## Seismic source characterization from GNSS data using deep learning

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#### Abstract

The detection of deformation in GNSS time series associated with (a)seismic events down to a low magnitude is still a challenging issue. The presence of a considerable amount of noise in the data makes it difficult to reveal patterns of small ground deformation. Traditional analyses and methodologies are able to effectively retrieve the deformation associated to medium to large magnitude events. However, the automatic detection and characterization of such events is still a complex task, because traditionally-employed methods often separate the time series analysis from the source characterization. Here we propose a first end-to-end framework to characterize seismic sources using geodetic data by means of deep learning, which can be an efficient alternative to the traditional workflow, possibly overcoming its performance. We exploit three different geodetic data representations in order to leverage the intrinsic spatio-temporal structure of the GNSS noise and the target signal associated with (slow) earthquake deformation. We employ time series, images and image time series to account for the temporal, spatial and spatio-temporal domain, respectively. Thereafter, we design and develop a specific deep learning model for each data set. We analyze the performance of the tested models both on synthetic and real data from North Japan, showing that image time series of geodetic deformation can be an effective data representation to embed the spatio-temporal evolution, with the associated deep learning method outperforming the other two. Therefore, jointly accounting for the spatial and temporal evolution may be the key to effectively detect and characterize fast or slow earthquakes.

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#### Key Points:

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10	We develop deep learning approaches on synthetics mimicking the spatio-temporal
11	structure of static deformation and realistic GNSS noise
12	We design three deep learning models and we test them over three GNSS data rep-
13	resentations
14	Transformers and image time series of deformation can effectively characterise small

deformation patterns associated with the seismic source

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#### 16 Abstract

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#### 37 Plain Language Summary

The continuous monitoring of ground displacement with Global Navigation Satel-38 lite System (GNSS) allowed, at the beginning of the 2000's, the discovery of slow earth-39 quakes – a transient slow slippage of tectonic faults that releases stress without gener-40 ating seismic waves. Nevertheless, the detection of small events is still a challenge, be-41 cause they are hidden in the noise. Most of the methods which are traditionally employed 42 are able to extract the deformation down to a certain signal-to-noise level. However, one 43 can ask if deep learning can be a more efficient and powerful alternative. To this end, 44 we address the problem by using deep learning, as it stands as a powerful way to autom-45 atize and possibly overcome traditional methods. We use and compare three data rep-46 resentations, that is time series, images and image time series of deformation, which ac-47 count for the temporal, spatial and spatio-temporal variability, respectively. We train 48 our methods on synthetic data, since real data sets are still not enough to be effectively 49 employed with deep learning, and we test on synthetic and real data as well, claiming 50 that image time series and its associated deep learning model may be more effective to-51 wards the study of the slow deformation. 52

#### 53 1 Introduction

Global Navigation Satellite System (GNSS) is one of the reference sources of in-54 formation in geodesy. Geodetic data can help analyze the ground displacement with mil-55 limeter precision as well as monitor its evolution through time (Blewitt et al., 2018). Such 56 data is commonly used to monitor the ground displacement as a response to environ-57 mental (e.g., tides, snow pack or hydrology), tectonic or seismic forcing, and to charac-58 terize the mechanical response of the Earth to these forcings. Notably, GNSS data has 59 been widely used to study the deformation associated with the different phases of the 60 earthquake cycle. This lead to a better understanding of the loading of faults between 61 earthquakes, of the seismic ruptures studied with either static or kinematic approaches, 62 and of the processes driving the post-seismic relaxation (Bock & Melgar, 2016; Bürgmann, 63 2018, and references therein). At the beginning of the 2000's, the discovery of slow slip 64 events (Lowry et al., 2001; Rogers & Dragert, 2003; Dragert et al., 2001; Ozawa et al., 65

2002) was made possible by the continuous monitoring of ground displacement with GNSS, 66 and constituted a paradigm shift in the understanding of fault mechanics and earthquake 67 physics. Since then, research on slow slip events has been very active. Several studies 68 focus on one particular event or tectonic area, involving visual inspection of the data and 69 dedicated modelling method with a fine-tuning of the parameters (Ozawa et al., 2001; 70 Hirose & Obara, 2005; Wallace & Beavan, 2010; Radiguet et al., 2011; Ozawa et al., 2013; 71 Radiguet et al., 2016; Socquet et al., 2017; Wallace et al., 2016; Wallace, 2020; Itoh et 72 al., 2022; ?, ?). Another research direction aims at performing a systematic character-73 ization of slow slip events (Michel et al., 2019; Frank & Brodsky, 2019; Takagi et al., 2019; 74 Nishimura, 2014, 2021; Donoso et al., 2021; Okada et al., 2022; Rousset et al., 2017), with 75 two underlying objectives: (1) construct catalogues of events allowing for an inter-event 76 comparison and extraction of scaling laws (Ide et al., 2007; Gomberg et al., 2016), and 77 (2) improve the signal over noise ratio in order to detect and characterize events that 78 are at the limit of detection capabilities. The present work is in line with this latter ap-79 proach. Our ultimate goal will be to develop a method able to systematically detect and 80 characterize potential Slow Slip Events (SSEs) on active faults, including small ones, by 81 taking advantage of the large availability of GNSS data. When looking for slow slip events, 82 analyzing their static deformation signature in GNSS time series becomes mandatory 83 since they are not associated with wave arrivals, and because longer time spans need to 84 be considered, with respect to regular earthquakes. To this end, as the static deforma-85 tion associated with regular or slow earthquakes can be approximated with a similar sim-86 ple dislocation model (Okada, 1985), we use GNSS data to characterize the static de-87 formation signature of earthquakes as a preliminary step towards the SSE analysis. Cat-88 alogues listing the source of all  $M_w$  earthquakes are made available by the routine anal-89 ysis of seismic recordings by seismological agencies, allowing for a benchmarking with 90 real GNSS data against an independent ground truth. 91

In this work we focus on the North Japan subduction, offshore Honshu, which is 92 one of the seismic regions that is best instrumented in the world. In this area, slow slip 93 phenomena are scarce compared to warmer subduction zones, such as Cascadia or South 94 Japan, and its kinematics is essentially associated with regular earthquakes (Fukuda, 2018; 95 Hirose et al., 2014). Therefore, new observations, such as seafloor data (Ito et al., 2013; 96 Nishikawa et al., 2019), or new methods applied to terrestrial GNSS records (Nishimura, 97 2014, 2021; Rousset et al., 2017; Khoshmanesh et al., 2020, e.g.) or to seismic records 98 (Marsan et al., 2013; Gardonio et al., 2018, 2019; Uchida et al., 2016, e.g.) are required 99 to detect potential bursts of slow slip. 100

Machine learning and deep learning methodologies have recently been successfully 101 applied to geosciences. In seismology, they have been used to address topics such as earth-102 quake detection and phase selection resulting in seismic catalogues of unprecedented den-103 sity (Zhu & Beroza, 2019; Mousavi et al., 2020; Ross et al., 2019; Kong et al., 2019; Zhu 104 et al., 2019; Seydoux et al., 2020), earthquake early warning (X. Zhang et al., 2021; Münchmeyer 105 et al., 2021; Saad et al., 2020), prediction of ground deformation (Kong et al., 2019; Mousavi 106 et al., 2020), earthquake magnitude estimation (Mousavi & Beroza, 2020; Münchmeyer 107 et al., 2020; Saad et al., 2020). However, machine learning techniques applied to the anal-108 ysis of geodetic time series are less numerous. Relevant applications in the frame of the 109 analysis of the slow slip events have been presented by Rouet-Leduc et al. (2019, 2020); 110 Hulbert et al. (2019, 2020); He et al. (2020), with notable applications to InSAR data 111 by Rouet-Leduc et al. (2021); Anantrasirichai et al. (2019). As we can remark from the 112 literature, seismic recordings are still the main source of information for the analysis of 113 surface ground movements, linked to either slow or regular earthquakes. Thus, this is 114 another motivation to explore the potential of machine learning to analyse GNSS times 115 series. We want to explore and test recent developments in machine learning applied to 116 time series or image analysis, to be able to mine the geodetic data and characterize the 117 events with a physics based approach. 118

In this paper, we address the problem of the fast seismic source characterization, 119 *i.e.*, estimating the location and magnitude of a "regular" seismic event, based on deep 120 learning applied to GNSS position time series. To the best of our knowledge, this is the 121 first attempt of using machine learning–based techniques in such a direction. Again, this 122 is not a goal *per se*, but must be rather seen as a first step towards the development of 123 methods dedicated to the detection of slow slip events. We solve our problem as a re-124 gression in the framework of supervised learning, meaning that the input data used dur-125 ing the training are labelled. The data ground truth comes from seismic catalogs, serv-126 ing as a benchmark for our analyses. We explore three different ways to represent GNSS 127 data (time series, images, image time series) taking into account both the spatial coherency 128 and the temporal variability of GNSS data. We associate a customized deep learning model 129 to each data representation either by re-adapting already existing methods or by design-130 ing it afresh. Training and testing of the different methods is first made on synthetics. 131 The performance of our methods is then evaluated against real GNSS data using an in-132 dependent benchmark coming from actual earthquakes catalogs. The strengths and the 133 pitfalls of the presented methods are discussed by envisioning some possible strategies 134 to improve the results. The same analysis, applied to SSEs, would not be as straight-135 forward, since SSE catalogs are still scarce. However, this methodology can serve as ba-136 sis for further development. 137

#### 138 2 Methods

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# 2.1 Background work and positioning

# 2.1.1 Machine learning and deep learning methods for the seismic source characterization

In the frame of the source characterization, deep learning has proven to be partic-142 ularly effective, as demonstrated by van den Ende and Ampuero (2020) and Münchmeyer 143 et al. (2021), among the most recent works. As pointed out, a multi-station approach 144 may more effectively locate the seismic source, in spite of other approaches using single-145 station waveforms, as (Mousavi & Beroza, 2020). Yet, combining observations from mul-146 tiple stations is indeed a non-trivial task. It is possible to assign a weight to each (seis-147 mic or GNSS) station which depends on certain metrics, as done by Rousset et al. (2017), 148 albeit addressing a different problem. van den Ende and Ampuero (2020) explicitly in-149 ject the location of each seismic stations in form of latitude and longitude coordinates, 150 while Münchmeyer et al. (2021) employ a sinusoidal embedding (*i.e.*, the position is en-151 coded through sinusoidal functions (Vaswani et al., 2017)) for the station locations, out-152 performing already existing methods and showing promising results in terms of earth-153 quake early warning and source characterization. Nevertheless, as a general remark, no 154 straightforward guideline is available to effectively take both the temporal and the net-155 work geometry into account. Therefore, exploiting the spatial distribution is indeed a 156 key problem which we are willing to address in this work. 157

#### 2.1.2 Followed approach

An overview of the proposed methodology is shown in Figure 1. As any standard 159 machine learning model, the pipeline consists in a training and an inference phase. Dur-160 ing the training process, a model is provided with data to learn from. In case of super-161 vised learning, a couple  $\langle$  input, desired output  $\rangle$  is presented to the model, which *learns* 162 by minimizing a certain error metric between the estimated output and the desired out-163 put, which serves as a reference. We use epicenter position and the magnitude of the event 164 as a target output for the characterization, with GNSS data as input. In the inference 165 phase, the trained model is used to make predictions on new data. We will test our meth-166 ods both with synthetic and with real data. We provide new input data to the trained 167 model and we compare the outcomes with the reference outputs, *i.e.*, the epicenter po-168



Figure 1. Schema of the workflow, summarizing the training and the inference phases. A given deep learning model is trained by providing an input and a desired output. Here we use GNSS data as input and a couple consisting of  $\langle$  epicenter position, magnitude  $\rangle$  for each event. During the training process, the model will learn a nonlinear function to map GNSS inputs to an approximation of its position and magnitude. Once trained, this model can be used to perform tests on new, independent data. Here we train on synthetic data and we test both on synthetic and real data.

sition and the event magnitude associated to this new input data. Training our models with supervised learning applied to earthquakes allows us to benefit from a benchmark coming from real earthquakes catalogs.

We make use of synthetic data to train and validate our deep learning models and we test on synthetic and real data afterwards. Japan is probably one of the best instrumented regions in the world, with GNSS data among the cleanest and the densest ones. Yet, we did not train our models with real data for the following two main reasons.

1. GNSS data suffers from the presence of data gaps and missing stations. They can 176 be associated to station inactivity (e.q., electricity blackouts) or to inconsistent 177 daily measurements, for example due to large earthquakes. Moreover, the num-178 ber of GNSS stations may evolve over time, due to the installation of new receivers 179 or to the temporary unavailability of certain ones. It can moreover make it hard 180 to collect regular and well formatted subsets of data to train on. This drastically 181 reduces the number of exploitable training samples, which is indeed a key issue 182 when training deep learning models (LeCun et al., 2015). 183

2. Real data is not uniformly distributed in terms of source parameters, most notably 184 position and magnitude. Since we are dealing with subduction events, most of the 185 actual epicenters will be located on the subduction interface. This can constitute 186 a limitation since a deep learning model trained on such a configuration might not 187 generalize well for events which would be located inshore or sufficiently far from 188 the training area. In addition, the magnitude distribution follows the Gutenberg-189 Richter scaling law (Gutenberg, 1956). As a consequence, the deep learning meth-190 ods would be biased because of the small magnitude events, which will be more 191 numerous, thus possibly resulting in worse performance on the larger ones. To this 192 end, we generate synthetic ruptures whose source parameters are assumed to be 193 random variables drawn from an uniform distribution. 194

By employing synthetic data, it is possible to generate as many samples as needed, overcoming the lack of data and exploiting the features of deep models. Nonetheless, the resemblance between the synthetic data and the real one plays a critical role, since it will have an impact on how well the deep learning model will perform on real data: we need to generate ultra realistic time series. To this end, we add realistic noise computed from actual GNSS data, as it will be detailed in section 2.2.

#### 2.2 Generation and representation of synthetic data

We generate synthetic data samples as the sum of a modeled displacement signal and a realistic noise sample. We rely on three data representations both for synthetic and for real samples and we associate each of them to a different deep learning model. More formally, the synthetic data set is represented as a set of N couples  $\{\mathbf{x}_n, \Theta_n\}_{n=1}^N$ , with  $\Theta$  a set of source parameters (epicenter position, magnitude, focal mechanism, etc.) and  $\mathbf{x}$  being the data following an additive model:

$$\mathbf{x} = \mathbf{s} + \boldsymbol{\varepsilon} \tag{1}$$

with s the synthetic signal (cf. section 2.2.1) and  $\varepsilon$  the noise term (cf. section 2.2.2).

#### 209 2.2.1 Synthetic displacement

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We obtain the synthetic displacement signals **s** by relying on Okada's dislocation model (Okada, 1985). The model input parameters are generated as follows. Earthquake hypocentral positions (longitude, latitude, depth) are assumed to be uniformly distributed random variables, with longitude  $x \sim \mathcal{U}(139^\circ, 146^\circ)$ , latitude  $y \sim \mathcal{U}(35^\circ, 41^\circ)$  and depth  $d \sim \mathcal{U}(2 \text{ km}, 100 \text{ km})$ . Event magnitudes are generated as  $m \sim \mathcal{U}(5.8, 8.5)$  and static moments  $M_0$  are computed accordingly, as (Hanks & Kanamori, 1979):

$$M_0 = 10^{1.5m + 9.1} \text{ N} \cdot \text{m} \tag{2}$$

Fault azimuth direction  $\phi_S$  (strike), dip angle  $\delta$  and slip angle  $\lambda$  (rake) are constrained to a thrust focal mechanism, by allowing for a certain variability of fault slip combinations:  $\phi_S \sim \mathcal{U}(160^\circ, 240^\circ), \delta \sim \mathcal{U}(20^\circ, 30^\circ), \lambda \sim \mathcal{U}(75^\circ, 100^\circ)$ . Static stress drop  $\Delta \sigma$ is assumed to be a lognormal random variable with an average value of 3MPa and a standard deviation of  $\pm 30$  MPa. A circular crack is assumed with radius *R* computed as (Aki & Richards, 2002):

$$R = \left(\frac{7}{16}\frac{M_0}{\Delta\sigma}\right)^{1/3} \tag{3}$$

which can be used to approximate a rectangular dislocation, having length L and width W, by imposing the equality of the surfaces:

$$\pi R^2 = L \cdot W \tag{4}$$

The fault aspect ratio is assumed such that the fault length L and width W satisfy: W = L/2, with L computed as  $L = \sqrt{2\pi}R$ . It should be noticed that the dislocation surface does not change as a function of the aspect ratio between L and W. The average slip  $\bar{u}$  is also derived for a circular crack and it is computed as (Aki & Richards, 2002):

$$\bar{u} = \frac{16}{7\pi} \frac{\Delta\sigma}{\mu} \tag{5}$$

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with  $\mu$  the shear modulus, assumed equal to 30 GPa.

Okada's dislocation model is applied to each one of this set of earthquake sources
to compute the predicted synthetic displacement at each GNSS station in Honshu from
the Earth Observation Network System in Japan (GEONET). Hence, the theoretical deformation field at all station locations in Honshu is obtained for each dislocation setting.

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#### 2.2.2 Realistic noise computation

Noise in GNSS time series constitutes one of the most critical issues, as it is spa-235 tially and temporally correlated (Ji & Herring, 2013; Dong et al., 2002). Here we define 236 noise as everything which is not the signal of interest, being the coseismic signal offsets. 237 At first approximation, its spectrum can be represented as a white noise at the lowest 238 frequencies, and a colored noise having a  $1/f^{\kappa}$  decay starting from a certain corner fre-239 quency, with the spectral index  $\kappa$  being usually fitted from the highest frequencies of the 240 periodogram (Williams et al., 2004; J. Zhang et al., 1997; Mao et al., 1999). The spa-241 tial distribution of such a noise is not random. On one hand, some common patterns must 242 be found among near stations, therefore it can be helpful to discriminate noise from other 243 types of signals. On the other hand, making this type of analysis is difficult, because of 244 the unpredictability of those spatial patterns as well as the intrinsic difficulty in handling 245 such topological consistency in a consistent manner. 246

Realistic perturbations, *i.e.*, noise, are needed to mimic real displacement data. Here
we rely on realistic noise samples computed from real GNSS time series by following an
existing approach for surrogate data generation (Schreiber & Schmitz, 1996; Prichard

& Theiler, 1994). By removing known signals (e.g. earthquakes, postseismic relaxation, 250 SSEs, jumps associated with antenna changes etc) from GNSS time series from a quadratic 251 trajectory model (Marill et al., 2021), we obtain GNSS residual time series that contain 252 the noise that we want to mitigate. Then, a Principal Component Analysis (PCA) is per-253 formed on 100-days windows, by taking into consideration all the stations at the same 254 time. Afterwards, a Fourier Transform (FT) is applied and the phase spectrum is ran-255 domized by picking a new phase  $\varphi \sim \mathcal{U}(0, 2\pi)$ . The same shuffling sequence is adopted 256 for the whole network in order to preserve the spatial coherency between stations. Af-257 ter this process, an Inverse FT and an Inverse PCA are performed. As a result, the trans-258 formed noise samples  $\varepsilon$  will have, on average, the same spatial covariance. Moreover, we 259 can build new noise samples by randomizing the phase, since the Power Spectral Den-260 sity (PSD) of the transformed samples and the actual ones will be asymptotically equiv-261 alent. 262

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#### 2.2.3 GNSS data representations

We build three data types: time series, images and image time series. The raw data come in the form of time series. Then, we derive images to take the spatial information into account, and image time series to take advantage of both the time and space patterns. A schematic view is provided in Figure 2. Moreover, here we do not aim at estimating the hypocentral depth, therefore we do not consider the vertical component of GNSS data, as it does not provide any additional constraint.

Time series. We build synthetic position time series by considering a noise window of 100 days (cf. section 2.2.2). We add a Heaviside step to simulate the coseismic displacement (Bevis & Brown, 2014), with the onset time (cf.  $t_c$  in Figure 2) being at the center of the window. The step amplitude for each station depends on the modeled displacement (cf. section 2.2.1). More formally, the time series structure is represented by a tensor  $\mathbf{X} \in \mathbb{R}^{L \times T \times 2}$ , with L the number of stations and T the number of time steps, the location (latitude, longitude) of the station being given by  $\mathbf{S} \in \mathbb{R}^{L \times 2}$ .

**Differential images.** Images of interpolated deformation field are computed as 277 follows. By assuming the coseismic onset at time  $t_c$ , we consider the difference between 278 the displacement at time  $t_c + 1$  day and  $t_c - 1$  day, namely the differential coseismic 279 displacement field for each station in the GNSS network. We interpolate the deforma-280 tion field in space as follows. We first employ a median anti-aliasing filter with a grid 281 spacing of 25 arc minutes ( $\approx 45$  km), then we interpolate the points in space by using 282 adjustable tension continuous curvature splines (with tension factor T = 0.25) (Smith 283 & Wessel, 1990). The resulting image dimensions are  $76 \times 36 \times 2$  pixels. Afterwards, 284 we mask the sea by forcing to zero all the offshore pixels, in order not to extrapolate off-285 shore, which may degrade the performance of the deep learning methods. Mathemat-286 ically, the differential images are obtained by rasterizing for a given time step  $t_c$  an im-287 age as a tensor  $\mathbf{D} \in \mathbb{R}^{I \times J \times 2}$  being  $I \times J$  the resolution of the image  $\mathbf{D}$  and  $\mathbf{D}(\mathbf{S}(k)) =$ 288  $\mathbf{X}(k, t_c + 1) - \mathbf{X}(k, t_c - 1)$  with  $\mathbf{S}(k)$  the position (latitude, longitude) of the k-th sta-289 tion and  $t_c$  the time of the coseismic offset. The value of I and J, as well as the content 290 of the pixels  $\mathbf{D}(\mathbf{S}(k))$ , for  $k \notin \mathbf{S}$ , have been described before. 291

**Image time series.** Image time series are built from position time series by in-292 terpolating the position information at each frame with the same approach employed for 293 the differential images. We consider 15 days of data, with the first 7 frames correspond-294 ing to the week before the coseismic displacement, the central frame corresponding to 295 the coseismic offset, and the remaining 7 days corresponding to the week after the co-296 seismic. Each frame of the image time series has dimensions  $76 \times 36 \times 2$  pixels. For-297 mally, an image time series is represented by tensor  $\mathbf{T} \in \mathbb{R}^{M \times I \times J \times 2}$ , with M the length 298 of the image time series and  $\mathbf{T}(t_i, \mathbf{S}(k)) = \mathbf{X}(k, t_c + i), i \in (-\lfloor \frac{M}{2} \rfloor, ..., 0, ..., \lfloor \frac{M}{2} \rfloor).$ 299



Figure 2. Outline of the three designed data representations. Each arrangement is designed for a specific deep learning model (cf. Figure 3 with corresponding colors). The data– arrangement procedure is shared between synthetic and real data, except for time series, which are directly available from GNSS recordings. (time series) is associated to the TS model. Synthetic position time series are built by adding a modeled signal (cf. section 2.2.1) to a realistic noise time series (cf. section 2.2.2) by imposing the time of the coseismic offset to be at the center of the window (cf. section 2.2.3). (differential images) is associated to the IMG model. Differential images of ground deformation are built by differentiating the GNSS displacement at the day following the coseismic time and the day before. Then, the differential deformation field is interpolated in space for each direction. (image time series) is associated to the TRA model. Image time series are the 3D–equivalent of position time series. A total of 15 days of deformation is collected, by selecting the week before and the week after the coseismic offset (included). For each day, a spatial interpolation is performed by employing the same method as for differential images to produce a couple of images representing a frame in the whole time series.

In all the three representations, we consider that the coseismic offset time  $t_c$  is known. Indeed, we focus here only on the characterization part, and not the detection.

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#### 2.3 Employed deep learning methods

We developed a deep learning method specifically designed for the characteristics of each chosen data representation. We designed three methods by adapting different stateof-the-art methods that were not originally designed for geodetic data, in order to best address our specific problem. A graphical outline of the methods is provided in Figure 3.

#### $_{308}$ 2.3.1 Time-series based CNN (TS)

Time series can be effectively processed by Convolutional Neural Networks (CNNs), 309 extracting succinct information coming from temporal domain, as reviewed by Bergen 310 et al. (2019); Kong et al. (2019). Here we rely on the architecture proposed by van den 311 Ende and Ampuero (2020), originally proposed for seismic data. Their model has been 312 selected as a potential candidate as it presents several interesting features that can be 313 leveraged also when dealing with geodetic data. The first portion of their network con-314 sists in three convolutional blocks with an increasing number of feature maps. In each 315 block, three convolutional layers are used for the feature extraction, followed by a max-316 pooling layer, employed for subsampling the data. Afterwards, the coordinates of every 317 station associated with an input waveform are injected into the model, as taking into ac-318 count the location of seismic stations can improve the performance, which is the key char-319 acteristic of the model. The max-reduce strategy helps in aggregating the features re-320 lated to the stations, in order to select the feature from the station corresponding to the 321 most relevant contribution for the prediction. We exploited these features in our re-elaboration. 322 Moreover, in order to further mitigate the vanishing gradient problem, the rectified lin-323 ear unit (ReLu) (Agarap, 2018) activation function has been chosen for the hidden lay-324 ers. Since the output variables are uniformly distributed, such an activation function would 325 not squash the predictions in the boundaries of the output range, possibly making the 326 model more flexible when predicting patterns laying outside of the ranges used in the 327 training process, notably when testing for very small or large magnitude earthquakes. 328 The injected horizontal coordinates (latitude, longitude) of GNSS stations are previously 329 scaled in [0, 1]. The original model is also equipped with weights associated to the wave-330 forms accounting for inactivity or missing data from a station. We set them to 1 as the 331 GNSS network in Japan is quite dense and all the stations in synthetic data were assumed 332 to be functioning. Yet, it can represent a further useful development, as it will make the 333 model more flexible when testing on actual data as well as testing against other regions. 334 A visual summary of the model is outlined in the first box of Figure 3. 335

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#### 2.3.2 Image-based CNN (IMG)

We use a 2D CNN to analyze and extract features from interpolated deformation images. They are an effective solution to leverage the spatial coherency and covariance of data structured as images (LeCun et al., 2015) and have become one of the reference architectures for image-based tasks (Goodfellow et al., 2016), also with relevant applications in the geosciences (Rouet-Leduc et al., 2020; Anantrasirichai et al., 2019).

Here we rely on the architecture of MobileNetV2 (Sandler et al., 2018) as the feature extractor. This particular architecture has been chosen as it is lighter (in terms of the number of parameters) with respect to other state-of-art models, such as the VGG family (Simonyan & Zisserman, 2014). Yet, it presents some interesting features, such as the linear bottleneck layers and the depth-wise convolutions. The architecture presents a first convolutional layer followed by seven bottleneck layers. These layers perform an efficient convolution by relying on point-wise and depth-wise convolutions, presenting



Figure 3. The three reference deep learning methods designed in this work. Shaded cyan rectangles represent existing state-of-the-art models. Such models have been slightly modified or adapted, where specified (cf. section 2.3). Further details, such as dropout layers, stride and activation functions, have not been depicted to facilitate the reading. Arrows represent the layers operating between the input (left) and the produced output (right). (TS) The network progressively computes features from convolutions and downsamplings in the time dimension. The latitude and longitude information is then injected. The resulting 2D-array is finally expanded and the contribution coming from the most informative GNSS station is taken (max-reduce operation in yellow). Model readapted from (van den Ende & Ampuero, 2020). (IMG) was inspired by the MobileNetV2 architecture (Sandler et al., 2018). The input two-channel image is processed with convolutions and downsamplings by employing bottleneck layers (cf. section 2.3.2) with and without residual connections (orange and yellow arrows, respectively). (TRA) The first part of the network exploits the feature extractor of IMG to compute spatial features for each frame, which are packed in a 2D-array. Then, a positional embedding enforces time sequencing and prepares the intermediate-level data for the sequential analysis performed by the Transformer (self-attention as in Mousavi et al. (2020)).

residual connections when there is not any stride in the convolutions. We use a global average pooling strategy after the feature extractor. A scheme of the architecture is provided in the second box of Figure 3.

352

#### 2.3.3 Image time series-based Transformer (TRA)

Image time series-based approaches are required to account for both the spatial 353 and the temporal variability into the input data. Deep sequence models such as LSTM 354 (Long-Short Term Memory) or GRU (Gated Recurrent Unit) have been successfully used 355 in geosciences to exploit the sequential behaviour of the data (Bergen et al., 2019; Wang 356 et al., 2017), as well as Transformers, which have overcome the former becoming the ref-357 erence methods in the state-of-art (Vaswani et al., 2017; Münchmeyer et al., 2021; Mousavi 358 et al., 2020). We tested both the LSTM and the Transformer approaches and we chose 359 the latter, whose complexity is justified by its better ability to constrain the spatio-temporal 360 evolution. 361

Here we design a relatively simple model to validate to consider both spatial and 362 temporal features jointly, which can serve as a baseline to add more complexity in the 363 future. We first use a feature extractor to compress the input data dimensionality to ob-364 tain a reduced representation. We use the same architecture of the IMG feature extrac-365 tor and we distribute it in time, *i.e.*, we use the same feature extractor for each frame 366 of the image time series. As a result, we obtain a feature vector for each frame of the 367 image time series. Afterwards, we stack all the feature vectors in one matrix to be ex-368 plotted by the Transformer layer, as shown in the third box of fig 3. Since the self-attention 369 is, in general, order agnostic, we apply a Positional Embedding layer to ensure that the 370 relative position of the frame information is correctly enforced (Chollet, 2021). We chose 371 not to have a fixed mapping, therefore the embedding weights are learnt during the train-372 ing phase. After the embedding layer, we use a Transformer equipped with additive self-373 attention, as in (Mousavi et al., 2020). For simplicity, we use only one global self-attention. 374 According to our preliminary tests, the performance is not considerably increasing when 375 adding a second level of attention, possibly because our model is still too simple to ben-376 efit from a hierarchical attention structure. After the self-attention, we apply another 377 dropout (dropout rate 0.5) layer (cf. section 2.3.1) followed by a one dimensional Global 378 Max Pooling. As a final remark, we train the model by enforcing the feature extractor 379 to evolve from weights already learnt by IMG. Therefore, we apply a sort of fine-tuning 380 which may be beneficial for the self-attention to reach some acceptable parameter con-381 figurations in the early stage of the training already. The TRA architecture is presented 382 in the third box of Figure 3. 383

384

#### 2.4 Implementation and training details

We enforced the mean squared error (squared L2 norm) as loss function, *i.e.*, the objective function which is minimized during the training, defined as follows:

$$\mathcal{L}(\mathbf{y}, \hat{\mathbf{y}}) = \frac{1}{N} \frac{1}{d} \sum_{i=1}^{N} \sum_{j=1}^{d} (y_{i,j} - \hat{y}_{i,j})^2$$
(6)

where  $\mathbf{y} \in \mathbb{R}^{N \times d}$  and  $\hat{\mathbf{y}} \in \mathbb{R}^{N \times d}$  represent the ground truth and the predicted 387 output, respectively, with N being the number of observations and d the number of di-388 mensions. Notably, d = 3, being latitude, longitude and magnitude the output variables. 389 Hence, the loss function jointly minimizes the error on both position and magnitude. Since 390 the ranges of the output variables are not comparable, they are first scaled in (0, 1). Thanks 391 to this transformation, the high-range variables do not prevail on the others, possibly 392 masking small variations on low-magnitude variables. As a result, the loss minimization 303 turns out to be more regular and effective. 394

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Table 1. Quantitative results of the tested methods.

Model	Position error (km)	Magnitude error (°)
TS IMG	$156.46 \pm 116.94$ $133.07 \pm 146.97$	$0.26 \pm 0.20 \\ 0.18 \pm 0.18$
TRA	$\textbf{105.44} \pm \textbf{128.84}$	$0.13 \pm 0.15$

All the three models have been provided with a last fully-connected layer with three 395 outputs and a linear activation function (linear combination). Since the output variables 396 are uniformly distributed, such an activation function would not squash the predictions 397 in the boundaries of the output range, possibly making the model more flexible when 398 predicting patterns laying outside of the ranges used in the training process. Thereafter, 300 we enforce a dropout regularization (Srivastava et al., 2014) in this final layer (dropout 400 rate 0.5) at training time, which helps prevent the models from overfitting the training 401 data, in addition to the dropout regularization which may already be enforced through-402 out the previous layers. 403

We performed the training of the three models by adopting a mini-batch stochastic gradient learning (Bottou et al., 2018) with a batch size of 128 samples and the ADAM method (Kingma & Ba, 2014) for the optimization. The learning rate was chosen according to a grid-search optimization and the best value was found at 0.001. We initialize all the network weights with an orthogonal initializer (Saxe et al., 2013) for TS and with a uniform Xavier initializer (Glorot & Bengio, 2010) for IMG and TRA.

We employ twenty thousand synthetic samples that we divide it into training, val-410 idation and test sets with proportions of 60%, 20% and 20% respectively. We used the 411 training and validation sets for the training phase. When the loss on the validation set 412 is not decreasing anymore in a certain number of training steps, the training is termi-413 nated and the model's weights are loaded with the ones associated to the best loss value. 414 Moreover, the validation set has been employed to tune the hyper-parameters of the mod-415 els (such as the learning rate, the best architecture, etc.) in order to prevent any over-416 fitting. The test set is used for the final inference and for the performance analysis. 417

The code was implemented in Python using the Tensorflow (Abadi et al., 2016) library as well as the higher–level package Keras (Chollet et al., 2015). The training was run on NVIDIA Tesla V100 Graphics Processing Units (GPUs).

#### <sup>421</sup> 3 Results on synthetic data and discussion

We first evaluate the performance of the three models on a synthetic test set, independent of the training and validation ones. In order to concretely compare the three methods, the synthetic and real data sets under consideration are the same for all the models and differ only in their input representation.

Table 1 shows quantitative results in terms of average error and standard deviation for the three models with respect to the synthetic test set.

The position error is assumed as the Euclidean distance and is computed for each sample as:

$$E_{p}^{i} = \sqrt{\left(x_{i} - \hat{x}_{i}\right)^{2} + \left(y_{i} - \hat{y}_{i}\right)^{2}} \tag{7}$$



Figure 4. Comparison of the performance of the tested models at inference time. TS, IMG and TRA models are shown on columns respectively. For each row, latitude, longitude and magnitude predictions are reported, respectively. Each point of the scatter plots represents a test sample, whose magnitude is indicated by the colorbar, and it is illustrated as a function of both its actual and predicted value. Black dashed lines represent the ideal prediction, while solid black lines represent the rolling median.

where  $x_i$  and  $y_i$  represent the actual longitude and latitude and  $\hat{x}_i$  and  $\hat{y}_i$  the predicted longitude and latitude, respectively. We adopt a Mean Absolute Error (MAE) for the magnitude, which is computed for each sample as:

$$E_m^i = |m_i - \hat{m}_i| \tag{8}$$

where  $m_i$  and  $\hat{m}_i$  are the actual and predicted magnitude, respectively. Then, the total position and magnitude errors are computed by averaging  $E_p^i$  and  $E_m^i$ .

The quantitative results evidence that the TRA method outperforms the other two, in terms of average error, both in position (105.44) and in magnitude (0.13), with a lower standard deviation in position (128.84) with respect to the IMG method (146.97) and slightly higher with respect to the TS method (116.94). We may expect TRA to have also a lower standard deviation, yet it depends on many factors which can be related to the type of data used, as well as randomness in the training.

#### **3.1** Analysis of the performance

Figure 4 shows the prediction of the three models on the synthetic test set color-442 coded by the actual magnitude of the test events. Indeed, the performance of all the mod-443 els depends on the magnitude, which is closely related to the Signal-to-Noise ratio (SNR). 444 As we can observe in the third row, low magnitudes tend to be overestimated by all mod-445 els, likely because there is an intrinsic resolution threshold preventing the models from 446 achieving good performance when the SNR is not sufficiently high. For the lower mag-447 nitude events (blue points), also the localization ability is poor, as the predictions of the 448 449 three models do not follow, in general, the ideal prediction line. This behaviour may thus be linked to an intrinsic limitation of data information. 450

The solid black lines in the plots show the rolling median on the scatter plot computed on 150 samples. They give the general trend of the predictions. At first order, they can help in individuating a tentative magnitude threshold value, that is the value under which the magnitude prediction is significantly degraded. We can derive the following resolution limits:  $M_w^{TS} = 7$ ,  $M_w^{IMG} = 6.3$ ,  $M_w^{TRA} = 6.2$ . Those quantities have to be taken as a general indication.

The trend of the magnitude prediction for TS deviates from the ideal prediction 457 line both for small and for large magnitudes, presenting a median saturation around  $M_w$  6.5 458 and  $M_w$  7.8, respectively (cf. black solid lines). The saturation for high magnitudes could 459 be due to the employed network architecture as well as to specific features associated to 460 the type of data. The magnitude prediction for the image-based methods, *i.e.*, IMG and 461 TRA, better adhere to the ideal prediction line, with a progressively smaller error vari-462 ance at larger magnitudes, in line with the SNR improvement. As for the magnitude res-463 olution, TRA is the method associated with smaller error variances and with a better 464 median trend. 465

From the latitude and longitude prediction, *i.e.*, the localization performance, we 466 can observe that the models do not treat similarly the low and high magnitudes. No-467 tably, for magnitudes smaller than the SNR limit, TS assigns them an average position 468 (*i.e.*, near 38 for the latitude and 142 for the longitude). This behaviour is clearly in-469 dicated by the horizontally-clustered blue points. This pattern is indeed coherent with 470 the choice of the quadratic loss function used to train the model. In fact, at first order, 471 the best guess is represented by the mean value of the output range subject to the pos-472 terior distribution (Haykin, 2008; Moon & Stirling, 2000). We can derive that, when the 473 SNR is below a certain resolution threshold, the model associates low-magnitude events 474 to average coordinates, which likely minimize the average error. For higher magnitudes, 475 the TS latitude predictions are more clustered around the ideal prediction line, although 476 a tendency towards the mean values is still present, while TS predicts the longitude of 477 high-magnitude events either in the proximity of the GNSS network (longitudes less than 478  $\sim 142$ ) or in far field (longitudes higher than  $\sim 144$ ). Conversely, image-based meth-479 ods characterize low-magnitude events as having a random position in the region of in-480 terest (cf. scattered blue points), while being able to precisely constrain higher magni-481 tude events, with predictions tightly clustered around the ideal line. Moreover, the me-482 dian prediction lines for IMG and TRA are more stable and significantly bends only in 483 correspondence of far field events (longitudes near  $145^{\circ}$ ), which is physically consistent. 484

485

#### 3.2 Spatial variability of the location error

Figure 5 shows the location error as a function of the ground truth spatial coordinates. The plot has been computed by interpolating the location error for each test data sample onto a grid, corresponding to the area of interest. This smoothed heatmap indicates the amount and the distribution of location errors all over the tested region, for different magnitude ranges. This type of representation can help to assess the physical consistency of the tested models, as well as revealing systematic biases in the error



Figure 5. Comparison of the location error of the tested models, reported in the columns. Each subplot shows the location error associated to the test samples, interpolated on a grid which corresponding spatial coordinates are indicated along the axes. Magenta data points represent the position of GNSS stations in Japan. The heatmap depicts the distribution of the error in position committed by the tested models, for different magnitude ranges, in rows. Arrows show the average direction of position error for patches of  $1 \times 1$  arc degree. The arrows have the same scale throughout all the subplots, making a comparison be possible among different models.

pattern, which can be more evident for specific magnitude ranges. Moreover, it is useful to compare and discuss how the error distribution of certain events can be influenced
by their relative position with respect to the GNSS network.

From the heatmaps of the first two lines, corresponding to the magnitude ranges 495 (5.8, 6.3) and (6.3, 6.8), we can see how the three methods handle the characterization 496 of low-magnitude events (cf. 3). We can remark that TRA is able to better resolve small 497 magnitude events in near field (*i.e.*, in proximity of the GNSS network). By increasing 498 the magnitude range, the error amplitude of IMG and TRA are decreasing, affecting only 499 500 the points which are far from the network (on the east side). For high magnitudes, TS tend to localize most events in far field, seemingly not taking advantage of the strong 501 GNSS signal in the near field. 502

The error pattern for image-based methods is, therefore, more physically consis-503 tent. The most reasonable explanation is that image-based models can better capture 504 the spatial information by extracting spatial features which are essential for the char-505 acterization. As a general comment, we do not see any clear bias and the error patterns 506 exhibit a correct behaviour, since, as the magnitude increases, highest errors are pushed 507 towards the far field. We notice that, for low magnitudes, the maximum error associated 508 to the TS is about 200 km less than the other models, as its bias correctly minimizes the 509 average error, yet without providing any discriminant ability to the model. By increas-510 ing the magnitude, errors become smaller and smaller, with the events contributing to 511 the largest errors being distributed on the east (offshore) side, in favor of TRA, which 512 is associated with the most reasonable error pattern. 513

514

#### 3.3 Influence of the distance from the GNSS network on the predictions

Figure 6 helps us in analyzing the dependency of errors to the relative position with respect to the GNSS network. Each scatter plot represents the error as a function of the distance to the nearest GNSS station. Such a distance is computed from the coordinate of a hypocenter as the 3D Euclidean norm, in order not to take into account the Earth curvature. This kind of representation is effective in revealing patterns of the position and magnitude errors as function of both distance, on the x axis, and magnitude, in color code.

In order to better summarize and understand this behaviour, we identified three regions, according to the relative distance to the nearest station: being d the distance to the nearest station, we will refer to near, intermediate and far field when  $d \leq 0.5$ ,  $0.5 \leq d \leq 3$  and  $d \geq 3$  arc degrees, respectively (see dashed lines in Figure 6). The dashed lines correspond to the median for several magnitude ranges (cf. Figure 5).

For the TS model, we can see in the first row a non negligible presence of errors 527 due to high magnitude events both in near field and intermediate field, while image-based 528 methods being able to correctly locate a larger number of high and even low magnitude 529 events. Looking at the magnitude estimation (second row), we can observe for TS a clus-530 ter of errors corresponding to very high magnitude events in near field in the upper part 531 (average error  $M_w = 0.8$ ), and a second cluster of errors associated to lower-intermediate 532 magnitudes affecting all the region. Conversely, image-based methods are more accu-533 rate in the magnitude estimation, with a less biased error pattern: the median curves 534 of errors increase with the distance, both for the magnitude and the position estimation. 535 Moreover, since the depth has been taken into account when computing the distance to 536 the nearest GNSS station, we also find that the underestimation of large magnitudes com-537 538 mitted by TS (cf. Figure 4) is affecting very shallow and near events, leading to the conclusion that image-based data representation can bring more exploitable information about 539 the deformation field. Therefore, more low-magnitude events are captured. 540



**Figure 6.** Comparison of errors as a function of the distance to the nearest GNSS station. The deep learning models are shown in columns, while the rows indicate position and magnitude errors, respectively. Each scatter plots depicts errors as a function of the Euclidean distance to the nearest GNSS station, expressed in arc degrees. Each data point, representing the position error and the absolute magnitude error between the test samples and the model predictions, is color coded by the actual magnitude of the event. Solid lines represent the median of subsets of the data points, filtered by magnitude ranges as indicated in the legend in the top right. Vertical dashed lines discriminate among near, intermediate and far field, respectively.

	depth $\leq 30~{\rm km}$	$30~{\rm km} < {\rm depth} \le 60~{\rm km}$	$60~{\rm km} < {\rm depth} \le 100~{\rm km}$
near field	6	6.2	6.5
interm. field	6.8	6.8	7
far field	7.5	7.5	7.8

Table 2. Magnitude thresholds of TRA estimated against the synthetic test set.



**Figure 7.** Position error, computed for each test sample, as a function of the magnitude (x axis), the depth range (columns) and the distance range (cf. Figure 6) with respect to the GNSS network (rows) for TRA. The orange solid line represents the result of a median smoothing by employing a kernel size of 15 points.

#### 541

#### 3.4 Magnitude threshold estimation from TRA localization error

<sup>542</sup> By positioning this deep learning pipeline in an operational framework, is it inter-<sup>543</sup>esting to ask whether a characterization coming from a learning model is reliable. In-<sup>544</sup>deed, we are interested in the confidence of the model subject to the physical constraints <sup>545</sup>that come from the tectonic context under consideration, notably the magnitude (SNR), <sup>546</sup>the distance and the depth, as seen in the previous sections.

Figure 7 shows the position error for the TRA method, computed for each test sam-547 ple, as a function of the magnitude, with each subplot corresponding to a different range 548 of hypocenter-station distances and hypocentral depths. The general idea is to get an 549 estimation of the magnitude threshold for different settings, *i.e.*, for different values of 550 depth and distance to the GNSS network. This can serve as a prior probability associ-551 ated to each new event that we are willing to characterize, such that we can assess, a pri-552 ori, if the deep learning model will be able to characterize it with an acceptable preci-553 sion. To keep it simple, in this study we will not estimate any probability but assign a 554 hard threshold (characterizable, non-characterizable). 555

As discussed in previous sections, as the depth increases, the magnitude detection limit also increases. For events having a depth  $d \leq 30$  km, we can set a magnitude thresh-



**Figure 8.** Histograms of the predicted magnitude (orange bars) and ground truth synthetic magnitude (blue bars) as a function of the distance range (cf. Figure 6) with respect to the GNSS network (rows) and for different magnitude ranges (columns), for the TRA model.

old at  $M_w$  6, by selecting a limit where there is an evident discontinuity and where the 558 error is reasonably low with respect to the general trend. As for intermediate and far 559 field, it is harder to find a clear discontinuity, as the interplay between magnitude, dis-560 tance and depth is generally nonlinear, yet a general tendency can be still observed. The 561 estimated thresholds will be  $M_w$  7 for the intermediate and far fields. We should also 562 consider that, to parity of depth range, the relative distance between the event and the 563 GNSS network strongly affects the probability of correct retrieval, making the magni-564 tude threshold larger and larger. This poses some limitations in the characterization of 565 deep and far offshore events, with only large magnitude earthquakes being characteri-566 zable in those conditions. A summary of the chosen magnitude thresholds for TRA can 567 be found in table 2. 568

569 570

#### 3.5 Trade-off between depth and magnitude in the TRA magnitude resolution

Figure 8 shows the histograms of the estimated (using TRA) and the real magni-571 tude as a function of the distance ranges (same ranges than in Figure 6). This plot helps 572 us further understand in which conditions the model predictions are reliable, by study-573 ing their statistics with respect to the magnitude posterior distribution. For sufficiently 574 high magnitudes (third and fourth column), the conditional posterior of the predicted 575 and the ground truth magnitude are in good accordance. This may be an indication that 576 the method has indeed learnt how to correctly characterize the high magnitudes based 577 on input image time series (see also Figure S1 in the Supporting Information). However, 578 the predicted magnitude distributions for lower magnitudes are far from the ground truth 579 values and seem to be drawn from a Gaussian distribution. We performed a statistical 580 normality test by following the approach of (R. B. D'Agostino, 1971; R. D'Agostino & 581 Pearson, 1973), confirming that the predicted magnitudes for TRA in low-SNR condi-582 tions are following a normal distribution. The same observation is found for IMG, while 583 the TS prediction is more difficult to interpret (cf. figures S2 to S7). 584

The deep learning models are not provided with any prior on the depth, therefore they cannot resolve the ambiguities coming from the interaction between magnitude, position and depth. Hence, by adding a prior knowledge on the depth, it may be possible to better resolve the magnitude, improving also the localization performance.

For the first two magnitude ranges (cf. first and second columns of Figure 8), the 589 distributions are centered around  $M_w = 6.3$ . Since their standard deviation is  $\sigma = 0.2$ , 590 the 99.7% of the realizations will fall within  $\pm 3 \sigma$ , *i.e.*, in the range (5.8,7). In fact,  $M_w =$ 591 7 is seemingly a threshold value beyond which the magnitude resolution ability of TRA 592 is exceptionally high (cf. Figure S1). Therefore, the method is resolving intermediate-593 low magnitude by drawing predictions according to a normal probability distribution cen-594 tered in the middle of the uncertainty range. Furthermore, some residuals of the Gaus-595 sian tail are visible in the third column, likely corresponding to deformation fields as-596 sociated to high-depth events, which thus have been wrongly characterized as  $M_w$ 597 6.4 events. 598

#### <sup>599</sup> 4 Application to real GNSS data

#### 4.1 Data processing

The data selection for real events in Japan has been conducted as follows. The F-601 Net catalog from NIED (cf. https://www.fnet.bosai.go.jp) has been exploited and events 602 ranging from 1998 to 2021 have been selected according to the studied range of charac-603 teristics (epicentral position, hypocentral depth, magnitude, see section 2.2.1) for a to-604 tal of 85 events. Magnitudes have been allowed to exceed the 8.5 limit in order to fur-605 ther test the models on high-magnitude events, even though it's out of the training range. 606 Since GNSS data is daily sampled, if more than one event is recorded in the same day, 607 only the maximum magnitude event is kept. All events in 2011 have been removed ex-608 cept the Tohoku event (11 March 2011). Indeed, the earthquake and subsequent tsunami 609 damaged several GPS stations, and the time series of the remaining ones are dominated 610 by a strong post-seismic relaxation effect making GNSS time series difficult to interpo-611 late and interpret on an automated manner. 612

Two GNSS data sets have been collected: the data processed in double difference 613 at ISTerre (Institut des Sciences de la Terre) that range from 1998 to 2019 (Marill et al., 614 2021; gnss products, 2019) and the data processed in PPP at NGL (Nevada Geodetic 615 Laboratory) (Blewitt et al., 2018). that range from 2009 to 2021. We performed outlier 616 detection and removal by processing the data with the *hampel filter* (Pearson et al., 2016) 617 with a window length n = 3. Thereafter, we extracted, for each date in the seismic cat-618 alog, a window of 100 days, centered onto the coseismic offset (cf. section 2.2). We con-619 sidered a 100-day stack of time series as valid if at least 60% of the stations are present 620  $(i.e., \sim 180)$  and if at least the 70% of the median number of data points in the 100-621 day window (i.e., 70) is not undefined (i.e., less than 30% of data gaps). The remain-622 ing data gaps are filled as follows. After centering the time window on the coseismic off-623 set date, we compute the linear trend in the first and the second half. Thanks to this 624 procedure, an approximation is provided for the small data gaps and also a first order 625 reconstruction of the coseismic offset when that information may be missing. Finally, 626 the data is detrended, *i.e.*, the linear trend is subtracted for every 100-day stack. 627

After the previous processing, the ISTerre/DD and the NGL/PPP data sets contain 70 and 52 labelled time series. We used the magnitude thresholds obtained for TRA (cf. table 2) to differentiate the theoretically characterizable events from the rest, as shown in Figure 9, that is if magnitude, depth and position of the events are such that they satisfy those experimentally-derived relationships. We found 8 and 5 characterizable events for ISTerre/DD and NGL/PPP data sets, respectively. The data is further rearranged



Figure 9. Seismic catalog associated to the ISTerre/DD and NGL/PPP data sets, respectively. ISTerre/DD data set contains 114 events ranging from 1998 to 2019, while NGL/PPP set contains 94 events ranging from 2009 to 2021. Focal mechanisms are depicted for each event and their size is proportional to the magnitude, according to the legend at the top left. Blue–colored focal mechanisms indicate the characterizable events according to table 2.

**Table 3.** Quantitative results of the tested methods on the characterizable events belonging tothe real data sets.

	ISTerre/DD		NGL/PPP	
Model	Position error (km)	Magnitude error (°)	Position error (km)	Magnitude error (°)
TS IMG TRA	$\begin{array}{c} 1422.53 \pm 2634.99 \\ \textbf{87.98} \pm \textbf{78.49} \\ 140.08 \pm 150.79 \end{array}$	$11.57 \pm 24.52 \\ 0.26 \pm 0.20 \\ 0.36 \pm 0.31$	$\begin{array}{c} 1536.11 \pm 2553.39 \\ 143.52 \pm 79.44 \\ \textbf{126.45} \pm \textbf{87.61} \end{array}$	$\begin{array}{c} 12.53 \pm 23.99 \\ 0.48 \pm 0.44 \\ \textbf{0.43} \pm \textbf{0.21} \end{array}$

into differential images and image time series and the performance of the three deep learning methods are evaluated.

#### 4.2 Results and discussion

The quantitative results are shown in table 3, while Figure 10 shows the performance of the tested methods on the two real data sets. The displacement fields associated to all the characterizable events in the ISTerre/DD dataset are represented in Figure 11.

The performance of the image-based models is more accurate than the TS model on both data sets, in line with the results obtained on synthetic data (cf. section 3). This is probably linked to the presence of a huge amount of data gaps and missing stations,



Figure 10. Performance plots on real data from ISTerre/DD and NGL/PPP data sets. Each subplot shows the *real vs predicted* comparison for the estimated parameters (Latitude, longitude and magnitude in each row) for each of the three methods (TS, IMG and TRA in each columns). For each scatter plot, circles represent mean predictions associated to events having thrust focal mechanism, with crosses indicating any other focal mechanism. The solid dashed line shows the line of perfect predictions. The data points are color–coded according to the time of occurrence. The predictions of the TS model associated to the events that occurred the 13 June 2008 and the 11 March 2011 (Tohoku) are not visible as the  $ir^2 predictions$  are located outside of the plot limits (*i.e.*, outliers, cf. section 4.2).

7 8 Actual value 7.0

6.5

7 8 Actual value

7.0

6.5

7.25

11 March 2011 (Tohoku)

Predic

6.6

6.4

6 75 7 00 Actual value

6.50



Figure 11. Displacement fields associated to the eight events of the ISTerre/DD data set. The deformation fields have been computed by subtracting the deformation at day  $t_c + 1$  and  $t_c - 1$ . In each subplot, the focal mechanism from the NIED catalog is shown as well as the magnitude and depth (in each title) with the yellow, brown and blue points representing the predictions for TS, IMG and TRA, respectively. The predictions of the TS model are not visible in the plot for the events occurred the 13 June 2008, the 19 July 2008 and the 11 March 2011 (Tohoku), as they have been located outside of the figure bounds (cf. section 4.2).

which worsen the resemblance between synthetic and real data, thus deteriorating the 644 performance of TS. As a result, image-based models can better deal with data gaps thanks 645 to the spatial interpolation. Hence, the amount and continuity of the data plays an es-646 sential role on the final prediction accuracy, which is indeed mitigated by the image and 647 image time series representations. It is also worth to notice that all the models have a 648 larger error associated to the NGL/PPP dataset, probably because of the Precise Point 649 Positioning solution, which is slightly noisier with respect to the DD approach. Since the 650 noise in the training samples is obtained from DD solution time series, it is not surpris-651 ing that the model may have a lower performance better on data obtained from a PPP 652 solution. Therefore, this constitutes a possible limitation of the method, which could be 653 overcome by applying a fine-tuning to improve the results on the PPP data set. For these 654 reasons, we will focus on the ISTerre/DD data set henceforward. 655

The events in Figure 10 have been marked with a different symbol if their rupture 656 has a thrust focal mechanism ( $\phi_S = 200 \pm 40^\circ, \ \delta = 25 \pm 20^\circ, \ \lambda = 90 \pm 45^\circ$ ). Differenti-657 ating thrust and non-thrust events is interesting to assess if the shape of the associated 658 deformation field plays a key role in the characterization performed by image-based mod-659 els, given that the model was trained on thrust events only. Indeed, the results shown 660 in Figure 10 seem to suggest that the shape of the deformation field (e.q., cf. outliers)661 having strike-slip focal mechanism) is not a relevant feature in the characterization of 662 the location and the magnitude, since the predictions of the image-based models do not 663 seem to depend on the nature of the focal mechanism, which indeed would be a key in-664 formation when inverting for the focal mechanism itself. Hence, the amplitude of the de-665 formation and the SNR (linked to the intensity of the interpolated image) are likely the 666 most informative characteristics to retrieve the epicenter and the magnitude of the earthquake, especially in the coastal stations, which often register the highest displacement 668 values. 669

Interestingly, IMG and TRA models seem to be complementary on some events, 670 as shown in figures 11 (d), (g) and (h). The TRA model is unable to separate the source 671 of deformation in the 19 July 2008 event (Figure 11 (d)) because of a persistent outlier 672 in the displacement field, whose influence is better mitigated by the differential approach 673 used for IMG (cf. Figure S9 - S11). On the contrary, TRA can effectively retrieve the 674 21 November 2016 event (11 (g)), likely thanks to the spatio-temporal approach (cf. Fig-675 ure S12 - S17), while IMG is not well performing. This seems to suggest that the two 676 different image-based data representations carry some particular characteristics com-677 ing from the network geometry and the spatio-temporal variability of the data. 678

As a further comment, we notice that the outlier displacement value north of the 679 epicenter of the 19 July 2008 event (cf. Figure 11 (d)) is actually an artifact introduced 680 by the linear interpolation performed on the time series in presence of a large data gap 681 (cf. Figure S8). Therefore, either a more efficient method should be set up for the miss-682 ing data interpolation, or artifacts should be taken into account in the training data base. 683 Accounting for the data gaps is not a trivial task and future developments should focus 684 on this aspect, since, as we saw, the larger the data gaps, the harder is the character-685 ization. 686

Finally, it is worth to mention the performance of the models on the Tohoku event (11 March 2011,  $M_w = 8.7$ , according to the NIED solution used in the paper), which is estimated as a  $M_w \sim 8.6$  and  $M_w \sim 8.5$  event by IMG and TRA, respectively, with the IMG model correctly exceeding the training upper bound ( $M_w$  8.5) on a pattern which has never been presented to the network during the training phase. Yet, it should be noted that this result should be taken carefully as the actual magnitude of the event is 9.1 (Lay, 2018).

#### <sup>694</sup> 5 Conclusions

We studied and developed an end-to-end framework for the seismic source characterization with GNSS data. We constructed three deep learning methods associated with three data representations: time series, differential images and image time series. We train our methods on synthetic data generated to be subduction events compliant with actual events occurring in the Japan subduction zone. We tested the methods both on synthetic and real GNSS data, and we studied the performance and the sensitivity of the three methods, evidencing their strengths and their limits.

Image-based methods outperform time series-based methods, possibly because their 702 associated data representations better exploit the topology of the GNSS network. The 703 wavelength of the deformation is seemingly better constrained with images with respect 704 to time series, the longitudinal extent of the deformation being more difficult to char-705 acterize by means of the temporal evolution only. Results on synthetic data clearly ev-706 idence a detection threshold associated to GNSS data, which is associated to the SNR, 707 and also dependent on the depth and position of events. This allows us to partition the 708 output space by identifying regions in which the source characterization can be performed 709 with confidence. 710

711 Performance on real data sets is consistent with the results obtained on synthetic data and shows accurate and reliable results. Image-based methods outperform the time-712 series based approach in both the real data sets, with image-time series and the TRA 713 model showing that the spatio-temporal approach proposed is crucial in resolving the 714 location and magnitude of most of the real events. However, the noise characterization 715 needs to be improved, in order to better account for outliers in GNSS time series, data 716 gaps and, possibly, common modes. By improving the simulation of the realistic noise, 717 we can produce more and more real-looking synthetic data, possibly having better re-718 sults on the characterization and a lower SNR threshold. Nonetheless, the results on real 719 data are promising and could potentially lead to an effective analysis of the slow defor-720 mation, which would benefit from the present work as well as from the potential refine-721 ments that we have listed before. 722

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# **@AGU**PUBLICATIONS

## Solid Earth

Supporting Information for

### Seismic source characterization from GNSS data using deep learning

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**Figure S2.** Position error, computed for each test sample, as a function of the magnitude (x axis), the depth range (columns) and the distance range (cf. fig. 6) with respect to the GNSS network (rows) for TS. The orange solid line represents the result of a median smoothing by employing a kernel size of 15 points. The red solid line represents the TRA median (cf. fig. 7).



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**Figure S8.** Interpolated time series (N-S component) associated to a 100—day window centered onto the 19 July 2008 for the **GAMIT** data set. Each line represents a different GEONET station. The red line is an artifact caused by a large data gap, producing a false westwards displacement, which hides the eastwards displacement due to the seismic signal.



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#### 21 November 2016, M<sub>w</sub> 7.0, depth: 11.0 km



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