# Multipole Transforms 

David Kirkby<br>Affiliation not available

December 16, 2019

## 1 Introduction

Our goal is efficient and robust numerical evaluation of Fourier integrals of the form
$f(\vec{r})=N_{n}(a, b) \int d^{n} k e^{+b i(\vec{k} \cdot \vec{r})} \tilde{f}(\vec{k}) \quad, \quad \tilde{f}(\vec{k})=\tilde{N}_{n}(a, b) \int d^{n} r e^{-b i(\vec{k} \cdot \vec{r})} f(\vec{r})$,
with normalization factors

$$
\begin{equation*}
N_{n}(a, b)=|b|^{n / 2}(2 \pi)^{-n(1+a) / 2} \quad, \quad \tilde{N}_{n}(a, b)=|b|^{n / 2}(2 \pi)^{-n(1-a) / 2} \tag{2}
\end{equation*}
$$

where the constants $a$ and $b$ establish our choice of Fourier convention ${ }^{1}$. We focus on two- and three-dimensional $(n=2,3)$ transforms of functions that can be adequately represented with a small number of (not necessarily low order) multipoles. Specifically, for $n=2$, we expand

$$
\begin{equation*}
f\left(r, \varphi_{r}\right)=\sum_{m=-\infty}^{+\infty} f_{m}(r) \Phi_{m}\left(\varphi_{r}\right) \quad, \quad \tilde{f}\left(k, \varphi_{k}\right)=\sum_{m=-\infty}^{+\infty} \tilde{f_{m}}(k) \Phi_{m}\left(\varphi_{k}\right) \tag{3}
\end{equation*}
$$

using the polar basis functions

$$
\begin{equation*}
\Phi_{m}(\varphi) \equiv \frac{1}{\sqrt{2 \pi}} e^{i m \varphi} \tag{4}
\end{equation*}
$$

with orthonormality ${ }^{2}\left(\delta_{D}\right.$ and $\delta$ are the Dirac and Kronecker delta functions, respectively):

$$
\begin{equation*}
\sum_{m=-\infty}^{+\infty} \Phi_{m}(\varphi) \Phi_{m}^{*}\left(\varphi^{\prime}\right)=\delta_{D}\left(\varphi-\varphi^{\prime}\right) \tag{5}
\end{equation*}
$$

[^0]and
\[

$$
\begin{equation*}
\int_{0}^{2 \pi} d \varphi \Phi_{m}(\varphi) \Phi_{m^{\prime}}^{*}(\varphi)=\delta_{m m^{\prime}} \tag{6}
\end{equation*}
$$

\]

Similarly, for $n=3$, we expand
$f\left(r, \theta_{r}, \varphi_{r}\right)=\sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{+\ell} f_{\ell m}(r) Y_{\ell m}\left(\theta_{r}, \varphi_{r}\right) \quad, \quad \tilde{f}\left(k, \theta_{k}, \varphi_{k}\right)=\sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{+\ell} \tilde{f_{\ell m}}(k) Y_{\ell m}\left(\theta_{k}, \varphi_{k}\right)$,
using the spherical-harmonic basis functions ${ }^{3}$ (with associated Legendre polynomials $P_{\ell}^{m}$ )

$$
\begin{equation*}
Y_{\ell m}(\theta, \varphi) \equiv \sqrt{\frac{2 \ell+1}{2} \frac{(\ell-m)!}{(\ell+m)!}} P_{\ell}^{m}(\cos \theta) \Phi_{m}(\varphi) \tag{8}
\end{equation*}
$$

with orthonormality ${ }^{4}$

$$
\begin{equation*}
\sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{+\ell} Y_{\ell m}(\theta, \varphi) Y_{\ell m}^{*}\left(\theta^{\prime}, \varphi^{\prime}\right)=\delta_{D}\left(\cos \theta-\cos \theta^{\prime}\right) \delta_{D}\left(\varphi-\varphi^{\prime}\right) \tag{9}
\end{equation*}
$$

$a n d{ }^{5}$

$$
\begin{equation*}
\int d \Omega Y_{\ell m}(\theta, \varphi) Y_{\ell^{\prime} m^{\prime}}^{*}(\theta, \varphi)=\delta_{\ell \ell^{\prime}} \delta_{m m^{\prime}} \tag{10}
\end{equation*}
$$

In the special case of a three-dimensional $f(\vec{r})$ that is cylindrically symmetric, i.e. has no $\phi_{r}$ dependence, only $m=0$ terms contribute to the multipole expansion equation 7 . Since

$$
\begin{equation*}
Y_{\ell 0}(\theta, \varphi)=\sqrt{\frac{2 \ell+1}{4 \pi}} L_{\ell}(\mu) \tag{11}
\end{equation*}
$$

with $L_{\ell}$ the Legendre polynomial and $\mu \equiv \cos \theta$, it is then convenient to replace equation 7 with the equivalent expansion

$$
\begin{equation*}
f\left(r, \mu_{r}\right)=\sum_{\ell=0}^{\infty} f_{\ell}^{(\mu)}(r) L_{\ell}\left(\mu_{r}\right) \quad, \quad \tilde{f}\left(k, \mu_{k}\right)=\sum_{\ell=0}^{\infty} \tilde{f}_{\ell}^{(\mu)}(k) L_{\ell}\left(\mu_{k}\right) \tag{12}
\end{equation*}
$$

in which the coefficient functions are simply rescaled

$$
\begin{equation*}
f_{\ell}^{(\mu)}(r)=\sqrt{\frac{2 \ell+1}{4 \pi}} f_{\ell 0}(r) \quad, \quad \tilde{f}_{\ell}^{(\mu)}(k)=\sqrt{\frac{2 \ell+1}{4 \pi}} \tilde{f}_{\ell 0}(k) \tag{13}
\end{equation*}
$$

Our general approach is to translate the multi-dimensional Fourier transform, equation 1, into a small number (one per multipole) of one-dimensional

[^1]Hankel ( $n=2$ ) or spherical Bessel $(n=3)$ transforms between the weight functions $f_{m}(r)$ and $\tilde{f_{m}}(k)$ or $f_{\ell m}(r)$ and $\tilde{f_{\ell m}}(k)$. We derive these transforms below then, in the following sections, describe our numerical algorithm for calculating them. Note that we rely on the fact that the Fourier transform does not mix the multipoles, so that there is a one-to-one correspondence between modes in r-space and k-space.

Using equation 6 , we find for $n=2$

$$
\begin{equation*}
f_{m}(r)=\int_{0}^{2 \pi} d \varphi_{r} f\left(r, \varphi_{r}\right) \Phi_{m}^{*}\left(\varphi_{r}\right) \quad, \quad \tilde{f_{m}}(k)=\int_{0}^{2 \pi} d \varphi_{k} \tilde{f}\left(k, \varphi_{k}\right) \Phi_{m}^{*}\left(\varphi_{k}\right) \tag{14}
\end{equation*}
$$

and similarly for $n=3$ using equation 10

$$
\begin{equation*}
f_{\ell m}(r)=\int d \Omega_{r} f\left(r, \theta_{r}, \varphi_{r}\right) Y_{\ell m}^{*}\left(\theta_{r}, \varphi_{r}\right) \quad, \quad \tilde{f_{\ell m}}(k)=\int d \Omega_{k} \tilde{f}\left(k, \theta_{k}, \varphi_{k}\right) Y_{\ell m}^{*}\left(\theta_{k}, \varphi_{k}\right) \tag{15}
\end{equation*}
$$

In the case of $n=3$ cylindrical symmetric, equations 11 and 13 yield the equivalent
$f_{\ell}^{(\mu)}(r)=\frac{2 \ell+1}{2} \int_{-1}^{+1} d \mu_{r} f\left(r, \mu_{r}\right) L_{\ell}\left(\mu_{r}\right) \quad, \quad \tilde{f}_{\ell}^{(\mu)}(k)=\frac{2 \ell+1}{2} \int_{-1}^{+1} d \mu_{k} \tilde{f}\left(k, \mu_{k}\right) L_{\ell}\left(\mu_{k}\right)$.
Next, we replace $f\left(r, \varphi_{r}\right)$ or $f\left(r, \theta_{r}, \varphi_{r}\right)$ in these equations with their Fourier expansions of equation 1 , obtaining

$$
\begin{equation*}
f_{m}(r)=N_{n}(a, b) \int_{0}^{\infty} k d k \int_{0}^{2 \pi} d \varphi_{r} \int_{0}^{2 \pi} d \varphi_{k} e^{+b i(\vec{k} \cdot \vec{r})} \tilde{f}\left(k, \varphi_{k}\right) \Phi_{m}^{*}\left(\varphi_{r}\right) \tag{17}
\end{equation*}
$$

and

$$
\begin{equation*}
f_{\ell m}(r)=N_{n}(a, b) \int_{0}^{\infty} k^{2} d k \int d \Omega_{r} \int d \Omega_{k} e^{+b i(\vec{k} \cdot \vec{r})} \tilde{f}\left(k, \theta_{k}, \varphi_{k}\right) Y_{\ell m}^{*}\left(\theta_{r}, \varphi_{r}\right) \tag{18}
\end{equation*}
$$

Substituting the multipole expansions of $\tilde{f}$ from equations 3 and 7 , we find

$$
\begin{equation*}
f_{m}(r)=N_{n}(a, b) \int_{0}^{\infty} k d k \int_{0}^{2 \pi} d \varphi_{r} \int_{0}^{2 \pi} d \varphi_{k} e^{+b i(\vec{k} \cdot \vec{r})} \sum_{m^{\prime}=-\infty}^{+\infty} \tilde{f_{m^{\prime}}}(k) \Phi_{m^{\prime}}\left(\varphi_{k}\right) \Phi_{m}^{*}\left(\varphi_{r}\right) \tag{19}
\end{equation*}
$$

and

$$
\begin{equation*}
f_{\ell m}(r)=N_{n}(a, b) \int_{0}^{\infty} k^{2} d k \int d \Omega_{r} \int d \Omega_{k} e^{+b i(\vec{k} \cdot \vec{r})} \sum_{\ell^{\prime}=0}^{\infty} \sum_{m^{\prime}=-\ell^{\prime}}^{+\ell^{\prime}} \tilde{\ell}_{\ell^{\prime} m^{\prime}}(k) Y_{\ell^{\prime} m^{\prime}}\left(\theta_{k}, \varphi_{k}\right) Y_{\ell m}^{*}\left(\theta_{r}, \varphi_{r}\right) \tag{20}
\end{equation*}
$$

Finally we expand the plane wave factor in terms of polar $(n=2)$ basis functions

$$
\begin{equation*}
e^{+b i(\vec{k} \cdot \vec{r})}=2 \pi \sum_{m=-\infty}^{+\infty}(b i)^{m} J_{m}(k r) \Phi_{m}^{*}\left(\varphi_{r}\right) \Phi_{m}\left(\phi_{k}\right) \tag{21}
\end{equation*}
$$

or spherical harmonic ( $n=3$ ) basis functions

$$
\begin{equation*}
e^{+b i(\vec{k} \cdot \vec{r})}=4 \pi \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{+\ell}(b i)^{\ell} j_{\ell}(k r) Y_{\ell m}^{*}\left(\theta_{r}, \varphi_{r}\right) Y_{\ell m}\left(\theta_{k}, \varphi_{k}\right), \tag{22}
\end{equation*}
$$

where $J_{m}$ and $j_{\ell}$ are the Bessel and spherical Bessel functions, respectively. Using equations 5 or 9 to simplify the resulting expressions, we find that the moments of $f$ and $\tilde{f}$ are related by a scaled Hankel transform when $n=2$,

$$
\begin{equation*}
f_{m}(r)=c \int_{0}^{\infty} k d k J_{m}(k r) \tilde{f_{m}}(k) \quad, \quad c=2 \pi N_{n}(a, b)(b i)^{m} \tag{23}
\end{equation*}
$$

and by a scaled spherical Bessel transform when $n=3$,

$$
\begin{equation*}
f_{\ell m}(r)=c \int_{0}^{\infty} k^{2} d k j_{\ell}(k r) \tilde{f_{\ell m}}(k) \quad, \quad c=4 \pi N_{n}(a, b)(b i)^{\ell} . \tag{24}
\end{equation*}
$$

Substituting $f \leftrightarrow \tilde{f}, k \leftrightarrow r, N_{n} \rightarrow \tilde{N}_{n}$, and $b \rightarrow-b$, we find similar results for the inverse transforms

$$
\begin{equation*}
\tilde{f_{m}}(k)=c \int_{0}^{\infty} r d r J_{m}(k r) f_{m}(r) \quad, \quad c=2 \pi \tilde{N}_{n}(a, b)(-b i)^{m}, \tag{25}
\end{equation*}
$$

and

$$
\begin{equation*}
\tilde{f_{\ell m}}(k)=c \int_{0}^{\infty} r^{2} d r j_{\ell}(k r) f_{\ell m}(r) \quad, \quad c=4 \pi \tilde{N}_{n}(a, b)(-b i)^{\ell} . \tag{26}
\end{equation*}
$$

The $n=3$ cylindrically symmetric $f_{\ell}^{\prime}(r)$ and $\tilde{f}_{\ell}^{\prime}(k)$ transform identically to $f_{\ell m}(r)$ and $\tilde{f_{\ell m}}(k)$, as given in equations 24 and 26 , since they are simply related by an $\ell$-dependent rescaling factor.

## 2 Spherical Bessel Transforms

Our algorithm for evaluating the scaled spherical Bessel transform equations 24 and $26^{6}$ follows earlier work in references $[3,2]$.

With the change to dimensionless variables

$$
\begin{equation*}
\kappa=\log \left(k / k_{0}\right) \quad, \quad \rho=\log \left(r / r_{0}\right), \tag{27}
\end{equation*}
$$

we can write equation 24 as a convolution suitable for FFT evaluation

$$
\begin{equation*}
\left(r / r_{0}\right)^{\alpha} \cdot f_{\ell m}(r)=\int_{-\infty}^{+\infty} G(\rho+\kappa) F(\kappa) d \kappa \tag{28}
\end{equation*}
$$

with dimensionless functions

$$
\begin{equation*}
G(s)=e^{\alpha s} j_{\ell}\left(k_{0} r_{0} e^{s}\right) \quad, \quad F(s)=c e^{(3-\alpha) s} k_{0}^{3} \tilde{f_{\ell m}}\left(k_{0} e^{s}\right), \tag{29}
\end{equation*}
$$

[^2]where the $r$-weighting parameter $\alpha$ is arbitrary at this point. The function $F$ depends on $f$, but $G(s)$ depends only on the values of $k_{0} r_{0}$ and $\alpha$ (see Fig. 1), and has asymptotes
\[

$$
\begin{equation*}
G_{+}(s)=\frac{1}{k_{0} r_{0}} e^{(\alpha-1) s} \quad, \quad G_{-}(s)=\frac{2^{-(\ell+1)} \sqrt{\pi}}{\Gamma(\ell+3 / 2)}\left(k_{0} r_{0}\right)^{\ell} e^{(\alpha+\ell) s} \tag{30}
\end{equation*}
$$

\]

where $G_{+}$is the envelope of oscillations with a rapidly decreasing period $2 \pi e^{-s} /\left(k_{0} r_{0}\right)$.
In order to simplify the calculations, we use our freedom to chose $k_{0} r_{0}$ and $\alpha$ to symmetrize $G(s)$ via

$$
\begin{equation*}
\alpha=\frac{1-\ell}{2}, \quad\left(\frac{k_{0} r_{0}}{2}\right)^{\ell+1}=\frac{\Gamma(\ell+3 / 2)}{\sqrt{\pi}} \tag{31}
\end{equation*}
$$

which yields

$$
\begin{equation*}
G_{+}(s)=G_{-}(-s)=\frac{1}{2}\left[\frac{\sqrt{\pi}}{\Gamma(\ell+3 / 2)}\right]^{1 /(\ell+1)} \cdot \exp (-(\ell+1) s / 2) \tag{32}
\end{equation*}
$$

We can write this using

$$
\begin{equation*}
s_{0}=\frac{2}{\ell+1} \quad, \quad G_{0}=\frac{1}{k_{0} r_{0}} \tag{33}
\end{equation*}
$$

as

$$
\begin{equation*}
G_{+}(s)=G_{-}(-s)=G_{0} e^{-s / s_{0}} \tag{34}
\end{equation*}
$$



Figure 1: Plots of $f(s) \cdot(\ell+1) / 2$ with $\alpha=0, k_{0} r_{0}=1$ and $\ell=0,2$ (dashed), 4 (dotted). Fainter curves show the large- $s$ asymptotic forms. Left- and righthand plots are for the spherical Bessel and Hankel transforms, respectively.

## 3 Hankel Transforms

We use a similar algorithm to calculate the scaled Hankel transforms of equations 23 and 25 , using primes to denote cases where the $n=2$ and $n=3$ cases differ.

We can transform eqn. 23 to the convolution equation 28 using

$$
\begin{equation*}
G^{\prime}(s)=e^{\alpha s} J_{\ell}\left(k_{0} r_{0} e^{s}\right) \quad, \quad F^{\prime}(s)=c e^{(2-\alpha) s} k_{0}^{2} \tilde{f_{\ell m}}\left(k_{0} e^{s}\right), \tag{35}
\end{equation*}
$$

where, as above, the $r$-weighting parameter $\alpha$ is arbitrary at this point. Fig. 1 compares this $f(s)$ with the spherical Bessel transform case and shows that they have similar asymptotic behavior:

$$
\begin{equation*}
G_{+}^{\prime}(s)=\left(\frac{2}{\pi k_{0}^{\prime} r_{0}^{\prime}}\right)^{1 / 2} e^{\left(\alpha^{\prime}-\frac{1}{2}\right) s} \quad, \quad G_{-}^{\prime}(s)=\frac{2^{-\ell}}{\Gamma(\ell+1)}\left(k_{0}^{\prime} r_{0}^{\prime}\right)^{\ell} e^{\left(\alpha^{\prime}+\ell\right) s} \tag{36}
\end{equation*}
$$

Again, we symmetrize via

$$
\begin{equation*}
\alpha^{\prime}=\frac{1-2 \ell}{4} \quad, \quad\left(\frac{k_{0}^{\prime} r_{0}^{\prime}}{2}\right)^{\ell+1 / 2}=\frac{\Gamma(\ell+1)}{\sqrt{\pi}} \tag{37}
\end{equation*}
$$

which yields

$$
\begin{equation*}
G_{+}^{\prime}(s)=G_{-}^{\prime}(-s)=\left[\pi^{\ell} \Gamma(\ell+1)\right]^{-1 /(2 \ell+1)} \cdot \exp (-(2 \ell+1) s / 4) \tag{38}
\end{equation*}
$$

We can rewrite this using

$$
\begin{equation*}
s_{0}^{\prime}=\frac{4}{2 \ell+1} \quad, \quad G_{0}^{\prime}=\left(\frac{2}{\pi k_{0}^{\prime} r_{0}^{\prime}}\right)^{1 / 2} \tag{39}
\end{equation*}
$$

as

$$
\begin{equation*}
G_{+}^{\prime}(s)=G_{-}^{\prime}(-s)=G_{0}^{\prime} e^{-s / s_{0}^{\prime}} \tag{40}
\end{equation*}
$$

## 4 Discretization

In order to perform the convolution using FFT, we discretize $s_{n}=n \Delta s$ with $|n| \leq N$, and require that both $F$ and $G$ go to zero outside of the sampled interval (large $s_{N}=N \Delta s$ ) and are sampled sufficiently finely (small $\Delta s$ ).

We then pick $s_{N}$ so that $G\left(s_{N}\right) \simeq \epsilon G(0)$, i.e.,

$$
\begin{equation*}
s_{N}=-s_{0} \log \epsilon+\delta \tag{41}
\end{equation*}
$$

where $\epsilon$ is a numerical precision parameter that specifies the fractional level at which we truncate $G(s)$ and we choose $\delta>0$ so that $s_{N}$ lands at the next zero of the oscillation (in the limit of small $\epsilon$ )

$$
\begin{equation*}
e^{\delta}=\lceil Y\rceil / Y \quad, \quad e^{\delta^{\prime}}=\frac{\left\lceil 1 / 8+Y^{\prime} / 4\right\rceil-1 / 8}{Y^{\prime}} \tag{42}
\end{equation*}
$$

where $\lceil Y\rceil$ is the smallest integer $\geq Y$ with

$$
\begin{equation*}
Y \equiv \frac{k_{0} r_{0}}{2 \pi} \epsilon^{-s_{0}} \tag{43}
\end{equation*}
$$

We meet the second condition by requiring $n_{s}$ samples ${ }^{7}$ of the large- $s$ oscillation at $s_{N}$. We accomplish this using $\Delta s \leq \Delta s_{\max }$ with

$$
\begin{equation*}
\Delta s_{\max }=h \epsilon^{s_{0}} \quad, \quad h \equiv \frac{2 \pi}{n_{s} k_{0} r_{0}} \tag{44}
\end{equation*}
$$

In order to preserve the node at $s_{N}$ and ensure that $N$ is integral, we pick

$$
\begin{equation*}
N_{G}=\left\lceil s_{N} / \Delta s_{\max }\right\rceil \quad, \quad \Delta s=s_{N} / N_{G} \tag{45}
\end{equation*}
$$

Fig. 2 shows examples of $G(s)$ tabulated with these prescriptions. Note that the symmetrized $G(s)$ peaks at small positive $s$ but that the oscillations for $s>0$ lead to a negative first moment.

We now turn to the discretization of $F(s)$, which depends on an unknown $f$. Since $s=\kappa$ when evaluating $F(s)$, our choice of $\Delta s$ above directly determines the sampling of $\tilde{f_{\ell m}}(k)$. To ensure that $f$ is sufficiently sampled, we take

$$
\begin{equation*}
\Delta s_{\max }=\min \left(\frac{\log 10}{N_{10}}, c \epsilon^{s_{0}}\right) \tag{46}
\end{equation*}
$$

where $N_{10}$ is the minimum allowed sampling per decade of $\tilde{f_{\ell m}}(k)$.
In practice, we want to know $f_{\ell m}(r)$ over some range $r_{\text {min }}<r<r_{\text {max }}$. Therefore, we set

$$
\begin{equation*}
r_{0}=\sqrt{r_{\min } r_{\max }} \tag{47}
\end{equation*}
$$

and calculate the corresponding $k_{0}$ using the value of $k_{0} r_{0}$ obtained in eqn. (31) or eqn. (37). We want a result that is free of aliasing artifacts over $2 N_{F}$ samples where

$$
\begin{equation*}
2 N_{F}=\log \left(r_{\max } / r_{\min }\right) / \Delta s \tag{48}
\end{equation*}
$$

so we pad the tabulated $G\left(s_{n}\right)$ with $2 N_{F}$ zeros and perform transforms of length $2\left(N_{F}+N_{G}\right)$. The resulting range of $k$ used in the calculation has limits

$$
\begin{equation*}
k_{\min }=k_{0} \exp \left(-\left(N_{F}+N_{G}\right) \Delta s\right) \quad, \quad k_{\max }=k_{0} \exp \left(+\left(N_{F}+N_{G}\right) \Delta s\right) \tag{49}
\end{equation*}
$$

## 5 Errors

We have framed our discretization in terms of $\epsilon$ but, in practice, a better choice would be $\varepsilon=\Delta s / s_{\max }$ since this more directly correlates with the processing

[^3]

Figure 2: Tabulated values of the symmetrized functions $G(s)$ for $\ell=0$ (topleft), 2 (top-right), 4 (bottom-left), and 6 (bottom-right). Curves show the functions $G(s)$ on a log scale, using dashing for negative segments of the oscillation, and points show the discretization $\left|G\left(s_{n}\right)\right|$ for $|n| \leq N$, using $n_{s}=2$ and $\varepsilon=0.01\left(\epsilon \simeq 1.1 \times 10^{-1}, 4.7 \times 10^{-3}, 3.2 \times 10^{-4}, 3.0 \times 10^{-5}\right.$ for $\ell=0,2,4,6$.)
and memory requirements. Therefore we convert a specified $\varepsilon$ into a value of $\epsilon$ by inverting

$$
\begin{equation*}
\varepsilon=h \epsilon^{s_{0}} / \log \epsilon^{-s_{0}} \tag{50}
\end{equation*}
$$

which can either be done numerically or else using the approximation

$$
\begin{equation*}
\epsilon^{s_{0}} \simeq-\frac{L_{0}}{6 L_{1}^{3}}\left[6 L_{1}^{4}+6 L_{1}^{2} L_{2}\left(L_{1}+1\right)-3 L_{1} L_{2}\left(L_{2}-2\right)+L_{2}\left(2 L_{2}^{2}-9 L_{2}+6\right)\right] \tag{51}
\end{equation*}
$$

where

$$
\begin{equation*}
L_{0} \equiv \varepsilon / h \quad, \quad L_{1} \equiv \log \left(L_{0}\right) \quad, \quad L_{2} \equiv \log \left(-L_{1}\right) \tag{52}
\end{equation*}
$$

which is valid for $L_{0} \lesssim 0.35$. The discretizations in Fig. 2 are calculated with a fixed $\varepsilon=0.01$, and so have corresponding values of $\epsilon$ that increase with $\ell$.

In order to provide more direct control of the transform accuracy, we implement a driver routine that starts from some initial $\varepsilon$ (nominally 0.01 ) and then reduces its value by factors until ${ }^{8}$

$$
\begin{equation*}
\left|f_{\ell}\left(r_{j} ; \varepsilon\right)-f_{\ell}\left(r_{j} ; 2 \varepsilon\right)\right|<\max \left(e_{\mathrm{abs}, \ell} \cdot r_{j}^{p}, e_{\mathrm{rel}, \ell} \cdot\left|f_{\ell}\left(r_{j} ; \varepsilon\right)\right|\right) \tag{53}
\end{equation*}
$$

for all $j$, where the $r_{j}$ are equally spaced over the range $r_{\text {min }} \leq r \leq r_{\max }$ and $p$ allows for an $r$-weighing of the 'constant' error criterion (but we normally use $p=0$ ). We abandon the halving of $\varepsilon$ once it reaches some prescribed minimum value (nominally $10^{-6}$ ), to protect against pathological cases.

[^4]For the purposes of transforming a set of multipoles that are subsequently recombined using equation 3,7 , or 12 , we require that ${ }^{9}$

$$
\begin{equation*}
e_{\mathrm{rel}, \ell}=\frac{e_{\mathrm{rel}}}{M} \frac{\left|\sum_{\ell^{\prime}} f_{\ell^{\prime}}(r) L_{\ell^{\prime}}\left(\mu_{r}\right)\right|}{\left|f_{\ell}(r) L_{\ell}\left(\mu_{r}\right)\right|} \quad, \quad e_{\mathrm{abs}, \ell}=\frac{e_{\mathrm{abs}}}{M} \tag{54}
\end{equation*}
$$

for all $\left(r, \mu_{r}\right)$ on a grid covering the region of interest, where $M$ is the number of multipoles included in the expansion.

## References

[1] Scott Dodelson. Inhomogeneities. Elsevier, 2003.
[2] A. J. S. Hamilton. Uncorrelated modes of the non-linear power spectrum. Monthly Notices of the Royal Astronomical Society, 312(2):257-284, Feb 2000.
[3] J.D. Talman. NumSBT: A subroutine for calculating spherical Bessel transforms numerically. Computer Physics Communications, 180(2):332-338, Feb 2009.
[4] David H. Weinberg, Michael J. Mortonson, Daniel J. Eisenstein, Christopher Hirata, Adam G. Riess, and Eduardo Rozo. Observational probes of cosmic acceleration. Physics Reports, 530(2):87-255, Sep 2013.

[^5]
[^0]:    ${ }^{1}$ Use $a=1$ and $b=1$ for the convention of references $[1,4]$.
    ${ }^{2}$ http://dlmf.nist.gov/1.17E12

[^1]:    ${ }^{3}$ http://dlmf.nist.gov/14.30E1
    ${ }^{4}$ http://dlmf.nist.gov/1.17E25
    ${ }^{5}$ http://dlmf.nist.gov/14.30E8

[^2]:    ${ }^{6}$ Since these are simply related by the substitution $r \leftrightarrow k$, we use the notation of equation 24 in the following, without any loss of generality.

[^3]:    ${ }^{7}$ Nyquist limit sampling corresponds to $n_{s}=2$.

[^4]:    ${ }^{8}$ We assume $n=3$ with spherical symmetry here for notational convenience.

[^5]:    ${ }^{9}$ We assume $n=3$ with spherical symmetry here for notational convenience.

