



## 11 **Key Points**

- 12 • Hunga Tonga-Hunga Ha'apai eruption produced vertically overlapping but slightly  
13 displaced mid-stratospheric anomalies in H<sub>2</sub>O and aerosols.
- 14 • IR cooling by enhanced H<sub>2</sub>O layer explains the observed ~ 1.5K mid-stratospheric  
15 temperature decrease following the eruption.
- 16 • A simple model of the eruption H<sub>2</sub>O enhancement combined with spreading of the plume  
17 explains the observations.

18

## 19 **Plain Language Summary**

20 Hunga Tonga-Hunga Ha'apai submarine volcano eruption on January 15, 2022 injected up to  
21 150 Tg of water into the stratosphere. A month after eruption, a distinct aerosol and water vapor  
22 layer formed in the tropical southern hemisphere (SH) stratosphere. The water vapor layer is  
23 slightly displaced above the aerosol layer at 26 km. These two layers persisted in the tropical SH  
24 stratosphere over the next four months while slowly moving apart in altitude. The isolation of  
25 the layers and their separate motion are consistent with our understanding of tropical  
26 stratospheric dynamics. A cold temperature anomaly forms coincident with the water vapor  
27 layer, which we show to be due to enhanced IR radiative cooling by water vapor. Using a simple  
28 model, we show how the water vapor layer forms slightly above the aerosol layer.

29

## 30 **Abstract**

31 On Jan. 15, 2022, the Hunga Tonga-Hunga Ha'apai eruption injected SO<sub>2</sub> and H<sub>2</sub>O into the  
32 middle stratosphere. The eruption produced a persistent mid-stratospheric sulfate aerosol layer,  
33 mostly below 26 km, confined to Southern Hemisphere (SH) tropics. Coincident with, and  
34 slightly above the aerosol layer, an enhanced H<sub>2</sub>O layer is also observed. The SH tropical  
35 confinement is simulated using a trajectory model. Measurements over several months following  
36 the eruption show that the H<sub>2</sub>O layer is slowly rising while the aerosol layer is descending. The  
37 H<sub>2</sub>O layer upward movement is consistent with the vertical velocity at these altitudes.  
38 Gravitationally settling explains the descent of the aerosol layer. A cold anomaly coincident  
39 with the H<sub>2</sub>O enhancement is observed and is caused by thermal adjustment to the additional  
40 H<sub>2</sub>O IR cooling. A simple model of volcanic water injection at the time of the eruption simulates  
41 the observed H<sub>2</sub>O.

42

## 43 **Index Terms**

44 0340 Middle atmosphere dynamics

45 0341 Middle atmosphere: constituent transport and chemistry

46 0370 Volcanic effects

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## 50 1. Introduction

51 The Hunga Hunga-Tonga Ha'apai (HT) (20.54°S, 178.3°W) submarine volcano violently  
52 erupted on Jan. 15, 2022. Since HT was a submarine volcano, it appears to have lofted a  
53 significant amount of water into the stratosphere. Indeed, Microwave Limb Sounder (MLS)  
54 measurements show that HT water enhancement was quite high relative to SO<sub>2</sub> (Millán et al.,  
55 2022) – hereafter M22. The MLS estimated water injection was up to 146 Tg (M22). The  
56 eruption plume was detected up to 57 km (Carr et al., 2022; Proud et al., 2022). The Ozone  
57 Mapping and Profile Suite – Limb Profiler (OMPS-LP) detected extinction enhancements above  
58 45 km (Taha et al., 2022).

59  
60 In this paper we will examine at the evolution of the water vapor and aerosol enhancements that  
61 followed the HT eruption. M22 noted that the amount of water deposited in the stratosphere by  
62 HT was unprecedented in the modern history of volcanic eruption observations. Several MLS  
63 water vapor profiles made shortly after the eruption show concentrations exceeding 300 ppmv  
64 against a normal stratospheric concentration of ~4 ppmv. As the eruption evolved, MLS water  
65 vapor maps show that above about 2 hPa (~43 km), the plume quickly spreads and that the water  
66 vapor enhancement disperses. A secondary maximum at about 25 hPa (~26 km) persists (M22).  
67 The aerosol field shows similar behavior with rapid dispersal at higher altitudes but persistent  
68 high levels of aerosol extinction below ~ 25 hPa (~26 km) (Taha et al., 2022). The aerosol  
69 extinction in this layer grows over the 30 days following the eruption presumably due to the  
70 conversion of SO<sub>2</sub> to sulfate aerosols (e.g. Zhu et al., 2020).

71  
72 There are several key questions concerning the HT eruption: Why did the unusual water vapor  
73 layer form and persist? Below we show that the water vapor enhancement overlaps the top of  
74 the extinction anomaly, but they are distinct, and furthermore the two enhancements vertically  
75 separate over time. We have also discovered a temperature anomaly in the 25-28 km region. We  
76 provide an explanation for the temperature anomaly as well as for the formation and evolution of  
77 the water vapor and aerosol layers.

## 78 79 2. Data sets

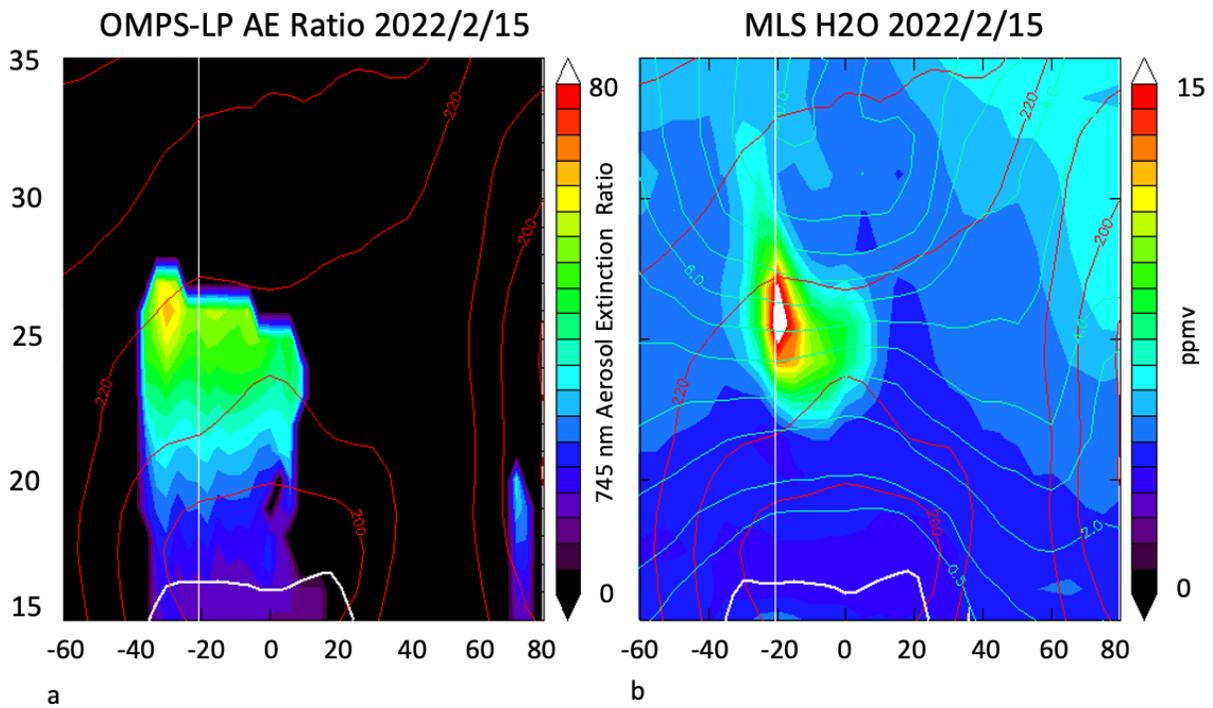
80  
81 We use MLS v5 for ozone, temperature and H<sub>2</sub>O, the quality of this data for the HT anomaly is  
82 discussed in M22. The MLS algorithm quality flags and convergence alerts were set for some  
83 plume profiles shortly after the eruption. Even with the quality flag and convergence filters set,  
84 the data looks reasonable and generally agrees with sonde and other validation data. For aerosols,  
85 we use OMPS-LP level-2 V2.1 745 nm extinction-to-molecular ratio data (AE) from all three  
86 OMPS-LP slits (see Taha et al., 2021). Taha et al. (2022) indicated that the standard V2.1  
87 released data (used in this study) provides the most accurate aerosol retrieval up to 36 km. Thus,  
88 we restrict our constituent analysis to below 35 km which contains the main locus aerosol plume  
89 (Taha et al. 2022; Fig. 4). The MLS and OMPS-LP extinction data sets are averaged over 4 days  
90 and then averaged onto a 5° latitude-longitude grid.

91  
92 To simulate the dispersal of the plume, we use the Forward Domain Filling (FDF) trajectory  
93 model (Schoeberl et al., 2018) modified to inject a dense column of parcels over the HT location

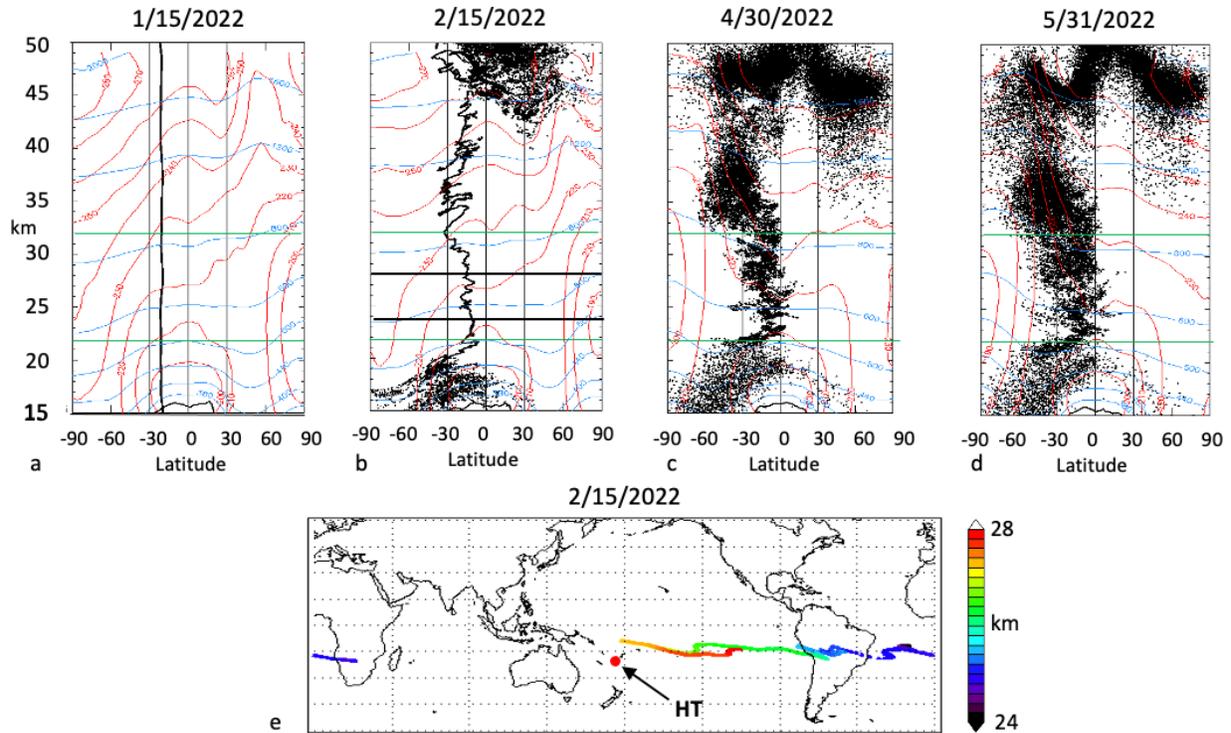
94 on Jan 15, 2022. This simulation uses MERRA-2 reanalysis winds, temperatures, and heating  
95 rates (Gelaro, et al., 2017).

96  
97 **3. Analysis**

98  
99 Figure 1 shows the zonal mean distribution of water vapor and aerosol extinction ratio on Feb  
100 15, 2022, a month after the eruption. The HT aerosol plume reaches 26 km in the region 30°S to  
101 about 5°N. The extinction data is quite sensitive to plumes extending outward from the tropics  
102 thus tends to show a wider distribution than the water vapor field. The water vapor plume is  
103 centered at 26 km and extends up to 30 km in the SH tropics. The water vapor plume mostly  
104 overlaps the aerosol plume while extending slightly above it.  
105  
106



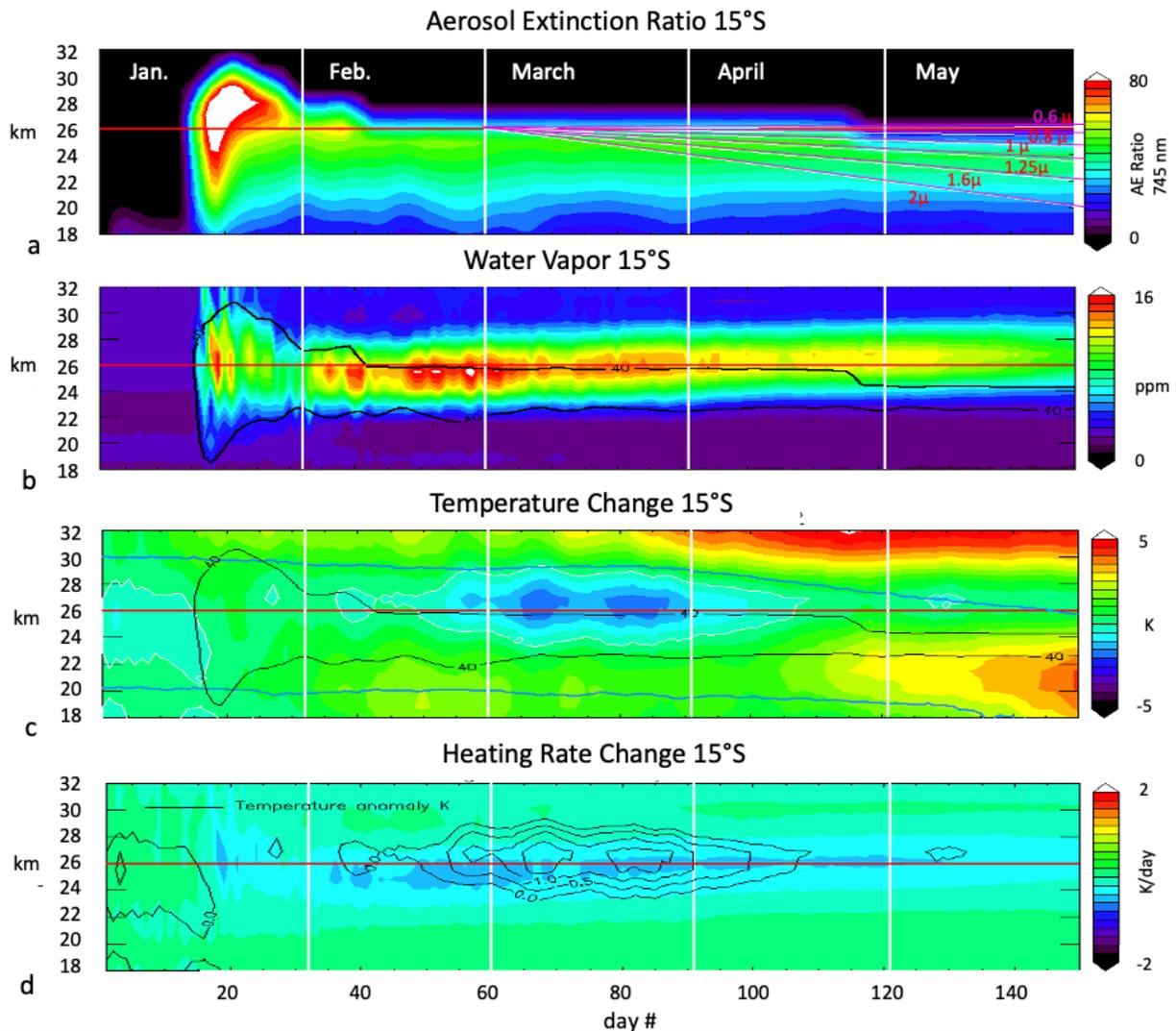
107  
108 *Figure 1 The zonal mean OMPS-LP 745 nm aerosol extinction/ molecular extinction ratio (Part*  
109 *a) and MLS water vapor (ppmv) (Part b) on Feb. 15, 2022. The red contours show the MLS*  
110 *temperature field. The thick white line is the zonal mean tropopause. The green contours are*  
111 *MLS ozone mixing ratio (ppm). The vertical white line denotes the latitude of the HT volcano.*



112  
 113 *Figure 2 Dispersal of HT plume simulated by the FDF model: (Part a) shows the initial parcel*  
 114 *distribution on Jan 15, (Part b) parcel distribution of Feb. 15, 2022, (Part c) shows the*  
 115 *distribution at the end of April, and (Part d) shows the distribution at the end of May. Along the*  
 116 *bottom (Part e), a map of parcels between 24-28 km with color scale indicating altitude. In Parts*  
 117 *a,b,c,d red contours are MERRA-2 temperatures, blue contours are potential temperature.*  
 118 *Horizontal green lines show the isolation region 22-32km. Horizontal black lines in Part b*  
 119 *indicate the domain in Part e. The red dot locates HT on the map.*

120 Figure 2 shows the dispersal of the plume using the FDF trajectory model. From the initial  
 121 distribution (Fig. 2a), the plume evolves slowly and is mostly confined to the region between the  
 122 equator and 30°S in the height range 22-32 km. This confinement is still somewhat evident at the  
 123 end of April. The isolation of the Northern Hemisphere (NH) tropics from the Southern  
 124 Hemisphere (SH) tropics in this region was first noted by Stolarski et al. (2014) when analyzing  
 125 the interhemispheric phasing of the tropical ozone concentration. Below ~20 km parcels are  
 126 dispersing mostly to the SH extra-tropics along the isentropes. Above about 35 km parcels are  
 127 also drifting together into the SH. At highest altitudes, parcels are moving out of the tropics into  
 128 the NH extra-tropics.

129  
 130 Timeseries of the zonal mean aerosols and water vapor at 15°S±2.5° are shown in Fig. 3a,b. We  
 131 also plot the temperature anomaly (Fig. 3c) as a departure of the zonal mean MLS temperature  
 132 on Jan 1 from the subsequent zonal mean temperatures. The perturbation heating rate shown is  
 133 Fig. 3d is computed using the AER longwave radiative transfer model (Mlawer et al., 1997).  
 134 The heating rate calculation uses MLS ozone, temperature, and water vapor. The anomaly is  
 135 computed by fixing the water vapor to the pre-eruption profile and computing the heating rates  
 136 over the period. We then subtract those heating rates from the heating rates computed using MLS  
 137 observed water data.



139  
 140 *Figure 3 Times series of aerosol extinction ratio (AE) (Part a), water vapor (Part b),*  
 141 *temperature anomaly (Part c), and heating rate anomaly (Part d) at 15°S. Parts b & c show*  
 142 *black contours of 40 AE that outline the aerosol anomaly. In Part d, the heating rate anomaly*  
 143 *has the Part c temperature contours superimposed. Pink lines in Part a represent the downward*  
 144 *gravitational settling of aerosols of different diameters ( $\mu\text{m}$ ) as labeled. Blue contours in Part c*  
 145 *indicates the altitude of the zero zonal wind lines at the equator showing the descent of the QBO.*  
 146 *The red line in all parts indicates 26km.*

147 First, comparing the aerosol extinction field (Fig. 3a) with the water vapor (Fig. 3b), we see that  
 148 the water vapor anomaly is slowly ascending whereas the aerosol concentration appears to be  
 149 descending. The simple explanation for this effect is that the water vapor is transported upward  
 150 with the diabatic circulation that gives rise to the tropical trace gas tape recorders (Schoeberl et  
 151 al., 2008a) whereas the aerosols are gravitationally settling. The 26 km water vapor anomaly  
 152 ascent rate is  $\sim 2$  km over 80 days or  $\sim 0.028$  cm/s. This vertical velocity is consistent with the  
 153 velocity of  $0.03 \pm 0.005$  cm/s at 26 km estimated by Schoeberl et al. (2008b). Using 0.028 cm/sec  
 154 as a background ascent velocity, Fig 3a shows the net settling rate for aerosols with different

155 sizes after day 60 (Stokes formulas in Pruppacher et al. 1998). The change in the aerosol height  
156 appears to match the settling for aerosol modal diameter of  $\sim 0.8 - 1 \mu\text{m}$ . Smaller particles are  
157 carried upward by the circulation into warmer, lower relative humidity environment, and will  
158 evaporate (Tsagkogeorgas et al., 2017).

159  
160 Second, Fig. 3c shows a cold temperature anomaly that begins to appear in early to mid-  
161 February, and the anomaly magnitude is consistent with radiosonde measurements (Vömel et al.,  
162 2022). This temperature anomaly is approximately coincident with the change in the cooling  
163 rate (Fig. 3d; correlation of  $r = 0.42$ ) due to enhanced water vapor. If we assume in the  
164 thermodynamic equation that the temperature change ( $\Delta T$ ) balances the change radiative heating  
165 ( $\Delta H$ ),  $\Delta T \sim \alpha \Delta H$ , then we compute a Newtonian cooling time scale ( $\alpha^{-1}$ ) of 5 days at 26 km. This  
166 time constant is consistent other estimates of the Newtonian cooling rate for this region (e.g.  
167 Newman and Rosenfield, 1997). Thus, the temperature changes observed in the mid stratosphere  
168 are a thermal adjustment to the increased IR cooling.

169  
170 By mid-March, the descending QBO temperature anomaly is evident in temperature increase at  
171 32 km as the westerly phase descends through this region. The equatorial zero wind line altitude  
172 is superimposed on Fig. 3c to indicate the descent (see [https://acd-  
173 ext.gsfc.nasa.gov/Data\\_services/met/qbo/qbo.html](https://acd-ext.gsfc.nasa.gov/Data_services/met/qbo/qbo.html)).

174  
175 Note that volcanic aerosols can also heat the stratosphere (Aubry et al., 2021 and references  
176 therein). Shortly after the eruption, sonde measurements show a  $< 2\text{K}$  increase in temperatures  
177 below 25 km that disappears by early February (Vömel et al., 2022). After February we see no  
178 evidence of a temperature change co-located with the aerosol layer probably because the  
179 dispersed aerosol layer is too attenuated.

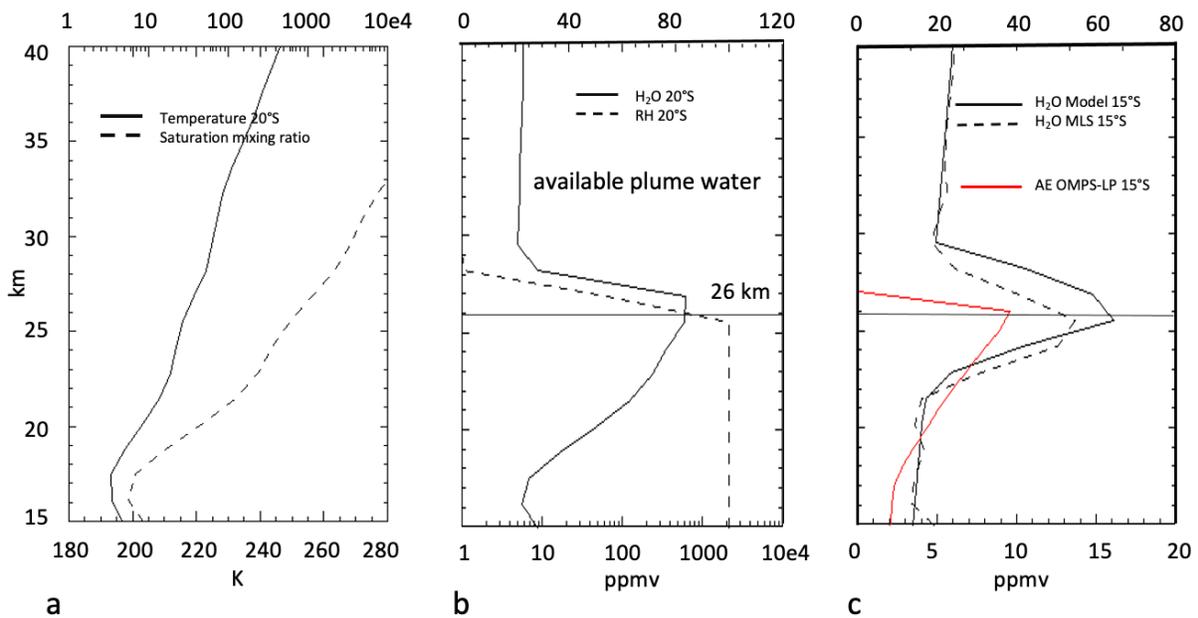
180  
181 Why does the water vapor anomaly extend above the aerosol anomaly? To explore this problem,  
182 we have constructed a very simple model of the HT plume. Initially, the eruption is propelled  
183 upward by the explosion and latent heat release by condensation of water vapor. The plume  
184 temperature is likely well above stratospheric ambient. Within days to weeks the plume shears  
185 out and plume temperature cools to ambient. First, we assume, for simplicity, that the amount of  
186 water vapor available is now limited by the saturation mixing ratio over ice; i.e. excess water  
187 forms ice particles that quickly fall out until the relative humidity is reduced to 100%. We then  
188 assume that the amount of HT water lofted decreases above the eruption centroid height in mid-  
189 February – prior to mid-February the system is still in adjustment. These assumptions define the  
190 available water. Finally, we assume that the volcanic cloud is  $\sim 950 \text{ km}$  in diameter at 16 km  
191 (somewhat larger than the GOES-17 image from NASA Worldview, but not unrealistic). This  
192 area is assumed to expand with altitude, conserving mass. As in Fig. 2e, the eruption cloud  
193 stretches out in longitude. Thus, the mid-February MLS zonal mean water vapor is the available  
194 water in the eruption cloud reduced by the ratio of the initial eruption cloud area to the tropical  
195 zonal mean area.

196  
197 The model uses the OMPS-LP mid-February aerosol extinction profile (Fig. 1a) to set the  
198 eruption centroid height. We center a Gaussian distribution of erupted material at the centroid.  
199 The temperature profile at  $20^\circ\text{S}$  is shown in Fig. 4a along with the saturation mixing ratio over  
200 ice (Murphy and Koop, 2005). Below the eruption centroid, the ice amount is equal to the

201 saturation mixing ratio; above the centroid the amount of water available is the saturation mixing  
 202 ratio times the normalized by the Gaussian centroid. We add the observed background pre-  
 203 eruption zonal mean MLS water vapor profile for realism. Fig. 4b shows the assumed eruption  
 204 available water vapor profile and relative humidity on Feb 15 using centroid height of 26 km and  
 205 a Gaussian standard deviation of 0.65 km. The available water reaches 600 ppmv at 26 km.  
 206 MLS did observe the water vapor mixing ratios over 300 ppm at 26 km Jan. 16 (M22) and there  
 207 are sonde measurements of even higher mixing ratios in this stratospheric region  
 208 (Vömel et al., 2022).

209  
 210 The Feb. 15 zonal mean water vapor field (Fig. 4c) is assumed to be 15° wide from 5°S to 20°S  
 211 and consists of the diluted plume shown in Fig. 3b. Fig. 4c also shows the observed aerosol  
 212 extinction profile. The extinction profile is only used to verify the height of the eruption centroid  
 213 and its width. The water vapor profile shows good agreement with zonal mean MLS data at  
 214 15°S. The extension of the water plume above the aerosol plume is also reproduced. The Fig 4c  
 215 column water vapor mass above 100 hPa is 32.8 Tg; the MLS mass is 31.2 Tg.

216  
 217 In summary, the simple model requires three factors to explain the water vapor anomaly that  
 218 extends above the aerosol anomaly: (1) the change in the saturation mixing ratio with altitude a  
 219 controlled by the tropical temperature profile, (2) a decrease in volcanic water injection above  
 220 the eruption centroid altitude, and (3) spatial dilution of the eruption plume. \



222  
 223 *Figure 4 Model of HT water vapor injection. Part a shows Jan 15, 2022 temperature and*  
 224 *saturation mixing ratio profile at the location of HT (20°S). Part b shows a model of the eruption*  
 225 *available water vapor and relative humidity. Black lines for centroid at 26 km. Part c shows the*  
 226 *zonal mean aerosol extinction ratio profile for Feb. 15, 2022 (red). Zonal mean water vapor*  
 227 *profile (black) for the model and MLS zonal mean water vapor (dashed).*

228  
 229 **5. Summary and Discussion**

230 The HT volcanic eruption produced stratospheric enhancements of both water and aerosols  
231 (sulfate after SO<sub>2</sub> oxidation). Our analysis shows that the aerosol and water vapor enhancements  
232 persisted from Jan 15 to May 31, 2022. Between 22-32 km the enhancements are confined  
233 mostly to the SH tropics as is evident from observations and consistent with a trajectory analysis.  
234 This isolation of the stratospheric SH tropics from the NH tropics is consistent with tropical  
235 ozone observations (Stolarski et al., 2014). Below about 20 km, the aerosol observations and  
236 trajectory analyses show that aerosols and water mostly disperse out of the SH tropics. The  
237 trajectories suggest that most of the aerosols move to the SH with a smaller amount moving into  
238 the NH. Above 40 km the trajectory model suggests that eruption material moves into the  
239 Northern Hemisphere as part of the cross-hemispheric upper stratospheric circulation (Schoeberl  
240 and Strobel, 1978; Holton and Wehrbein, 1980).

241  
242 By mid-February, the tropical mid-stratosphere aerosol and water vapor enhancements are  
243 slightly offset from each other, with the water vapor anomaly about 1 km higher. The two  
244 distinct layers continue separate over the 4 ½ month period after the eruption. The ascent speed  
245 of the water vapor anomaly is consistent with the magnitude of the upward branch of the large  
246 scale diabatic circulation. The descent of the aerosol layer is consistent with the gravitational  
247 settling of particles between 0.6 and 1 µm. Smaller particles will be carried upward and  
248 evaporate.

249  
250 Tropical temperatures at 26 km, 15°S show anomalous decrease about a month after the eruption  
251 at the water vapor enhancement altitude. This temperature decrease is also seen in sonde  
252 measurements (Vömel et al., 2022). IR radiative transfer computations show that the  
253 temperature decrease is associated with enhanced water vapor IR cooling as might be expected  
254 (de F. Forster and Shine, 1999). The Newtonian cooling rate calculated from observed  
255 temperature and cooling rate changes is consistent with previous computations (Newman and  
256 Rosenfield, 1997).

257  
258 To explore the formation of the water vapor anomaly, we use a simple model of the eruption. In  
259 the model we define an eruption centroid altitude, we assume that there is a decreasing amount  
260 of water injected above that altitude and the relative humidity below that altitude is 100%. The  
261 water vapor then disperses zonally. Our model water vapor matches the zonal mean MLS  
262 measurements one month after the eruption and is consistent with the range of MLS H<sub>2</sub>O  
263 measurements made shortly after the eruption (M22).

264  
265 Our simple model suggests that even larger water vapor anomalies would have formed if the  
266 volcanic eruption had lofted water into higher, warmer stratospheric air. On the other hand,  
267 smaller water vapor anomalies would have occurred for lower altitude injections or higher  
268 latitude injections into colder stratospheric air. This, along with the fact that most volcanic  
269 eruptions in the recent past were not submarine may explain why water vapor enhancements  
270 have not been as evident in previous eruptions (M22).

## 271 272 **Acknowledgements**

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274 authors would like to thank Natalya Kamarova for discussions.

275

276 **Open Research**

277 MERRA-2 Reanalysis data. Gelaro et al. (2017). MERRA-2 data are obtained from the Global  
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280 [10.5067/WWQSXQ8IVFW8](https://doi.org/10.5067/WWQSXQ8IVFW8). Data is public, unrestricted access (registration required).

281  
282 OMPS-LP data, Taha et al. (2021), is available at  
283 [https://disc.gsfc.nasa.gov/datasets/OMPS\\_NPP\\_LP\\_L2\\_AER\\_DAILY\\_2/summary](https://disc.gsfc.nasa.gov/datasets/OMPS_NPP_LP_L2_AER_DAILY_2/summary) DOI: [10.506](https://doi.org/10.5067/7/CX2B9NW6FI27)  
284 [7/CX2B9NW6FI27](https://doi.org/10.5067/7/CX2B9NW6FI27). The algorithm is documented in. Data is public, unrestricted access  
285 (registration required).

286  
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289 Meteorological Products (DMPs): <https://mls.jpl.nasa.gov/eos-aura-mls/dmp> (registration  
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