

# Age-Independent Oceanic Plate Thickness and Asthenosphere Melting from SS Precursor Imaging

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## Key Points:

- SS precursors reveal depths of the lithosphere-asthenosphere boundary independent of seafloor age
- The oceanic lithosphere-asthenosphere boundary is a strong reflector and can be explained by 1.5%-2% of partial melt
- Large depth variations of the top and the bottom boundaries of the asthenosphere suggest a heterogeneous melting process

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**Abstract**

The Earth's asthenosphere is a mechanically weak layer characterized by low seismic velocity and high attenuation. The nature of this layer has been strongly debated. In this study, we process twelve years of seismic data recorded at the global seismological network (GSN) stations to investigate *SS* waves reflected at the upper and lower boundaries of this layer in global oceanic regions. We observe strong reflections from both the top and the bottom of the asthenosphere, dispersive across all major oceans. The average depths of the two discontinuities are 120 km and 255 km, respectively. The *SS* waves reflected at the lithosphere and asthenosphere boundary are characterized by anomalously large amplitudes, which require  $\sim 12.5\%$  reduction in seismic velocity across the interface. This large velocity drop can not be explained by a thermal cooling model but indicates 1.5%-2% localized melt in the oceanic asthenosphere. The depths of the two discontinuities show large variations, indicating that the asthenosphere is far from a homogeneous layer but likely associated with strong and heterogeneous small-scale convections in the oceanic mantle. The average depths of the two boundaries are largely constant across different age bands. In contrast to the half space cooling model, this observation supports the existence of a constant-thickness plate in oceanic regions with a complex and heterogeneous origin.

## Plain Language Summary

In the plate tectonic theory, the outermost shell of the Earth consists of a small number of rigid plates (lithosphere) moving horizontally on the mechanically weak asthenosphere. In oceanic regions, the lithosphere is thought to be formed by gradual cooling of the hot mantle. Therefore, the thickness of the plate depends on the age of the seafloor. The problem with the classic half space cooling model is that bathymetry and heat flow measurements at old seafloors do not follow its predicted age dependence. A modified theory, called the plate model, can better explain those geophysical observations by assuming additional heat at the base of an oceanic plate with a constant thickness of about 125 km. However, such a constant-thickness plate has not been observed in seismology. In this study, we image the asthenosphere boundaries using a global dataset of seismic waves reflected off the Earth's internal boundaries. We observe strong reflections from both the top and the bottom of the asthenosphere, across all major oceans. The amplitudes of these waves can be explained by 1.5%-2% of partial melt and the average boundary depths are independent of seafloor age. This observation supports the existence of a constant-thickness plate in the global oceans with a complex origin.

## 1 Introduction

2 In the plate tectonic theory, the outermost shell of the Earth consists of a small number  
3 of rigid plates (lithosphere) moving horizontally on the mechanically weak asthenosphere.  
4 The origin of the asthenosphere as well as the defining mechanism of its top and bottom  
5 rheological interfaces have been highly controversial (Fischer et al., 2010; Rychert et al.,  
6 2012; Kawakatsu & Utada, 2017; Fischer et al., 2020; Rychert et al., 2020; Karato, 1992;  
7 Gaherty & Jordan, 1995). The oceanic plates make up  $\sim 70\%$  of the Earth's surface and  
8 they have a relatively simple geological and tectonic history and therefore they are ideal for  
9 resolving these fundamental questions. The classic half-space cooling model predicts that  
10 the thickness of the thermal boundary layer as well as the depth of the ocean increase pro-  
11 portionally with the square root of the sea-floor age. While this simple conductive cooling  
12 model successfully explains the first-order observations in the oceans, bathymetry and heat  
13 flow measurements at seafloor older than  $\sim 70$  million years do not follow the age dependence  
14 predicted by the half space cooling model. The plate model, which assumes additional heat  
15 at the base of an oceanic plate with a constant thickness, successfully explains the flattening  
16 of sea floor depth and heat flow observations (Parsons & Sclater, 1977). However, such a  
17 constant-thickness plate has not been observed in seismology, and the exact source of the  
18 additional heat remains unclear, probably associated with small-scale convections (Richter,  
19 1973; Richter & Parsons, 1975; Parsons & Sclater, 1977; Richards et al., 2020) or oceanic  
20 hotspots (Korenaga & Korenaga, 2008).

21  
22 It is a general feature in global seismic surface wave studies that the high-velocity lid in  
23 oceanic regions becomes thicker with sea-floor age (French et al., 2013; James et al., 2014;  
24 Godfrey et al., 2017; Ma & Dalton, 2019). It has also been suggested that a plate model with  
25 additional heat at a constant depth of about 125 km fits surface wave observations (Maggi et  
26 al., 2006). Recent surface-wave studies suggest a small amount of melt is trapped within the  
27 entire low-velocity oceanic asthenosphere (Debayle et al., 2020), which might provide the  
28 additional heat required by the plate model. The depths of the lithosphere-asthenosphere  
29 boundary (LAB hereinafter) have also been studied using secondary seismic phases reflected  
30 or converted at the interface, including *SS* precursors (Rychert & Shearer, 2011; Schmerr,  
31 2012; Tharimena et al., 2017), receiver functions (Li et al., 2000, 2004; Kawakatsu et al.,  
32 2009; Rychert & Shearer, 2009) and active source studies (Stern et al., 2015; Mehouchi  
33 & Singh, 2018). Those studies suggest large variations as well as an origin of the astheno-

34 sphere much more complex than gradual thermal variations with depth as predicted in the  
35 half space cooling model (D. Turcotte & Oxburgh, 1967). A variety of rheological mecha-  
36 nisms have been proposed, including a change in grain size (Faul & Jackson, 2005), seismic  
37 anisotropy (Karato & Wu, 1993; Beghein et al., 2014; Auer et al., 2015), elastically accom-  
38 modated grain boundary sliding (Karato, 2012) and near melting (Yamauchi & Takei, 2016).  
39 The large velocity drop and high attenuation also make partial melt a dominant mechanism  
40 in many studies (Fischer et al., 2020; Rychert et al., 2020; Debayle et al., 2020; Stern et  
41 al., 2015; Mehouachi & Singh, 2018; Schmerr, 2012; Tharimena et al., 2017; Li et al., 2000,  
42 2004; Kawakatsu et al., 2009). For example, a thin sublithosphere melt channel beneath the  
43 normal oceanic seafloor has been proposed for the equatorial Atlantic Ocean (Mehouachi &  
44 Singh, 2018).

45  
46 The velocity increase at the base of the asthenosphere has been incorporated in the  
47 widely used 1-D global reference earth model, PREM, with about  $\sim 7\%$  of velocity increase  
48 across the 220-km discontinuity (Dziewonski & Anderson, 1981). A discontinuity at about  
49 this depth was first observed in Europe and North America (Lehmann, 1959, 1961; Hales et  
50 al., 1980) and later confirmed by studies using surface-wave dispersion measurements, un-  
51 derside seismic reflections, ScS reverberations and receiver functions (Goncz & Cleary, 1976;  
52 Vidale & Benz, 1992; Revenaugh & Jordan, 1991; Sacks et al., 1979). This discontinuity has  
53 been reported in continental regions and is also called the Lehmann discontinuity. However,  
54 many studies have concluded that the 220-km discontinuity is not global in nature and a  
55 reflection from this depth is missing in the global long-period stacks (Shearer, 1991; Gu  
56 et al., 2001; Deuss & Woodhouse, 2002, 2004; Schmerr & Garnero, 2006), which indicates  
57 that the existence of the discontinuity is either absent in oceanic regions, or there are large  
58 variations in the depth of this discontinuity.

59  
60 In this paper, we analyze twelve years of long-period transverse component seismograms  
61 recorded at 151 GSN stations (Fig. S1) to investigate  $SS$  waves reflected at the top and  
62 the bottom of the asthenosphere in oceanic regions, namely, the  $S_{\text{LAB}}S$  wave reflected at  
63 the LAB and the  $S_{220}S$  reflected at the bottom of the asthenosphere. We will interpret  
64 the low velocity zone (LVZ) between the two discontinuities observed in oceanic regions as  
65 the asthenosphere (“weak layer”) as it roughly corresponds to the depth range of the global

66 LVZ in model PREM. This is the depth range where geotherm may exceed mantle solidus  
 67 and partial melt occurs (D. L. Turcotte & Schubert, 2002).

## 68 **2 Data & Methods**

69 *SS* precursors are underside reflections from internal discontinuities and they arrive  
 70 before the main *SS* waves which are reflected at the surface of the Earth. *SS* precursors are  
 71 very sensitive to the structures of the interfaces at the reflection midpoints, about halfway  
 72 between the earthquake and the station. They provide good data coverage in the global  
 73 oceanic areas where seismic stations are sparse. In this study, we examine 32,369 broad-  
 74 band seismograms at 151 GSN station from 543 earthquakes occurred between January 2009  
 75 and December 2020 with *SS*-wave reflection points in oceanic regions (Figs 1 & 2). The  
 76 moment magnitude of the earthquakes ranges from 6.0 to 8.0 such that *SS* precursors are  
 77 excited by potentially sufficient seismic energy (Schmerr, 2012). We download seismograms  
 78 from the Data Management Center at the Incorporated Research Institutions for Seismology  
 79 (IRIS). The instrument responses are removed and the East-West and North-South compo-  
 80 nent displacement seismograms are rotated to obtain the radial and transverse component  
 81 seismograms.

82  
 83 The seismograms are band-pass filtered between 10 and 80 mHz and decimated to 0.1  
 84 second sampling interval. Seismograms with weak or complex *SS* waves due to source pro-  
 85 cesses are discarded (Figs S2 & S3). Seismograms with noisy *SS* waves and precursors are  
 86 also discarded (Fig. S3). This leaves 6,143 sets of transverse component seismograms with  
 87 epicentral distances greater than  $80^\circ$ . We visually inspect seismograms for *SS* precursors,  
 88 including  $S_{\text{LAB}}S$ ,  $S_{220}S$ ,  $S_{410}S$  and  $S_{660}S$ . Clear  $S_{\text{LAB}}S$  arrivals are identified on 1,380  
 89 seismograms (about 22.5% of the entire dataset) from 144 stations and 395 earthquakes  
 90 (Figs 1 & 2). The majority of the data with strong  $S_{\text{LAB}}S$  waves (981 out of 1380) have  
 91 focal depths shallower than 45 km, and the epicentral distance varies between  $80.1^\circ$  and  
 92  $176.3^\circ$  with the majority larger than  $100^\circ$  (Fig. S1).

93  
 94 *SS* precursors from the 220-km discontinuity ( $S_{220}S$ ) are observed on 2,756 seismo-  
 95 grams. We are able to pick more  $S_{220}S$  phases than  $S_{\text{LAB}}S$  phases from seismograms partly  
 96 because many  $S_{\text{LAB}}S$  arrivals are too close to the main *SS* wave arrivals and they are not

97 used in this study to avoid strong phase interferences. The  $S_{\text{LAB}}S$  and  $S_{220}S$  signals can be  
 98 both clearly observed on 1,021 of seismograms. There is no apparent geographic pattern as-  
 99 sociated with the presence (or absence) of the SS precursors (Fig. 2). The amplitudes of the  
 100 secondary reflected waves are expected to be small, and the SS precursors are often below  
 101 the noise level due to weak source radiation, small reflection coefficient and/or defocusing  
 102 caused by mantle heterogeneities. For example, only about 30% of the recorded SS waves  
 103 have clear SS precursors from the 410-km and the 660-km discontinuities in recent global  
 104 studies (Guo & Zhou, 2020, 2021).

105  
 106 The observed SS precursors from the LAB ( $S_{\text{LAB}}S$ ) are characterized by large ampli-  
 107 tudes that are comparable to the amplitudes of the mantle transition zone SS precursors  
 108  $S_{410}S$  and  $S_{660}S$  (Fig. 1). To investigate the velocity reduction across the LAB disconti-  
 109 nuity, we construct 1-D reference models modified from PREM to include a large velocity  
 110 drop in the asthenosphere (Fig. 3 & Fig. S4). We calculate synthetic seismograms based on  
 111 traveling-wave mode summation (Liu & Zhou, 2016). The global centroid-moment-tensor  
 112 (CMT) solutions and the USGS Preliminary Determination of Epicenters (PDE) source lo-  
 113 cations and origin times are used in the calculations of the 1-D synthetic seismograms. The  
 114 synthetic seismograms are complete, including all seismic phases with exact amplitudes for  
 115 earthquakes in 1-D earth models. The effects of incident angles on seismic amplitudes have  
 116 been automatically accounted for. The synthetic seismograms are then processed using the  
 117 same bandpass filter as applied to the observed seismograms. The differences in SS precur-  
 118 sors between different models facilitate the identification of the  $S_{\text{LAB}}S$  and  $S_{220}S$  waves on  
 119 the observed seismograms (Fig. 3).

120  
 121 We measure the amplitudes of the SS precursors  $S_{\text{LAB}}S$ ,  $S_{410}S$  and  $S_{660}S$  at a period  
 122 of 25 seconds. The measurements are made in the frequency domain using a 40-second  
 123 window centered at the arrival time of the SS precursors (Fig. 4). We use a short time  
 124 window to limit the interference between the  $S_{\text{LAB}}S$  wave and the main SS wave. Example  
 125 amplitude measurement experiments using synthetic data show that amplitude ratios at  
 126 the measurement frequency as well as their frequency-dependent variations can be well cap-  
 127 tured using a 40-second time window (Fig. 5). We will focus on measurements at a period  
 128 of 25 seconds in this study. The longest half duration of the earthquakes used in this study  
 129 is about 25 seconds and seismic energy often decreases rapidly at frequencies higher than

130 the earthquake corner frequency. In addition,  $SS$  precursors at higher frequencies can be  
 131 heavily contaminated by noises (meteorological and multiple scattering). At periods much  
 132 longer than 25 seconds, seismograms are not suitable for  $SS$  precursor studies as the precursors  
 133 are not well separated and surface-wave overtone dispersion also becomes a problem.  
 134 The frequency dependence of the measurements and their corresponding finite-frequency  
 135 sensitivities will be documented in a separate paper. Amplitude ratios  $S_{LAB}S/S_{410}S$  and  
 136  $S_{LAB}S/S_{660}S$  are calculated for the observed datasets as well as synthetic seismograms in  
 137 1-D reference models (Fig. 6).

138  
 139 We measure the differential arrival times between the  $SS$  waves and their precursors  
 140 in the dataset. The observed and synthetic seismograms are aligned using the  $SS$  travel-  
 141 time measurements, and the residue arrival times of the  $SS$  precursors are then calculated  
 142 in the frequency domain at a period of 25 seconds (Xue et al., 2015) (Fig. 7). The time  
 143 shifts due to uncertainties in source origin times do not affect the final measurements as only  
 144  $\delta t|_{S_{LAB}S} - \delta t|_{SS}$  and  $\delta t|_{S_{220}S} - \delta t|_{SS}$  differential traveltimes are used to determine the depths  
 145 of the discontinuities. The length of the measurement windows ranges from 42 to 117 sec-  
 146 onds for  $SS$  waves, 37 to 69 seconds for  $S_{LAB}S$  waves and 38 to 75 seconds for  $S_{220}S$  waves.  
 147 The length of a measurement window depends on the arrivals of the neighboring phases,  
 148 and the measurement windows are chosen to minimize possible phase interferences. The re-  
 149 lation between time delays and depth perturbations of the discontinuities depends on their  
 150 finite-frequency sensitivity kernels. We calculate finite-frequency traveltime sensitivities to  
 151 boundary depth perturbations in the framework of travelling-wave mode coupling, which  
 152 fully account for source radiation patterns, phase interactions as well as time-domain win-  
 153 dowing and tapering applied in making frequency-dependent measurements (Deng & Zhou,  
 154 2015; Zhou, 2009; Zhou et al., 2005). Example finite-frequency boundary sensitivity kernels  
 155 for traveltime measurements are plotted in Fig. 4. The finite-frequency sensitivities display  
 156 a typical X shape due to the minimax-time nature of the reflected waves (Dahlen, 2005).  
 157 Unlike direct body waves which have minimum-time ray paths, surface-reflected phases are  
 158 minimax waves in that the reflection point is a stationary maximum for perturbations in  
 159 the source-receiver ray plane and a minimum for perturbations perpendicular to the plane.

160  
 161 Seismic waves at different frequencies are sensitive to different regions (Fresnel zones)  
 162 and this introduces frequency-dependent time shifts when lateral variations exist in discon-

163 tinity topography, as a result,  $SS$  precursors do not always have the same polarities as  
 164 expected for 1-D earth models (Guo & Zhou, 2020). The interferences between different  
 165 waves within the measurement window are accounted for in the calculation of the finite-  
 166 frequency sensitivities, including the travel time curves of the interference phases as well  
 167 as their amplitudes. We use travelttime sensitivity kernels to identify possible phase inter-  
 168 ferences in the measurement windows, including interferences with the main  $SS$  waves and  
 169 other phases such as the precursors of depth phases (Fig. S5), precursors and/or multiples of  
 170  $S_{cS}$  waves (Fig. S6). Sensitivity kernels with abnormal values indicate strong interferences  
 171 and those measurements are excluded (Figs S5 & S6). In addition, we exclude measure-  
 172 ments out of the two standard deviations. This leaves 1274 and 929 sets of measurements for  
 173  $S_{LAB}S$  and  $S_{220}S$ , respectively. The thicknesses of the asthenosphere at the 921 locations  
 174 are calculated as the depth difference between the 220-km discontinuity and the LAB.

### 175 **3 Results**

176 We process a total number of 32,369 transverse component seismograms that have re-  
 177 flection points in oceanic regions and observe clear  $SS$  waves on 6,143 seismograms with  
 178 simple source time functions.  $SS$  precursors are secondary reflected waves, their amplitudes  
 179 are much smaller than the main  $SS$  waves and they are often heavily influenced by scat-  
 180 tered waves and phase interactions. As expected, the majority of the seismograms with  
 181 clear  $SS$  waves do not show strong  $SS$  precursors from the two discontinuities. The most  
 182 striking observation from this dataset is the anomalously large amplitudes of the  $S_{LAB}S$   
 183 phases on 1,380 seismograms, with the majority (1,021 out of 1,380) accompanied by strong  
 184  $S_{220}S$  phases. The  $SS$  precursors are well separated from the main  $SS$  waves, and their  
 185 amplitudes are comparable to the amplitudes of the mantle transition zone phases,  $S_{410}S$   
 186 and  $S_{660}S$  (Fig. 1). The geographic distribution of the reflection points is dispersive across  
 187 major oceans, including the Pacific, the Atlantic and the Indian ocean, with seafloor age  
 188 spanning from 10 to 170 million years old (Fig. 1).

#### 190 **3.1 Large Amplitudes of the LAB $SS$ precursors**

191 To quantify the observed large amplitudes of the  $SS$  precursors associated with the  
 192 LAB discontinuity, we calculate the amplitude ratios between the  $S_{LAB}S$  phase and two ref-  
 193 erence phases,  $S_{410}S$  and  $S_{660}S$ . The amplitude measurements are made in the frequency

194 domain based on spectra division at a period of 25 seconds (Guo & Zhou, 2020; Xue et  
 195 al., 2015). In Fig. 6, we plot the histograms of the minimum amplitude ratios, defined as  
 196  $\gamma = \text{minimum}[\log(A_{S_{\text{LAB}}S}/A_{S_{410}S}), \log(A_{S_{\text{LAB}}S}/A_{S_{660}S})]$  (Table S1). We have used the min-  
 197 imum values to avoid over amplification when the amplitude of  $S_{410}S$  or  $S_{660}S$  is small due  
 198 to scattering or defocusing. The precursors  $S_{\text{LAB}}S$  and  $S_{410}S$  (or  $S_{660}S$ ) have similar ray  
 199 paths in the bulk mantle, the use of amplitude ratios minimizes the impact of possible focus-  
 200 ing and defocusing effects. In addition, the main SS waves travel through the same regions  
 201 sampled by the SS precursors, and the distribution the SS amplitude measurements for the  
 202 entire dataset is very similar to that for the subset in which anomalously large  $S_{\text{LAB}}S$  were  
 203 observed, indicating that the observed large amplitudes of the  $S_{\text{LAB}}S$  waves are not a result  
 204 of focusing caused by mantle heterogeneities (Fig. S7). To investigate possible interference  
 205 from the main  $SS$  waves and other phases on the frequency-domain  $S_{\text{LAB}}S$  amplitude mea-  
 206 surements, we make additional time-domain amplitude measurements using the maximum  
 207 value of the envelope function for each  $S_{\text{LAB}}S$  measurement window. The measurements  
 208 made in the frequency domain based on spectra division and those made in the time domain  
 209 using envelope functions in general agree well (Fig. S8). The observed mean value of the  
 210 minimum amplitude ratio  $\gamma$  is close to zero (Fig. 6), indicating that the reflection coeffi-  
 211 cients at the LAB are about the same as those at the 410 and the 660. The corresponding  
 212 velocity contrasts across the LAB are expected to be larger than the contrasts across the 410  
 213 and the 660 due to their smaller incident angles at shallower depths. The geographic dis-  
 214 tribution of the amplitude ratios do not show dependence on the age of the sea floor (Fig. 6).

215

216 To estimate the velocity change across the LAB, we construct a 1-D reference model  
 217 with the depths of the LAB and the 220-km discontinuity at 130 km and 250 km, respec-  
 218 tively (Fig. 3). The velocity jump across the LAB is 12.5% in the reference model (Model  
 219 I).  $SS$  precursors from all upper mantle discontinuities can be clearly identified on the ob-  
 220 served seismograms when the observed and synthetic seismograms are filtered to the same  
 221 frequency band (Fig. 1). We make amplitude measurements and calculate the relative am-  
 222 plitude ratios  $\gamma = \text{minimum}[\log(A_{S_{\text{LAB}}S}/A_{S_{410}S}), \log(A_{S_{\text{LAB}}S}/A_{S_{660}S})]$  using the synthetic  
 223 seismograms following the same process as applied on the observed seismograms. The am-  
 224 plitude ratios between  $S_{\text{LAB}}S$  and  $S_{410}S$  (or  $S_{660}S$ ) calculated for Model I are very close  
 225 to the observations (Fig. 6). To better constrain the velocity drop across the LAB required  
 226 to produce the large amplitudes of  $S_{\text{LAB}}S$ , we introduce two additional models, Model II

227 with 6% of velocity drop and Model III with 11% of velocity drop across the LAB. Example  
 228 synthetic seismograms are plotted in Fig. 3 and the amplitude measurements are plotted  
 229 in Fig. 6. The mean logarithm amplitude ratio  $\gamma$  calculated using Model II as a reference  
 230 model is -0.3, meaning that the corresponding average amplitudes of the LAB precursors  
 231 are about 50% of the amplitudes of the 410 (or the 660) precursors, much smaller than  
 232 the observations. The mean logarithm amplitude ratio  $\gamma$  calculated using Model III as a  
 233 reference model is slightly smaller than the observed value. Based on the calculations, we  
 234 conclude that 12.5% of velocity drop across the LAB is necessary in our model to explain  
 235 the observed large amplitude of the  $S_{\text{LAB}}S$  waves (Fig. 6).

### 237 3.2 Depths of the LAB and the 220-km Discontinuity

238 To constrain the depths of the LAB and the 220-km discontinuity, we measure the  
 239 differential traveltimes  $\delta t|_{S_{\text{LAB}}S} - \delta t|_{SS}$  and  $\delta t|_{S_{220}S} - \delta t|_{SS}$  with respect to synthetic seis-  
 240 mograms calculated in Model I (Fig. 7), similar to the studies of Guo & Zhou (2020, 2021)  
 241 in which  $S_{410}S$  and  $S_{660}S$  traveltime measurements were used to investigate the depths of  
 242 the 410-km and the 660-km discontinuities at a global scale. The time shifts due to uncer-  
 243 tainties in source origin times do not affect the final measurements as we use differential  
 244 traveltimes. We apply 3-D crust and mantle wave speed corrections using global models  
 245 CRUST1.0 (Laske et al., 2013) and S40RTS (Ritsema et al., 2011). Model I is constructed  
 246 as the reference model for the oceanic regions where large-amplitude SS precursors have  
 247 been observed. As it is not a global reference model, there is overall about 5 seconds of  
 248 traveltime shifts in  $\delta t|_{S_{\text{LAB}}S} - \delta t|_{SS}$  after 3-D wavespeed and crustal corrections (Fig. 7 &  
 249 Fig. S9). The mean  $\delta t|_{S_{220}S} - \delta t|_{SS}$  traveltime delay before and after the corrections remains  
 250 approximately the same, with an average value close to zero. This indicate that the average  
 251 velocity in the uppermost mantle in the reference model is close to the global average.

253 To obtain depth perturbations of the LAB and the 220-km discontinuity, we calculate  
 254 the sensitivities of the differential traveltimes to depth perturbations of the interfaces by  
 255 integrating the finite-frequency sensitivity kernels over the surface of the boundary (Fig. 4).  
 256 The LAB depths obtained from this study varies between  $\sim 70$  and  $\sim 160$  km (Fig. 8) and  
 257 the mean LAB depths obtained using the same data with and without the corrections are  
 258 120 km and 125 km, respectively (Figs. 8 & S10). The depth of the 220-km discontinuity

259 varies between 180 and 340 km with a mean value of  $\sim 255$  km, and it does not change  
 260 with wavespeed and crustal corrections. We calculate the thickness of the asthenosphere  
 261 in regions where both the LAB and the 220-km discontinuities are well defined by strong  
 262 SS precursors from both discontinuities. The thickness of the asthenosphere ranges from 50  
 263 km to 220 km with an average of 140 km. Large depth variations of the LAB and the 220-  
 264 km discontinuity are observed across the global oceanic regions and the depth can change  
 265 abruptly over small geographic distances.

266  
 267 The depths of the LAB and the 220-km discontinuity are plotted as a function of seafloor  
 268 age in Fig. 9. The depths of the two discontinuities obtained using the same dataset but  
 269 without the 3-D crustal and mantle wave speed corrections are plotted in Fig. S11. The  
 270 average depth of the two discontinuities are independent of seafloor age, regardless of the  
 271 corrections applied. To quantify uncertainties in traveltime measurements and discontinu-  
 272 ity depth perturbations, we calculate frequency-dependent traveltime measurements at five  
 273 different periods ranging from 23 to 27 seconds. The standard deviation of those traveltime  
 274 measurements are then converted to uncertainties in discontinuity depth using the corre-  
 275 sponding finite-frequency sensitivities. The depth uncertainties are plotted as error bars in  
 276 Fig. 9, they are generally small, with an average of 1.1 km for the LAB and 1.2 km for the  
 277 220-km discontinuity.

## 278 4 Discussions

### 279 4.1 Melt Spots in the Oceanic Asthenosphere

280 The observed large amplitudes of the *SS* precursors require a large velocity change  
 281 across the LAB. The presence of a small amount of melt may significantly reduce seismic  
 282 velocity. It has been suggested that S-wave velocity reduction is about 7.9% for every  
 283 percent of melt in realistically shaped melt in the upper mantle based on finite element  
 284 calculations for shear modulus reduction (produced by the presence of a connected network  
 285 of realistically shaped and naturally organized melt inclusions), with the geometries of the  
 286 inclusions taken directly from laboratory calculations (Hammond & Humphreys, 2000). The  
 287 finite element model predictions are also in general agreement with recent experimental re-  
 288 sults (Chantel et al., 2016). The large amplitudes of the *SS* precursors observed in this  
 289 study can be explained by 1.5%-2% of melt in the asthenosphere. This melt concentration is

comparable to observations at mid-ocean ridges, for example, the East Pacific Rise (Team, 1998). In a recent surface-wave study (Debayle et al., 2020), a large melt fraction of up to 1% beneath the entire oceanic lithosphere has been suggested, in general agreement with the overall estimation of melt (0.3-2%) from electrical conductivity study (Ni et al., 2011).

A sharp increase in the water content across the LAB has been proposed as a possible candidate for significant wave speed reduction through enhanced anelasticity (Karato, 2012; Karato & Jung, 1998). To explore the effect of anelasticity (seismic quality factor  $Q$ ) on the amplitudes of  $SS$  precursors, we calculate synthetic seismograms in models with and without strong anelasticity in the asthenosphere and compare the amplitudes of the  $SS$  precursors. In Fig. 10, the velocity and density structures in Model I and Model IV are identical but their  $Q$  values in the asthenosphere are different,  $Q=80$  in Model I and  $Q=20$  in Model IV. The low  $Q$  value in the asthenosphere in Model IV results in a much smaller  $SS$  amplitude but the amplitude reduction on the  $S_{LAB}S$  wave is very limited. This is because while both the  $SS$  wave and the  $S_{LAB}S$  wave experience more attenuation due to enhanced anelasticity, anelasticity also reduces the effective wave speed in the low  $Q$  region. Therefore, the effective velocity contrast across the LAB increases, resulting a larger reflection coefficient and increased amplitude of the  $S_{LAB}S$  wave. The amplitude ratios calculated in Model I (Fig. 6) and Model IV (Fig. 10) do not show significant differences in their histograms. We conclude that the large amplitudes of  $S_{LAB}S$  waves therefore can not be explained by a change in anelasticity.

The observed large amplitudes of the  $SS$  precursors can not be explained by seismic anisotropy. Strong radial anisotropy (up to 10%) has been observed in the oceanic asthenosphere with SH waves travelling faster than SV waves (Dziewonski & Anderson, 1981; Zhou et al., 2006; Nettles & Dziewoński, 2008; Burgos et al., 2014; Beghein et al., 2014). This radial anisotropy would lead to larger SH wave velocity beneath the LAB, and therefore reduced velocity contrast across the LAB and smaller precursor amplitudes, while the observed large amplitudes of the  $SS$  precursors require anomalously large velocity drop (12.5%) across the boundary. Frozen-in radial anisotropy in the oceanic lithosphere from petrological fabrics or melt ponding has been suggested (Auer et al., 2015). In this case, the SH wave speed would become faster in the oceanic lithosphere, which may cause a larger SH velocity contrast across the LAB but a reduced velocity contrast in SV velocity. In this study, strong

323  $S_{\text{LAB}}S$  phases are also observed on the radial component seismograms (Fig. S12). The ob-  
 324 served  $SS$  precursors with large amplitudes also display a good azimuth coverage (Fig. 1),  
 325 indicating that the observed large amplitudes of the  $S_{\text{LAB}}S$  waves are unlikely a result of  
 326 azimuth anisotropy in the lithosphere (Beghein et al., 2014).

327  
 328 It is important to emphasize that we have interpreted the low velocity zone (LVZ) be-  
 329 tween the two discontinuities observed in oceanic regions as the asthenosphere, and we have  
 330 modeled the wavespeed structure associated with reflected waves as first-order discontinu-  
 331 ities. In 1-D earth models (e.g., PREM), a first-order discontinuity is an equivalent mathe-  
 332 matical representation of the earth structure. The same applies to other discontinuities in  
 333 the Earth including the Moho, the 410 and the 660. Synthetic seismograms calculated based  
 334 on the equivalent first-order discontinuities can explain the observed seismograms. The large  
 335 amplitude of the LAB phase observed in this study requires about 12.5% of velocity jump  
 336 across a first-order discontinuity. The same velocity change over a gradient zone of 5 km or  
 337 less may also explain the observed  $S_{\text{LAB}}S$  amplitudes, with less than 2% of difference in their  
 338 average amplitudes (Fig. S13). If the velocity change occurs over a much larger gradient zone  
 339 of 20 km, the average  $S_{\text{LAB}}S$  amplitude will decrease by  $\sim 17\%$ , and the required velocity  
 340 increase would be larger in order to produce the same peak amplitude (Deng & Zhou, 2015).

## 342 4.2 Age-Independent Thickness of the Oceanic Plate

343 The observed LAB  $SS$  precursors characterized by large amplitudes can be modeled as  
 344 waves reflected off a first-order discontinuity with a large velocity contrast. The strong  $SS$   
 345 precursors from the LAB are observed across the global oceanic regions, with an average  
 346 depth of 120 km that is independent of seafloor age (Fig. 9). This observation supports the  
 347 plate model for the oceanic bathymetry and heat flow measurements, in which a reheating  
 348 process is required at the base of the oceanic plate with a constant thickness of about 100-  
 349 125 km. The reheating process is probably caused by small-scale mantle convection beneath  
 350 the large oceanic plate (Richards et al., 2020).

351  
 352 The observed depths of the LAB in this study are characterized by substantial local  
 353 variations. The standard deviation of the LAB depths calculated for the entire dataset is

354 at  $\sim 15$  km. We calculate the standard deviation of the LAB depths at different length  
355 scales (Fig. S14). The standard deviation can be up to  $\sim 17$  km at small length scales and it  
356 becomes consistent with the standard deviation of the entire dataset when the length scale  
357 exceeds  $\sim 1000$  km. Reflectors at depths between 120 and 180 km have been detected across  
358 the Pacific ocean in a *SS* precursor stacking study, though they were found in only 16%  
359 of the stacks (Schmerr, 2012). LAB reflectors at depths of about 100-140 km have been  
360 reported in Hawaii where high-resolution receiver function studies are available with the  
361 deployment of local stations (Li et al., 2000, 2004). The large variation of the LAB depths  
362 is also consistent with surface waves studies in which individual transects often display con-  
363 siderable depth variability in age-averaged profiles (Rychert et al., 2020; French et al., 2013).  
364

365 In seismic studies, age-dependent reflectors have been reported in several oceanic re-  
366 gions, especially under the young seafloors (Rychert et al., 2021; Rychert & Shearer, 2011;  
367 Tharimena et al., 2017; Wang et al., 2020). For example, a recent receiver function study  
368 from ocean bottom seismometers in the equatorial Mid Atlantic Ocean discovered that the  
369 LAB depth increases from about 30 km at the mid-ocean ridge to about 80 km beneath 30  
370 million years old seafloors in some locations (Rychert et al., 2021). The relation between  
371 the crust age and the LAB depth is much less clear across the Pacific Ocean and it has been  
372 suggested that regional dynamical processes may play an important role in asthenospheric  
373 melt production (Schmerr, 2012). In this study, we use long period *SS* waves, as a result,  
374 signals from a very shallow reflector ( $< 40$  km) will interfere with the main *SS* waves and  
375 will not be picked up. It is also possible that there may be multiple reflectors in the litho-  
376 sphere in some regions and what we observe in this study represents a strong deep reflector  
377 associated with melting not secular cooling. In general, velocity change associated with a  
378 thermal lithosphere is expected to be smaller and much more gradual than the reflectors  
379 associated with the chemical differentiation (melting) process.  
380

381 Similar to the LAB, the depths of the 220-km discontinuity also do not depend on the  
382 age of the seafloor and are characterized with large local variations. It is understood that  
383 the smoothness (roughness) of a same discontinuity in different studies often depends on the  
384 regions of study as well as smoothing applied in some of the inversions. Reflectors at depths  
385 of 250-300 km (the X discontinuity) have been observed in many oceanic regions, including  
386 the South Pacific and the Indian Ocean (Deuss & Woodhouse, 2002, 2004). Large depth

387 variations up to  $\sim 100$  kilometers on the 220-km discontinuity have been report over length  
388 scales of several hundred kilometers beneath the northwestern Pacific in a short-period array  
389 study (Rost & Weber, 2001). The large local variations in the depths of the discontinuities  
390 are expected to generate significant variability in both the waveforms and travel times of the  
391 *SS* precursors. When stacking is applied to *SS* precursors with reflection points in regions  
392 where large depth variations occur over very short distances, it may be possible that the  
393 large amplitudes of *SS* precursors could be effectively averaged out in stacking results due  
394 to phase equalization (Gu et al., 2001). In Fig. S14, we show a simple example to illustrate  
395 the concept that, in some cases, the large amplitudes of *SS* precursors may not be picked  
396 up in stacking results when large variations in *SS* precursor amplitudes and arrival times  
397 are present.

## 399 5 Conclusions

400 The thermal boundary as predicted by the half space cooling model is not observed in  
401 *SS* precursors in this study. Instead, we observe anomalously large amplitudes of *SS* pre-  
402 cursors reflected off the lithosphere-asthenosphere boundary (LAB), which can be explained  
403 by  $\sim 12.5\%$  of velocity drop across the boundary. This indicates 1.5%-2% of localized melt  
404 across the global oceanic regions. The large variability in the depths and amplitudes of the  
405 *SS* precursors observed across the global oceanic region suggests a heterogeneous melting  
406 process in the oceanic asthenosphere. The majority of the LAB *SS* precursors are accom-  
407 panied by strong reflections from the 220-km discontinuity. This indicates that the 220-km  
408 discontinuity may define the lower boundary of the local asthenosphere where melting occurs.

410 The plate model, which requires additional heat at the base of a constant-thickness  
411 oceanic plate, explains the bathymetry and heat flow observations that do not follow half  
412 space cooling predictions. While such a constant-thickness plate has not been reported in  
413 seismic studies, the oceanic plate as defined by the strong LAB reflector in this study does  
414 not thicken with age but show an average depth of 120 km across different age bands. This  
415 observation supports the existence of a constant-thickness plate in oceanic regions. The  
416 localized melt spots distributed over the global oceanic regions may be essential to decouple  
417 the oceanic plates from the underlying mantle by dramatically reducing the mantle viscosity

418 (Debayle et al., 2020; Holtzman, 2016).

419

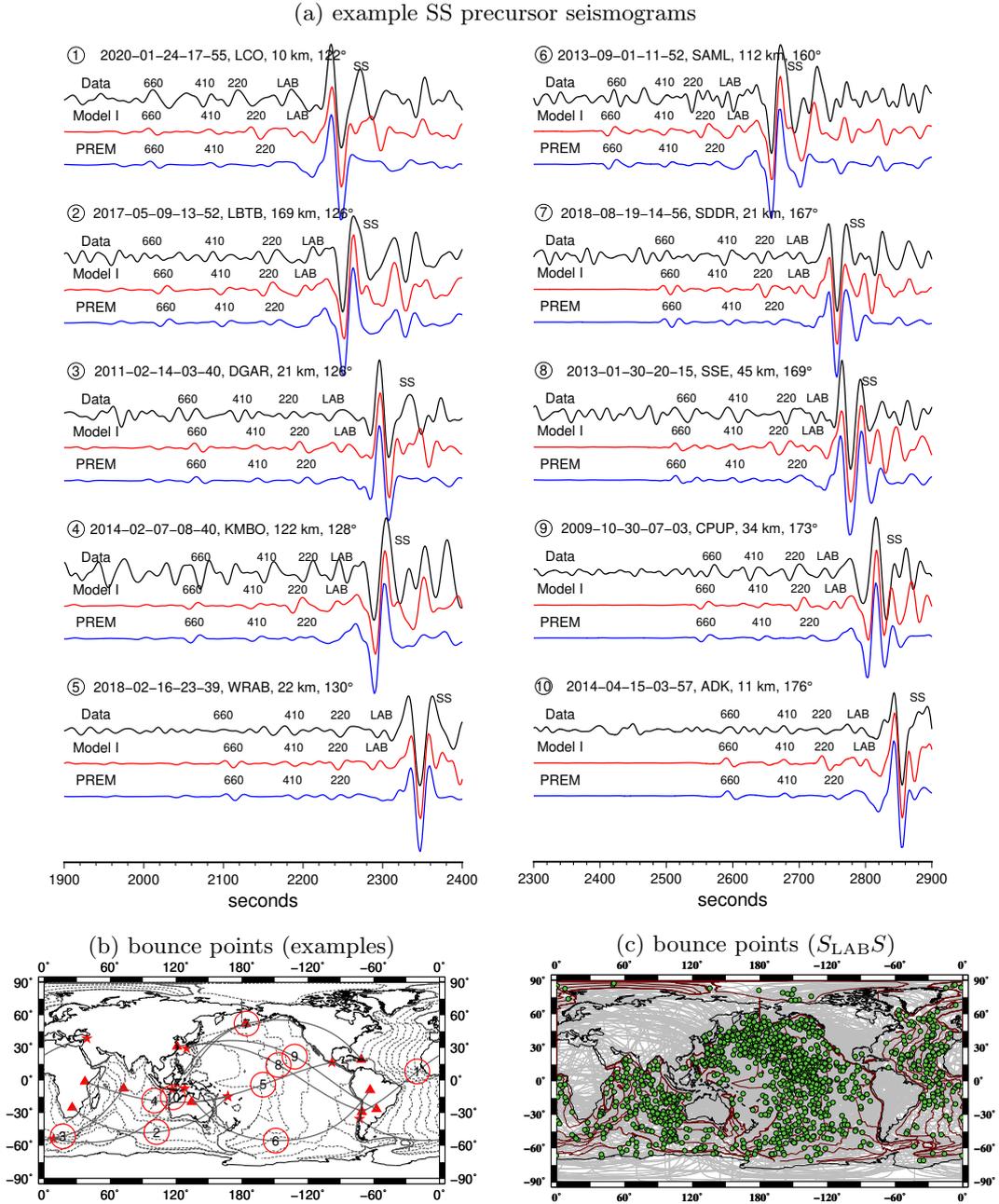


Figure 1: (a) example transverse-component seismograms with large-amplitude  $S_{LAB}S$  and  $S_{220}S$  phases. The black seismograms are data, the red seismograms are synthetics calculated in a reference 1-D model (Model I) with 12.5% of the velocity jump across the LAB, and the blue synthetics are calculated in PREM in which there is no discontinuity at the LAB depth. The reference models are plotted in Fig. 3. The seismograms have been band-pass filtered between 10 and 80 mHz and aligned using their SS arrivals for better illustration. The precursor closest to the main SS wave is labeled as  $S_{LAB}S$  as we investigate possible reflections from the top of the asthenosphere. The arrivals of the  $S_{220}S$ ,  $S_{410}S$  and  $S_{660}S$  waves are also labeled. The earthquake event date/time and station name as well as depth and epicentral distance are denoted on each seismogram. The corresponding geographic ray paths (black lines) and bounce points (red circles) of the  $S_{LAB}S$  waves are plotted in (b). Ray paths and bounce points of the 1,380  $S_{LAB}S$  waves with anomalously large amplitudes are plotted in (c). Seafloor age contours are plotted at 20, 60, 100 and 140 Ma.

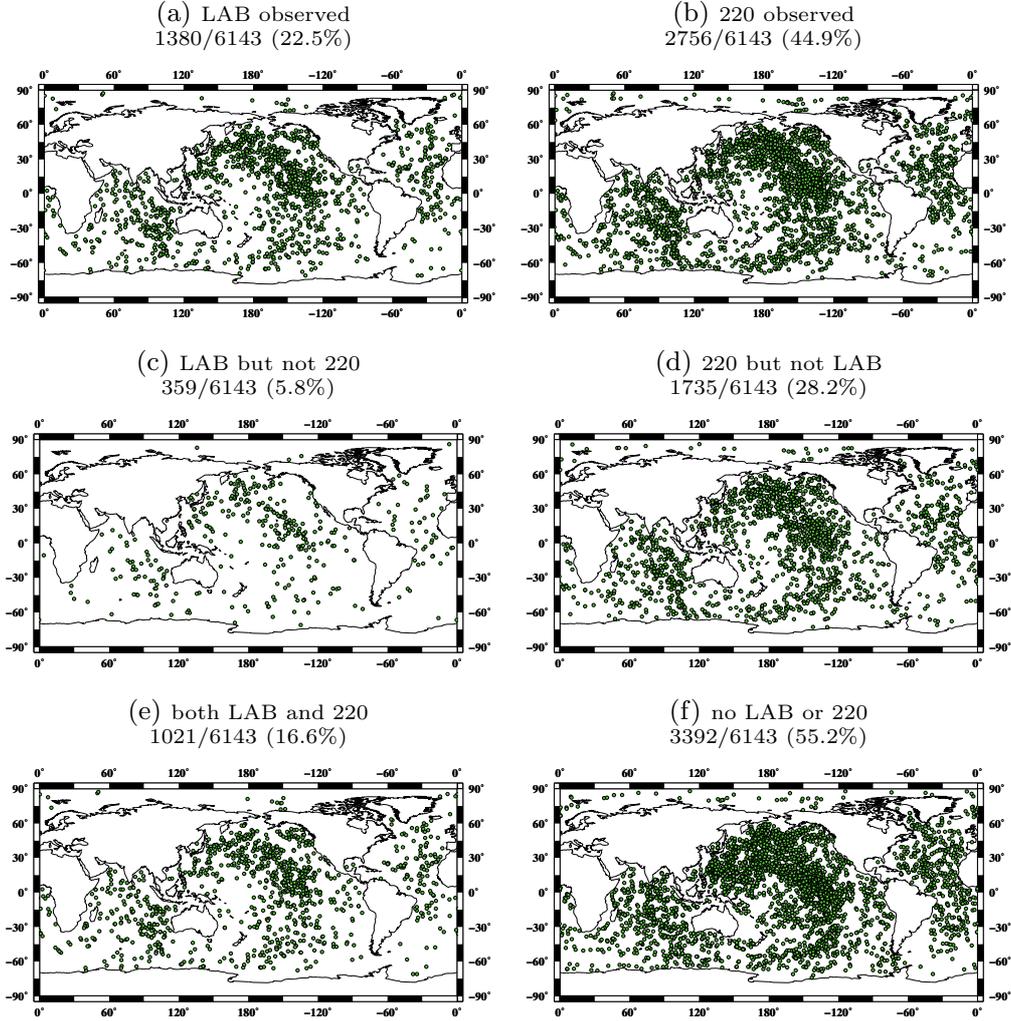


Figure 2: Geographic distribution of subsets of the data plotted at  $SS$  bounce points. (a) clear  $S_{LAB}S$  observed, (b) clear  $S_{220}S$  observed, (c)  $S_{LAB}S$  observed but not  $S_{220}S$ , (d)  $S_{220}S$  observed but not  $S_{LAB}S$ , (e) both  $S_{LAB}S$  and  $S_{220}S$  observed, (f) no  $S_{LAB}S$  or  $S_{220}S$  observed. There is no apparent geographic pattern associated with either the presence or absence of the  $S_{LAB}S$  and  $S_{220}S$ .

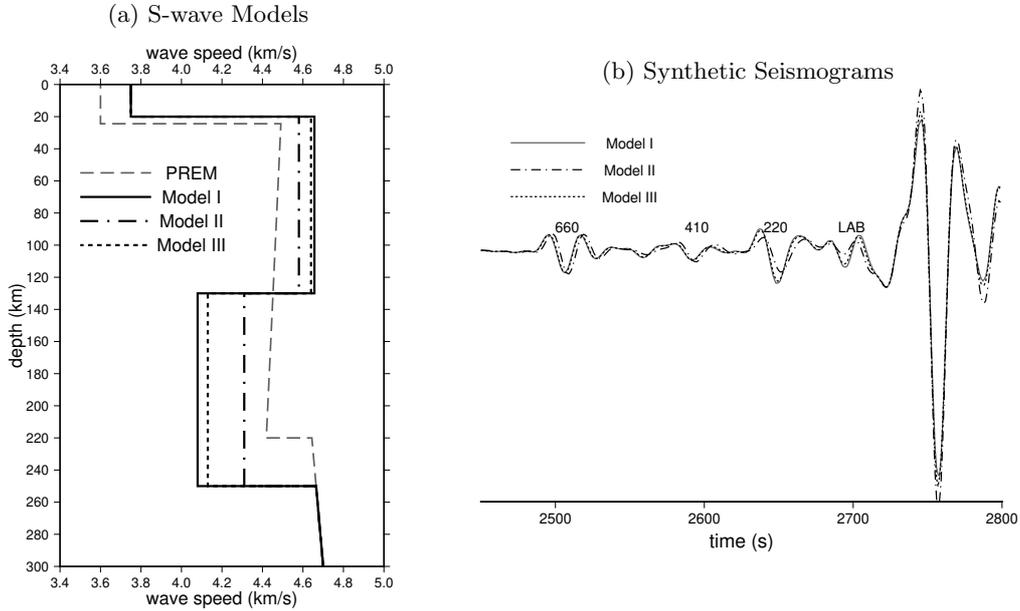


Figure 3: (a) Three 1-D reference earth models Model I, Model II and Model III used in this study and their corresponding example synthetic seismograms are plotted in (b). Model PREM is also plotted in (a) for reference. The synthetic seismograms are calculated using the three reference models for a magnitude 6.9 Indonesia earthquake occurred in August 2018 recorded at a GSN station SDDR (<https://doi.org/10.7914/SN/CU>). The observed seismogram is plotted in Fig. 1. The seismograms have been band-pass filtered between 10 and 80 mHz and have been aligned using the main  $SS$  wave arrivals for better illustration. The velocity contrast across the LAB at a depth of 130 km is 12.5% in Model I, 6% in Model II and 11% in Model III. The corresponding  $S_{LAB}S$  wave amplitude in Model I is much larger than that in Model II and slightly larger than that in Model III, as expected.

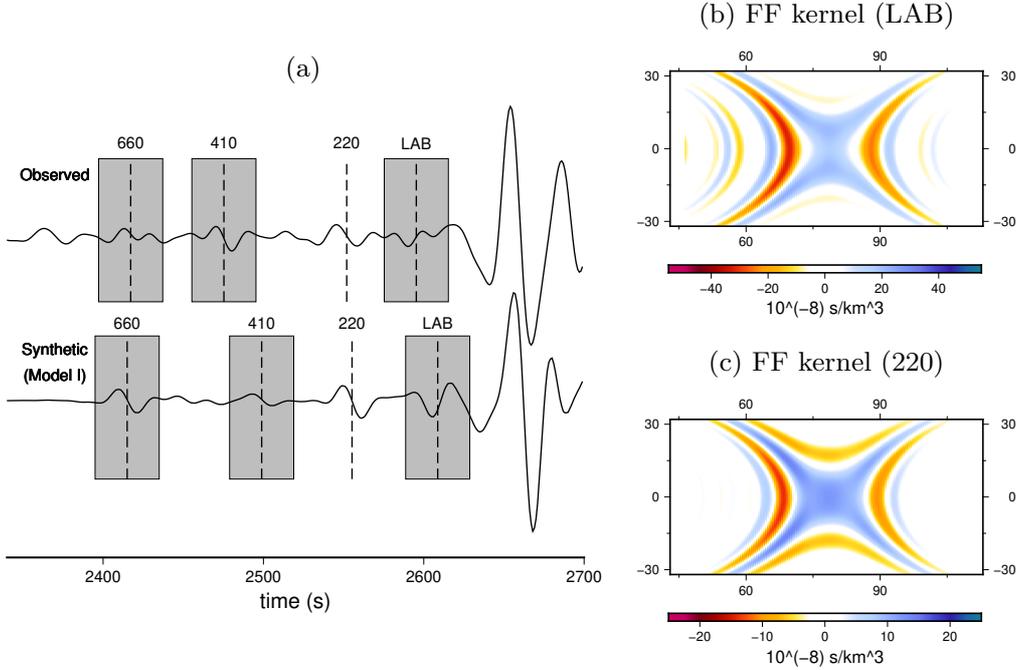


Figure 4: (a) shows example measurement windows used for LAB amplitude ratio measurements  $A_{S_{LAB}S}/A_{S_{410}S}$  and  $A_{S_{LAB}S}/A_{S_{660}S}$ . The seismograms are for a magnitude 6.6 Mexico earthquake occurred in January 2016 recorded at a GSN station MSEY (<https://doi.org/10.7914/SN/II>), and the epicentral distance is  $158^\circ$ . The synthetic seismogram is calculated for Model I as in Fig. 3 and both seismograms have been band-pass filtered between 10 and 80 mHz. (b) and (c) are finite-frequency traveltime boundary sensitivity kernels for the  $S_{LAB}S$  and  $S_{220}S$  waves, respectively. The sensitivity kernels are plotted in map view in the ray coordinates, centered at the bounce point which is about  $79^\circ$  away from the source and the receiver.

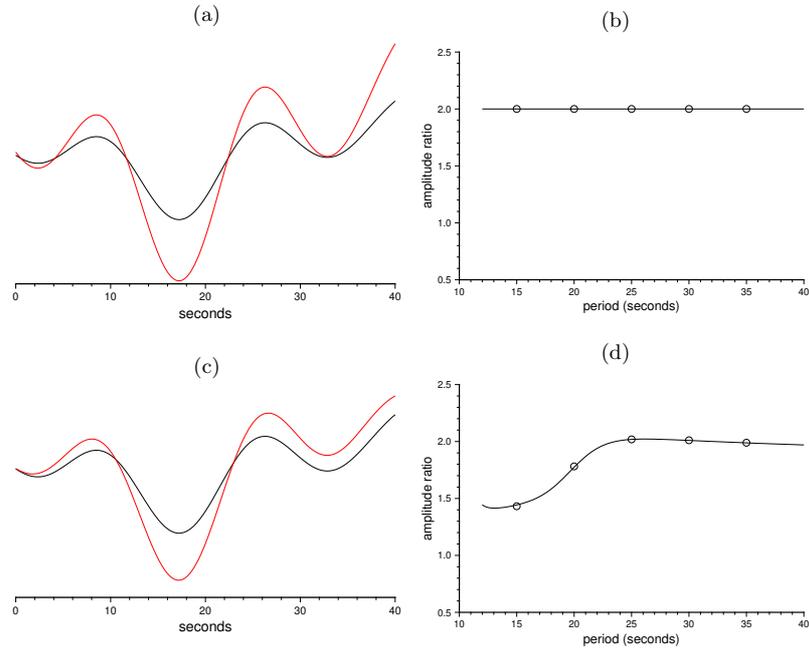


Figure 5: Synthetic amplitude measurement experiment using the same 40-second time window as applied in data. (a) The black seismogram is the LAB phase from the observed seismogram in Fig. 4, and we multiply the seismogram by a constant factor of two (frequency-independent) to obtain the red seismogram. (b) Theoretical amplitude spectral ratio (black line) and measurements made at periods of 15, 20, 25, 30 and 35 seconds (circles). (c) The black seismogram is the same as in (a) and the red seismogram is obtained by multiplying the amplitude spectrum of the black seismogram with a frequency-dependent function. The corresponding theoretical amplitude ratios and measurements made at 15, 20, 25, 30 and 35 seconds period are plotted in (d). The experiment confirms that amplitude ratios at the measurement frequency (25 seconds) can be determined using a 40-second time window.

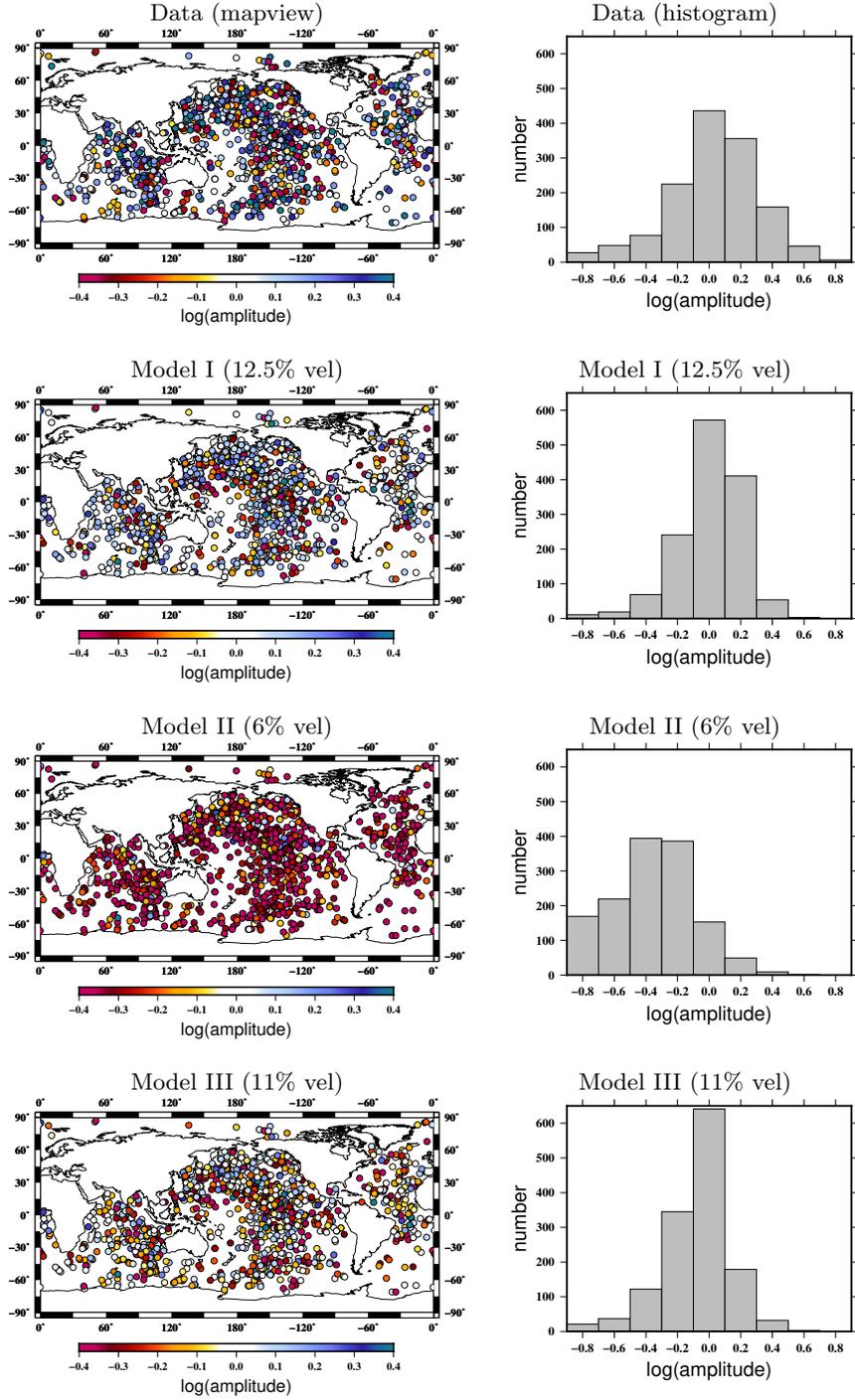


Figure 6:  $S_{LAB}S$  amplitude measurements  $\gamma = \text{minimum} [\log(A_{S_{LAB}S}/A_{S_{410}S}), \log(A_{S_{LAB}S}/A_{S_{660}S})]$  obtained using observed seismograms (top) as well as synthetic seismograms calculated for three reference models (Model I, II and III) plotted at the bounce points in mapviews and histograms. The observed  $S_{LAB}S$  amplitudes show a similar distribution (histogram) to the amplitude ratios calculated for Model I (12.5% velocity drop across the LAB). The amplitude ratios calculated in Model II (6% velocity drop across the LAB) are overall much smaller than the observations, and the amplitude ratios calculated in Model III (11% velocity drop across the LAB) are slightly smaller than the observed amplitude ratios. We conclude that the observed large amplitude of the  $S_{LAB}S$  waves can be explained by 12.5% of velocity drop across the LAB.

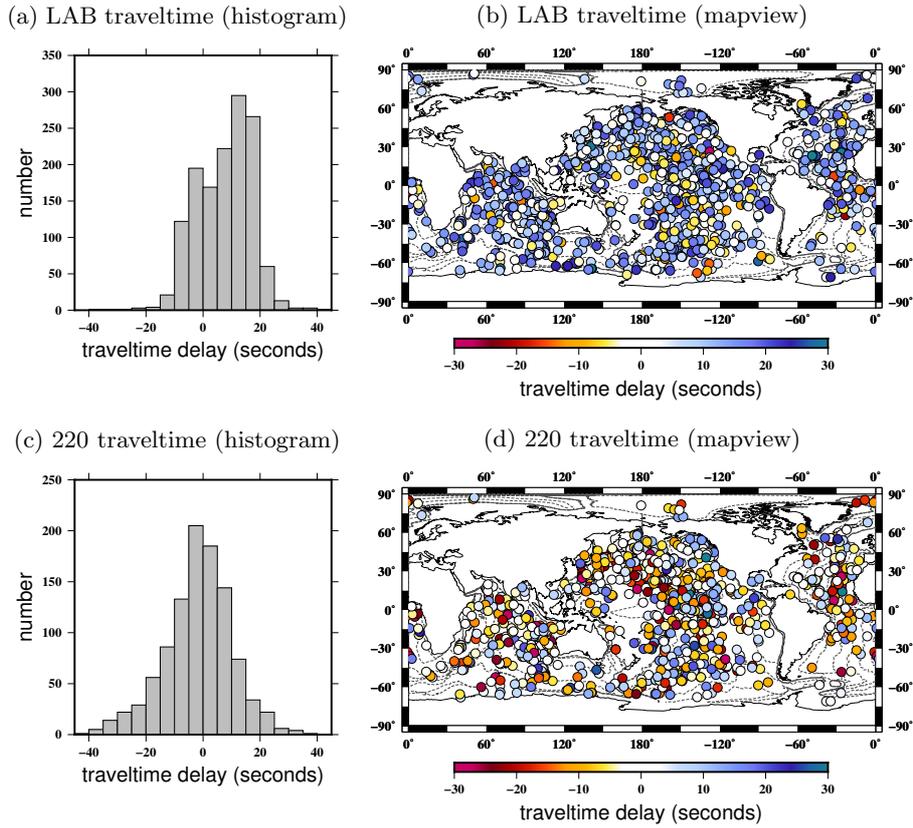


Figure 7: (a) and (b) are  $S_{LABS}$  traveltime measurements plotted in histograms and mapviews at their bounce points. The measurements are made with respect to Model I synthetic seismograms. 3-D mantle wavespeed and crustal corrections have been applied. (c) and (d) are the same as (a) and (b) but for  $S_{220S}$ .

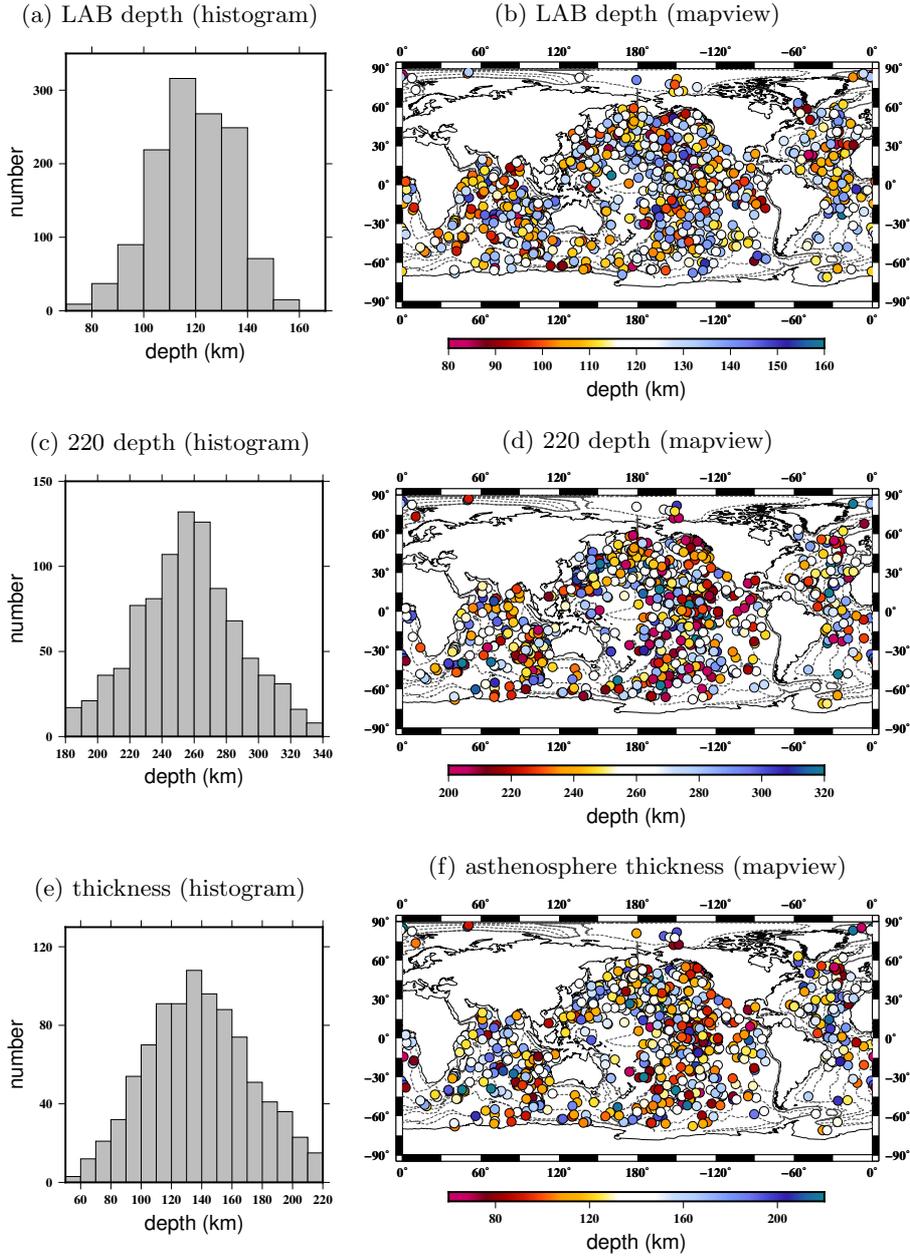


Figure 8: (a) and (b) are the depths of the LAB calculated from traveltimes measurements, plotted in histogram and mapview at  $S_{LAB}S$  bounce points. (c) and (d) are the depths of the 220-km discontinuity. (e) and (f) are the asthenosphere thicknesses calculated from the depths of the LAB and the 220.

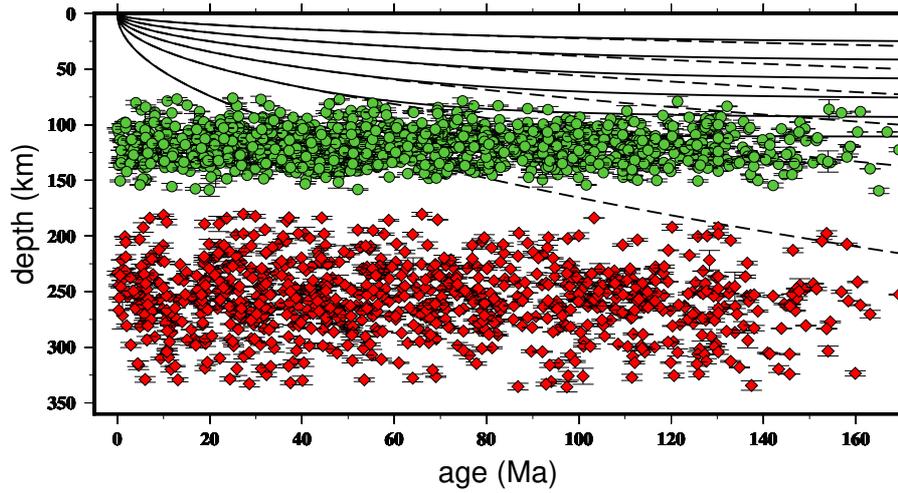


Figure 9: **Age-independent thickness of the oceanic plate.** Green circles and red diamonds are depths of the LAB and the 220-km discontinuity obtained from this study, plotted as a function of the seafloor age. Isotherms at an interval of  $200^{\circ}\text{C}$  (starting at  $300^{\circ}\text{C}$ ) from the half space cooling model (dashed line) and the plate model (solid line) are plotted for reference. The observed depths of the two discontinuities show significant local variations. The average depths of the LAB and the 220-km discontinuity are at 120 km and 255 km, independent of seafloor age. The depth uncertainties estimated from frequency-dependent measurements are plotted as error bars.

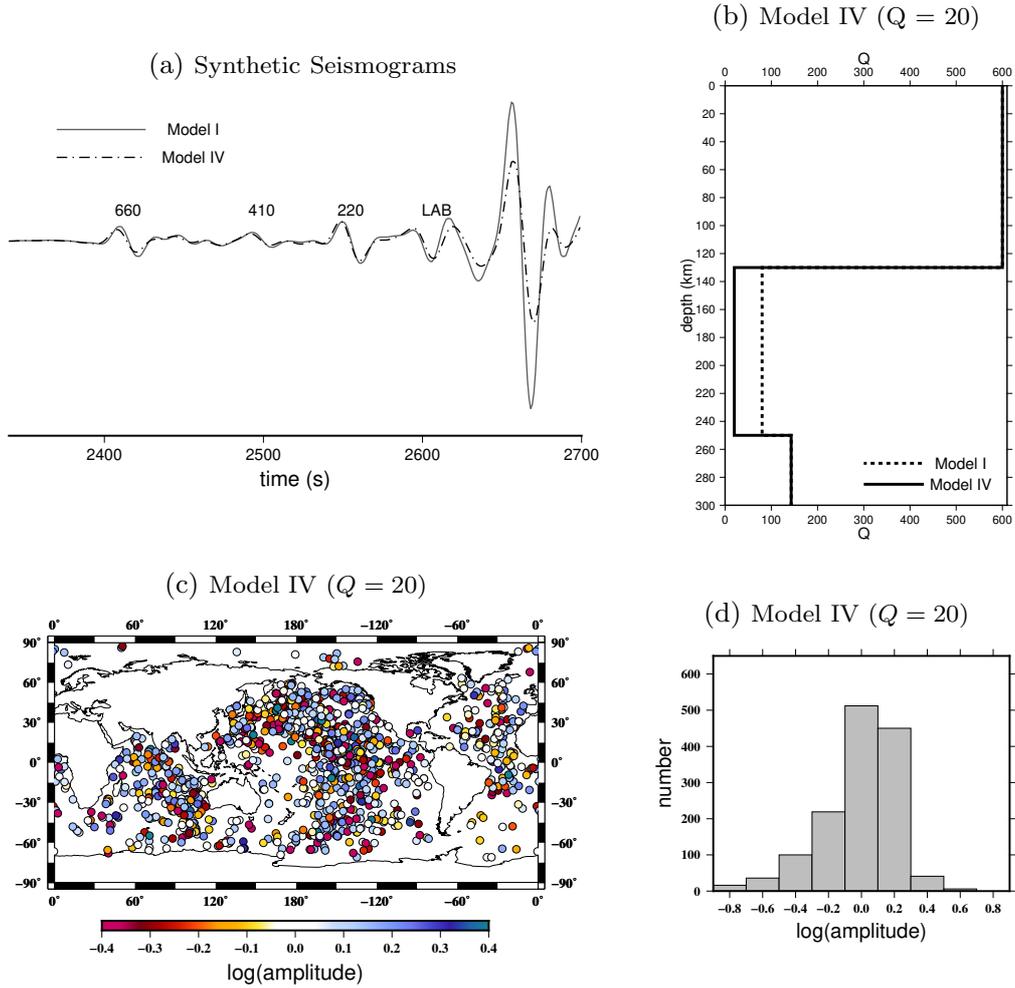


Figure 10: (a) Synthetic seismograms calculate for Model I and Model IV as in Fig. 3. (b)  $Q$  structure in the 1-D reference models Model I and Model IV. Model IV is identical to Model I in velocity and density but has a much smaller  $Q$  value ( $Q=20$ ) in the asthenosphere than in Model I ( $Q=80$ ). The amplitude of the main  $SS$  wave becomes smaller in Model IV synthetics due to the overall stronger attenuation associated with the lower  $Q$  value in the asthenosphere but its impact on the amplitude of the  $SS$  precursor  $S_{LAB}S$  is very limited. This is because anelasticity also reduces the effective wave speed in the low  $Q$  region. Therefore, velocity contrast across the LAB increases, resulting a larger reflection coefficient which increases the amplitude of the  $S_{LAB}S$  wave. The observed large amplitudes of the  $S_{LAB}S$  waves therefore can not be fully explained by a reduction of  $Q$  values in the asthenosphere.

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## **Data Availability Statement**

Raw seismic waveforms used in this study are available from the IRIS platform (<https://ds.iris.edu/ds/nodes/dmc/>). All GSN stations were used (Network codes are CU, GT, IC, II, IU). Data preparation and preprocessing were performed using the Seismic Analysis Code (<https://ds.iris.edu/ds/nodes/dmc/software/downloads/sac/>). Figures were made with the Generic Mapping Tools (GMT) package (Wessel & Smith, 1995).

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