velocities across the deforming zone by assuming that the fault can be approximated as a vertical buried 1-D screw dislocation (Savage & Burford, 1973) defined by Equation 5:

$$v_{para}(x) = \left(\frac{s}{\pi}\right) arctan\left(\frac{x}{d}\right) + c \tag{5}$$

where $v_{para}(x)$ is the horizontal surface velocity parallel to the fault, x is the perpendicular distance from the fault, s is the interseismic slip rate, d is the locking depth, and c is a scalar offset. We calculate fault-parallel velocities using the decomposed east InSAR velocities and the north component of the GNSS velocity field, using the local strike within each profile. We take three 70 km wide fault-perpendicular profiles across the MRF (A-A', B-B', and C-C'), avoiding areas with post-seismic signals (Figure 4). We solve for the interseismic slip rate, the locking depth, and a scalar offset to the velocities (Equation 5), assuming strikes of 300°, 310°, and 315° for A-A', B-B', and C-C', respectively. We fix the fault location based upon the intersection of the profile lines with the fault trace from Walker et al. (2010), and apply a scalar offset to the velocities so that the profiles are centred on approximately 0 mm/yr where they intersect the fault trace.

We estimate best fit values for each model parameter by implementing an affine-invariant ensemble Markov Chain Monte Carlo (MCMC) sampler, developed by Goodman and Weare (2010). This Bayesian approach uses semi-random walks of a given number of walkers to explore the posterior probability distribution of the model, based upon known prior constraints. Solutions are ranked using the weighted misfit between observed and model velocities. It demonstrates improved performance over traditional Metropolis-Hasting algorithms, especially in the presence of complex parameter spaces (Goodman & Weare, 2010). This method has been widely used for tectonic applications (Hussain, Wright, et al., 2016; Hussain, Hooper, et al., 2016; Szeliga & Bilham, 2017; Aslan et al., 2019; Goto et al., 2019; Tesson et al., 2021). Our MCMC sampler uses 600 walkers and runs for 300,000 iterations. We remove the first 20% of solutions as burn-in, producing 48,000 valid solutions. From these we calculate the maximum a posteriori probability (MAP) solution - i.e. the most likely solution based upon the prior probabilities - and uncertainties for each model parameter. We assume a uniform prior for all model parameters based upon limits of $-10 \leq s \leq 20$ (mm/yr), $1 \leq d \leq 50$ (km), and $-10 \leq c \leq 10$ (mm/yr).

To account in the inversion for the noise of the data and correlation between nearby pixels, we calculate the spatial covariance function of the data after removing tectonic and anthropogenic signals (e.g. Hussain, Hooper, et al., 2016; Weiss et al., 2020). We

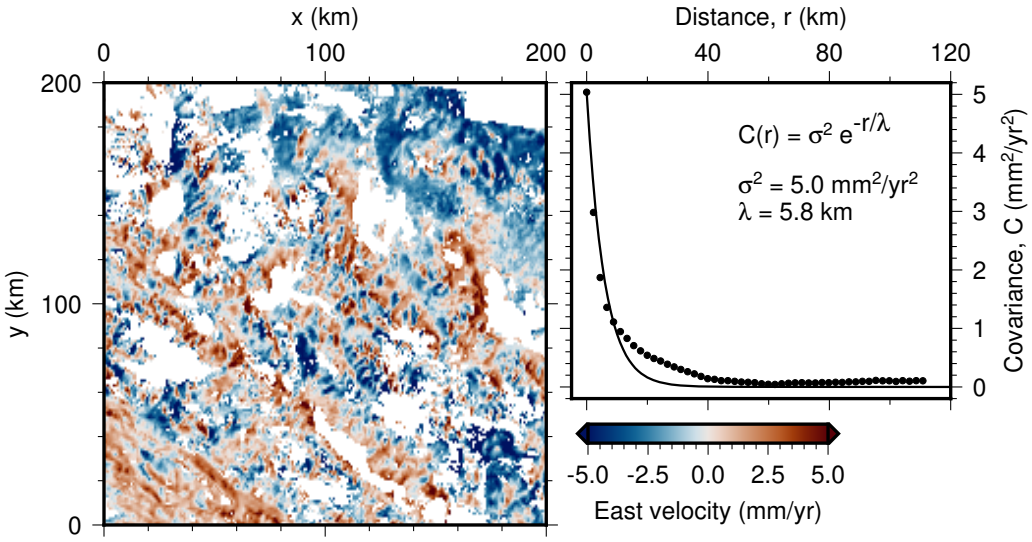

Figure 6. Autocorrelation function and best-fitting exponential (Equation 6) (right) based upon isolated non-tectonic and non-anthropogenic noise in the East InSAR average velocities (left).

take a 200×200 km region of the decomposed East velocities, avoiding the post-seismic signals in the west of the study area, and mask out any pixels associated with a vertical rate greater than ±5 mm/yr; these signals are largely correlated with basins and are likely due to groundwater subsidence. We forward model the interseismc slip on the MRF using Equation 5, assuming a slip rate of 3 mm/yr below a locking depth of 20 km, and subtract the East component of this model from our decomposed east velocities. Finally, we remove a first-order polynomial plane from the velocities to account for any residual tectonic signal or orbital ramps. The resulting velocities (Figure 6) should contain negligible tectonic and anthropogenic signals, with the majority of the velocity field consisting of short-wavelength residual atmospheric noise that remained after the GACOS correction (Murray et al., 2019). We fit an exponential radial covariance function (Hussain, Hooper, et al., 2016) to these data of the form:

$$C(r) = \sigma^2 e^{-\frac{r}{\lambda}} \tag{6}$$

where $C(r)$ is the covariance between two velocity measurements at a distance of r, σ^2 is the variance, and λ is the exponential length scale. We solve Equation 6 for σ^2 and λ, estimating values of 5.0 mm^2/yr^2 and 5.8 km, respectively (Figure 6). The misfit between r values of 10–40 km, where the covariance model underestimates the observed decay in noise at mid distances, relates to asymmetry within the noise structure. This is potentially as a result of NW-SE aligned topographic structures (mountain ranges and interleaved valleys) with similar widths and lengths. We generate a variance-covariance matrix for all pixels within each profile using Equation 6 and our estimated exponential parameters, where r is the 2-D distance between pixels, and use this to weight our Bayesian inversion.

4 Results

Figure 7 shows the decomposed East, vertical, and fault-parallel velocities, the latter of which was calculated assuming a fixed strike of 310°. We observe higher velocities (red) to the northeast in both the East and fault-parallel velocities, with a veloc-

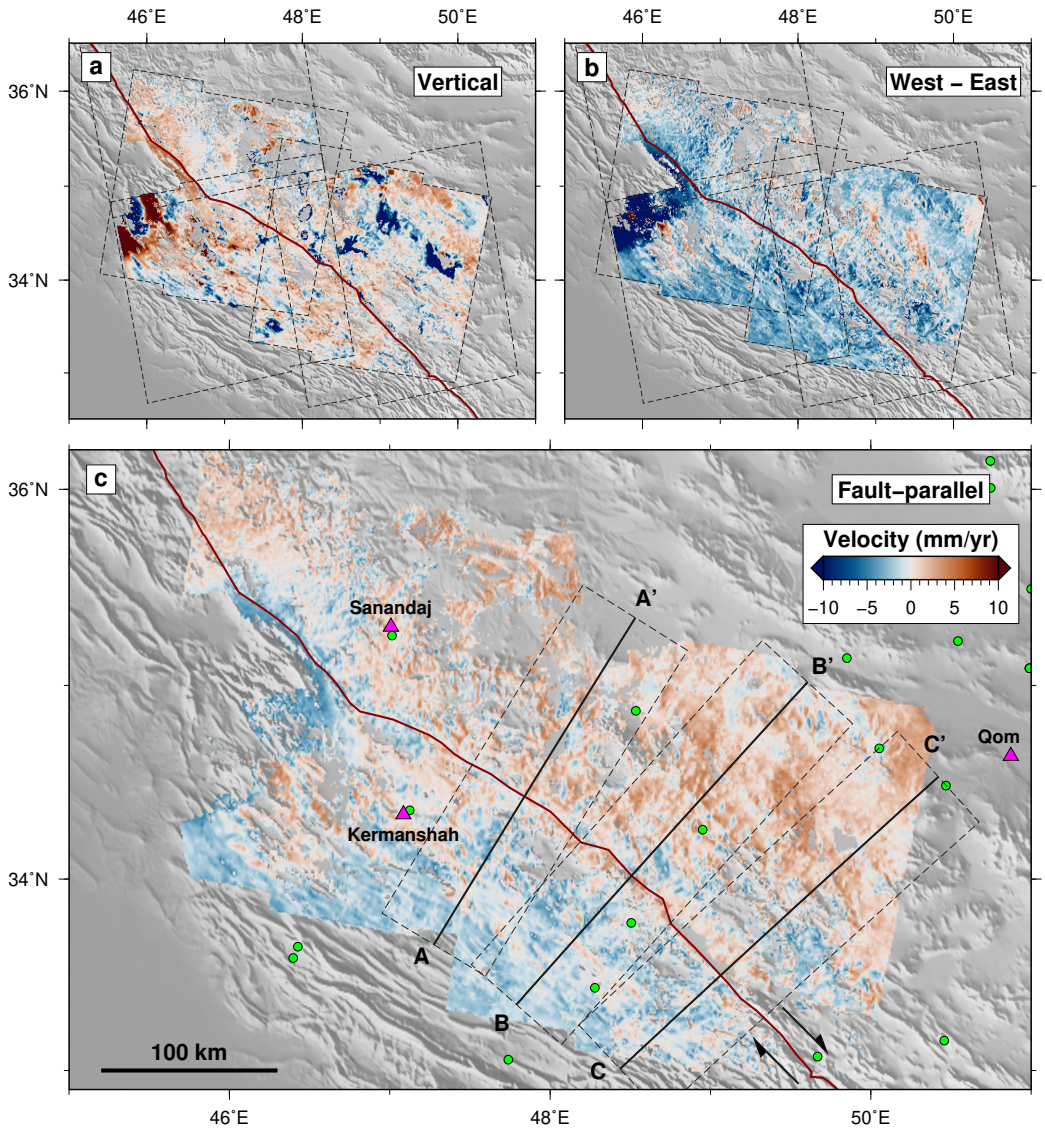

Figure 7. Decomposed Vertical (a) and East (b) velocities, and fault-parallel (c) velocities assuming an overall regional fault strike of 310°, relative to the MRF. Positive values (red) indicate motion upwards, to the east, and in a right-lateral sense, respectively. Fault-perpendicular profiles (black lines) are taken perpendicular to the projected MRF trace (red line). Velocities are overlain onto hill-shaded SRTM topography, with city locations (pink triangles) and GNSS locations (green circles) marked for reference.

Table 3. Maximum a posteriori probability estimates and uncertainties of the parameters in Equation 5 for the three profiles shown in Figure 7, for both independent (Ind) profiles and the three-profile joint model with a shared locking depth.

		A-A'		B-B'		C-C'	
Ind	s (mm/yr)	2.6	(1.2–4.6)	2.9	(1.6–4.7)	3.5	(1.0–4.7)
	d (km) (95% IQR)	18.4	(5.6–48.9)	19.4	(7.1–48.6)	49.7	(unbounded)
	c (mm/yr)	-0.0	(-0.5–0.4)	0.2	(-0.3–0.6)	-0.8	(-1.2– -0.2)
Joint	s (mm/yr)	2.8	(1.2–4.8)	3.1	(1.5–4.7)	2.6	(1.1–4.7)
	d (km)			29.0	(18.7–41.9)		
	c (mm/yr)	-0.1	(-0.8–0.8)	-0.2	(-0.6–0.9)	-0.7	(-1.4–0.3)

Note. Results are given with 95% confidence intervals with the exception of the locking depths for the individual A-A' and B-B' results, which are given with 95% interquantile ranges (IQR), as these distributions are significantly non-Gaussian (Figures 9).

ity difference across the MRF of several millimeters per year. The velocity difference is lower west of 47.5° E where we expect the velocities to be biased by the Sarpol-e Zahab earthquake sequence (Figure 4). No equivalent velocity contrast can be seen in the vertical velocities. The large negative (blue) velocities previously observed in the LOS velocities are now only present in the vertical velocities, supporting the idea that these are subsidence signals related to groundwater extraction.

Figure 8 shows the results of the MCMC inversion, and Figure 9 shows the marginal posterior probability distributions for each parameter. MAP solutions and uncertainties for each parameter are given in Table 3. We include the mean and standard deviation of the fault-parallel velocities (Figures 8 and 10), weighted using the variance of the East velocities (Figure S6) and calculated in a 10 km window moved in 1 km increments along the profiles.

Our MAP estimates of slip rate are consistent between profiles A-A' (2.6 mm/yr), B-B' (2.9 mm/yr), and C-C' (3.5 mm/yr), to within the 95% confidence intervals (1.2–4.6, 1.6–4.7, and 1.0–4.7 mm/yr, respectively) (Table 3). We see similar consistency between our MAP estimates of locking depth for profiles A-A' (18.4 km) and B-B' (19.4 km). The posterior distributions for both are either skewed (B-B') or non-Gaussian (A-A'), and so we provide the 95% interquantile range (IQR), defined as the difference between the 0.025 and 0.975 quantiles, for A-A' (5.6–48.9 km) and B-B' (7.1–48.6 km) as this is more representative. The upper locking depth for B-B' may also be considered unbounded, given that the distribution levels off above 20 km. Profile C-C' may be considered unbounded within the parameter range we have chosen, tending towards a very deep locking depth.

To better constrain the locking depth, we subsequently model the velocities across all three profiles simultaneously. We estimate a shared locking depth across all three profiles and keep separate solutions of profile-specific slip rates and offsets, otherwise repeating the previous setup. We would expect the locking depth to be approximately constant across relatively short section of a major fault. Figure 10 shows the results of the MCMC inversion, Figure 11 shows the marginal posterior probability distributions for each parameter, and Table 3 summarises the MAP estimates and uncertainties. Our MAP estimates of slip rate are again consistent across profiles A-A' (2.8 mm/yr), B-B' (3.1 mm/yr), and C-C' (2.6 mm/yr) to within the 95% confidence intervals (1.2–4.8 mm/yr, 1.5–4.7 mm/yr, and 1.1–4.7 mm/yr, respectively). The posterior distribution for our joint estimate of

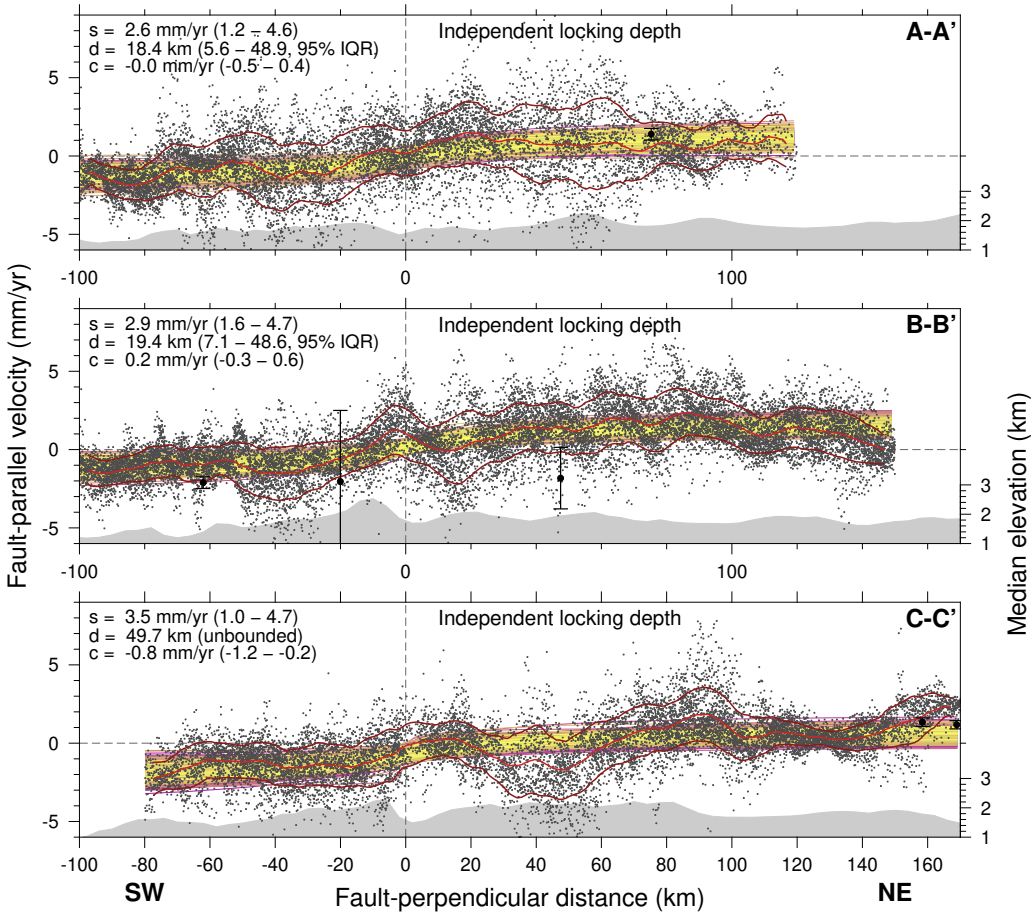

Figure 8. Northeast-southwest profiles of fault-parallel (strike direction) velocity (with positive motion to the SE shown by the black dots) relative to the surface trace of the MRF (set at zero distance), from Figure 7, modelled with independent locking depths. The light and dark red lines show the weighted average and weighted standard deviation of the velocities, respectively. A uniformly randomly selected 1% of modelled solutions are shown ranked by the posteriori probability, from yellow (best) to pink (worst). The MAP solution is given in the top left of each panel with 95% confidence intervals (CI) centred on the mean, with the exception of the locking depths which are given with the 95% interquartile range (IQR) centred on the median. The median elevation and GNSS velocities within each profile are shown in grey and as black circles with error bars, respectively.

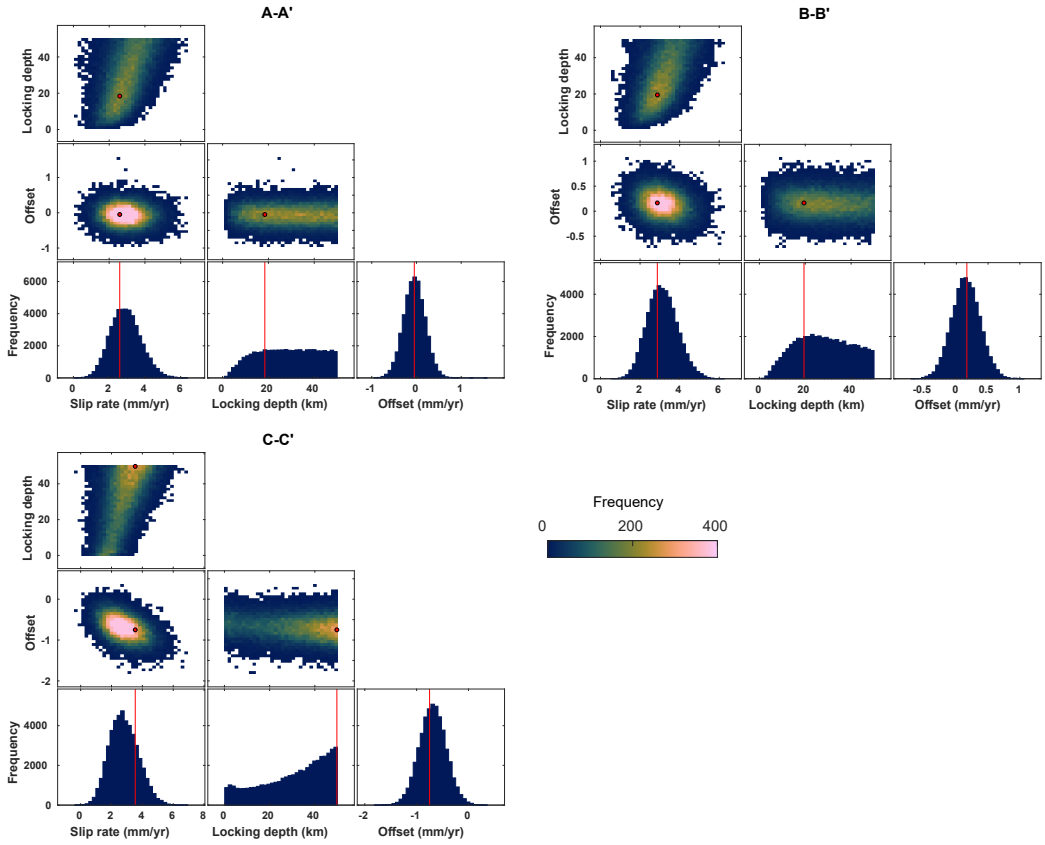

Figure 9. Marginal probability distribution for the profiles in Figure 8, including the MAP solution (red line and dot), for the MCMC inversion (48,000 solutions) under the assumption of independent slip rates, locking depths, and offsets for each profile. We assume a uniform prior for all model parameters based upon limits on the values of slip rate (-10–20 mm/yr), locking depth (1–50 km), and offset (-10–10 mm/yr).

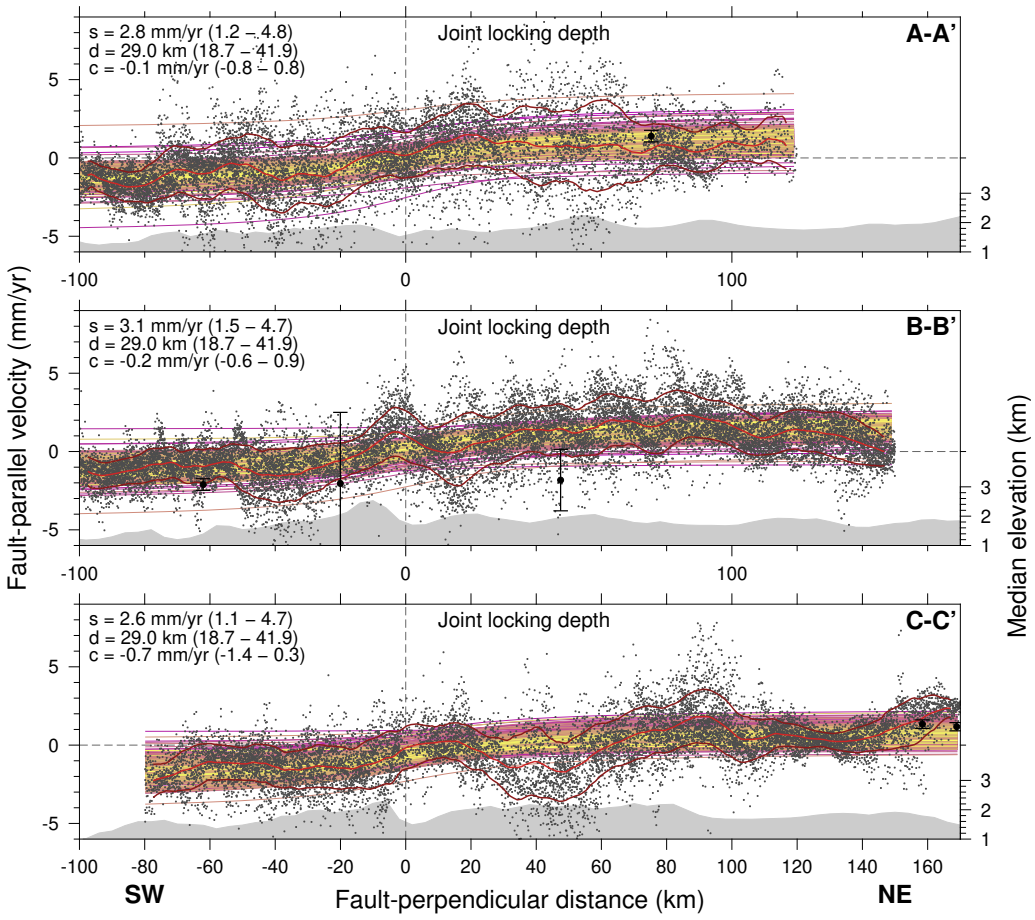

Figure 10. Northeast-southwest profiles of fault-parallel (strike direction) velocity (with positivie motion to the SE shown by the black dots) relative to the surface trace of the MRF (set at zero distance), from Figure 7, modelled with a joint locking depth and individual slip rates and offsets. The light and dark red lines show the weighted average and weighted standard deviation of the velocities, respectively. A uniformly randomly selected 1% of modelled solutions are shown ranked by the posteriori probability, from yellow (best) to pink (worst). The MAP solution is given in the top left of each panel with 95% confidence intervals (CI) centred on the mean. The median elevation and GNSS velocities within each profile are shown in grey and as black circles with error bars, respectively.

locking depth (29.0 km) is now Gaussian, giving a 95% confidence interval of 18.7–41.9 mm/yr. This interval only encompasses the individual locking depth estimate for profile B-B' (19.4 km), although the joint estimate of 29.0 km is within the 95% IQR for both profiles A-A' and B-B'. While not directly comparable, the uncertainties on the joint estimate of locking depth are narrower than those on the individual estimates. The distribution appears more Gaussian, with a small skew to higher values. While the bounds are more defined than the individual estimates, a high kurtosis value of 5.5 (compared to an average of 3.6 for the slip rate distributions) indicates long tails on the distribution.

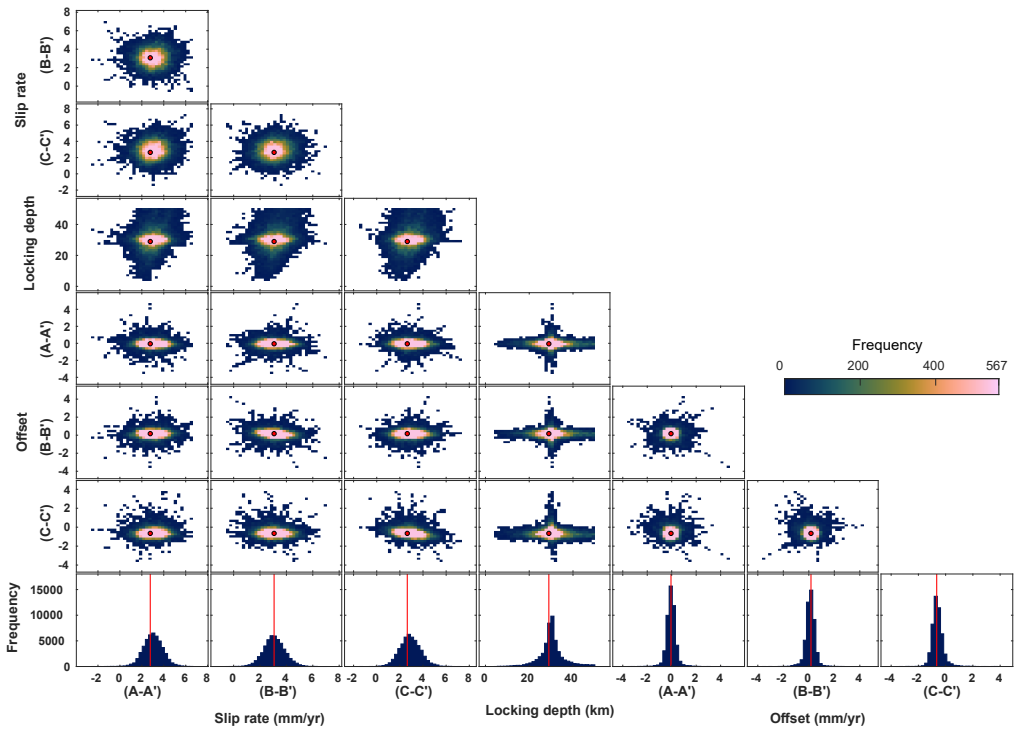

Figure 11. Marginal probability distribution for the profiles in Figure 10, including the MAP solution (red line and dot), for the MCMC inversion (48,000 solutions) under the assumption of a single locking depth along strike of the MRF (i.e. a common joint locking depth for all three profiles). We assume a uniform prior for all model parameters based upon limits on the values of slip rate (-10–20 mm/yr), locking depth (1–50 km), and offset (-10–10 mm/yr).

5 Discussion

We have resolved a small rate of strain accumulation at the millimeter per year level across spatial distances of 100 km over the Main Recent Fault from 5.5 years of Sentinel-1 InSAR times series. Furthermore, we have placed geodetic bounds for the first time on the depth below which the fault is slipping in this interseismic period. Below we compare our results to previous estimates of interseismic slip rate, discuss the model uncertainties, which remain high despite the large volume of data used here, and also discuss challenges in constraining the locking depth for relatively noisy interseismic InSAR datasets.

5.1 Estimation of the Locking Depth

Our results provide the first geodetic estimate for the locking depth of the Main Recent Fault, SW Iran. Our three estimates of locking depth from profiles A-A' (18.4 km), B-B' (19.4 km), and the joint model (29.0 km), all lie within each other's large uncertainties (5.6–48.9 km, 7.1–48.6 km, and 18.7–41.9 km). While the upper limit is poorly constrained for both A-A' and B-B', and C-C' is unbounded within our prior limits, we show that a joint inversion of multiple along-strike profiles with variable strike angles can better constrain the locking depth, both in terms of the upper bound and the width of the uncertainties. The improved upper bound on our joint estimate of the locking depth may be a result of the averaging of the long wavelength trend either side of the fault, which eliminates some unrealistically deep values. Our 95% confidence interval for the joint estimate of 18.7–41.9 km, equal to a standard deviation of 5.8 km, is in line with typical locking depth uncertainties from InSAR (e.g. Karimzadeh et al., 2013; Walters et al., 2011; Karimzadeh et al., 2013), although narrower uncertainties are achievable with higher slip rates (e.g. H. Wang et al., 2009; Fattahi & Amelung, 2016). Calibrated earthquake locations for the MRF from Karasözen et al. (2019) show a maximum centroid depth of roughly 20 km, suggesting that the fault is locked to around this depth. Examining the fault-parallel projected GNSS velocities from Khorrami et al. (2019) shown in Figures 8 and 10, we can see that the local GNSS network lacks suitable station density around the fault trace to capture the velocity gradient, and thus estimate the locking depth.

The poorly constrained upper bounds on the locking depth for profiles A-A' and B-B' suggest that the fit of our screw dislocation model to our fault-parallel velocities is relatively insensitive to the choice of locking depth. Similarly, the MAP estimate of slip rate for C-C' (3.5 mm/yr) is within the 1σ uncertainties of the slip rate estimates from A-A' and B-B', despite the large differences in locking depth (49.7 km versus 18.4 km and 19.4 km, respectively).

The wavelength of the velocity gradient across the fault trace is primarily controlled by the locking depth (Equation 5). Most of the strain is accommodated within a distance either side of the fault that is similar to a few times the locking depth. Figure 12 shows the weighted least squares estimates of slip rate for fixed values of locking depth between 1 and 50 km, along with the normalised RMS misfit between the model (Equation 5) and the fault-parallel velocities, for each individual profile and with a joint locking depth as in Figure 11. We weight the least squares with the same variance-covariance matrix as used for the Bayesian analysis. For all three profiles and the joint model, we observe a strong trade-off between the interseismic slip rate on the fault and the locking depth, from 1–1.5 mm/yr at 1 km to 3.4–3.7 mm/yr at 50 km. The magnitude of this trade-off is greatest for locking depths below 10 km. The normalised RMS misfit shows large minimums for profiles A-A' (7–20 km), B-B' (14–35 km), and the joint model (18–50 km). The misfit for C-C' does not define a flat minimum within the 1–50 km locking depth range, in agreement with the unbounded distribution show in Figure 11. Comparing Figure 12 to the a posteriori distributions shown in Figure 9, we can see that the choice of a deeper locking depth has little impact on the overall model fit, and that MAP estimates of locking depth are strongly controlled by the slip rate.

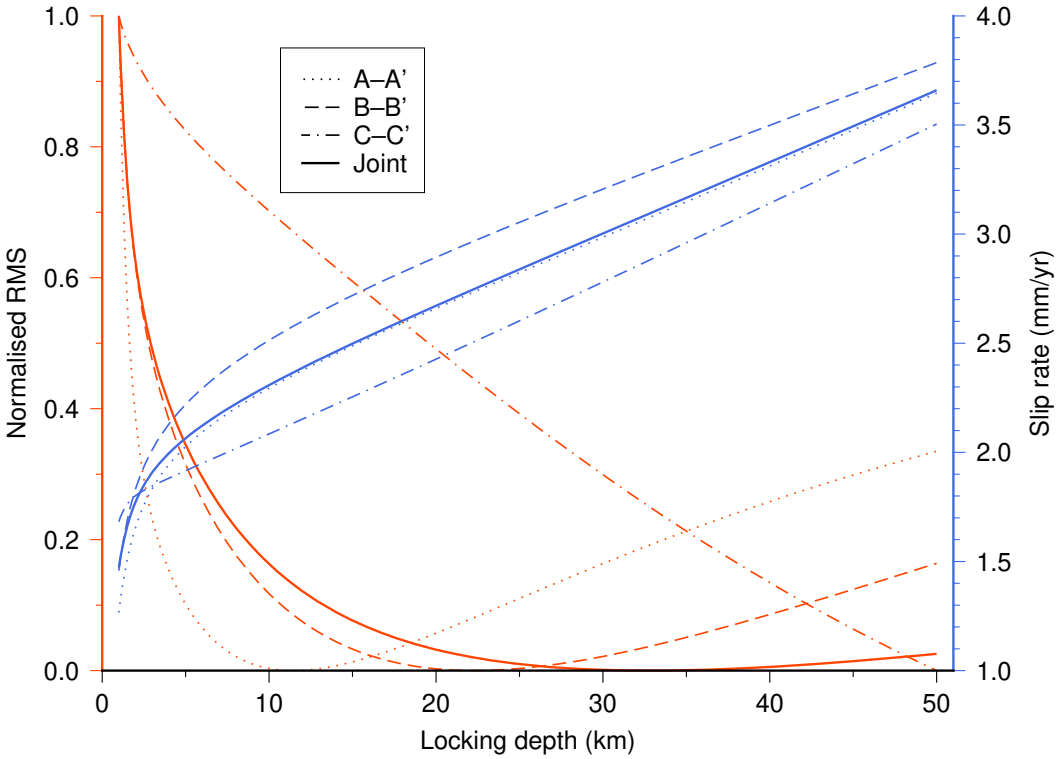

Figure 12. Trade-off between the weighted least-squares estimate of slip rate (blue lines) and fixed values of locking depth for profiles A-A' (dotted line), B-B' (dashed line), C-C' (dot-dashed line), and the joint profiles (solid line). In the case of the joint profiles, where we solve for three slips and three offsets, the average slip for all three profiles is shown. The normalised root mean square (RMS) misfit (red lines) between the forward model and the observations highlights a broad minimum for both B-B' and the joint profiles upwards of 15 km, while A-A' gives a lower and narrowed minimum, and C-C' does not reach a minimum.

5.2 Previous Estimates of the Interseismic Slip Rate

We produce two MAP estimates of the interseismic slip rate for each profile, using both individual and joint locking depths. These range between 2.6–3.5 mm/yr and 2.6–3.1 mm/yr, respectively. The narrower range for the joint model relates primarily to profile C-C', for which the slip rate for the individual profile is shifted to a higher value by the deep locking depth. The confidence intervals are almost identical between the individual and joint estimates. In both cases, the minimum of the confidence interval is above 0 mm/yr, and so we can be confident that the fault is actively accumulating strain in a right-lateral sense. For the individual profile estimate, we calculate a mean slip rate of 3.0 ± 1.0 mm/yr (2σ). We do not calculate a mean and uncertainty from the joint slip rates, as the shared locking depth means that these estimates are not independent. However, the range of estimates (2.6–3.1 mm/yr) is in agreement with the average individual rate.

These slip rate estimates are comparable to some of the slowest geodetically observed interseismic slip rates in the literature. Bell et al. (2011) estimated a rate of 3± 2 mm/yr for the Manyi fault, Tibet, using 10 ERS images over 5 years to form long-period interferograms. Mousavi et al. (2015) estimated a similar rate of 4.75±0.8 mm/yr for the Shahroud fault, northeast Iran, from 45 Envisat images over a 7 year period. Both studies highlight the importance of the length of the time series over the number of images when resolving slow slip rates. The low slip rate estimate for the Manyi fault was possible partly due to the lack of significant atmospheric noise, both because of low topographic variation along the profile, and because of the arid climate of the high Tibetan Plateau. The latter also allowed for long temporal baseline interferograms with sufficient coherence to obtain reasonable coverage of the tectonic signal, although we note that the velocity field is patchy. Bell et al. (2011) also highlight the difficulty in constraining the locking depth for signals of this magnitude, with their Monte Carlo solutions reaching their limits of 0–40 km. Mousavi et al. (2015) provides a narrower locking depth estimate of 10±4 km (66% CI), although the result is still relatively insensitive to the choice of locking depth in comparison to the choice of slip rate.

Our range of slip rate estimates is in agreement with previous GNSS derived estimates for the Main Recent Fault from Vernant et al. (2004, b) (3±2 mm/yr), Walpersdorf et al. (2006) (4–6 mm/yr), and Khorrami et al. (2019) (2.7–4 mm/yr), along with the geological/geomorphological-derived estimates from Alipoor et al. (2012) (1.6–3.2 mm/yr) and Copley and Jackson (2006) (2–5 mm/yr). Our estimates are lower than other geological/geomorphological-derived estimates from Talebian and Jackson (2002) (10–17 mm/yr) and Bachmanov et al. (2004) (10 mm/yr), and comparable to the lower end of estimates from Authemayou et al. (2009) (3.5–12.5 mm/yr). Whilst our InSAR results are not entirely independent of the GNSS data, as the GNSS data from Khorrami et al. (2019) are used to constrain the north-south velocities, the InSAR data does provide significant additional constraint. Our new InSAR-derived interseismic slip estimate affirms that the MRF is slipping at a geodetic rate of a few millimeters per year.

The broad range of slip rates estimates from geological and geomorphological methods can be tied to differences in two parameters: the average offset on the fault, and the time over which the offset occurred. Talebian and Jackson (2002), Copley and Jackson (2006), and Alipoor et al. (2012) all use the inception of the MRF as their timescale. Both Talebian and Jackson (2002) and Copley and Jackson (2006) assume an age of 3–5 ma for the MRF, based on the onset of shortening in the Simply Folded Belt of the Zagros, but measure markedly different offsets of 50 km and 10-15 km, respectively. Alipoor et al. (2012) estimate an offset of 16 km on the MRF, and use an initiation time of 5–10 ma based on the timing of slab break-off below the Zagros. Bachmanov et al. (2004) and Authemayou et al. (2009) both measure offsets over a period of thousands of years, the former using the beginning of the Holocene, and the latter using the exposure ages of samples from various geomorphic features. Copley and Jackson (2006) suggest that, based on differ-

ing offsets measured along adjacent fault sections, that either the average slip rate on the MRF varies along strike, or that the onset of faulting was heterogeneous along the length of the fault. In addition, Austermann and Iaffaldano (2013) proposed that the total convergence rate may have decreased by 30% since 5 Ma, which would reduce the slip rate along the MRF. However, the overlap between the lower end of the geological slip rate estimates and the geodetic estimates could suggest that the slip rate on the MRF may have been largely consistent through time.

Our InSAR-derived velocity fields and fault modelling suggest that the component of motion parallel to the tectonic boundary between the Arabian plate and Iran in our study area is localised onto the MRF. We expand on this by calculating the percentage of partitioned strike-slip motion accommodated by the MRF across the northwestern Zagros. We take a subset of GNSS velocities from Khorrami et al. (2019), located either side of the northwestern Zagros (Figure S10), and calculate the difference between the weighted mean velocities. We assume a single range-strike of 310° and decompose the velocity difference into range-parallel and range-perpendicular components with values of 4.3 mm/yr and 5.3 mm/yr, respectively. Our interseismic slip rate estimate of 3 ± 1 mm/yr for the MRF would account for 46–93% of the overall strike-slip component. At the upper end, this suggests that movement along the MRF alone is sufficient to account for the strike-slip component across this part of the Zagros. At the lower end, this would suggest that strike-slip motion also occurs on adjacent faults, either outside of our study area or with slip rates below our sensing limit, or that a component of the range-parallel motion is accommodated by off-fault deformation. This is in agreement with block modelling from Khorrami et al. (2019), who estimate 2.7–4 mm/yr of fault-parallel motion along the MRF, and a smaller 0.5 mm/yr fault-parallel component in the Frontal fault zone, which shares a similar strike to the MRF between 46°–50° E. The localisation of strike-slip motion paired with distributed thrust faulting has also been observed in the Qilian Shan, northeastern Tibetan Plateau (Allen et al., 2017). Allen et al. (2017) suggest that this arrangement, previously observed in oblique oceanic subduction zones (McCaffrey et al., 2000), may apply generally to oblique convergence zones. Slip may have become localised on the MRF because of an existing weakness in the form of the Main Zagros Thrust Fault, the trace of which the MRF partially inherited.

The localisation of slip on the MRF poses a greater risk of higher magnitude earthquakes than if the deformation was distributed on smaller faults. Given the simplistic scenario that all accumulated strain on the MRF is released seismically by only M_w 7 earthquakes, we can calculate an average recurrence interval for a M_w 7 earthquake similar to the 1909 Dorud earthquake (Ambraseys & Moinfar, 1973) occurring on the fault (e.g. Walters et al., 2014) Assuming an interseismic slip rate of 2–4 mm/yr, on a vertical strike slip fault with a locking depth of 22 km (average of our estimates from A-A', B-B', and the joint model), a shear modulus of $3{\times}10^{10}$ Pa and a slip-to-length ratio of 3×10^{-5}, we estimate a recurrence time of 315–631 years.

Our estimates for both the interseismic slip rate and the locking depth of the MRF must be kept in the context of our modelling limitations. The poor signal-to-noise ratio of our tectonic signal limited the number of model parameters we were able to constrain. Our screw dislocation model (Equation 5) assumes that all interseismic slip can be reasonably approximated as being localised onto a single vertical plane. Source modelling of the 2006 M_w6.1 Silakhour earthquake (Peyret et al., 2008; Ghods et al., 2012), which ruptured two segments of the MRF, suggests that the dip of the MRF may be as low as 60°, in agreement with teleseismic crustal imaging (Dashti et al., 2020), and that the shear zone is located up to 10 km north of the fault trace. We also assume identical rheological parameters either side of the MRF, which in reality forms the boundary between the High Zagros and the Central Iranian Plateau (Allen et al., 2013). Future work may incorporate longer InSAR time series, including images for other InSAR satel-

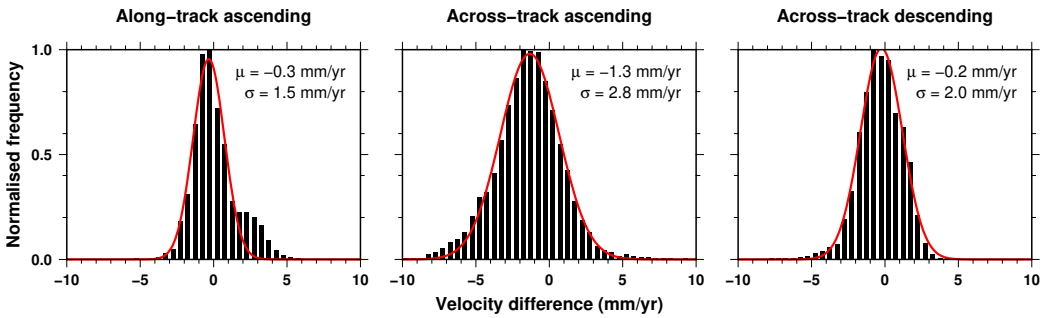

Figure 13. Histograms for the difference between projected horizontal velocities in frame overlaps along ascending tracks (174A-N and 174A-S), across ascending tracks (174A-N, 174A-S, and 101A), and across descending tracks (006D and 108D). Best fitting Gaussians are shown in red, along with their mean, μ, and standard deviation, σ.

lites such as Envisat and ERS, to improve the signal-to-noise ratio and to attempt to model a more complex fault geometry and rheological contrasts (e.g. Jolivet et al., 2008)

5.3 Modelling Velocity Uncertainties

Given our low interseismic slip rate estimate for the MRF, one of the slowest InSAR-derivied fault slips rates published so far, it is valuable to assess the uncertainties associated with our InSAR velocities. Bootstrapped estimates of the InSAR velocity uncertainty have been shown to decrease with increasing time series length (Morishita et al., 2020), with a 2 mm/yr standard deviation achievable with 1.4 and 1.8 years of 6 and 12 day acquisitions, respectively. Uncertainties for combined GNSS and InSAR velocities, however, may level off around 2–3 mm/yr as the uncertainty on the GNSS velocities becomes dominant (Weiss et al., 2020). As another estimate of the uncertainties, and to examine the effectiveness of the InSAR referencing, we calculate the difference in velocities (Figure 13) between overlapping frames (Walters et al., 2014). We assume that all velocities are purely horizontal by dividing by the sine of the incidence angle and then multiplying by the incidence angle at the centre of each frame. The frame overlap between 174A-N and 174A-S covers part of the Sarpol earthquake cluster, meaning our assumption of only horizontal velocities is poor. Despite this, we still observe 1σ uncertainties less than 3 mm/yr, inline with those observed in previous studies (Hussain, Hooper, et al., 2016; Liu et al., 2018; Weiss et al., 2020). This implies that the referencing of the InSAR LOS velocities to a Eurasia-fixed frame has been reasonably successful. These standard deviations can be considered as $\sqrt{2}\times$ the velocity uncertainty for each frame. This gives a 1σ uncertainty of 1.36 mm/yr for 006D and 108D. For the ascending frames we take the mean of the along-track and across-track values, giving a 1σ uncertainty of 1.41 mm/yr for 174A and 101A.

6 Conclusion

We have used over 5.5 years of Sentinel-1 SAR images across two ascending and two descending tracks to produce the first InSAR-derived estimate of interseismic slip rate for the Main Recent Fault, SW Iran. We combine InSAR LOS velocities with GNSS to estimate the fault-parallel velocity for three across-fault profiles which we model both individually and together to better constrain the fault parameters. Our estimated rate of 3.0 ± 1.0 mm/yr for the MRF between 47° E and 50° E is in agreement with previous geodetic rates from GNSS studies, and is one of the slowest interseismic slip rates measured using InSAR. We provide the first estimate of the locking depth of the fault

(18–30 km) while highlighting the difficulties of modelling tectonic signals close to the current InSAR noise level. We show that the strike-slip component of the overall plate motion in our study area is localised onto the MRF, with a slip rate of 3.0±1.0 mm/yr accounting for 46–93% of the strike-slip component across the western Zagros. Our results show that the MRF is an important major crustal structure that shows a localisation of strain at depth and which accommodates an appreciable portion of the relative motion between Arabia and Eurasia.

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The Sentinel-1 InSAR data are copyrighted by the European Space Agency and provided freely through the Copernicus Open Access Hub (https://scihub.copernicus.eu/). Sentinel-1 InSAR data are also freely distributed by the Alaska Satellite Facility (https://asf.alaska.edu/). Processed SAR images, wrapped interferograms, and unwrapped interferograms used in this work can be found on the COMET-LiCS Sentinel-1 InSAR portal (https://comet.nerc.ac.uk/COMET-LiCS-portal/). Information on accessing the LiCSBAS software can be found in Morishita et al. (2020). Information on accessing the GACOS atmospheric corrections can be found in Yu, Li, Penna, and Crippa (2018).

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Figure 1.

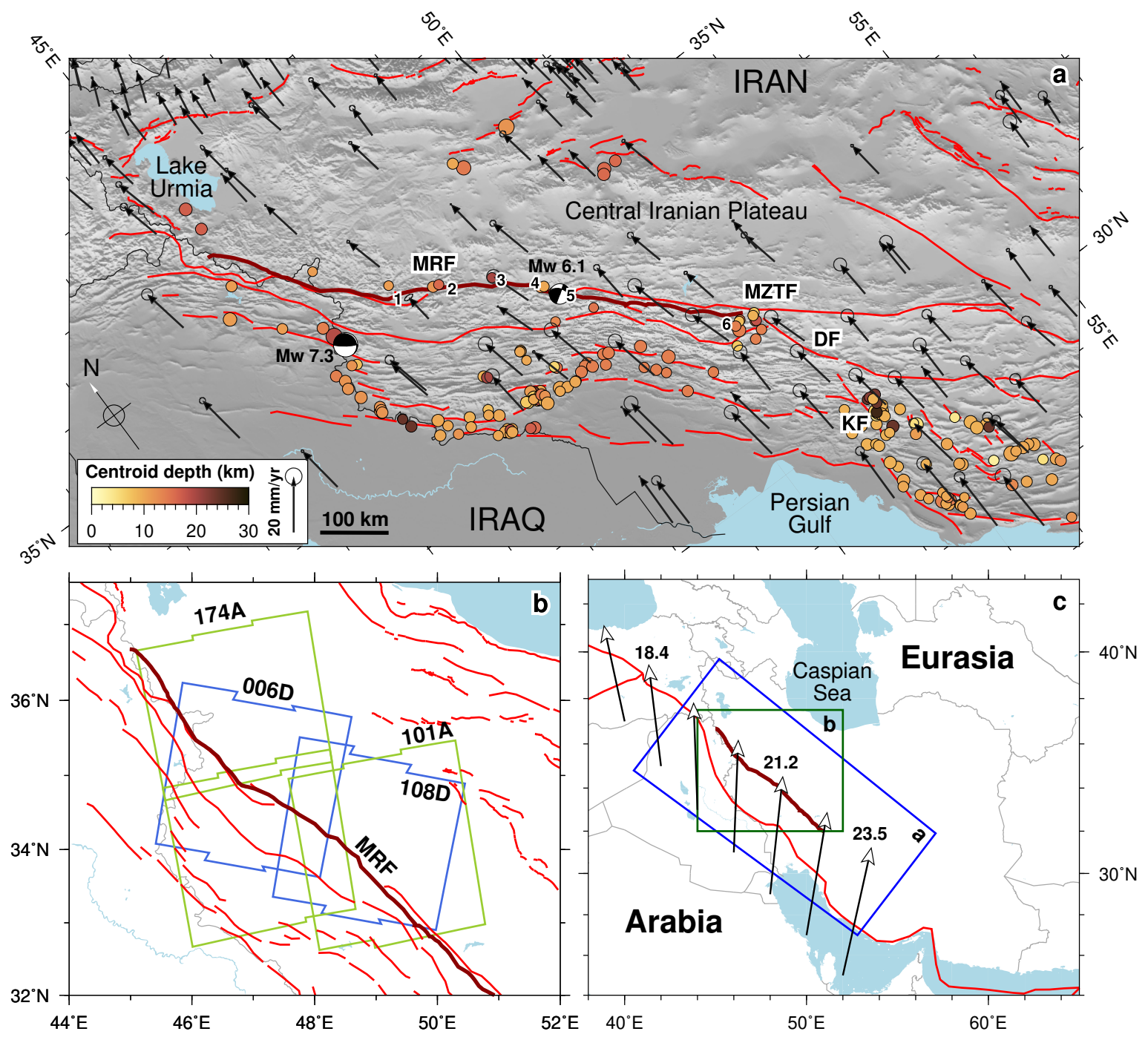

a
IRAN
Lake
Urmia
Central Iranian Plateau
MRF
Mw 6.1
MZTF
DF
KF
Mw 7.3
1
2
3
4
5
6
N
Centroid depth (km)
0
10
20
30
20 mm/yr
100 km
IRAQ
Persian
Gulf
45°E
50°E
35°N
55°E
30°N
55°E
35°N
b
174A
006D
101A
108D
MRF
36°N
34°N
32°N
44°E
46°E
48°E
50°E
52°E
c
18.4
Caspian
Sea
Eurasia
b
21.2
23.5
a
Arabia
40°N
30°N
40°E
50°E
60°E

Figure 2.

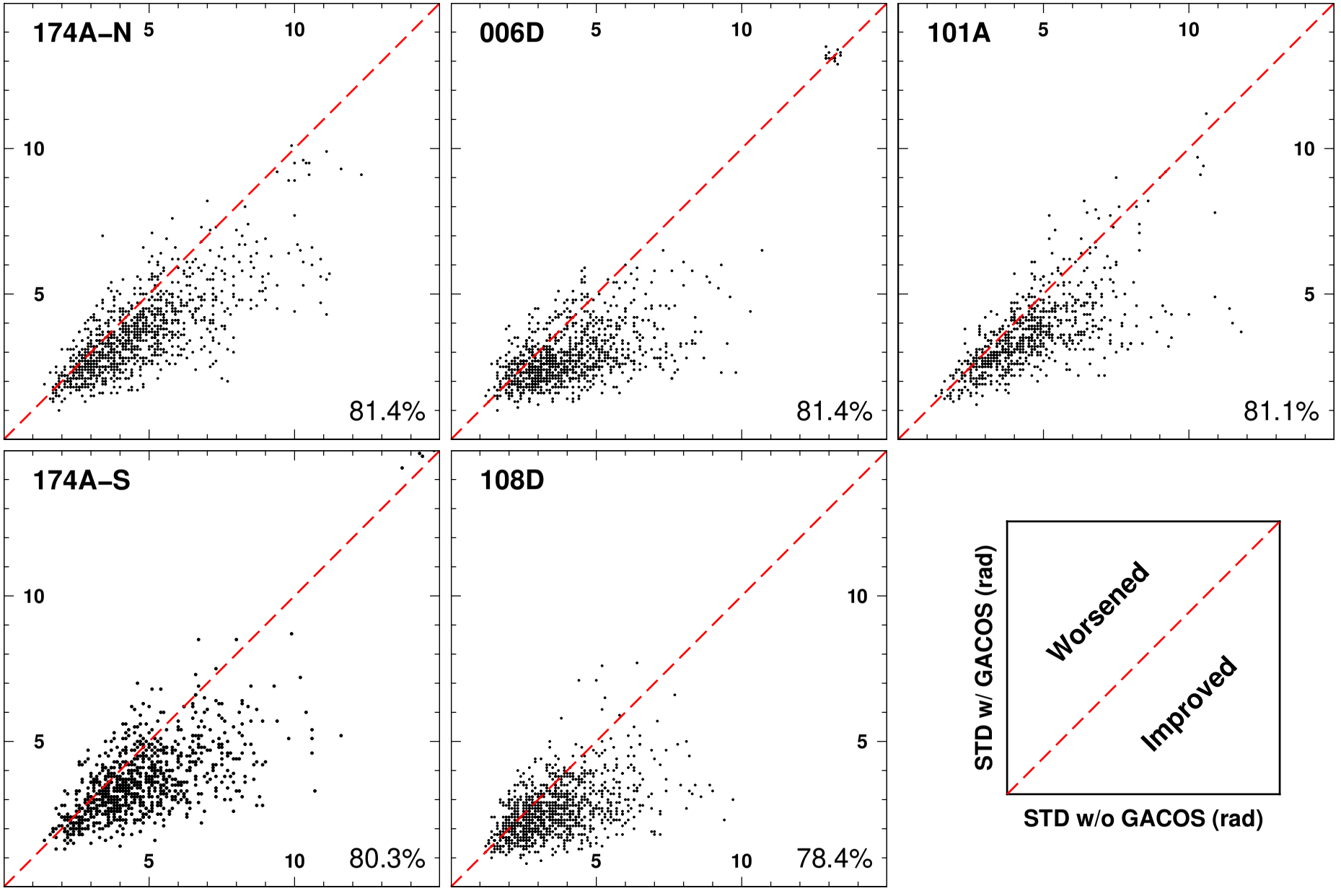

174A–N
5
10
10
5
81.4%
006D
5
10
81.4%
101A
5
10
10
5
81.1%
174A–S
10
5
5
10
80.3%
108D
10
5
5
10
78.4%
STD w/ GACOS (rad)
Worsened
Improved
STD w/o GACOS (rad)

Figure 3.

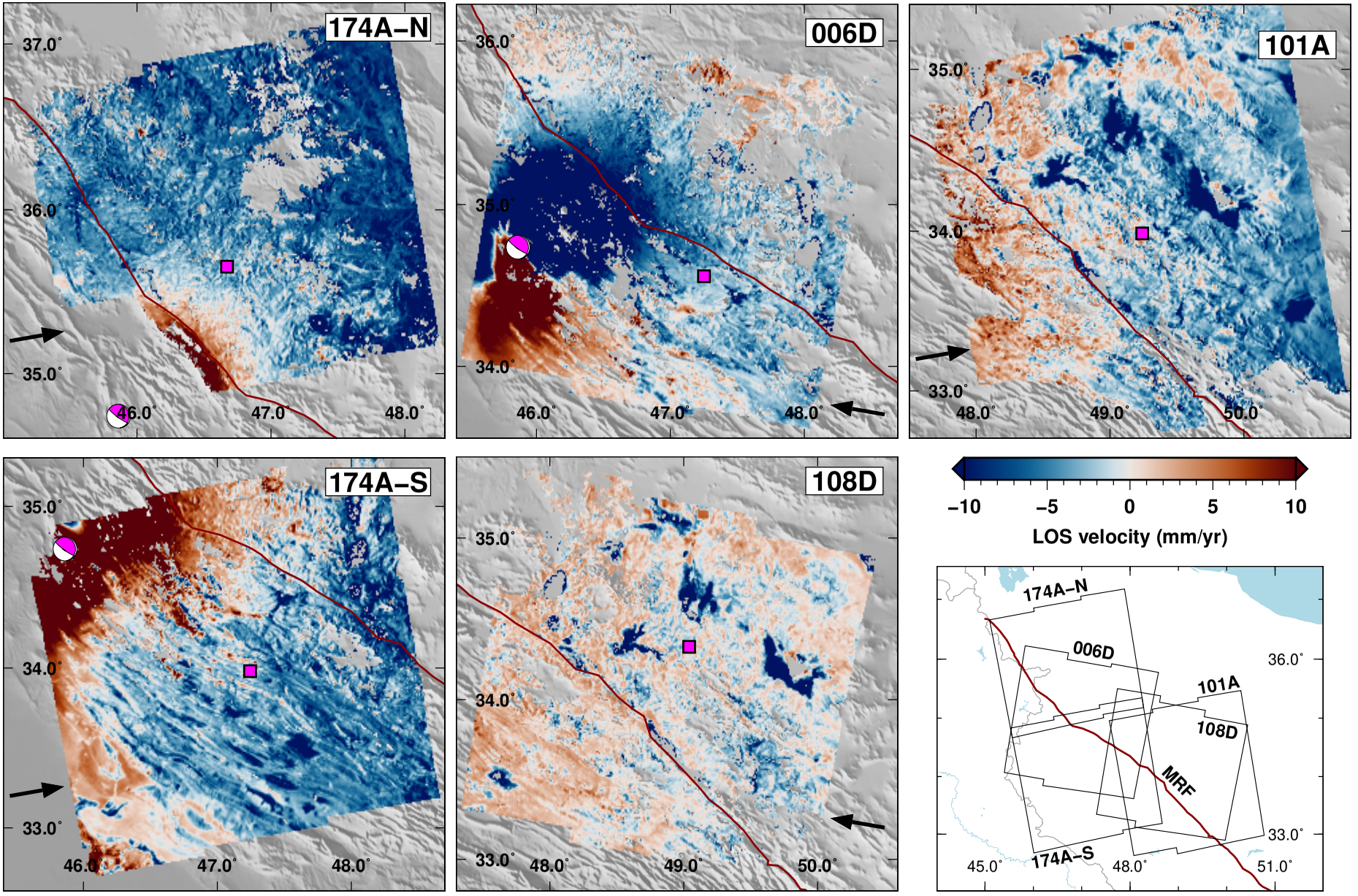

174A–N
006D
101A
174A–S
108D
37.0˚
36.0˚
35.0˚
34.0˚
33.0˚
46.0˚
47.0˚
48.0˚
49.0˚
50.0˚
−10
−5
0
5
10
LOS velocity (mm/yr)
MRF
45.0˚
51.0˚

Figure 4.

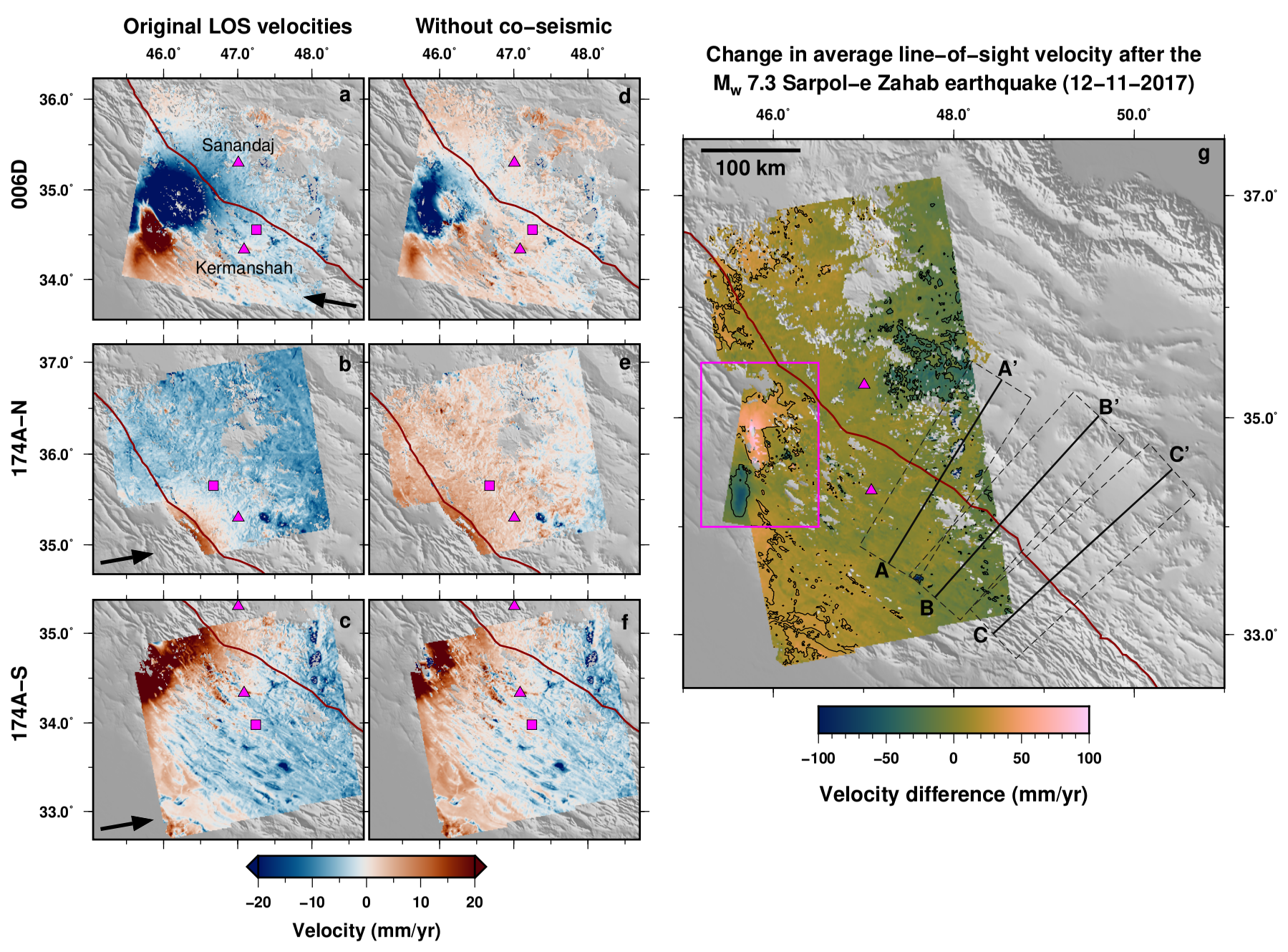

Original LOS velocities
Without co-seismic
Change in average line-of-sight velocity after the
M_w 7.3 Sarpol-e Zahab earthquake (12-11-2017)
006D
174A-N
174A-S
Sanandaj
Kermanshah
100 km
A
A'
B
B'
C
C'
Velocity (mm/yr)
Velocity difference (mm/yr)

Figure 5.

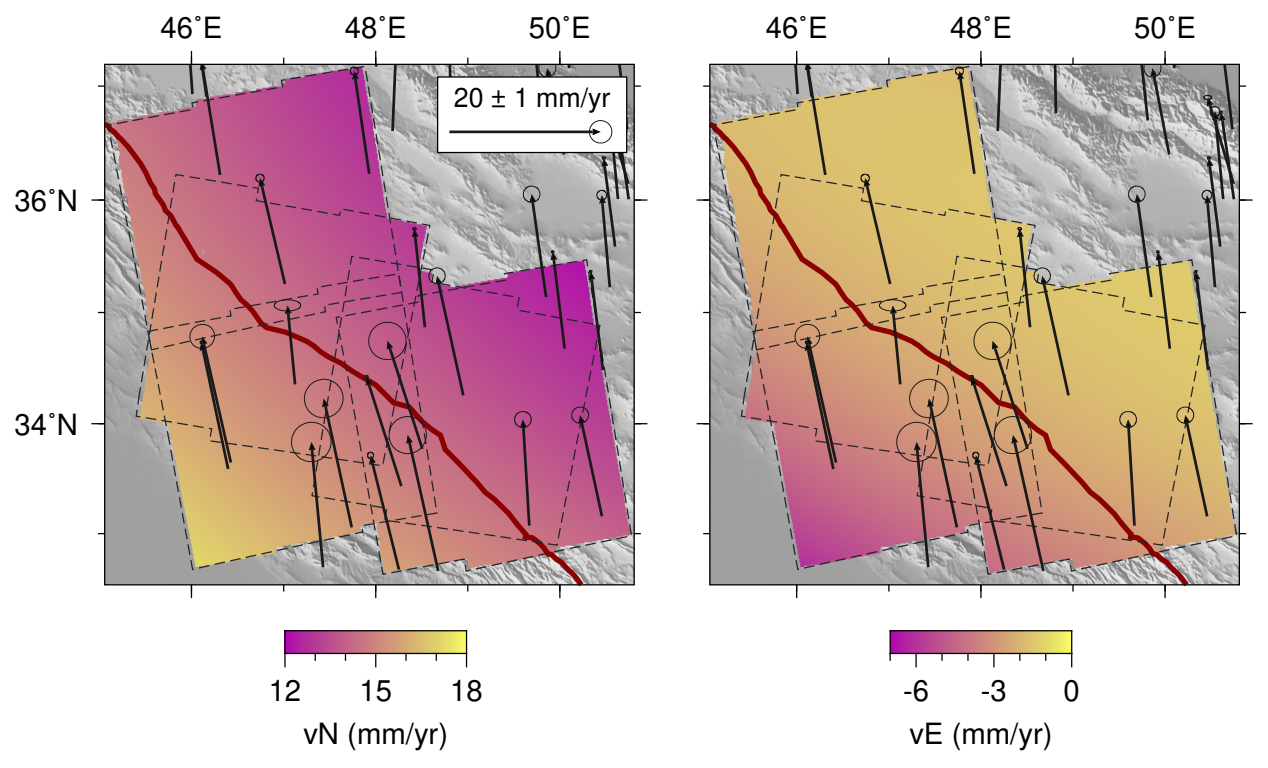

46˚E
48˚E
50˚E
36˚N
34˚N
20 ± 1 mm/yr
12
15
18
vN (mm/yr)
46˚E
48˚E
50˚E
-6
-3
0
vE (mm/yr)

Figure 6.

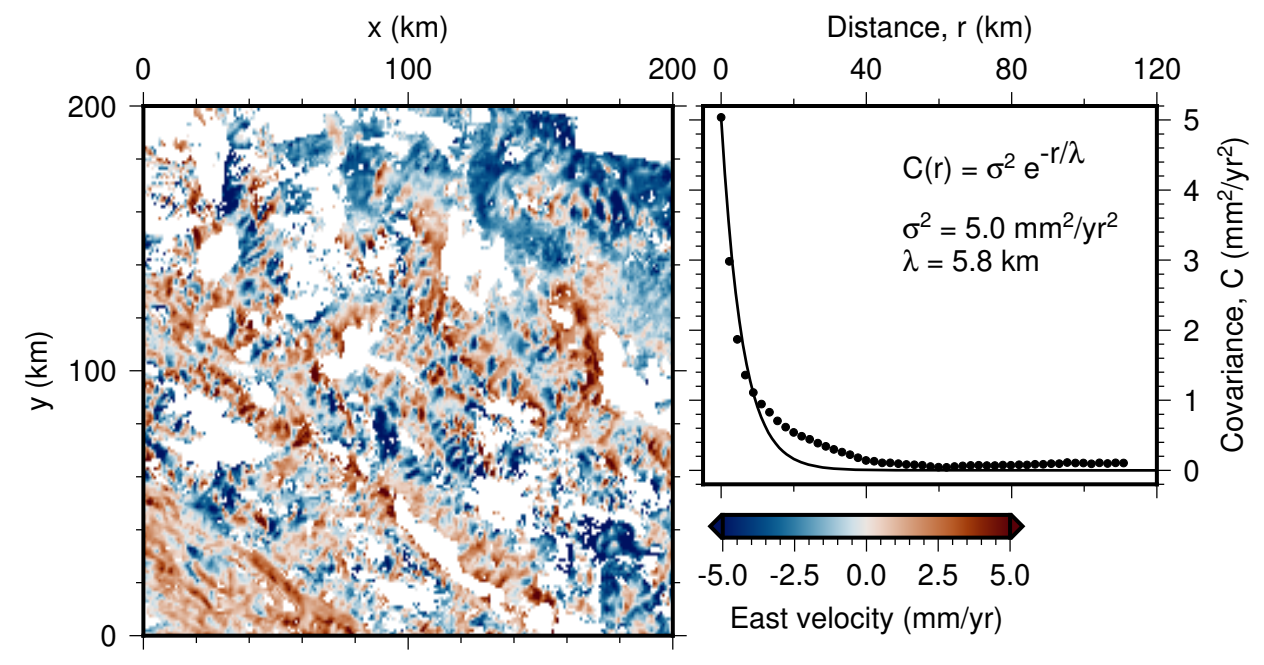

x (km)
0
100
200
y (km)
200
100
0
Distance, r (km)
0
40
80
120
C(r) = σ² e^-r/λ
σ² = 5.0 mm²/yr²
λ = 5.8 km
Covariance, C (mm²/yr²)
5
4
3
2
1
0
-5.0
-2.5
0.0
2.5
5.0
East velocity (mm/yr)

Figure 7.

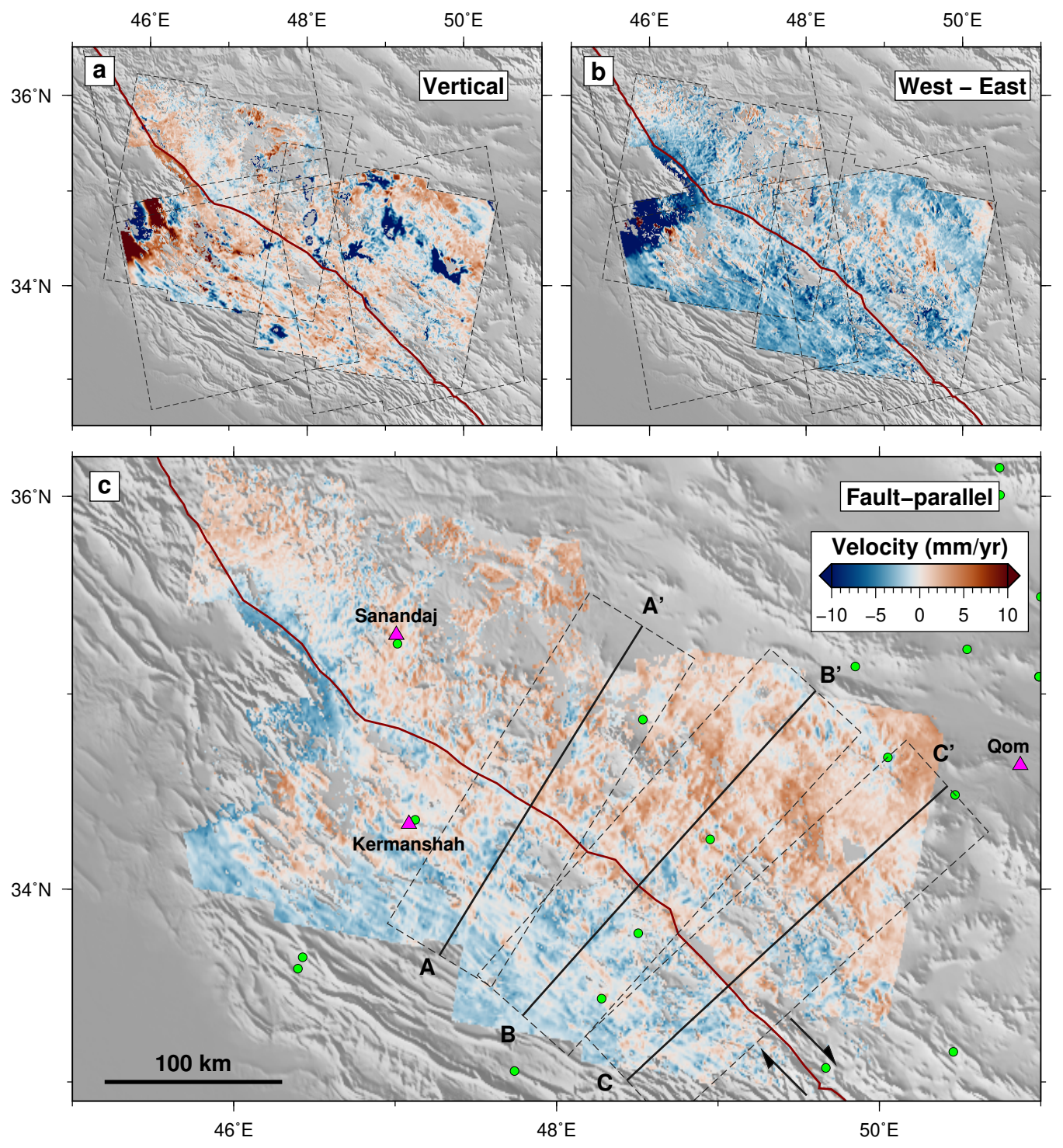

a
Vertical
b
West – East
c
Fault–parallel
Velocity (mm/yr)
−10 −5 0 5 10
Sanandaj
Kermanshah
Qom
A
A’
B
B’
C
C’
100 km
46˚E
48˚E
50˚E
36˚N
34˚N

Figure 8.

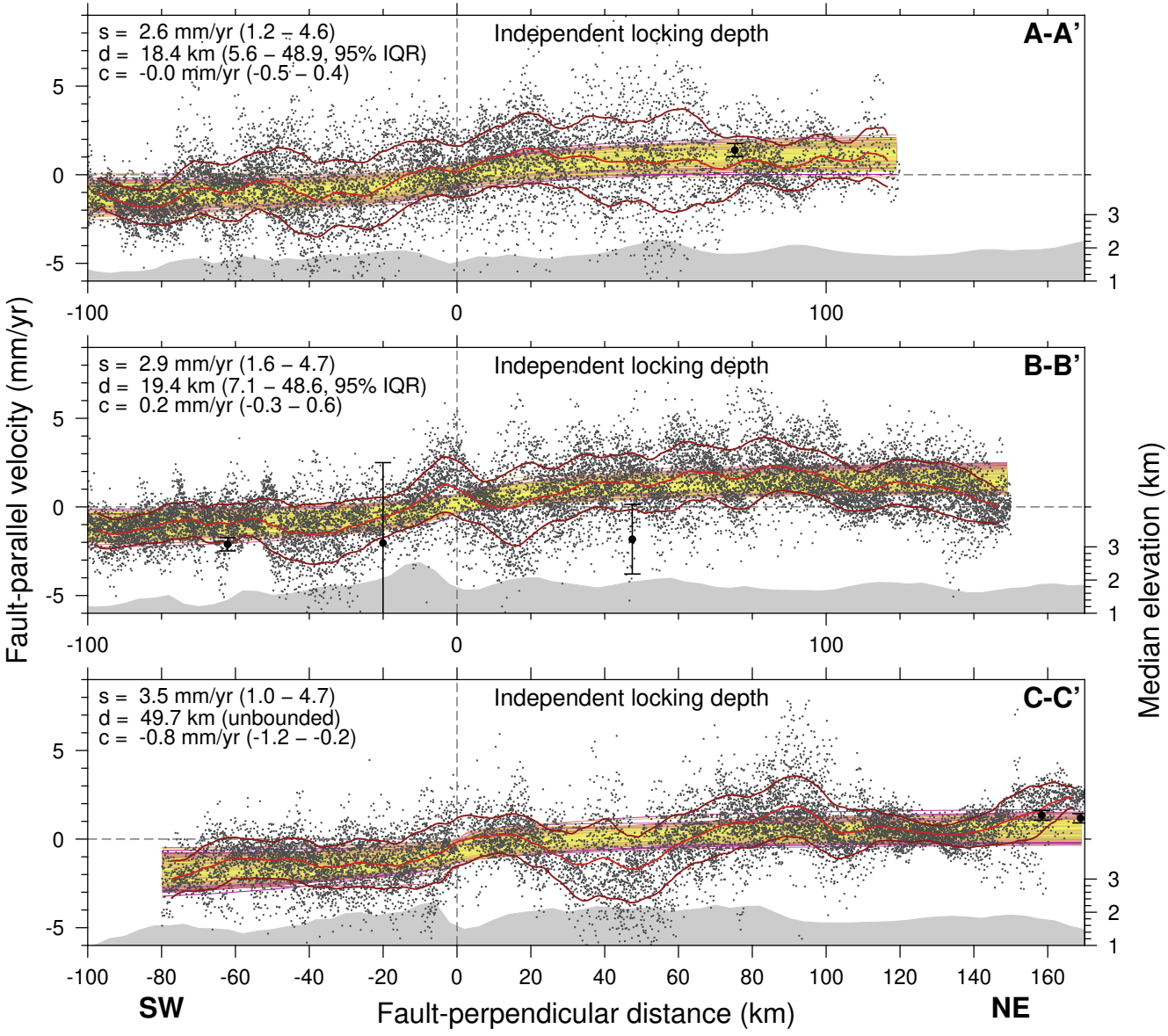

s = 2.6 mm/yr (1.2 – 4.6)
d = 18.4 km (5.6 – 48.9, 95% IQR)
c = -0.0 mm/yr (-0.5 – 0.4)
Independent locking depth
A-A'
s = 2.9 mm/yr (1.6 – 4.7)
d = 19.4 km (7.1 – 48.6, 95% IQR)
c = 0.2 mm/yr (-0.3 – 0.6)
Independent locking depth
B-B'
s = 3.5 mm/yr (1.0 – 4.7)
d = 49.7 km (unbounded)
c = -0.8 mm/yr (-1.2 – -0.2)
Independent locking depth
C-C'
Fault-parallel velocity (mm/yr)
Median elevation (km)
Fault-perpendicular distance (km)
SW
NE

Figure 9.

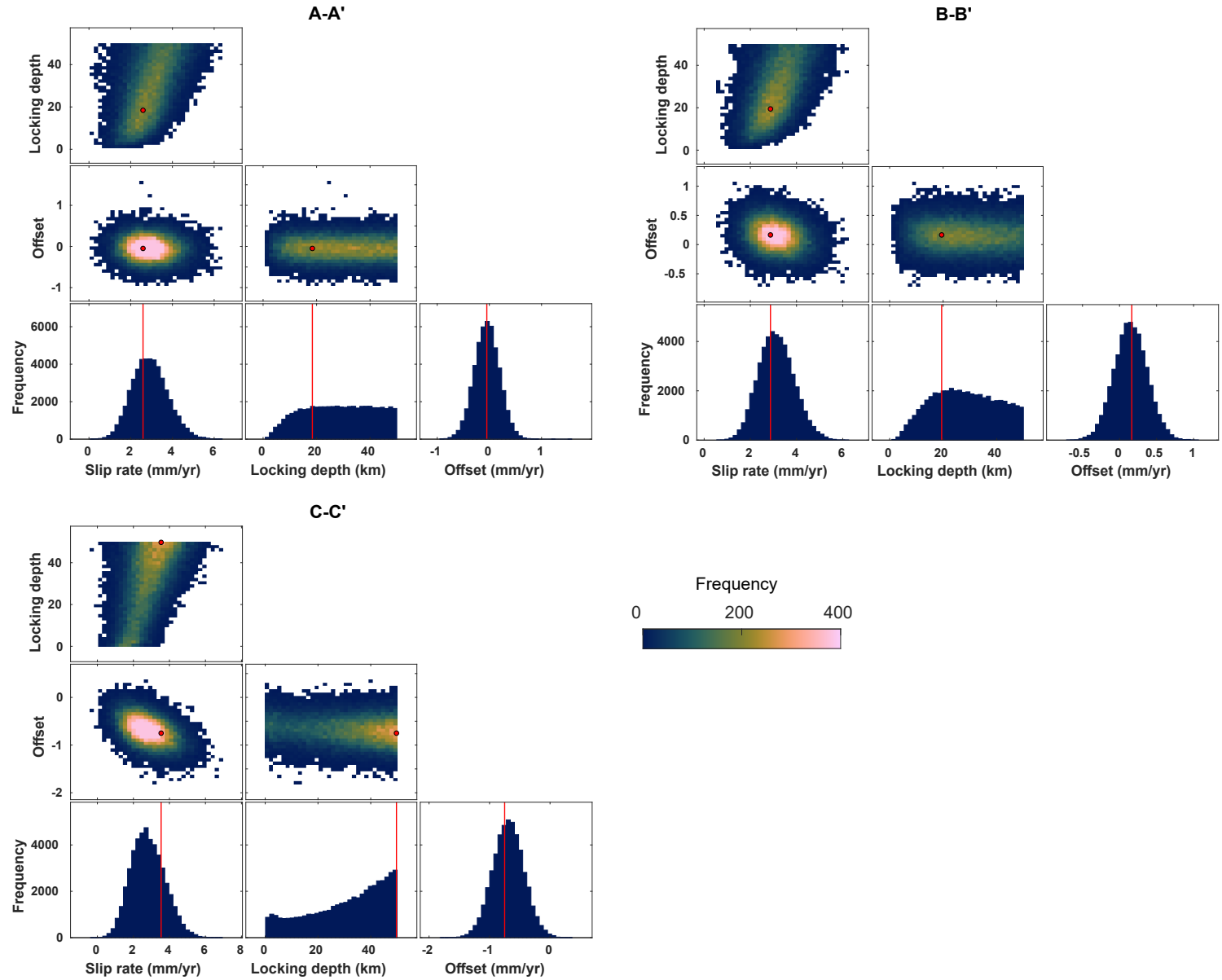

A-A'
B-B'
C-C'
Locking depth
Offset
Frequency
Slip rate (mm/yr)
Locking depth (km)
Offset (mm/yr)
Frequency
0
200
400

Figure 10.

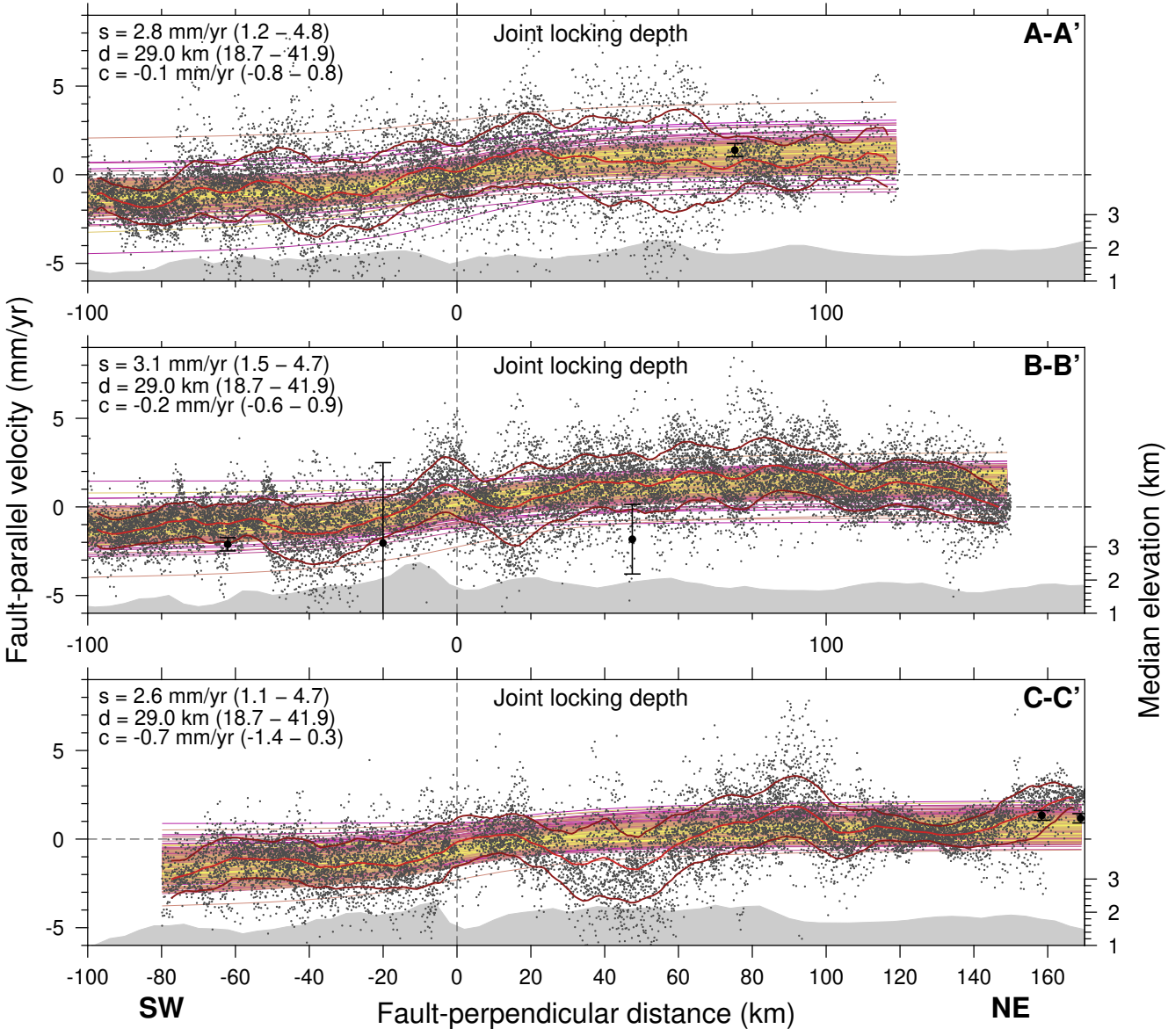

s = 2.8 mm/yr (1.2 – 4.8)
d = 29.0 km (18.7 – 41.9)
c = -0.1 mm/yr (-0.8 – 0.8)
Joint locking depth
A-A'
s = 3.1 mm/yr (1.5 – 4.7)
d = 29.0 km (18.7 – 41.9)
c = -0.2 mm/yr (-0.6 – 0.9)
Joint locking depth
B-B'
s = 2.6 mm/yr (1.1 – 4.7)
d = 29.0 km (18.7 – 41.9)
c = -0.7 mm/yr (-1.4 – 0.3)
Joint locking depth
C-C'
Fault-parallel velocity (mm/yr)
Median elevation (km)
Fault-perpendicular distance (km)
SW
NE

Figure 11.

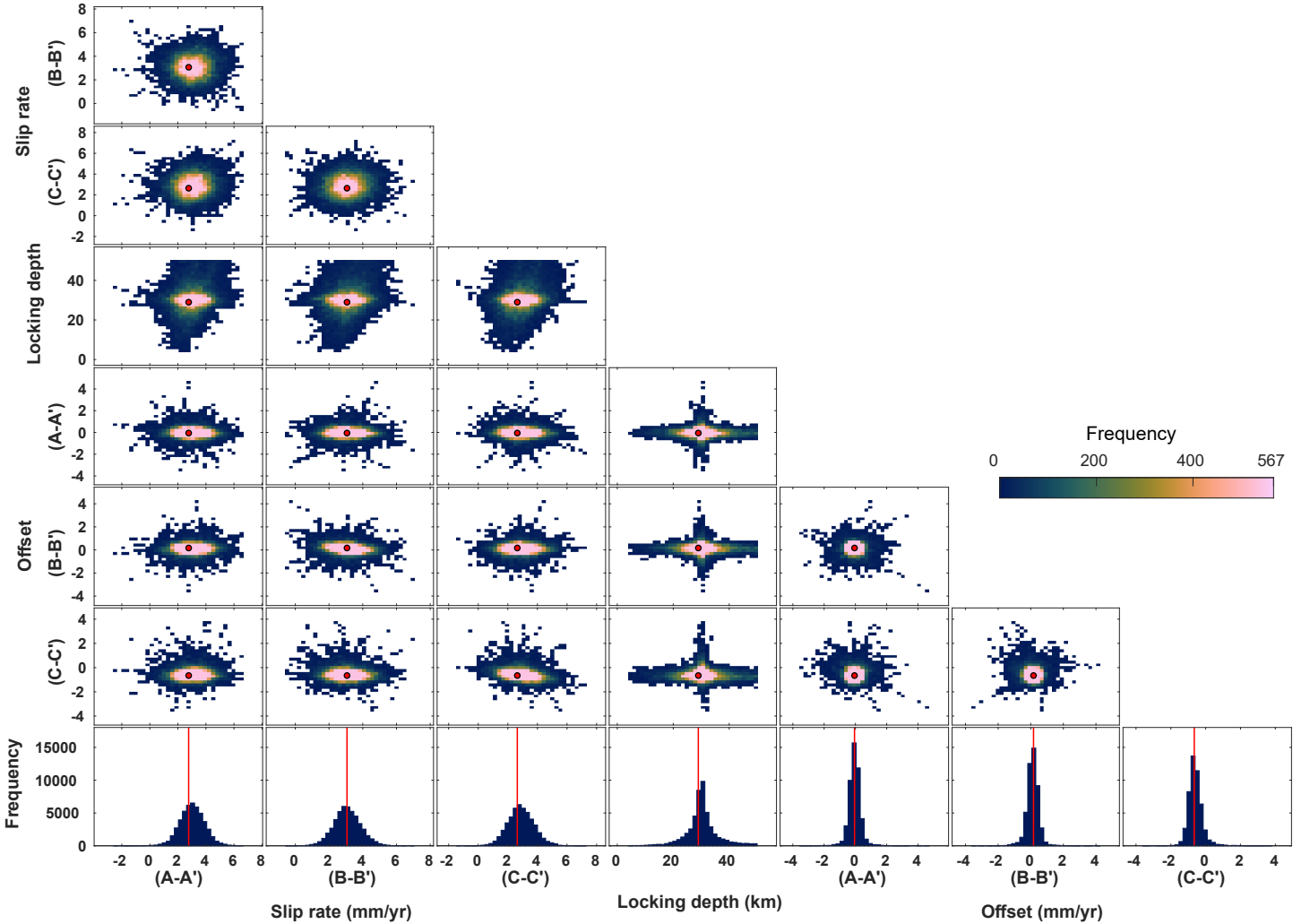

Slip rate
(B-B')
(C-C')
Locking depth
(A-A')
Offset
(B-B')
(C-C')
Frequency
Frequency
0
200
400
567
Slip rate (mm/yr)
Locking depth (km)
Offset (mm/yr)

Figure 12.

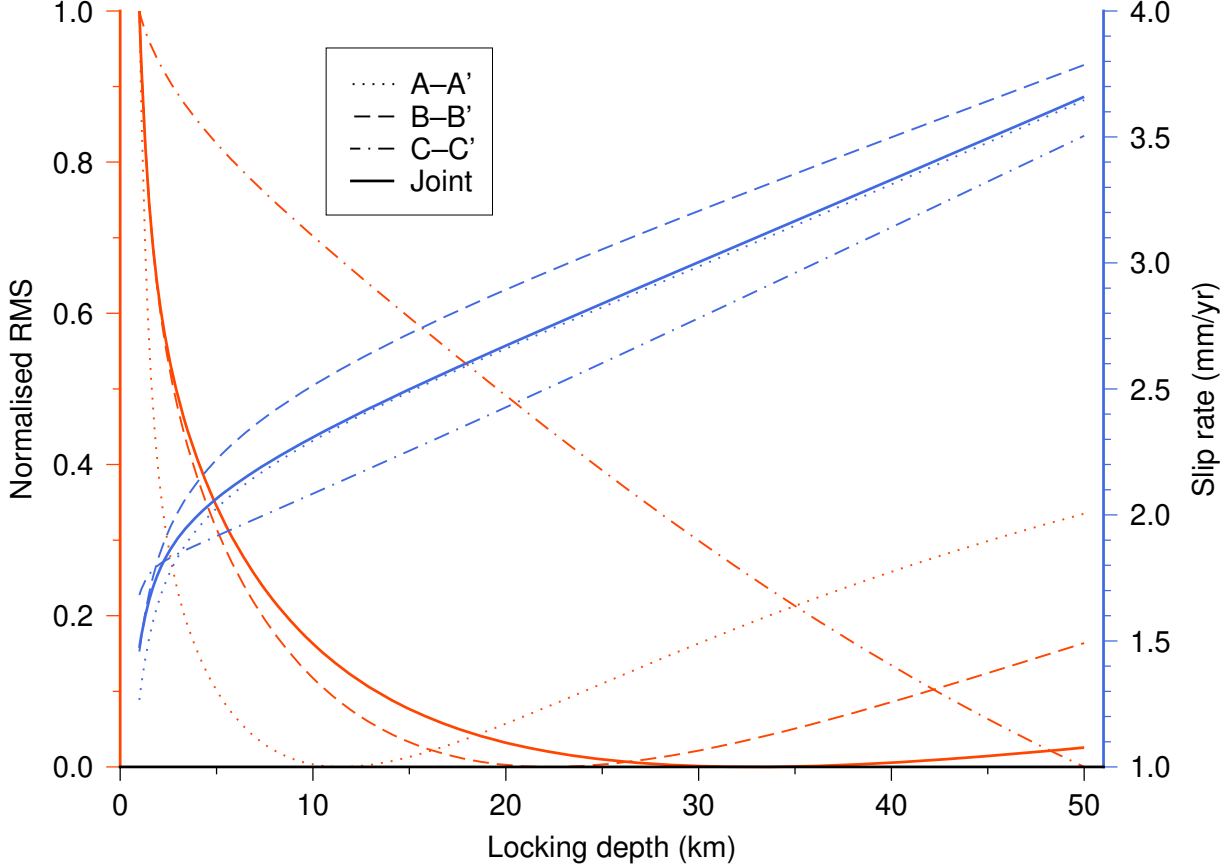
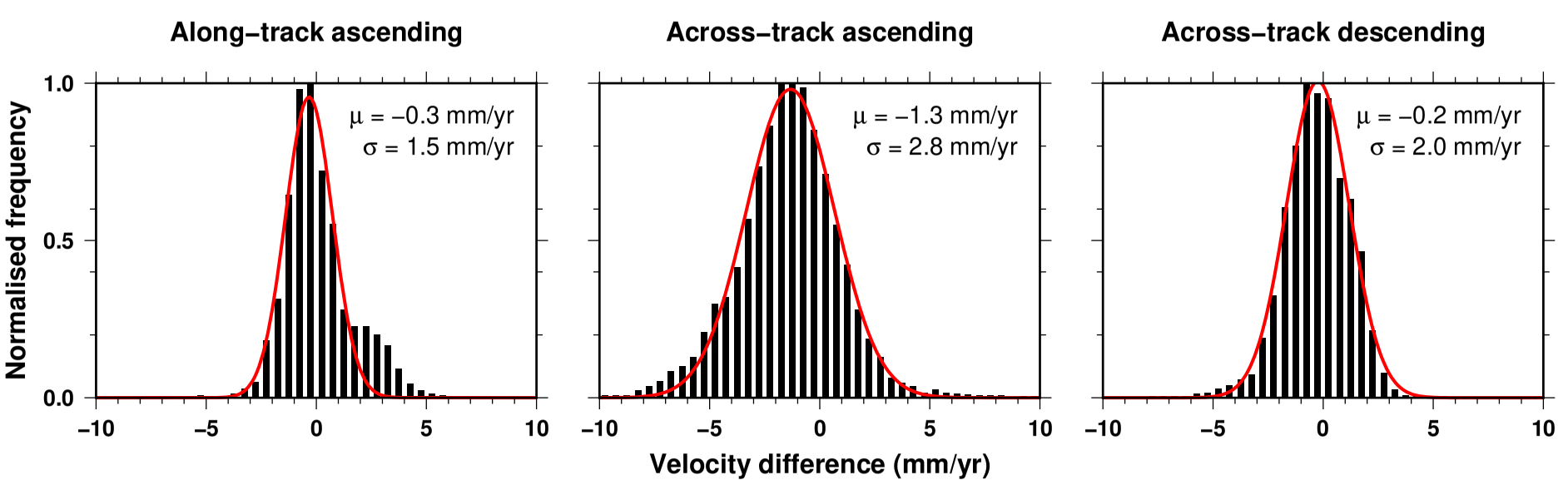

A–A'
B–B'
C–C'
Joint
Normalised RMS
Slip rate (mm/yr)
Locking depth (km)
0.0
0.2
0.4
0.6
0.8
1.0
1.0
1.5
2.0
2.5
3.0
3.5
4.0
0
10
20
30
40
50

Figure 13.


Along–track ascending
Across–track ascending
Across–track descending
Normalised frequency
1.0
0.5
0.0
μ = –0.3 mm/yr
σ = 1.5 mm/yr
μ = –1.3 mm/yr
σ = 2.8 mm/yr
μ = –0.2 mm/yr
σ = 2.0 mm/yr
–10
–5
0
5
10
Velocity difference (mm/yr)