# Semi-analytic approach to the background evolution of f(R)-gravity theories in the metric formalism

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#### Abstract

 $\mathit{Context:}\ \ldots$ 

Aims: An approximation of the modified Friedmann equation in the metric f(R)-gravity context is proposed. Methods: It is applied the Differential Transformation Method to obtain the approximation to the modified Friedmann equations. Results: ...

 $Conclusions: \ \ldots$ 

### Introduction

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## Background evolution for f(R)-gravity in the metric formalism

. . .

### The Differential Transformation Method

The differential transform of some function f(x) is given by

$$F(k) \equiv D[f(x)](k) = \frac{1}{k!} \left. \frac{\mathrm{d}^k f(x)}{\mathrm{d}x^k} \right|_{x_i} \,. \tag{1}$$

The inverse differential transform is then

$$f(x) = D^{-1}[F(k)](x) = \sum_{k=0}^{\infty} F(k)(x - x_i)^k = \sum_{k=0}^{\infty} \frac{(x - x_i)^k}{k!} \left. \frac{\mathrm{d}^k f(x)}{\mathrm{d}x^k} \right|_{x_i} \,. \tag{2}$$

It can be clearly seen that the DTM is the Taylor expansion of a function f(x) around  $x_i$ . In practice, since it is impossible to extend the summation to  $\infty$ , one must break the summation at some desired N terms so that the expansion above converges. Table 1 shows some properties and particular cases of the differential transform.

Table 1: Some DTM properties.

Function $f(x)$	Transform $F(k)$
$\alpha u(x) + \beta v(x)$	$\alpha U(k) + \beta V(k)^1$
$\frac{\mathrm{d}^n u(x)}{\mathrm{d}^n u(x)}$	$\frac{(k+n)!}{U(k+n)}$
$dx^n$	$\sum_{k=1}^{k} \sum_{k=1}^{k_{n-1}} \sum_{k=1}^{k_{3}} \sum_{k=1}^{k_{2}} \prod_{k=1}^{k} (l_{k}) \prod_$
$u_1(x)u_2(x)\ldots u_{n-1}(x)u_n(x)$	$\sum_{k_{n-1}=0} \sum_{k_{n-2}=0} \cdots \sum_{k_{2}=0} \sum_{k_{1}=0} U_{1}(k_{1})U_{2}(k_{2}-k_{1}) \cdots U_{n-2}(k_{n-2}-k_{n-1})U_{n}(k-k_{n-1})$
$x^n$	$\frac{n!}{(n-k)!k!}x_i^{n-k}(=x_i, 1 \text{ and } 0 \text{ for } k=0, k=n \text{ and } k>n, \text{ respectively})^2$

#### Applying DTM to the modified Friedmann equation

The transformed terms up to k = 3 are:

$$U(0) = u_i \tag{3}$$

$$U(1) = u_i^{\prime}(4)$$

$$U(2) = -\frac{1}{36x_i^6 U(0)D[f_{vv}](0)} \left\{ x_i^4 D[f](0) - 3x_i^4 \left[ x_i U(1) + 2U(0) \right] D[f_v](0) + 90x_i^5 U(0)U(1)D[f_{vv}](0) - 6(\Omega_{m,0}x_i + \Omega_{r,0}) \right\} (5)$$

$$U(3) = -\frac{1}{108x_i^6 U(0)D[f_{vv}](0)} \left\{ 4x_i^3 D[f](0) + x_i^4 D[f](1) - 3x_i^3 \left[ 8U(0) + 7x_i U(1) + 2x_i^2 U(2) \right] D[f_v](0) - 3x_i^4 \left[ 2U(0) + x_i U(1) \right] \right\}$$

$$V(k) = 12U(k) + 3\sum_{k_1=0}^{k} \frac{x_i^{1-k_1}}{(1-k_1)!k_1!} (k-k_1+1)U(k-k_1+1)$$
(7)

$$V(1) = 15U(1) + 6x_i U(2) \tag{8}$$

 $^{1}\alpha$  and  $\beta$  are constants.

<sup>2</sup>The result is  $\delta(k)$  when n = 0, where  $\delta(k)$  is the Kronecker delta symbol.

$$D[f](1) = V(1)f_v(v_i)$$
(9)

$$D[f_v](1) = V(1)f_{vv}(v_i)$$
(10)

 $D[f_{vv}](1) = V(1)f_{vvv}(v_i)$ (11)

# Some f(R) examples

...

# Conclusions

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