

Bremsstrahlung Electron Crossing Different Mediums: Extension of Asymmetric Backward Peaking Radiation Pattern from a Relativistic Particle Accelerated by Lightning Leader Tip Electric Field

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November 22, 2022

Abstract

Previously the radiation patterns of combined parallel and perpendicular motions from the accelerated relativistic particle at low and high frequencies of the bremsstrahlung process with an external lightning electric field were explained. The primary outcome was that radiation patterns have four relative maxima with two forward peaking and two backward peaking lobes. The asymmetry of the radiation pattern, i.e., the different intensities of forwarding and backward peaking lobes, is caused by the Doppler effect. A novel outcome is that bremsstrahlung has an asymmetry of the four maxima around the velocity vector caused by the curvature of the particle's trajectory as it emits radiation. Previously stated bremsstrahlung asymmetry, R was an asymmetry in the radiation lobe pairs about particles velocity vector. Bremsstrahlung asymmetry used to occur at the same level in both forward radiation lobe pairs and backward radiation lobe pairs. However, in high-density mediums where the emitted wave can lag behind the speed of the particle, symmetry of the magnitude of bremsstrahlung asymmetry, R differs between forward peaking radiation lobe pairs relative to backward peaking radiation lobe pairs. This is another novel asymmetry and it causes bremsstrahlung asymmetry, R to be larger in the forward peaking compared to backward peaking radiation. The outcome is the shrink in radiation length that occurs in the backward peaking lobes. This extended work reports, changes in the radiation pattern as the emitted wave propagates through different mediums. Two novel formulas are derived from Snell's law for a particle entering the medium horizontally and from any other angle between $\pi/2$ and $-\pi/2$ radians. The novel outcome is the change in angle between forward peaking radiation lobe pair and backward peaking radiation lobe pair defined as bremsstrahlung angle, θ_{brem} . When the bremsstrahlung particle crosses different mediums, change in angle between the forward and backward radiation lobe pairs, bremsstrahlung angle, θ_{brem} breaks into its components as each lobe changes angle at different magnitudes from the particle's velocity vector. Therefore, bremsstrahlung angle, θ_{brem} between forward-backward peaking lobes transforms into individual angles $\Omega_{\{1\}}$, $\Omega_{\{2\}}$, $\Omega_{\{3\}}$. $\Omega_{\{4\}}$ all measuring from the particle's velocity vector.

Bremsstrahlung Electron Crossing Different Mediums: Extension of Asymmetric Backward Peaking Radiation Pattern from a Relativistic Particle Accelerated by Lightning Leader Tip Electric Field

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Key Points:

- Bremsstrahlung asymmetry, R causes Bremsstrahlung angle, θ_{brem} between two forward and backward peaking lobe pairs to get either larger or smaller as the radiation crosses different mediums
- More mediums the bremsstrahlung electron crosses, forward-backward peaking radiation pattern tends to either asymmetric, distorted Dipole radiation pattern or asymmetrically oriented forward backward peaking radiation pattern.
- Asymmetrically oriented forward backward peaking radiation pattern means each four lobe would have different angle from the particle's velocity vector. For bremsstrahlung in vacuum, all four lobes had same angle from the particle's velocity vector, which was half of the bremsstrahlung angle, $\frac{\theta_{brem}}{2}$ when particle was in vacuum. Bremsstrahlung particle crossing different mediums have each of its peaking lobe at different angle from the velocity vector where each is $\Omega_1, \Omega_2, \Omega_3, \Omega_4$
- Bremsstrahlung asymmetry, R causes a difference in top and bottom forward peaking lobes which in turn causes different magnitudes of diffraction
- Doppler effect causes different diffraction between forward and backward peaking lobes. Backward peaking lobes diffract more compared to forward peaking lobes
- Novel formula extended Snell's Law to diffraction of bremsstrahlung asymmetric, R waves and predicting final direction of the forward and backward lobe pairs
- As the bremsstrahlung electron follows spiral curved trajectory, in theory, all peaking lobes enter new medium at different angles. Hence, all diffract at different magnitudes. This is another reason that contributes to the increase and decrease in Bremsstrahlung angle, θ_{brem} between forward-backward peaking lobes

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Abstract

Previously the radiation patterns of combined parallel and perpendicular motions from the accelerated relativistic particle at low and high frequencies of the bremsstrahlung process with an external lightning electric field were explained. The primary outcome was that radiation patterns have four relative maxima with two forward peaking and two backward peaking lobes. The asymmetry of the radiation pattern, i.e., the different intensities of forwarding and backward peaking lobes, is caused by the Doppler effect. A novel outcome is that bremsstrahlung has an asymmetry of the four maxima around the velocity vector caused by the curvature of the particle's trajectory as it emits radiation. Previously stated bremsstrahlung asymmetry, R was an asymmetry in the radiation lobe pairs about particles velocity vector. Bremsstrahlung asymmetry used to occur at the same level in both forward radiation lobe pairs and backward radiation lobe pairs. However, in high-density mediums where the emitted wave can lag behind the speed of the particle, symmetry of the magnitude of bremsstrahlung asymmetry, R differs between forward peaking radiation lobe pairs relative to backward peaking radiation lobe pairs. This is another novel asymmetry and it causes bremsstrahlung asymmetry, R to be larger in the forward peaking compared to backward peaking radiation. The outcome is the shrink in radiation length that occurs in the backward peaking lobes. This extended work reports, changes in the radiation pattern as the emitted wave propagates through different mediums. Two novel formulas are derived from Snell's law for a particle entering the medium horizontally and from any other angle between $\Pi/2$ and $-\Pi/2$ radians. The novel outcome is the change in angle between forward peaking radiation lobe pair and backward peaking radiation lobe pair defined as bremsstrahlung angle, θ_{brem} . When the bremsstrahlung particle crosses different mediums, change in angle between the forward and backward radiation lobe pairs, bremsstrahlung angle, θ_{brem} breaks into its components as each lobe changes angle at different magnitudes from the particle's velocity vector. Therefore, bremsstrahlung angle, θ_{brem} between forward-backward peaking lobes transforms into individual angles $\Omega_1, \Omega_2, \Omega_3, \Omega_4$ all measuring from the particle's velocity vector.

1 Introduction

Bremsstrahlung radiation patterns were predicted to be forward and backward peaking with associated novel bremsstrahlung asymmetry, R (Yücemöz & Füllekrug, 2021). In addition, the time evolution of dipole radiation pattern into forward-backward peaking was demonstrated. The reasoning for the existence of forward-backward peaking due to the collapse and separation of the lobes of the dipole radiation pattern was associated with the conservation of symmetry axes (Yücemöz & Füllekrug, 2021). Furthermore, the bremsstrahlung process of high-density mediums revealed a new outcome that symmetry of the bremsstrahlung asymmetry, R about the axis perpendicular to the direction of particle's motion, between the forward and backward peaking side is broken. Increasing refractive index causes bremsstrahlung asymmetry, R to exist at different ratios between forward and backward peaking radiation lobe pairs. Increasing refractive index increases the difference in bremsstrahlung asymmetry between the front radiation lobe pairs compared to backward radiation lobe pairs. Furthermore, increasing the refractive index shortens the radiation length in the backward peaking radiation side. This extended modelling predicts the changes in the emitted radiation pattern as the accelerating particle transits to different mediums. Novel equations 12 and 13 extended from Snell's law combining previous knowledge of the bremsstrahlung asymmetry parameter, R predicts how the symmetric bremsstrahlung angle, θ_{brem} between forward and backward peaking lobe pairs and how symmetric half bremsstrahlung angle, $\frac{\theta_{brem}}{2}$ from the particle's velocity vector within each lobe pair breaks. Novel symmetry break of bremsstrahlung angle, θ_{brem} into individual angles $\Omega_1, \Omega_2, \Omega_3, \Omega_4$, all describing the angle for each radiation lobe from particle's velocity vector. This novel symmetry break causes the ra-

83 diation pattern to be a distorted combination of both either forward or backward peak-
84 ing radiation pattern and the distorted dipole radiation pattern.

85 1.1 Aims & Objectives

86 This theoretical approach aims to extend the previous bremsstrahlung model to
87 different mediums. Previously, the bremsstrahlung electron was propagating inside a vac-
88 uum. This report puts bremsstrahlung electron into crossing different mediums and math-
89 ematically models changes in the radiation pattern as a result of wave refraction by ex-
90 tending Snell's law for the waves that have larger wavelengths than the particles present
91 in that specific medium. Therefore, mediums where wave refraction dominates wave scat-
92 tering.

93 2 Equation Relating Top and Bottom Radiation lobes Together Us- 94 ing Bremsstrahlung Asymmetry, R

95 Considering forward peaking part of the overall radiation pattern. Radiation in-
96 tensity, I of top lobe (I_T) can be related to the bottom lobe (I_B) with bremsstrahlung
97 asymmetry, R mathematically by

$$I_T = \frac{I_B}{(1 - R)} \quad (1)$$

98 This information can be used to find the bremsstrahlung angle, θ_{Brem} between the
99 two forward peaking radiation lobes that are bremsstrahlung asymmetric, with asym-
100 metry value R.

101 By knowing the radiation length difference or in other terms, linear distance in the
102 form of radiation intensity, $\Xi [Js^{-1}]$ between the maximum points of two radiation in-
103 tensities, I in the forward peaking lobes, bremsstrahlung angle between the two forward
104 peaking lobes can be written as,

$$\cos(\theta_{Brem}) = \frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B} \quad (2)$$

105 The equation 1 and 2 can also be used for backward peaking radiation lobes.

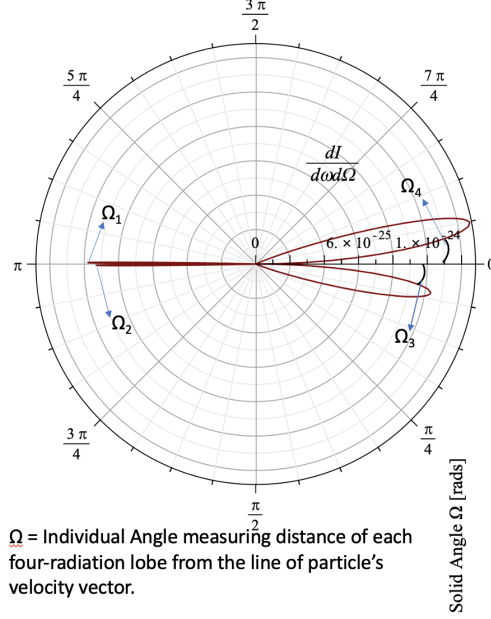


Figure 1. The Radiation patterns are emitted by the anti-clockwise rotating charged particle - bremsstrahlung process. High density (particle travelling inside the water) medium causes a novel asymmetry about a line perpendicular to the direction of motion of a particle. This novel asymmetry causes bremsstrahlung asymmetry, R to occur at different proportions in forward and backward peaking lobe pairs. Therefore, bremsstrahlung asymmetry, R is higher in forward peaking radiation and lower in backward peaking radiation. Moreover, a high-density medium also causes radiation length to shorten in backward peaking radiation. Plot is in Polar co-ordinates. Horizontal axis gives the radiation intensity per Solid angle, Ω , per emitted angular radiation frequency, ω . In addition, angle of the Polar plot is the Solid angle, Ω . The values used for plotting are: mean free time $\tau = 30 \mu\text{s}$, number of charges $z = 1$, $a = 100 \mu\text{m}$, $b = 1 \text{ nm}$ (a and b are related to mean free path), $s_{ft} = 1$, $s_f = 1$, $S_{SpecialR} = 1$, velocity-time scaling factor $s_{ftv} = 1 \times 10^9$ and velocity scaling factor $s_{fv} = 8.19 \times 10^{-11}$. Finally, the bremsstrahlung asymmetry is $R = 1/8$. In addition, $\frac{1}{9} \leq R \leq \frac{1}{3}$, medium conductivity, $\sigma = 0.005$, relative permeability, $\mu_r = 0.99$

3 Application of Snell's Law on Bremsstrahlung Asymmetric Radiation Passing Through a Medium with Positive Refractive Index

3.1 Two different Case Studies

3.1.1 Horizontal Incoming Particle

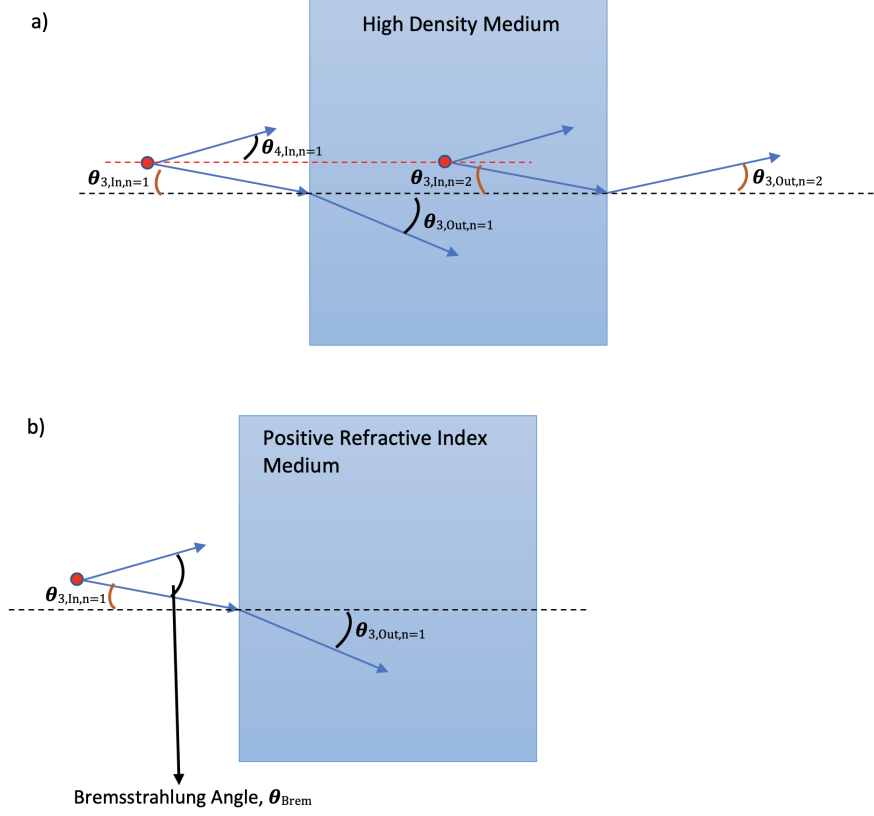


Figure 2. Case study A: the medium is high density such that the bremsstrahlung electron catches its own radiation and continues to emit radiation in the same medium. In this case, the bremsstrahlung electron could gain energy from its own radiation due to a curved trajectory that increases the likelihood of interaction with its own radiation. Case study B: the medium is not high density, the particle still lags its own emitted electromagnetic wave speed. In both cases bremsstrahlung angle, θ_{brem} between two forward peaking lobes increases. This is because of the bremsstrahlung asymmetry, R which changes the radiation intensity, I . Hence, causing a frequency difference between two forward peaking lobes which in turn causes the low-frequency wave to diffract more than the high-frequency lobe. More importantly, as the bremsstrahlung electron crosses more different mediums, the bremsstrahlung angle, θ_{brem} continues to grow such that eventually, the whole radiation pattern can transform back to a dipole radiation pattern. Extremely high frequencies such as X-rays diffract only by a tiny angle. However, forward peaking radiation patterns can start as early as in the MHz frequency range and peak in forward direction as the frequency increases.

When accelerated relativistic bremsstrahlung electron with bremsstrahlung asymmetric "R" forward-backward peaking radiation pattern enters into a medium with a refractive index of η_2 from a medium with refractive index η_1 , the final direction of the emit-

ted radiation in the final medium of η_2 can be predicted using Snell's law in combination with bremsstrahlung asymmetry, R.

$$\sin(\Omega_{4,Out,n=1}) = \frac{\sin(\Omega_{4,In,n=1})\eta_1}{\eta_2} \quad (3)$$

$$\sin(\Omega_{3,Out,n=1}) = \frac{\sin(\Omega_{3,In,n=1})\eta_1}{\eta_2} \quad (4)$$

Introducing the bremsstrahlung asymmetry, R parameter into equations 3 and 4.

From equation 1, angle between two forward peaking lobes can be written as,

$$\Omega_{4,In,n=1} + \Omega_{3,In,n=1} = \theta_{Brem} = \cos^{-1} \left[\frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B} \right] \quad (5)$$

Hence,

$$\sin(\Omega_{4,Out,n=1}) = \frac{\sin \left(\cos^{-1} \left[\frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B} \right] - \Omega_{3,In,n=1} \right) \eta_1}{\eta_2} \quad (6)$$

$$\sin(\Omega_{3,Out,n=1}) = \frac{\sin \left(\cos^{-1} \left[\frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B} \right] - \Omega_{4,In,n=1} \right) \eta_1}{\eta_2} \quad (7)$$

As the emitted radiation could go through many different mediums until it is detected by the detectors, to find the final direction of the emitted radiation through all of its journey from source particle to receivers, equations 6 and 7 can be written in the form of series. For the detection after the " n^{th} " medium refractive index

$$\sin(\Omega_{TopLobeakaNo.4,Out,nth}) = \sum_{n=1}^{\infty} \frac{\sin \left(\cos^{-1} \left[\frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B} \right] - \Omega_{Top,In,1+n} \right) \eta_n}{\eta_{1+n}} \quad (8)$$

$$\sin(\Omega_{BottomLobeakaNo.3,Out,nth}) = \sum_{n=1}^{\infty} \frac{\sin \left(\cos^{-1} \left[\frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B} \right] - \Omega_{Bottom,In,n} \right) \eta_n}{\eta_{1+n}} \quad (9)$$

122

3.1.2 Incoming Particle at an Angle

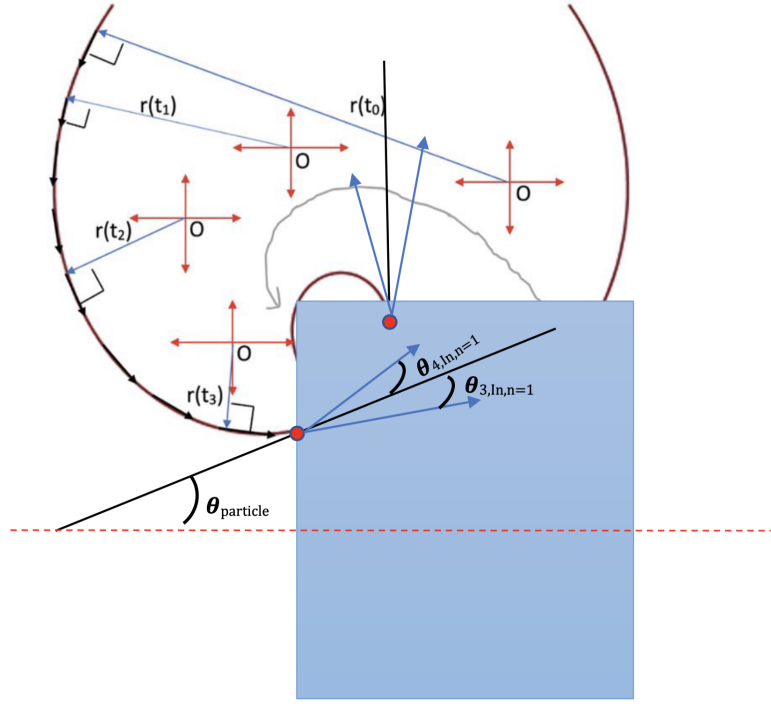


Figure 3. Accelerating bremsstrahlung electron following curved spiral trajectory transiting into a new medium coloured in blue. Orientation of the bremsstrahlung electron from the normal line (Horizontal line) of the new medium is described by the angle $\theta_{particle}$. When radiation is first emitted from the bremsstrahlung particle, each forward peaking lobe in the pair is symmetric about the particle's velocity vector, hence each forward lobe has an equal distance away from the particle's velocity vector. Hence, $\Omega_{3,ln,n=1} = \frac{\theta_{Brem}}{2}$ and $\Omega_{4,ln,n=1} = \frac{\theta_{Brem}}{2}$ from the particle's tangent velocity vector. Therefore, once the $\theta_{particle}$ is known, $\Omega_{3,ln,n=1}$ and $\Omega_{4,ln,n=1}$ are just $\mp \frac{\theta_{Brem}}{2}$ away from the $\theta_{particle}$. This symmetry of the angle of each lobe from the particle's velocity vector is broken once the radiation crosses from one medium into another. This is because of asymmetric refraction that occurs due to asymmetric radiation intensity caused by the bremsstrahlung asymmetry, R , and the curved trajectory.

123

Slope of the line tangent to the particle's spiral trajectory given by

$$\frac{dy}{dx} = \frac{r(t)' \cos(t) + r(t)'' \sin(t)}{-r(t)' \sin(t) + r(t)'' \cos(t)} \quad (10)$$

124

125

Angle, $\theta_{particle}$ of an incident electron into a medium with respect to the horizontal line is,

$$\theta_{particle} = \arctan\left(\frac{r(t)' \sin(t)}{r(t)' \cos(t)}\right) \quad (11)$$

126

$r(t)'$ is particle's velocity vector and is given by, $\frac{dr}{dt} = \frac{b^R(\omega')^R \cos(\theta_{n,r(t)})^R c}{c^R \omega' \cos(\theta_{n,r(t)})}$

$$\left(\frac{\tau^{2R} \left[t^{2R} \frac{2R}{t} + t^{2R} 2R' \ln(t) \right] - t^{2R} \left[\tau^{2R} \frac{2R}{\tau} + \tau^{2R} 2R' \ln(\tau) \right]}{\tau^{4R}} \right) - \frac{a\tau - at\tau'}{\tau^2} \quad (\text{Yücemöz \& Füllekrug, 2021, Eq. 3})$$

The bremsstrahlung particle, hence the emitted radiation enters different mediums at different angles following a spiral trajectory, bremsstrahlung asymmetry, R and Doppler effect cause differences in frequency for each radiation lobe. Therefore, Each radiation lobe would have distinctive incoming $\Omega_{In,n=1}$ and exit $\Omega_{Out,n=1}$ angles. Incoming angles are given formulated in equations 12, 13, 14, and 15. For backward peaking waves, incoming angles to the different mediums are given in equations 14 and 15. In addition, the forward peaking radiation lobe with angle number 4 has the same entry angle as the backward peaking radiation lobe with angle number two. However, the radiation of angle number four and angle number two propagates in the opposite direction to each other. Similar is also true for the forward peaking radiation lobe with angle number three and backward peaking radiation angle number one. All these peaking radiation lobes with associated angle numbers are shown in Figure 4.

$$\Omega_{4,In,n=1} = \theta_{particle} + \frac{\theta_{Brem}}{2} = \arctan\left(\frac{r(t)' \sin(t)}{r(t)' \cos(t)}\right) + \frac{\cos^{-1}\left[\frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B}\right]}{2} \quad (12)$$

$$\Omega_{3,In,n=1} = \theta_{particle} - \frac{\theta_{Brem}}{2} = \arctan\left(\frac{r(t)' \sin(t)}{r(t)' \cos(t)}\right) - \frac{\cos^{-1}\left[\frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B}\right]}{2} \quad (13)$$

Bringing back the backward peaking radiation pattern,

$$\Omega_{2,In,n=1} = \Omega_{4,In,n=1} \quad (14)$$

and

$$\Omega_{1,In,n=1} = \Omega_{3,In,n=1} \quad (15)$$

At the point of the boundary of the medium transition region, the bremsstrahlung angle, θ_{brem} changes exactly the opposite but at the same rate and magnitude between pairs. Hence, if the bremsstrahlung angle, θ_{brem} increases between the forward peaking pairs it would decrease between backward pairs as the radiation crosses between the same mediums but in the opposite order (i.e one from Medium A to B other from medium B to A).

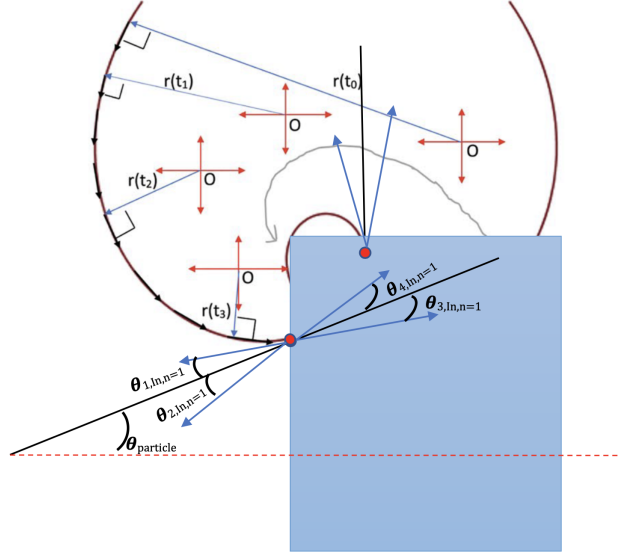


Figure 4. In addition to the Information presented in figure 3, This figure brings back the backward peaking radiation pattern into the equation. As can be seen, backward peaking radiation lobes propagate in the exact opposite directions compared to the forward peaking radiation lobes when radiation is first emitted from the particle. Backward peaking lobes have lower radiation intensity, hence, lower frequency and higher wavelengths due to the Doppler effect. This causes backward peaking radiation lobes to refract less than forward peaking radiation lobes. Which introduces another asymmetry about the line perpendicular to the particle's velocity vector and distortion in the overall radiation pattern.

By substituting equations 12, 13, 14 and 15 into equations 3 and 4, final exit angle can now be written for the forward peaking as follows

$$\sin(\Omega_{4,Out,n=1}) = \frac{\sin\left(\arctan\left(\frac{r(t)'}{r(t)'\cos(t)}\right) + \frac{\cos^{-1}\left[\frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B}\right]}{2}\right)\eta_1}{\eta_2} \quad (16)$$

$$\sin(\Omega_{3,Out,n=1}) = \frac{\sin\left(\arctan\left(\frac{r(t)'}{r(t)'\cos(t)}\right) - \frac{\cos^{-1}\left[\frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B}\right]}{2}\right)\eta_1}{\eta_2} \quad (17)$$

and exit angle for the backward peaking radiation lobes as follows

$$\sin(\Omega_{2,Out,n=1}) = \frac{\sin\left(\arctan\left(\frac{r(t)'}{r(t)'\cos(t)}\right) + \frac{\cos^{-1}\left[\frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B}\right]}{2}\right)\eta_2}{\eta_1} \quad (18)$$

$$\sin(\Omega_{1,Out,n=1}) = \frac{\sin\left(\arctan\left(\frac{r(t)'}{r(t)'\cos(t)}\right) - \frac{\cos^{-1}\left[\frac{I_T^2 + I_B^2 - \Xi^2}{2I_T I_B}\right]}{2}\right)\eta_2}{\eta_1} \quad (19)$$

151 The Refractive index for the refraction of backward peaking radiation is inverted
 152 as the radiation propagates in opposite direction and out of every new medium the bremsstrahlung
 153 electron gets into. $n = 1$ means bremsstrahlung electron only crossed into one differ-
 154 ent medium. $n = nth$ meaning particle crossed nth different mediums.

155 To predict radiation refraction right from the particle, we only need equations 12
 156 and 13 with only one medium property. However, to track radiation pattern crossing from
 157 different mediums over time since its first emission from the particle at medium $n = 1$,
 158 the equation needs to be used with multiple $n = nth$ mediums.

159 Moreover, for wave tracking from medium $n = 1$, until the $n = nth$ medium, an
 160 important step is to calculate the wave refraction pattern at medium $n = 1$ consider-
 161 ing all asymmetries, Doppler effect, the Bremsstrahlung asymmetry in the travelling wave
 162 pattern. Once this information is known, this would be the starting wave incident di-
 163 rection for the second medium. After the calculations at the $n = 1$ medium with equa-
 164 tions 16, 17, 18, and 19, until the $n = nth$ medium only Snell's law needs to be used
 165 which is given in equations 3 and 4 for the forward peaking radiation pattern. This is
 166 also true for refracted backward peaking radiation patterns where the only difference in
 167 the equation is the inverted refractive index as the radiation propagates in the opposite
 168 direction with respect to the forward peaking lobes.

4 Expected Radiation Patterns as a result of Wave Medium Crossing

When bremsstrahlung electron emits radiation patterns while transiting from one medium to another medium, the radiation pattern of the forward and backward side of the particle refracts in exactly the opposite directions. Therefore, if bremsstrahlung angle, θ_{brem} between forward peaking radiation lobes increases where bremsstrahlung angle, θ_{brem} separates into its components Ω_3 , Ω_4 and $\Omega_3 \neq \Omega_4$. In addition, $\Omega_4 > \Omega_3$ as radiation lobe that is described by the angle Ω_4 in bremsstrahlung asymmetric, R hence it has a lower frequency, therefore higher wavelength that makes it refract more compared to its other forward peaking pair described by the angle Ω_3 . The bremsstrahlung angle, θ_{brem} between backward peaking lobes should decrease as opposed to the increase in bremsstrahlung angle, θ_{brem} between forward peaking radiation lobes. This is because backward peaking radiation lobes are propagating in the opposite direction outside of the medium whereas, forward peaking lobes propagate into the new medium.

Overall, as sketched in figure 5, when forward peaking lobes tend towards forming a distorted dipole pattern because of $\Omega_3 \neq \Omega_4$. Backward peaking lobes tends to form distorted further backward peaking as a result of the decrease in bremsstrahlung angle, θ_{brem} where components that make bremsstrahlung angle, θ_{brem} are again not equal to each other. Hence, $\Omega_1 \neq \Omega_2$. Similarly, $\Omega_1 > \Omega_2$ because it has a larger wavelength and hence refracts at a larger angle.

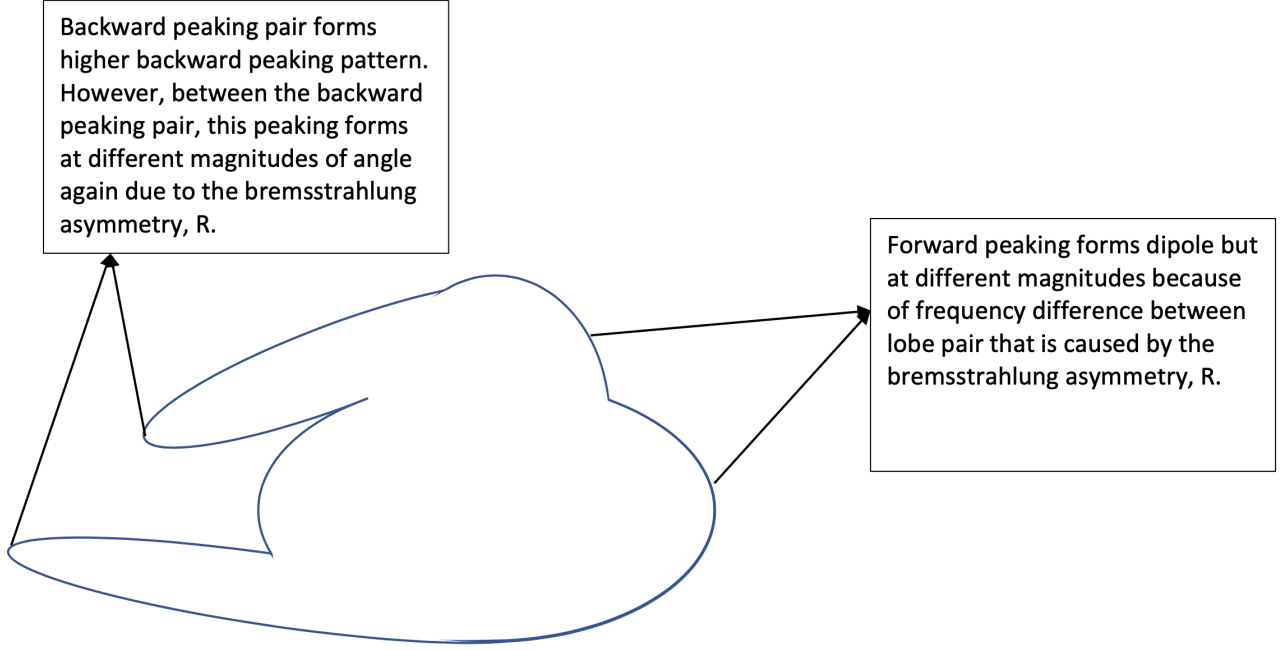


Figure 5. Sketch to show the expected overall refracted radiation pattern of the bremsstrahlung electron after transiting through one different medium. This sketch presents new generic information about the refracted radiation pattern that is expected to apply at all times for different mediums. This is not the wave track but rather important initial information required for the wave tracking. As can be seen, it is expected that forward and backward peaking radiation lobes should behave in opposite directions. So, when one leads toward dipole radiation other pair should lead more towards the direction of the particle's velocity vector (peaking). In this sketch, forward peaking tends towards a distorted dipole radiation pattern and backward lobes tend towards more peaking in the direction of the particle's velocity vector. Opposite tendencies are the result of pair lobes moving in the opposite direction where one enters into the new medium while the other lobe pair exits the new medium. Moreover, the distorted dipole tendency and backward distorted peaking tendencies are result novel symmetry break of bremsstrahlung angle, θ_{brem} into individual different angles $\Omega_1, \Omega_2, \Omega_3, \Omega_4$ where, $\Omega_3 + \Omega_4 = \theta_{brem-BetweenForwardPair} \neq \Omega_1 + \Omega_2 = \theta_{brem-BetweenBackwardPair}$. The cause of this new novel symmetry break of bremsstrahlung angle, θ_{brem} is the bremsstrahlung asymmetry parameter, R , curved trajectory, and the Doppler effect.

5 Summary

When the bremsstrahlung particle was in a vacuum, peaking lobes were all equally away from the particle's velocity vector by half of the bremsstrahlung angle, $\frac{\theta_{brem}}{2}$. When bremsstrahlung particle transits between multiple different mediums, peaking lobes are all at different angles away from the particle's velocity vector where each is defined as, $\Omega_1, \Omega_2, \Omega_3, \Omega_4$ from the particle's velocity vector. This is a novel bremsstrahlung angle, θ_{brem} symmetry break. They are defined in equations 12 and 13 that apply to all cases.

Most importantly, the whole radiation emission process should be divided into frames where each frame would represent the emission of radiation for each tangential point of the particle's trajectory. Therefore, even though the emitted radiation would be distorted and would have different angles from the particle's velocity vector, all are described by the exit angles given in equations 16, 17, 18, and 19. Because of the frame perspective, the whole process restarts again with new input angles given in equations rather than using the output angles of the distorted radiation pattern from the previous frame as the new input angles for the next frame.

Finally, from the example in this report, forward peaking tends toward a distorted dipole radiation pattern and backward lobes tend more peaking in the direction of the particle's velocity vector. The opposite tendencies should always be the case, however, it is medium and particle trajectory that defines whether either forward or backward peaking lobes would tend towards refracted dipole radiation pattern or peak more in the direction of the particle's velocity vector.

Acknowledgments

EPSRC and MetOffice sponsor my PhD project under contract numbers EG-EE1239 and EG-EE1077. I would like to thank a lot to my family for their support and good wishes. The Maple worksheets used to simulate the particle trajectory, external lightning leader tip electric field, particle velocity, bremsstrahlung radiation pattern in high density medium and the radiation patterns are openly available from the University of Bath Research Data Archive.

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RESEARCH ARTICLE

10.1029/2020JD033204

Key Points:

- A novel formula was developed to calculate bremsstrahlung radiation patterns of a charged non/relativistic particle with the Doppler effect
- The bremsstrahlung radiation pattern of a charged relativistic particle exhibits forward and backward peaking due to symmetry conservation
- The two forward and backward peaking lobes are asymmetric with respect to the velocity vector due to the curved particle trajectory

Supporting Information:

Supporting Information may be found in the online version of this article.

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Citation:

Yücemöz, M., & Füllekrug, M. (2021). Asymmetric backward peaking radiation pattern from a relativistic particle accelerated by lightning leader tip electric field. *Journal of Geophysical Research: Atmospheres*, 126, e2020JD033204. <https://doi.org/10.1029/2020JD033204>

Received 1 JUL 2020

Accepted 8 JUN 2021

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Asymmetric Backward Peaking Radiation Pattern From a Relativistic Particle Accelerated by Lightning Leader Tip Electric Field

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Abstract Terrestrial Gamma ray Flashes exhibit slopes of ionizing radiation associated with bremsstrahlung. Bremsstrahlung has a continuous spectrum of radiation from radio waves to ionizing radiation. The Poynting vector of the emitted radiation, that is, the radiation pattern around a single particle under the external lightning electric field during interaction with other particles or atoms, is not quite well known. The overall radiation pattern arises from the combination of radiation of parallel and perpendicular motions of a particle caused by the acceleration from the lightning electric field and the bremsstrahlung. The calculations and displays of radiation patterns are generally limited to a low-frequency approximation for radio waves and separate parallel and perpendicular motions. Here, we report the radiation patterns of combined parallel and perpendicular motions from accelerated relativistic particles at low and high frequencies of the bremsstrahlung process with an external lightning electric field. The primary outcome is that radiation patterns have four relative maxima with two forward peaking and two backward peaking lobes. The asymmetry of the radiation pattern, that is, the different intensities of forward and backward peaking lobes, are caused by the Doppler effect. A novel outcome is that bremsstrahlung has an asymmetry of the four maxima around the velocity vector caused by the curvature of the particle's trajectory as it emits radiation. This mathematical modeling helps to better understand the physical processes of a single particle's radiation pattern, which might assist the interpretation of observations with networks of radio receivers and arrays of γ -ray detectors.

1. Introduction

It was recently suggested that high-frequency radiation emissions observed in the atmosphere could originate from muons interacting with electric fields inside thunderclouds. This novel idea is based on a reduction of the muon detection during thunderstorm occurrences by the ground based telescope GRAPES-3 located in Ooty, India (Hariharan et al., 2019). Gamma-Ray Bursts (GRBs) are commonly thought to result from the interaction of neutron stars in outer space or comet collisions. GRBs emit photons in the energy range from keV to MeV that last ~ 10 s. However, a ~ 90 min long GRB was detected with photon energies ~ 18 GeV (Hurley et al., 1994). When Terrestrial Gamma-Ray Flashes (TGFs) were first observed by detectors of the Compton Gamma Ray Observatory (Fishman et al., 1994), their association with bremsstrahlung was demonstrated by the observation of the characteristic slopes of ionizing radiation (Dwyer et al., 2012), supported by Monte Carlo simulations that included the bremsstrahlung process (Dwyer, 2007). Another example of bremsstrahlung associated with lightning discharges is the detection of ultra-low frequency and very low frequency radio emissions of the same electrons that are also responsible for emitting terrestrial gamma ray flashes (Connaughton et al., 2013). TGFs are associated with low-frequency radio emissions, and these observations were used to identify their source altitude (Cummer et al., 2014; Pu et al., 2019). The source altitude was located to lie between two charged cloud layers in a thunderstorm. All the above discoveries offer experimental evidence for the continuous radiation spectrum of bremsstrahlung to occur. Relativistic runaway electrons are the source of high-frequency X- and γ -ray emissions observed in the upper troposphere at altitudes from ~ 12 to 14 km height (Celestin, 2016). High energy relativistic electrons have a larger mean free path such that they can attain larger velocities until they collide with an atom or molecule in the atmosphere. As these electrons are capable of reaching large velocities, they can emit ionizing radiation through the bremsstrahlung process. Low energy electrons are much more likely to collide with atmospheric atoms or molecules, leading to an increase in the number of free electrons in

the atmosphere (Celestin, 2016). Another working hypothesis is that bremsstrahlung radiation is emitted by thermal runaway electrons accelerated by intra-cloud lightning leader tips (Xu et al., 2015). Bremsstrahlung has a continuous electromagnetic spectrum. Low-frequency radio and optical emissions could also be due to fluorescence, where high-frequency TGFs are absorbed by air molecules (Xu et al., 2015). Numerical Monte Carlo simulations demonstrated the significance of the bremsstrahlung process as the primary process behind high-frequency emissions (Dwyer et al., 2012). Bremsstrahlung electrons emit radiation in forward peaking radiation patterns with an angle that scales with the inverse of the Lorentz factor of the relativistic electrons (Koch & Motz, 1959).

Asymmetric signal of γ -ray bursts measured by the GRB Monitor on the Fermi Gamma ray Space Telescope reveal the lightning leader charge structure. Asymmetric γ -ray pulses indicate the lightning leader charge flux, which exhibits a fast rise and slow decay of the leader tip electric field (Foley et al., 2014). The asymmetries in γ -ray pulses are thought to be caused by Compton scattering (Xu et al., 2019). The rise to decay time ratio of single γ -ray pulses was measured to be approximately 0.67 (Nemiroff et al., 1994). Data from the Burst and Transient Source Experiment reveal two different types of spectra of γ -ray bursts known as bright and dim GRBs. It was found that dim GRBs have less photon energy than bright GRBs (Norris et al., 1994). It was observed that as time passes, overall γ -ray photons transit from bright to dim photons as a photon bunch due to a time delay of approximately 100 μ s between the peaks arising from hard and soft photons (Grefenstette et al., 2008).

Experimental measurements of ionizing radiation and optical emissions by the Atmosphere Space Interactions Monitor on the International Space Station recently reported the detection of 217 TGFs from June 2, 2018, to April 1, 2019 (Østgaard et al., 2019), some associated with radio emissions from charged particles that are observed on the ground. All these measurements reveal the properties of γ -ray bursts. After the combination of the measurements from ground-based radio receivers and spacecraft, it was found that TGFs are produced at the very beginning of the lightning discharge process. It is well known that the observed γ -rays originate from the bremsstrahlung process (Xu et al., 2015). There are approximately 10^{17} – 10^{19} GRBs emitted during the bremsstrahlung process. It is well known that the initially emitted ionizing radiation is not the same in terms of energy and direction compared to the radiation measured by sensors. This difference is because the emitted radiation loses energy by back-scattering and interacting with other air molecules. The interaction causes an ionization and releases more electrons, which can explain why 10^{17} – 10^{19} γ -rays are emitted (Dwyer, 2008). Another theory explains γ -ray bursts to originate from the large electric fields of leader tip streamers producing $\sim 10^{12}$ electrons which then increase the number of electrons within the relativistic runaway electron avalanche (RREA) process that emits γ -ray photons (Babich et al., 2014, 2015; Celestin & Pasko, 2011; Chanrion & Neubert, 2010; Moss et al., 2006; Skeltved et al., 2017).

This contribution reports the modeling of an asymmetric forward peaking radiation pattern and an asymmetric backward peaking radiation pattern of a single particle bremsstrahlung process. The asymmetry occurs around the horizontal axis parallel to the direction of motion of the charged particle, and it is unique to the bremsstrahlung process as the particle continuously follows a curved trajectory of an anticlockwise rotation. Radiation patterns are calculated for both relativistic and non-relativistic velocities. The main asymmetry with four radiation peaks is unique to the bremsstrahlung process and occurs when the particle radiation transits from a dipole toward forward and backward peaking radiation patterns.

The particle trajectory is mostly determined by the Coulomb field and modified by the presence of the external lightning leader tip electric field. The change of the particle trajectory caused by the external leader tip electric field is accounted for by only adjusting the radius of the curvature of the particle's trajectory over time. This is possible because, when the leader tip electric field is resolved into its vector components, one vector component is tangent to the trajectory, whereas, the other vector component is perpendicular to it. The perpendicular component points either in the same or opposite direction to the centripetal acceleration caused by the target particle depending on the polarity of the leader tip electric field. The tangent component of the leader tip electric field contributes to determining the amount of angular rotation that can be covered by the particle during Coulomb interaction. The perpendicular component of the external leader tip electric field contributes to the radius of curvature of the curved trajectory caused by Coulomb interaction.

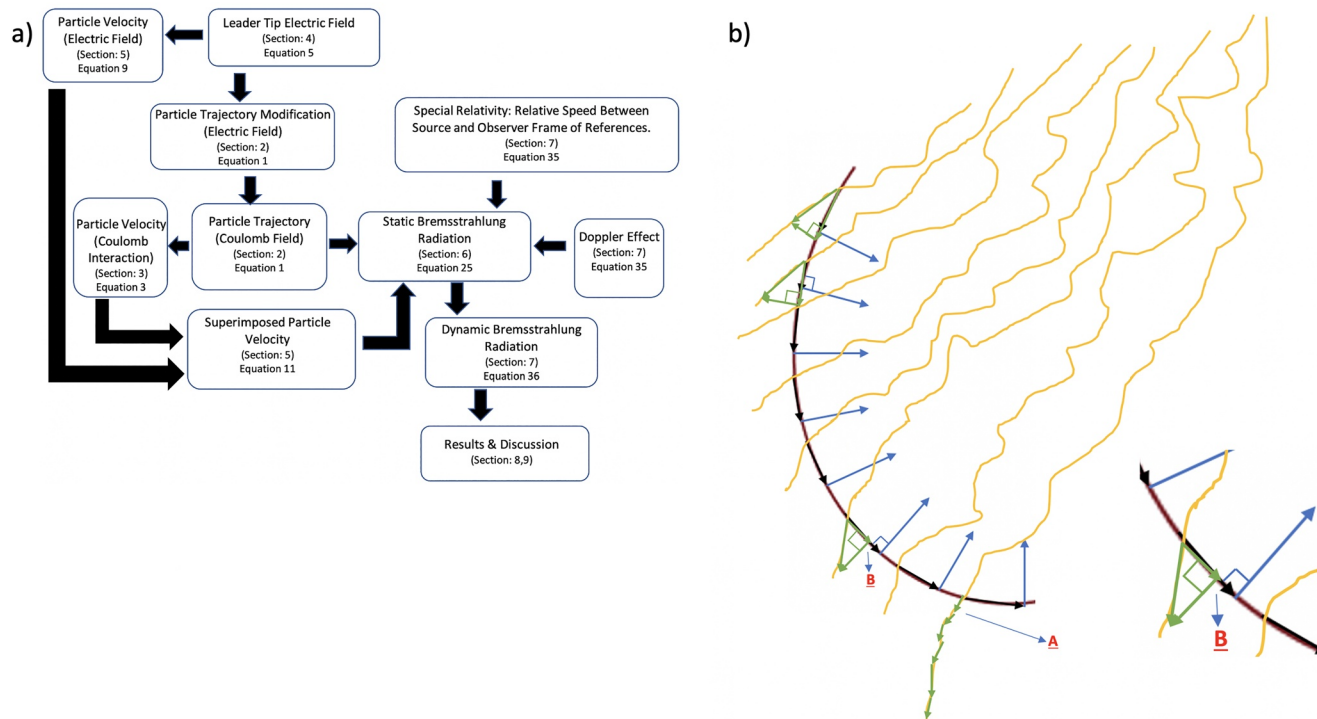


Figure 1. (a) The flow chart shows the general structure of the theory as it is developed throughout the text. (b) Explains how particle's curved trajectory caused by Coulomb field changes as a result of the external non-uniform leader tip electric field and how it is accounted for in the theoretical model. The yellow lines are the external electric field, black arrows are the tangential Coulomb velocity, thick blue arrows are the centripetal acceleration. Finally, green lines are the acceleration caused by the external electric field. As can be seen in region B, one component of the leader tip electric field acceleration is in the direction of tangential Coulomb velocity. Whereas, the other component is in the opposite direction of the centripetal acceleration. Leader tip electric field acceleration in the opposite direction of the centripetal acceleration of the Coulomb field only changes the radius of curvature of the trajectory. Hence, it can be accounted with parameter “b” in the trajectory Equation 1, defining the radius of curvature. The acceleration of the leader tip electric field tangent to the trajectory is accounted with the leader tip electric field velocity Equation 8. Finally, region A explains how the bremsstrahlung process ends, when the leader tip electric field acceleration component opposite to centripetal acceleration (contributing to the centrifugal acceleration) is larger than the centripetal acceleration of the target particle. The condition at region A allows an incoming particle to escape by overcoming the Coulomb force.

The particle trajectory defined by the Coulomb field and the particle velocity arising from the leader tip electric field are two independently derived equations that serve as ingredients that contribute to the bremsstrahlung radiation pattern (Figure 1a). This independence is important because, under a repulsive or attractive Coulomb force, the target particle mostly determines the trajectory of the incoming particle. Also, parameters defining particle trajectory by the target particles Coulomb field can account for the trajectory changes caused by the external leader tip electric field. Hence, when the radius of curvature modification caused by the external leader tip electric field is excluded from the particle trajectory, the temporal derivative of the particle trajectory is the velocity that an incoming particle experiences from the Coulomb force. The overall particle velocity is the superposition of the velocity arising from both the leader tip and the target particle's electric field. The particle velocity is considered to be dominated by the external lightning leader tip electric field.

1.1. Aims and Objectives

The primary aim of this contribution is to construct a generic mathematical model of electromagnetic radiation emitted by a single relativistic charged particle via the bremsstrahlung process that applies to high and low-frequency radiation. Moreover, a generic mathematical model that is also capable of explaining a particle motion whose velocity and the acceleration vectors have both parallel (particle following linear path) and perpendicular (particle following circular trajectory) orientations with each other, at the same time. The secondary aim is to use the model to investigate the details of the actual particle motion and ra-

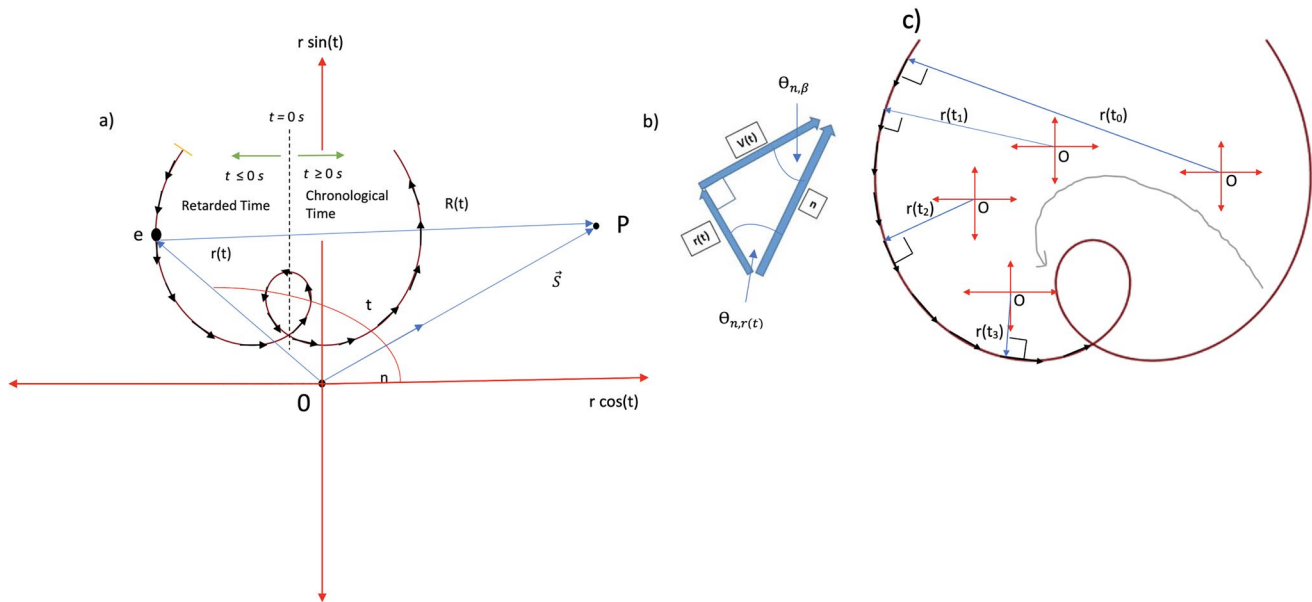


Figure 2. (a) Trajectory of the bremsstrahlung electrons given by Equation 1 in a polar coordinate system and radiation emissions by the change in velocity over time by a Coulomb force of other charges. O represents the target particle that defines the electron's spiral trajectory due to the Coulomb force (Equation 1). The tangential component of the electron acceleration and corresponding velocity vectors are displayed with black arrows and are tangential to the spiral trajectory (red line) and perpendicular to the position vector $r(t)$. $R(t)$ is the distance between the accelerated electron and the observer, which is a function of both retarded and chronological time (t_r , t). P is the position of an observer. \vec{S} is the Poynting vector, or radiant energy flux, which determines the direction of the energy flow per area of an emitted electromagnetic wave. The dimensionless unit vector n points in the direction of the Poynting vector \vec{S} . (b) The position vector $r(t)$, velocity vector $v(t)$ and the unit radiation vector n form a rectangular triangle. (c) Co-ordinate system and all the associated vectors are drawn from the target particle which is positioned at the origin, O . As the target particle is not a rigid body and displaces under the incoming particle's Coulomb field, the moving co-ordinate system is constructed (Brock, 2019). Displacement of the moving co-ordinate system is described with the collision cross-section and form factors (Supplemental information, Text S2). Position vector, $r(t)$ connects both target and incoming particle together. Time derivative of this position vector is the tangential velocity of the incoming particle and it is perpendicular to the position vector in the moving co-ordinate system. Without radius of trajectory changes caused by the external leader tip electric field, time derivative of the position vector is the tangential component of the Coulomb velocity. With radius of trajectory changes caused by the external leader tip electric field, time derivative of the position vector is the tangential component of the overall velocity.

radiation mechanisms during the bremsstrahlung process. Initial assumptions for the particle position vector and lightning leader tip electric field were made to achieve the stated aims. In addition, all the necessary physical details of the interaction, such as form factors and collision cross-sections, are included and investigated. They both play a crucial role in explaining high-frequency radiation. Moreover, the Doppler effect was added to transform the model from a static to a dynamic model. Transformation enables an explanation of the asymmetry of the radiation in forward and backward direction relative to the axis perpendicular to the particle's direction of motion. Finally, the second-order differential equation (Equation 13) was solved as an initial step to establish a generic model to explain radiation patterns using algebra, calculus, geometry, and the table of integral transforms.

2. Particle Position Vector

Starting with defining a curved path for a bremsstrahlung electron.

The position vector is formulated for a particle trajectory that is an anti-clockwise rotating spiral as a function of the retarded time characteristic for bremsstrahlung radiation.

The position vector in Equation 1 defines a spiral trajectory for an incoming particle, that is, an electron, induced by the Coulomb force of the target particle that causes the emission of bremsstrahlung radiation (Figure 2a). The spiral trajectory in Figure 2a and mathematically defined in Equation 1 is realistic even though the mean free path is quite short, e.g., nm- μ m in the atmosphere with a high recombination rate. For example, a circle with a radius of 1 m could also have a radius of 2 μ m, depending on the medium and

the recombination rate. It is still a circle but a scaled microscopic version of the initial macroscopic circle. Preserving geometry at different scales is also true in the formulated spiral trajectory (Equation 1). The decision on a specific particle trajectory considers the ratio of particle size to a curved trajectory radius. If the particle size is larger than the curvature radius, the particle trajectory is approximately a straight line. Therefore, a spiral particle trajectory is realistic because an electron has a size of $<2.8 \times 10^{-19}$ m as measured by the Hadron-Electron Ring Accelerator in Hamburg, Germany at the Deutsches Elektronen Synchrotron facility (Bourilkov, 2000).

Overall, acceleration caused by the Coulomb field can be resolved into its orthogonal components as centripetal and tangent Coulomb accelerations. The centripetal acceleration caused by the target particle defines the curved trajectory of the incoming particle and it is this centripetal acceleration that forces the incoming particle to stay on its curved trajectory. The tangent Coulomb velocity and acceleration components caused by the target particle due to the Coulomb field are perpendicular to the position and centripetal acceleration vectors. Hence, tangent to the particle's curved trajectory.

Similar to Coulomb field, leader tip electric field can also be resolved into its orthogonal vector components. As shown in Figure 1b, tangent to the particle trajectory component of the leader tip electric field increases the incoming particle's tangential Coulomb velocity, hence acceleration. The remaining acceleration component of the external lightning electric field acts in the opposite direction to centripetal acceleration and it only changes the radius of curvature of the incoming particle's trajectory which can be accounted with mathematical variable "b" in Equation 1 representing particles radius of curvature.

$$r(t) = \frac{(t^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R c}{\tau^{2R} c^R \omega' \cos(\theta_{n,r(t)})} - \frac{at}{\tau}, \quad (1)$$

where $r(t)$ is the position vector as a function of time t in s, R is the dimensionless bremsstrahlung asymmetry index, τ is the mean free time in s. Also, ω' is the angular frequency of the emitted electromagnetic wave in the frame of reference of the particle in rads/s, c is the speed of light, $\theta_{n,r(t)}$ is the angle between the emitted radiation unit vector n and the particle's position vector. The time range of the position vector is $-\infty < t < +\infty$. In addition, the factor b in m describes the interaction distance between the incoming particle and the target particle, which is the radius of the time-dependent position vector. The radius of the position vector is directly proportional to the parameter b . This radius of curvature is related to the mean free path because the curvature increases with time, contributing to the overall arc length, i.e., the mean free path of the accelerated particle. In other words, $b \propto \lambda_{v/c}$, where $\lambda_{v/c}$ is the mean free path of a particle at the velocity v which is given as a percentage of the speed of light v/c . The relation between a and $\lambda_{v/c}$ will be derived in Section 3. The factor a in m is an arbitrary adjustment parameter. It is introduced to correct the radius of the curvature of a particle during the bremsstrahlung process. This correction is required because the trajectory of a relativistic particle shrinks in size over time, and a propagation close to the speed of light introduces significant changes in the mean free path. Finally, Equation 1 is a function of time t . Hence, the particle will only cover some segment, or arc length, of the spiral, or the complete arc length of a particle's spiral trajectory when $t = \tau$.

3. Interaction Gap

Defining the mean free path, $\lambda_{v/c}$ and the particle velocity arising only from the Coulomb interaction, $\frac{dr}{dt}$.

The arc length, or the mean free path, of a curve in polar coordinates is given by

$$\lambda_{v/c} = \int_0^{\theta_f} \sqrt{r^2 + \left(\frac{dr}{d\theta}\right)^2} d\theta, \quad (2)$$

where $\lambda_{v/c}$ is the mean free path, r is the position vector given in Equation 1, and the derivative with respect to the polar angle is $\frac{dr}{d\theta}$. Equation 2 is a geometric equation used to calculate the arc length of any curve

in polar coordinates. As the particle covers its mean free path, it follows the curved spiral trajectory in small segments with time t . In this case, Equation 2 has to be formulated with the parameter t rather than θ . To satisfy the dimensional accuracy, the Jacobian is used to carry out a variable transform from θ to t to preserve the dimension m^2 in the square root of the integrand. In geometry, the unit of angle in rad is considered to be dimensionless. An example of this is the arc length formula of a circle ($s = r\theta$). Hence, $\theta = \frac{t}{\tau}$. The mean free time τ is a function of time due to its dependence on particle velocity that changes over time.

Therefore, the quotient rule $\left(\frac{d\theta}{dt} = \frac{\Lambda \frac{du}{dt} - u \frac{d\Lambda}{dt}}{\Lambda^2} \right)$ has to be used to find the derivative $\frac{d\theta}{dt}$. Input parameters

to the quotient rule are, $u = t$, $\Lambda = \tau$, $\frac{du}{dt} = 1$ and $\frac{d\Lambda}{dt} = \tau'$ where τ' is the first derivative of the mean free time

with respect to time t . The mean free time τ of the two identical particles is defined as $\tau = \frac{1}{\pi n v d^2}$, where n is the particle or molecule number density per unit volume in m^{-3} and d is the diameter in m of both an incoming and a target particle or molecule. The basic mean free time is defined between two identical particles with same diameter d where effective collision area is defined as a function of the two identical particle's diameter as $A_c = \pi d^2$. Finally, $\frac{d\theta}{dt} = \frac{\tau - t\tau'}{\tau^2}$ and the limits of the integration are $0 \leq \frac{t}{\tau} \leq \theta_f = 1$

rad. The maximum limit of the integral is 1, which means that the ratio of t to τ as time progresses should not exceed allowed mean free time τ for the particle to radiate.

Coulomb velocity of an incoming particle arising from electrostatic interaction between a target particle can be written as

$$\frac{dr}{dt} = \frac{b^R (\omega')^R \cos(\theta_{n,r(t)})^R c}{c^R \omega' \cos(\theta_{n,r(t)})} \left(\frac{\tau^{2R} \left[t^{2R} \frac{2R}{t} + t^{2R} 2R' \ln(t) \right] - t^{2R} \left[\tau^{2R} \frac{2R}{\tau} + \tau^{2R} 2R' \ln(\tau) \right]}{\tau^{4R}} \right) - \frac{a\tau - t\tau'}{\tau^2}. \quad (3)$$

Second terms, $t^{2R} 2R' \ln(t)$ and $\tau^{2R} 2R' \ln(\tau)$ inside the square brackets of Equation 3 represents a trajectory of varying radius of curvature, that is, spiral. Changing radius of curvature with time is directly linked to the changing bremsstrahlung asymmetry, R with time, R' . In the case of $R' = 0$, particle follows a constant radius trajectory where the bremsstrahlung asymmetry, R remains at a constant value. Such trajectory could be a circle.

This leads to

$$\lambda_{v/c} = \int_0^1 \sqrt{r^2 + \left(\frac{dr}{dt} \frac{\tau^2}{(\tau - t\tau')} \right)^2} \frac{\tau - t\tau'}{\tau^2} dt. \quad (4)$$

Equation 4 means that by knowing the mean free path of an electron in the atmosphere, the parameters a and b of a position vector $r(t)$ in equation one can be calculated.

4. Lightning Leader Tip Electric Field

In this section, the lightning leader tip electric field is formulated to define the particle velocity for the bremsstrahlung radiation. The velocity defined by an integration of the electric field can be equated to the formulated velocity (Equation 3) from the particle position vector to calculate the unknown parameters a and b . Moreover, this is an important step in defining particle velocity parameters such as the mean free path and time as the leader tip electric fields can be measured experimentally.

The measured lightning discharge electric field has a characteristic behavior of a sharp rise and a slow decay. This characteristic information was used to mathematically model the lightning leader tip electric field, E in Equation 5. The electric field is mathematically modeled using the ansatz of a sharp rise and a slow decay

$$E(t) = s_f \left[\frac{A n (s_{ft})^{(n-0.489)}}{(t / t_0)^{(n-0.489)}} - \frac{2 B (s_{ft})^m}{(t / t_0)^m} \right], \quad (5)$$

where $A = 4.083 \times 10^7 \left[\frac{\text{J}}{\text{Cm}} \right]$ or $\frac{\text{V}}{\text{m}}$, $B = 3.840 \times 10^7 \left[\frac{\text{J}}{\text{Cm}} \right]$ or $\left[\frac{\text{V}}{\text{m}} \right]$, $n = 1.95$ is a dimensionless constant, and

$m = 1.5$ is a dimensionless constant. The scaling factor $s_f = 1$ scales the lightning leader tip electric field magnitude, i.e., it is a dimensionless constant that can be used with different constant values to adjust the leader tip electric field to the required strength, $s_{ft} = 1$ is a dimensionless constant scaling factor of time, and the time t/t_0 is the relative time with respect to $t_0 = 1$ s where $t[s]$ is an independent time variable. The time range of the electric field is $s_{ft} 0.39694 \mu\text{s} \leq t \leq \infty$. Values closer to $0.397 \mu\text{s}$ describe the initial state and values closer to $72.2 \mu\text{s}$ describe the final state of the lightning leader tip electric field shown in Figure 3. The minimum value of time was determined such that it is the minimum value of the electric field before it goes to $-\infty$ at $t = 0$ s. The scaling factor determines the acceleration of the particle. As a result, the magnitude of the emitted radiation intensity scales in an indirect way by scaling the electric field which scales the force and hence the acceleration.

The lightning leader tip electric field in Equation 5 is derived such that it goes to zero at infinite time, as shown in Figure 3b. Any remaining residual electric field integrated over infinite time would accelerate a particle to velocities exceeding the speed of light.

The strength and duration of the electric field affect the magnitude and pattern of the emitted radiation. To adjust the electric field's strength and duration conveniently without affecting the graphical shape of Figure 3a, the scaling factor s_f for magnitude and s_{ft} for the duration are introduced.

5. Velocity Function

The velocity of the particle accelerated by the leader tip electric field is found by integrating the derived electric field (Figure 3, Equation 5). The upper and lower limits of the integration are chosen such that the result of the integral gives a velocity equation as a function of time. Next, the solution of integral is scaled to accelerate the particle up to 94% of the speed of light, i.e., 280,000 km/s, and thereby prevents the particle from exceeding the speed of light in vacuum, c . Finally, the particle velocity function should decay to the thermal velocity. Afterward, the particle should either remain constant at a thermal velocity or slowly decay to 0 m/s and remain stationary at an infinite time for the particle to stop radiating. A stationary or constant velocity at infinite time is important because Equation 13 requires an integral with limits between $\mp\infty$, and the particle cannot radiate for an infinite time. The derived velocity function (Equation 9) presented in Figure 4 accelerates the particle from a thermal velocity (92.493 km/s) up to 94% of c within ~ 12.427 ns and then decelerates back to a thermal velocity. The whole process takes ~ 24.854 ns.

The velocity function is the integral of the electric field function with respect to time, multiplied with the constant charged particle properties, which are elementary charge e , charge number z , and the inverse of accelerated electron mass $\frac{1}{m_e}$.

$$v(t) = \frac{ez}{m_e} \int_{s_{ft} 0.39694 \mu\text{s} + \left| \frac{t[s]}{t_0[s]} \right|}^{\infty} E(t) dt. \quad (6)$$

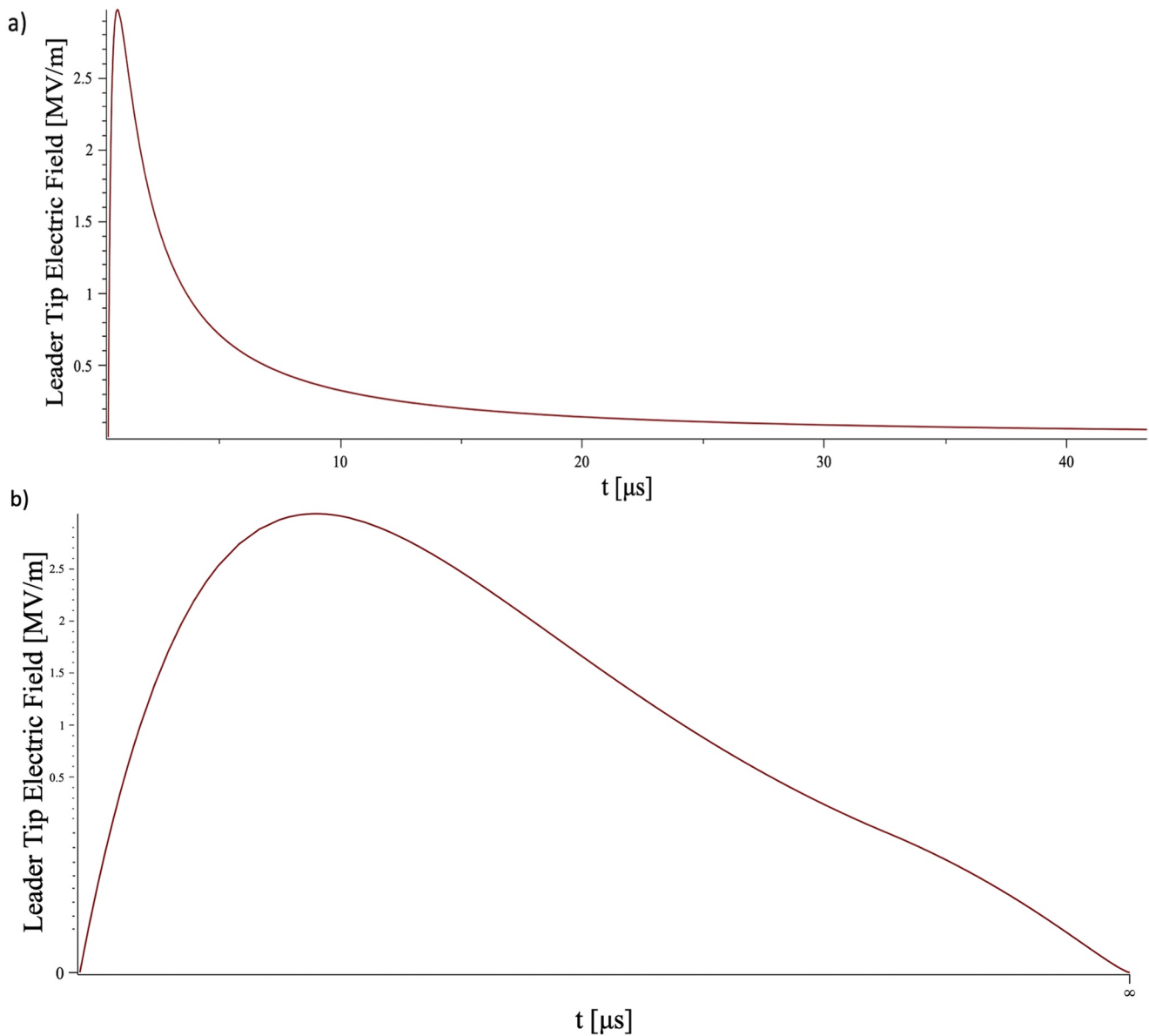


Figure 3. Derived electric field in Equation 5. (a) The leader tip electric field is modeled to fit observed electric field characteristics of lightning discharges. Lightning discharges exhibit a sharp rise and slowly decaying electric field. The peak value of the electric field was chosen to be ~ 3 MV/m, which is the approximate electric field for the conventional dielectric breakdown of air. For a RREA to develop, the electrons need to be relativistic such that bremsstrahlung occurs. In this case, the leader tip electric field is on the order of ~ 26 MV/m at 101 kPa of atmospheric pressure (Babich et al., 2015). (b) The derived electric field from 0 s to ∞ is plotted on logarithmic time scale. It is important that the leader tip electric field is 0 V/m at infinity. Any function used to describe the leader tip electric field that never goes to zero (i.e., an exponential) integrated up to infinity would cause a particle to exceed the speed of light.

Where $t_0 = 1$ s. The upper limit was set to be infinite in order to prevent the lower boundary of the integral from exceeding the upper limit. The lower limit involves a time parameter to prevent the indefinite integral from being a definite integral in order to create a velocity function. The absolute time is used here to create a symmetric velocity behavior in retarded time to explain the particle motion and also to satisfy the integral limits between $\mp\infty$ in Equation 13. The factorial was incorporated in order to prevent the electric field from becoming singular at $t = 0$ s.

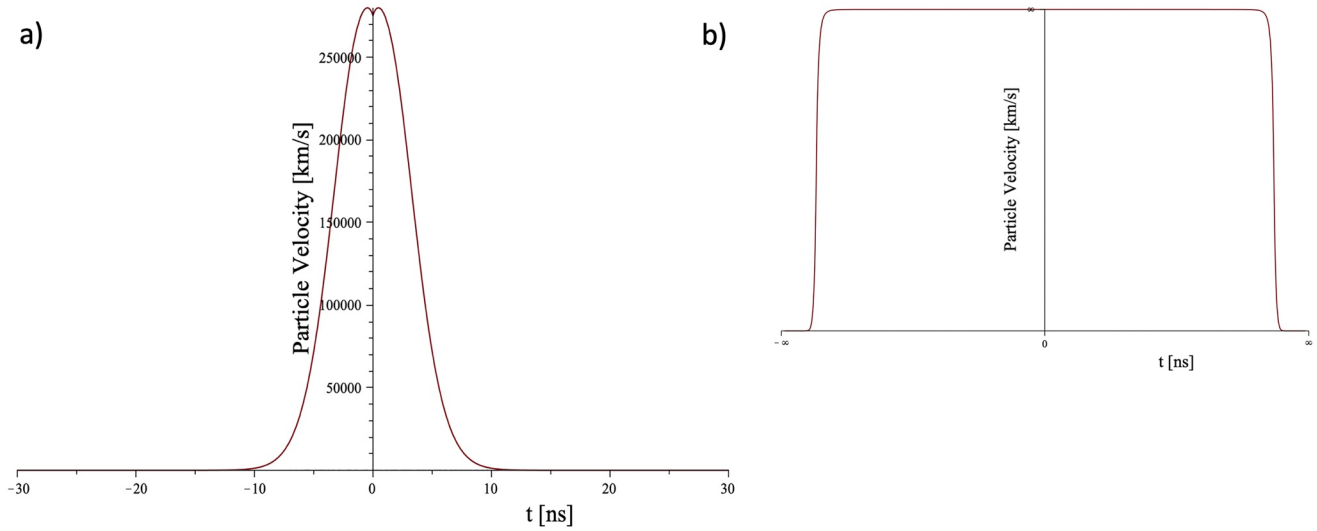


Figure 4. Velocity function (Equation 9) of the derived electric field. (a) There is no negative chronological time. However, negative time has a meaning in physics that it describes an off the record process of an experimentally observed electromagnetic pulse. Negative time is the retarded time in Liénard–Wiechert potentials that describes the radiation emission by the acceleration of a single particle. Retarded time in Liénard–Wiechert potentials is also the reason why Equation 13 has integral limits between $-\infty$ and ∞ . Negative time describes the duration of the process until the pulse detection time of the observer. Therefore, starting from ~ -10 ns up to ~ 10 ns, the particle gains velocity through an acceleration by the external leader tip electric field, and as it accelerates, it radiates electromagnetic radiation. When the particle velocity is peaking around $t = 0$ s, the observer starts receiving a signal and can measure the pulse as the process of particle acceleration, and the propagation of the electromagnetic wave from the source to an observer requires some time - known as the retarded, or negative, time. (b) The electron's velocity on an infinite time scale displays the end of a particle acceleration outside the range of $\sim \pm 11$ ns until $\pm \infty$. The peak velocity remains the same at 94% of the speed of light.

Separating integral (Equation 6) into the two terms of the electric field function (Equation 5) gives

$$v(t) = \frac{ez}{m_e} s_f A n s_{ft}^{(n-1.489+1)} \int_{s_{ft} 0.39694 \mu s + |t|!}^{\infty} \frac{1}{(t)^{(n-1.489)} t} dt - \frac{ez}{m_e} s_f 2 B s_{ft}^m \int_{s_{ft} 0.39694 \mu s + |t|!}^{\infty} \frac{1}{t^m} dt \quad (7)$$

and

$$v(t) = \frac{ez}{m_e} s_f A n (s_{ft})^{(n-1.489+1)} \left[\frac{(125) 2^{2.511+n} 3^{n-0.489} (9.657 \times 10^6)^{n-0.489} (23 s_{ft} + 5.794 \times 10^7 |t|!)^{-n+0.489} (3.969 \times 10^7 s_{ft} + |t|!)}{1000n - 1489} \right] - \frac{ez}{m_e} s_f 2 B (s_{ft})^m \left[\frac{(3.969 \times 10^{-7} s_{ft} + |t|!)^{-m+1}}{-1 + m} \right]. \quad (8)$$

The factorial of time only works for an integer number of time values. However, this is only a problem when the velocity function is used with non-integer values. To approximate velocities with a factorial of non-integer time values, Stirling's approximation can be used. The velocity function can be written with the substitution of all the constant values that do not require to be changed in order to scale. Also, to preserve the physical meaning of the characteristic lightning leader tip electric field feature with a sharp rise and a slow decay taking place in the atmosphere. These constants are A , B , n , and m . Substituting and simplifying gives

$$v(t) = s_{fv} s_f z (s_{ft})^{1.461} \frac{4.365 \times 10^{26}}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{fv} t|!)^{0.461}} - s_{fv} s_f z (s_{ft})^{1.5} \frac{1.565 \times 10^{27}}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{fv} t|!)^{0.5}}. \quad (9)$$

As the particle has a short mean free path with a short mean free time \sim ns, the time scale in the velocity function (8) was scaled accordingly to ns with $s_{fv} = 1 \times 10^9$ to reflect atmospheric electrodynamic conditions. Previous to the scaling, as the time t is integrated up to ∞ , the time scaling in the velocity function (8) has grown and it took ~ 12.427 s for the electron to reach 94% of the speed of light from its thermal velocity $v_{rms} = \sqrt{\frac{3kT}{m}}$ of 92.493 km/s at an example ambient air temperature of 188.15 K, when considering the mesosphere at an altitude of 80 km. Similarly, as the time t is integrated up to ∞ , the particle attained a large velocity with the multiplication of the constant term $\frac{ez}{m_e}$ in front of the integral. This is solved with a new velocity scaling factor $s_{fv} = 8.19 \times 10^{-11}$.

By knowing the particle velocity, the particle velocity equation (Equation 9) can be scaled directly by using the corresponding scaling factors for the required velocity. The mean free path (Equation 4) of a particle is a function of the total particle velocity. Although the derivative of the position vector (Equation 3) alone determines the velocity arising from the Coulomb force, Equation 3 is the total velocity when used for the mean free path (Equation 4). Hence, if the particle velocity is known, it can be equated to the derivative of the position vector on a spiral trajectory (Equation 3) to determine the parameters a (the relativistic mean free path correction parameter) and b (the interaction distance between an incoming and a target particle). This algebraic calculation transforms the velocity arising from the Coulomb field, that is, the derivative of the position vector in Equation 1, to an overall velocity, that is, the sum of the velocity arising from the Coulomb field and the velocity originating from the external lighting leader tip electric field for a particle on a spiral trajectory.

As can be seen from Figure 4, the above assumptions help to construct a velocity function (Equation 9). The complete radiation model is based on the fact that the particle reaches approximately 94% of the speed of light. Then the particle decelerates back to a thermal velocity as bremsstrahlung is the braking radiation due to obstacle particles or atoms, and as the source electric field from the leader tip starts to decay.

The relative particle velocity is

$$\beta(t) = \frac{v(t) + \frac{dr}{dt}}{c}. \quad (10)$$

Hence,

$$\beta(t) = \frac{s_{fv}s_f z(s_{fi})^{1.461}}{c} \frac{4.365 \times 10^{26}}{(1.333 \times 10^9 s_{fi} + 3.357 \times 10^{15} |s_{fv}t|!)^{0.461}} - \frac{s_{fv}s_f z(s_{fi})^{1.5}}{c} \frac{1.565 \times 10^{27}}{(1.333 \times 10^9 s_{fi} + 3.357 \times 10^{15} |s_{fv}t|!)^{0.5}} + \frac{dr}{dt} \frac{1}{c}, \quad (11)$$

where t is the time in s and the range of the time changes in the velocity function (Equation 9), hence in β (Equation 11) function. The time range in velocity function differs from the electric field function given in Equation 5 as a result of the integration process. The time range of the velocity function is $-\infty \leq t \leq \infty$. An important information to note about the indicated range of time t values separately in both, the scaled velocity function (Equation 9) and the electric field function (Equation 5), is that they are specific for this scaled velocity and non-scaled electric field function separately. There is no time parameter in any other equation nor in the final equation due to the nature of the definite integral of time. However, as indicated in Equation 13, the definite integral between the limits of $-\infty$ and ∞ of time results in a time t parameter to vanish in the final equation (Equation 36). Therefore, the range of time t values presented in Sections 4 and 5 can be neglected for the final Equation 36.

Before moving to the next section, the concepts explained so far make up the base of the complete theory. The next section uses the information of Sections 2, 4, and 5 to construct the core of the theory and to derive the novel generic radiation intensity pattern of a single particle unique to the bremsstrahlung process. Figure 5 shows how this theory fits into practical observations.

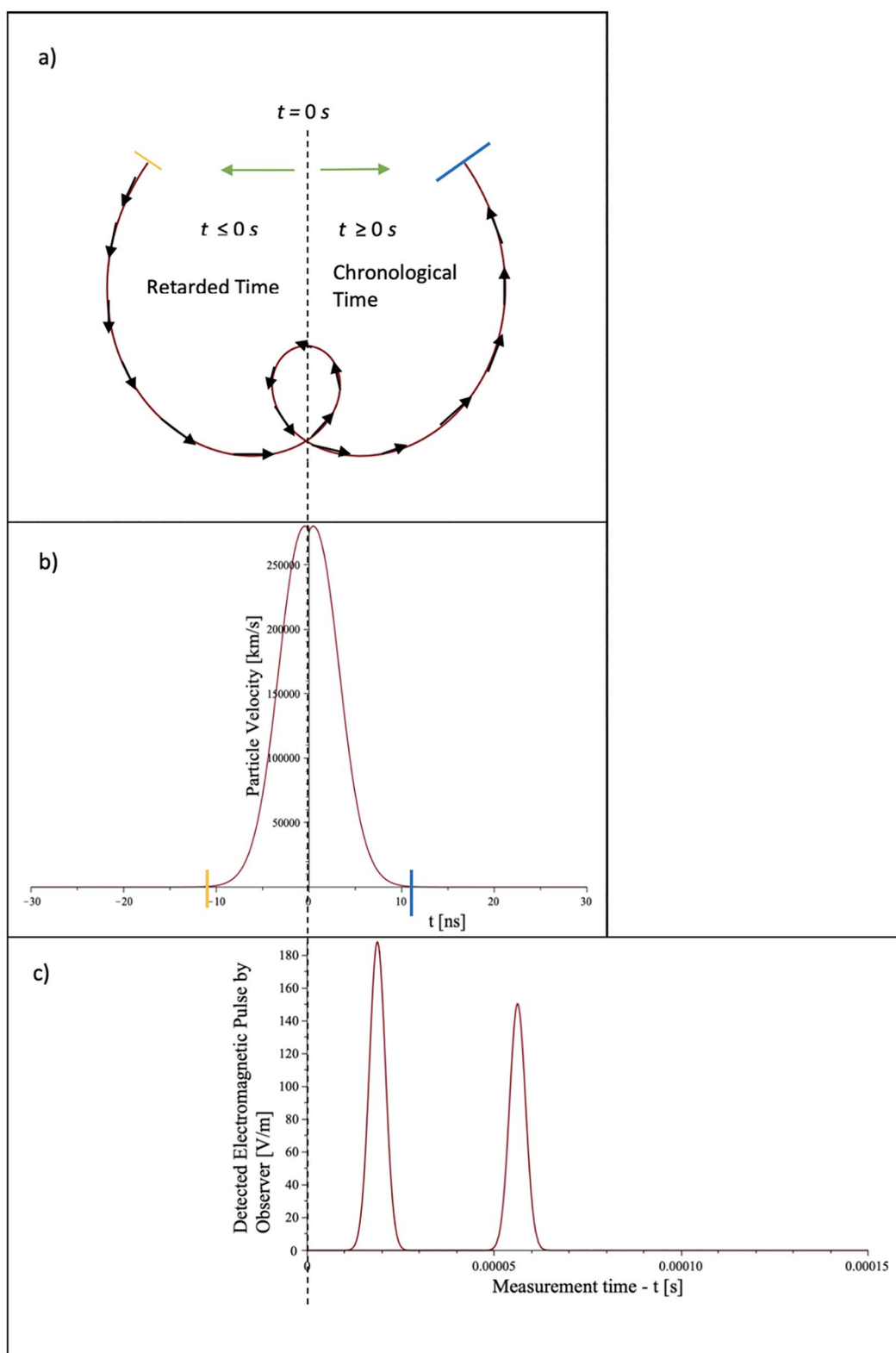


Figure 5

6. Mathematical Model of Combined High and Low-Frequency Bremsstrahlung Radiation Pattern

Deriving the bremsstrahlung radiation pattern.

This section introduces collision cross-sections as a part of explaining complete radiation pattern. Further information on collision cross-section and mathematical definitions consisting of form factors can be found in supplemental information file in the form of Text S2 (Punjabi & Perdrisat, 2014). Collision cross-section and form factors are very important in defining the final form of the radiation patterns at high frequency (Perdrisat et al., 2007). Form factors are not used in the final equation and in simulations of radiation patterns. For high-frequency emissions, the details of collision and interaction are crucial. They are accounted for by adding a cross-section of a particle of interest.

Hence, it is assumed that

$$\frac{d^3\chi}{d\omega'd\Omega_{rad}d\Omega_{par}} = \frac{d^2I}{d\omega'd\Omega_{rad}} \frac{d\sigma}{d\Omega_{par}}. \quad (12)$$

Dimensionally, $\chi = I \times \sigma$, where, σ in m^2 is the Coulomb scattering cross section, χ in $\text{Js}^{-1} \text{m}^2$ is the radiation cross section, and I in Js^{-1} is the radiation intensity. As a result of Liénard–Wiechert potentials, the radiation emitted by a single charged particle (Jackson, 1999, p. 675) is given by

$$\frac{d^2I}{d\omega'd\Omega_{rad}} = \frac{z^2 e^2 (\omega')^2}{4\pi^2 c \epsilon_0} |\zeta|^2, \quad (13)$$

where $\zeta = \int_{-\infty}^{\infty} n \times (n \times \beta) e^{i\omega'(t - \vec{n} \cdot \vec{r}(t)/c)} dt$. In addition, z is the dimensionless charge number, c is the speed of light, ϵ_0 is the permittivity of free space, ω' is the angular frequency of the emitted wave in rad/s in the particle's frame of reference and t is the time in s.

The radiation unit vector n , the position vector $r(t)$ in Equation 1, the particle motion, the position of an observer, and the overall theoretical concept about the bremsstrahlung radiation emission modeling that the equations are built upon are shown in Figure 2.

As n is a unit vector in the direction of the emitted radiation, the integrand of the integral in Equation 13 can be re-written in the sinusoidal form of the cross-product as

$$\zeta = \int_{-\infty}^{\infty} |\beta| \sin(\theta_{n,\beta}) e^{i\omega(t - \vec{n} \cdot \vec{r}(t)/c)} dt, \quad (14)$$

or in the sinusoidal form of dot product as

$$\zeta = \int_{-\infty}^{\infty} |\beta| \sin(\theta_{n,\beta}) e^{i\omega(t - r(t)\cos(\theta_{n,r(t)})/c)} dt. \quad (15)$$

Substituting Equation 1 and the first term of Equation 11 into Equation 14 and taking constants out results in

Figure 5. Complete order of events from initial particle acceleration until pulse detection by an observer. (a) The yellow line indicates the particle's starting position. It starts accelerating from -10 ns until $t = 0$ s as a result of an external leader tip electric field and Coulomb electric field (Bremsstrahlung) of the target particle in the atmosphere. (b) shows how the particle's velocity changes as a function of time. The particle is still allowed to propagate to positive time values. However, the starting time is from negative time values (retarded time). Negative time is required to reflect the reality that when the signal is measured by an observer at $t = 0$ s, it has to travel some time from the source to the observer. The acceleration process that causes the emission also requires some time to take place. The observer in chronological time does not observe all of these processes. Hence, they occur in negative retarded time. (c) sketches when the observer would receive the signal. It is plotted using the Dirac Delta function to sketch when the observer would receive the signal, which does not reflect the actual timing. The first signal could also be received at $t = 0$ s, however not in retarded time as it would then contradict the observations where there is no negative time.

$$\zeta = \frac{s_{fv} s_f z(s_{ft})^{1.461} 4.365 \times 10^{26}}{c} \sin(\theta_{n,\beta}) \int_{-\infty}^{\infty} \frac{1}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{fv} t|!)^{0.461}} e^{i\omega(t) \left[\frac{(t^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R c}{\tau^2 R_c^R \omega' \cos(\theta_{n,r(t)})} - \frac{at}{\tau} \right] \cos(\theta_{n,r(t)})/c} dt. \quad (16)$$

A variable transformation from t to s requires the Jacobian. Assuming that $s = -t$ it follows that $\frac{ds}{dt} = -1$ such that

$$\zeta = -\frac{s_{fv} s_f z(s_{ft})^{1.461} 4.365 \times 10^{26}}{c} \sin(\theta_{n,\beta}) \int_{-\infty}^{\infty} \frac{1}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{fv} (-s)|!)^{0.461}} e^{i\omega(-s) \left[\frac{((-s)^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R c}{\tau^2 R_c^R \omega' \cos(\theta_{n,r(t)})} - \frac{a(-s)}{\tau} \right] \cos(\theta_{n,r(t)})/c} ds. \quad (17)$$

The exponential term can be simplified such that

$$e^{-i\omega s} e^{-i\omega \frac{((-s)^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R c}{\tau^2 R_c^R \omega' \cos(\theta_{n,r(t)})} - i\omega \frac{as}{\tau} \cos(\theta_{n,r(t)})/c}. \quad (18)$$

Further simplification of the common parameters results in

$$e^{-i\omega s} e^{-i \frac{((-s)^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R}{\tau^2 R_c^R} - i\omega \frac{as}{\tau} \cos(\theta_{n,r(t)})/c}, \quad (19)$$

where $e^{-i\omega s}$ can be neglected as there is no $\frac{1}{c}$ term to reduce the rate of exponential decay. Therefore, this term goes to zero quickly with time such that

$$e^{-i \frac{((-s)^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R}{\tau^2 R_c^R} - i\omega \frac{as}{\tau} \cos(\theta_{n,r(t)})/c}. \quad (20)$$

Bringing the whole integral in Equation 16 together with the simplified exponential gives

$$\zeta = -\frac{s_{fv} s_f z(s_{ft})^{1.461} 4.365 \times 10^{26}}{c} \sin(\theta_{n,\beta}) \int_{-\infty}^{\infty} \frac{1}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{fv} (-s)|!)^{0.461}} e^{-i \frac{((-s)^R)^2 b^R (\omega')^R \cos(\theta_{n,r(t)})^R}{\tau^2 R_c^R} - i\omega \frac{as}{\tau} \cos(\theta_{n,r(t)})/c} ds. \quad (21)$$

Finally, the velocity Equation 9 converging to 0 m/s at $\mp\infty$ (see Figure 4b) enables the integral in Equation 21 to be evaluated. Convergence is due to the nature of derived electric field function, which goes to zero at infinite time and does not accelerate the particle any more. Furthermore, the use of functions such as Planck's radiation curve, Heidler current (Heidler et al., 2013), or Poisson distribution function with a variable transform to represent electric field would not be beneficial in deriving a velocity function as they diverge due to their exponential nature. The solution to the integral is approximated by the use of a formula in the book of Fourier transforms of exponential functions (Bateman et al., 1954, p. 121, Equation 23). The integral can be approximated by seeking a solution from the table of integral transforms by Harry Bateman for the exponential Fourier transform of the function $f(x) = f(s)$ being equal to

$$f(s) = \frac{1}{(1.333 \times 10^9 s_{ft} + 3.357 \times 10^{15} |s_{fv} (-s)|!)^{0.461}} e^{i\alpha^2 s^2}. \quad (22)$$

The solution of the integral requires index ν in the variable $(ix)^\nu$ that multiplies the exponential function in Exponential Fourier Transform of function $f(x) = f(s)$ to be $\nu > -1$ (Bateman et al., 1954, p. 121, Equation 23). The derivation of the velocity (Equation 9) results in the beta function (Equation 11) satisfying this requirement.

As the integrand of the integral in Equation 14 multiplies with the β function consisting of 2 terms given in Equation 11, the integral can be separated and solved individually for each term such that all terms can be added together to give the final result.

Hence, the approximated solution of the integral ζ for the first term of the scaled velocity function (11) β is $\nu_1 = -0.461 > -1$ such that

$$\frac{d^2 I}{d\omega \Omega_{rad}} = -\frac{s_{fv} s_f z(s_{ft})^{1.461} 4.365 \times 10^{26}}{c} \left[\pi^{1/2} 2^{-(1/2)\nu_1} \alpha^{-\nu_1-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{\nu_1}(2^{-1/2} \alpha^{-1} y) \right]. \quad (23)$$

The approximated solution of the integral ζ for the second term of the scaled velocity function (11) β is $\nu_2 = -0.5 > -1$ such that

$$\frac{d^2 I}{d\omega \Omega_{rad}} = \frac{s_{fv} z s_f(s_{ft})^{1.5} 1.565 \times 10^{27}}{c} \left[\pi^{1/2} 2^{-(1/2)\nu_2} \alpha^{-\nu_2-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{\nu_2}(2^{-1/2} \alpha^{-1} y) \right], \quad (24)$$

where $\alpha^2 = \frac{b^R(\omega')^R (\cos(\theta_{n,r(t)}))^R}{(\tau^R)^2 c^R} [s^{-2}]$ and $y = \frac{\omega' \cos(\theta_{n,r(t)}) a}{c \tau} [s^{-1}]$.

Therefore, the final radiation pattern for both high and low-frequency emission is the sum of the two terms

$$\begin{aligned} \frac{d^2 I}{d\omega \Omega_{rad}} = & \frac{z^2 e^2 (\omega')^2}{4\pi^2 c \epsilon_0} \left| \sin(\theta_{n,\beta}) \right| \left[-\frac{s_{fv} s_f z(s_{ft})^{1.461} 4.365 \times 10^{26}}{c} \right. \\ & \left. \left[\pi^{1/2} 2^{-(1/2)\nu_1} \alpha^{-\nu_1-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{\nu_1}(2^{-1/2} \alpha^{-1} y) \right] \right. \\ & \left. + \frac{s_{fv} z s_f(s_{ft})^{1.5} 1.565 \times 10^{27}}{c} \left[\pi^{1/2} 2^{-(1/2)\nu_2} \alpha^{-\nu_2-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{\nu_2}(2^{-1/2} \alpha^{-1} y) \right] \right] \Bigg|^2, \quad (25) \end{aligned}$$

where R is the bremsstrahlung asymmetry control parameter, which plays a crucial role in distinguishing the parameter y from α , which is a requirement in order to be able to approximate the solution to an integral (dimensionless). $D_\nu(z)$ is the parabolic cylinder function.

The parabolic cylinder function is given by (Whittaker & Watson, 1927, p. 347)

$$D_\nu(z) = 2^{v/2+1/4} z^{-1/2} W_{v/2+1/4,1/4}(1/2 z^2), \quad (26)$$

where, $W_{v/2+1/4,1/4}(1/2 z^2)$ is a Whittaker function (Whittaker & Watson, 1927, p. 346) and

$$W_{\kappa,\mu}(1/2 z^2) = \frac{\Gamma(-2\mu) M_{\kappa,\mu}(1/2 z^2)}{\Gamma(1/2 - \mu - \kappa)} + \frac{\Gamma(2\mu) M_{\kappa,-\mu}(1/2 z^2)}{\Gamma(1/2 + \mu - \kappa)}, \quad (27)$$

where, $M_{\kappa,\mu}(1/2 z^2)$ is another Whittaker function (Kiyosi Ito and The Mathematical Society of Japan, 1993; Whittaker & Watson, 1927, p. 347) and

$$M_{\kappa,\mu}(1/2 z^2) = {}_1F_1(1/2 + \mu - \kappa; 2\mu + 1; 1/2 z^2) (1/2 z^2)^{1/2+\mu} e^{-1/2 z^2}. \quad (28)$$

For the second term of Equation 27 with Whittaker M function of negative μ

$$M_{\kappa, -\mu}(1/2z^2) = {}_1F_1(\mu - \kappa; 2\mu; 1/2z^2)(1/2z^2)^{1/2+\mu} e^{-1/2z^2}, \quad (29)$$

where ${}_1F_1(1/2 + \mu - \kappa; 2\mu + 1; z)$ is a confluent hypergeometric function of the first kind (Abramowitz & Stegun, 1972) and

$${}_1F_1(1/2 + \mu - \kappa; 2\mu + 1; 1/2z^2) = \frac{U_{\kappa, \mu}(1/2z^2)}{e^{-1/2z^2/2}(1/2z^2)^{\mu+1/2}}, \quad (30)$$

where, $U_{\kappa, \mu}(\frac{z^2}{2})$ is the confluent hypergeometric function of the second kind, named Kummer's U function such that

$$U_{\kappa, \mu}\left(\frac{z^2}{2}\right) = \left(\frac{z^2}{2}\right)^{\mu+1/2} e^{-\frac{z^2}{4}} \sum_{n=0}^{\infty} \frac{(m - k + 1/2)_n}{n!(2\mu + 1)_n} \left(\frac{z^2}{2}\right)^n. \quad (31)$$

By direct comparison of Equation 26 with 25 and Equation 26 with 27, the constant quantities z , κ and μ are defined as $z = 2^{-1/2}\alpha^{-1}y$, $\kappa = v/2 + 1/4$, $\mu = 1/4$.

Finally, in order to complete Equation 25, the observation angles ($\theta_{n,\beta}$ and $\theta_{n,r(t)}$) need to be reduced from two to one as they are related to each other.

Currently, the radiation pattern is observed from $\cos(\theta_{n,r(t)})$ and $\sin(\theta_{n,\beta})$. In addition, $\cos(\theta_{n,r(t)})$ coming from substitute parameter α and y in Equation 25 and $\sin(\theta_{n,\beta})$ being a variable parameter in Equation 25 defining the radiation around the particle, also known as the solid angle. In addition, by looking at Figure 2b, it can be seen that the velocity vector, which is the derivative of the position vector $r(t)$ is always perpendicular to the position vector where radiation unit vector pointing in the direction of the emitted radiation makes up the hypotenuse of the formed rectangular triangle. Hence, the angles of the rectangular triangle can be equated to each other.

The velocity function defined in Equation 9 provides the particle speed attained by the acceleration due to the external leader tip electric field, and it is scalar. The derivative of the position vector $r(t)$ (Equation 1) provides the velocity caused by the Coulomb interaction of an incoming and target particle, which also determines the curved trajectory of the bremsstrahlung given by the position vector (Equation 1). Hence, the particle's overall velocity moving perpendicular to the position vector $r(t)$ is the sum of the Coulomb velocity arising from the derivative of the position vector $r(t)$ plus the scalar speed attained by the external leader tip electric field given in Equation 9. Equation 9 is not the velocity, but the speed. It only provides magnitude but gives no information about the direction of the particle's motion. The direction of the particle's motion is already provided by the position vector $r(t)$ given by Equation 1. As a result

$$\sin(\theta_{n,\beta}) = \frac{r(t)}{n} \quad (32)$$

and

$$\cos(\theta_{n,r(t)}) = \frac{r(t)}{n}, \quad (33)$$

such that

$$\sin(\theta_{n,\beta}) = \cos(\theta_{n,r(t)}). \quad (34)$$

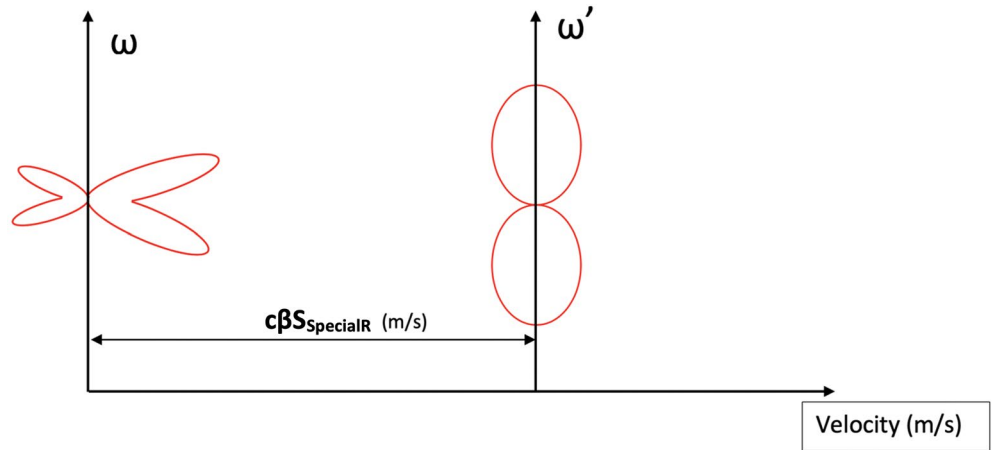


Figure 6. Particle radiation at two different frames of references. Laboratory (stationary) frame of reference ω and moving frame of reference ω' with the velocity $c \frac{v(t)}{c} S_{SpecialR}$. In addition, $S_{SpecialR} = 0$ meaning the observer is moving with the particle at the same velocity at all times, hence would see no radiation emission as the particle is not accelerating. As $S_{SpecialR} \rightarrow 0$, the observer would see dipole radiation pattern. As $S_{SpecialR} \rightarrow 1$, the observer would approach rest and would see the particle propagating at its own velocity defined by Equation 9 with the corresponding radiation pattern.

7. Dynamic Bremsstrahlung – Special Relativity and Doppler Effect

Now the Doppler shift is added to the emitted radiation.

Figure 6 displays a stationary laboratory frame ω where the relativistic effects and the Doppler shift is observed, particle's moving frame of reference ω' and the relative velocity $c\beta S_{SpecialR}$ between the observer and the two different frames that determines what an observer would detect. In addition, $S_{SpecialR}$ (dimensionless) is a scaling factor of the relative velocity between the two different frames of references. In particular, $S_{SpecialR} = 1$ means the observer is stationary, and the particle is propagating at its velocity. As the $S_{SpecialR}$ approaches zero, the relative velocity between the observer and the particle falls, meaning that the observer is getting closer to the particle's velocity and would start to observe or detect a non-relativistic dipole radiation pattern.

The equation for translating the angular frequency of the emitted wave into the laboratory frame for the Doppler shift is given by (Jackson, 1999, p. 720, Equation 15.40).

$$\omega' = \gamma\omega(S_{SpecialR} - \beta S_{SpecialR} \cos(\theta_{n,\beta})). \quad (35)$$

As shown in Figure 2b, $\theta_{n,\beta}$ is the angle between the emitted radiation unit vector and the particle velocity vector.

Substituting Equations 34 and 35 into the Equation 25 gives

$$\begin{aligned} \frac{d^2 I}{d\omega \Omega_{rad}} = & \frac{z^2 e^2 \left[\gamma\omega(S_{SpecialR} - \beta S_{SpecialR} \cos(\theta_{n,\beta})) \right]^2}{4\pi^2 c \epsilon_0} \left| \sin(\theta_{n,\beta}) \left[-\frac{s_{fi} s_f z (s_{fi})^{1.461} 4.365 \times 10^{26}}{c} \right. \right. \\ & \left. \left[\pi^{1/2} 2^{-(1/2)v_1} \alpha^{-v_1-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{v_1}(2^{-1/2} \alpha^{-1} y) \right] + \right. \\ & \left. \left. \frac{s_{fi} z s_f (s_{fi})^{1.5} 1.565 \times 10^{27}}{c} \left[\pi^{1/2} 2^{-(1/2)v_2} \alpha^{-v_2-1} e^{-\frac{y^2 \alpha^{-2}}{8}} \times D_{v_2}(2^{-1/2} \alpha^{-1} y) \right] \right] \right|^2, \end{aligned} \quad (36)$$

where the Lorentz factor γ is $\gamma = \frac{1}{\sqrt{1-\beta^2}}$ (dimensionless), and ω is the received angular frequency in the laboratory (stationary) frame of reference in rad/s. In addition, the new definition of α is $\alpha^2 = \frac{b^R(\gamma\omega(S_{SpecialR} - \beta S_{SpecialR}\cos(\theta_{n,\beta})))^R(\sin(\theta_{n,\beta}))^R}{(\tau^R)^2 c^R} [s^{-2}]$, and the new definition of y is $y = \frac{\gamma\omega(S_{SpecialR} - \beta S_{SpecialR}\cos(\theta_{n,\beta}))\sin(\theta_{n,\beta})a}{c\tau} [s^{-1}]$.

The scaling factor s_f in Equation 36 is carried over from the definition of the electric field Equation 5. As a result of the integration process, s_f scales the magnitude of the emitted radiation, which is directly linked with its previous purpose of scaling the electric field, which defines the magnitude of the emitted radiation. The scaling factor s_f can be approximated at high frequency with a scalar equation that gives the magnitude of the emitted high-frequency radiation from a single particle in an instantaneous circular motion (Jackson, 1999, p. 679, Equation 14.84).

Hence,

$$s_f \approx \frac{\frac{d^2 I}{d\Omega d\omega} \big|_{\theta=0}}{\left| \frac{d^2 I}{d\Omega d\omega} \right|} \approx \frac{\frac{3e^2 \gamma^2 \omega e^{\frac{-\omega}{\omega_c}}}{4\pi c \omega_c}}{\left| \frac{d^2 I}{d\Omega d\omega} \right|} \quad (37)$$

and the critical angular frequency (Jackson, 1999, p. 679, Equation 14.81) is

$$\omega_c = \frac{3\gamma^3}{2} \left(\frac{c}{p} \right), \quad (38)$$

where p is the radius of curvature of instantaneous circular motion in m , ω_c is the critical angular boundary frequency. Beyond critical angular boundary frequency, the emitted radiation would have a minimal value in all directions such that it can be neglected.

8. Results

This section presents predictions of radiation patterns of a single electron accelerated under an external lightning leader tip electric field using the derived Equation 36.

Radiation patterns change as a result of changing particle velocity, and acceleration as a particle get relativistic. As the acceleration of a charged particle increases, it emits higher intensity, hence higher frequency radiation. Therefore, the input parameter, emitted radiation frequency represents particle acceleration. To simulate relativistic and non-relativistic radiation patterns, the range of input velocity and frequency values were chosen and tabulated in Table 1.

The new radiation patterns reveal that as well as forward peaking, there are also backward peaking lobes as the particle gains speed. The magnitude of the acceleration determines the magnitude of the radiation intensity. The horizontal axis is the radiation intensity per solid angle per emitted frequency range in eight different radiation patterns displayed in Figure 7. Although there is overall more energy in the forward peaking lobes, the peak radiation intensity in the backward direction is slightly higher in patterns 5, 6, and 7 as they are more horizontally projected. However, the total radiation intensity is higher in the forward direction as the peaking lobes have larger beam areas due to the Doppler effect's presence.

The peak intensity of the backward lobes starts to decay as a result of the Doppler shift from pattern seven onwards. The Doppler shift effect can be observed from the consideration of both radiation length in the horizontal direction (as the horizontal axis is a measure of the radiation intensity per solid angle, i.e., more horizontally projected) and the total area of the emitted radiation pattern. The larger the area of the

Table 1

Radiation Pattern Input Parameters are Velocity as a Ratio to the Speed of Light β , and Emitted Radiation Frequency by the Particle for Eight Different Radiation Patterns

Velocity is shown as ratio with the speed of light β and emitted radiation frequency (Hz)		
Radiation patterns	Scaled bremsstrahlung electron velocity β (dimensionless)	Emitted radiation frequency (Hz)
Pattern 1	0.02	1 k
Pattern 2	0.36	10 k
Pattern 3	0.67	1 M
Pattern 4	0.69	7 M
Pattern 5	0.77	10 M
Pattern 6	0.85	100 M
Pattern 7	0.87	500 M
Pattern 8	0.90	1 G

Note. Emitted radiation frequency by a single particle is an input parameter to the final Equation 36. It plays a crucial role in delivering valuable information and determining the particle's total acceleration, hence the total velocity is indirect as there is no input acceleration parameter in the final Equation 36. In addition, β function determines the particle's total velocity arising significantly from the external lightning leader tip electric field. The remaining contribution to the overall velocity comes from the Coulomb force between an incoming and target particle, where the target particle causes an incoming particle to follow and covers some arc length of the spiral trajectory described in the position vector (Equation 1) during its mean free time of the collision.

radiation pattern, the more the beam extends in the horizontal direction or, when it is more projected toward the horizontal axis, the more energy the emitted wave has. An increasing particle velocity translates into increasing radiation frequency. Therefore an increase in energy in the forward direction of the particle lowers the frequency and hence lowers the energy at the backward direction of the particle due to the Doppler effect. For example, in the eighth radiation pattern, which is at $\sim 94\%$ of the speed of light and has 1 GHz frequency, increasing the frequency of the observed radiation leads to an increase in the Doppler effect that reduces the frequency, hence the energy ($E = hf$) of the beam at the back of the particle. As a consequence, backward radiation lobes start to shrink such that overall radiation pattern looks like a basic forward peaking radiation pattern (the pattern 8).

Radiation patterns on the polar plot have to start from 2π rad until 0 rad to show the particle's complete radiation. Progression toward the zero is because a particle starts from the retarded time and radiates until the chronological time $t = 0$ s and beyond, as shown in Figures 2 and 5. If the polar plot were made with time as a representation of the solid angle, the plot would have started from ~ -10 ns (when the particle starts accelerating, see Figures 2 and 5) until 0 s. In addition, this is the same as starting from positive time and ending at $t = 0$ s. Therefore, the final result is independent of positive or negative start time as long as the time flow is toward zero seconds. The independence of the sign is caused by the absolute time and symmetry of the velocity function. Progression toward zero seconds is caused by the initial start time being retarded negative time.

The derived mathematical model (Equation 36) is specifically for the bremsstrahlung process. Unlike many of the other radiation emission processes of charged particles (i.e., from linear acceleration or crossing the boundary between two different dielectric media), the bremsstrahlung process affects the shape of the emitted radiation. This bremsstrahlung effect causes an asymmetry of the emitted radiation about the particle's velocity vector or, in other terms, the direction of motion. To understand this effect, we can compare an electron to a car that travels in the dark with the headlight turned on. The headlights are the emitted electromagnetic radiation by the car. When the car gets into the bend, like the bremsstrahlung process of an electron, an observer outside the car can immediately tell the radiation shape of the headlight would be asymmetric by looking at the reflections on the road compared to the case when the car follows a straight path. This novel effect is clearly shown in supporting information in Figure S1, which displays a real-life example of visible asymmetry of the car headlights (electromagnetic radiation) on a bend.

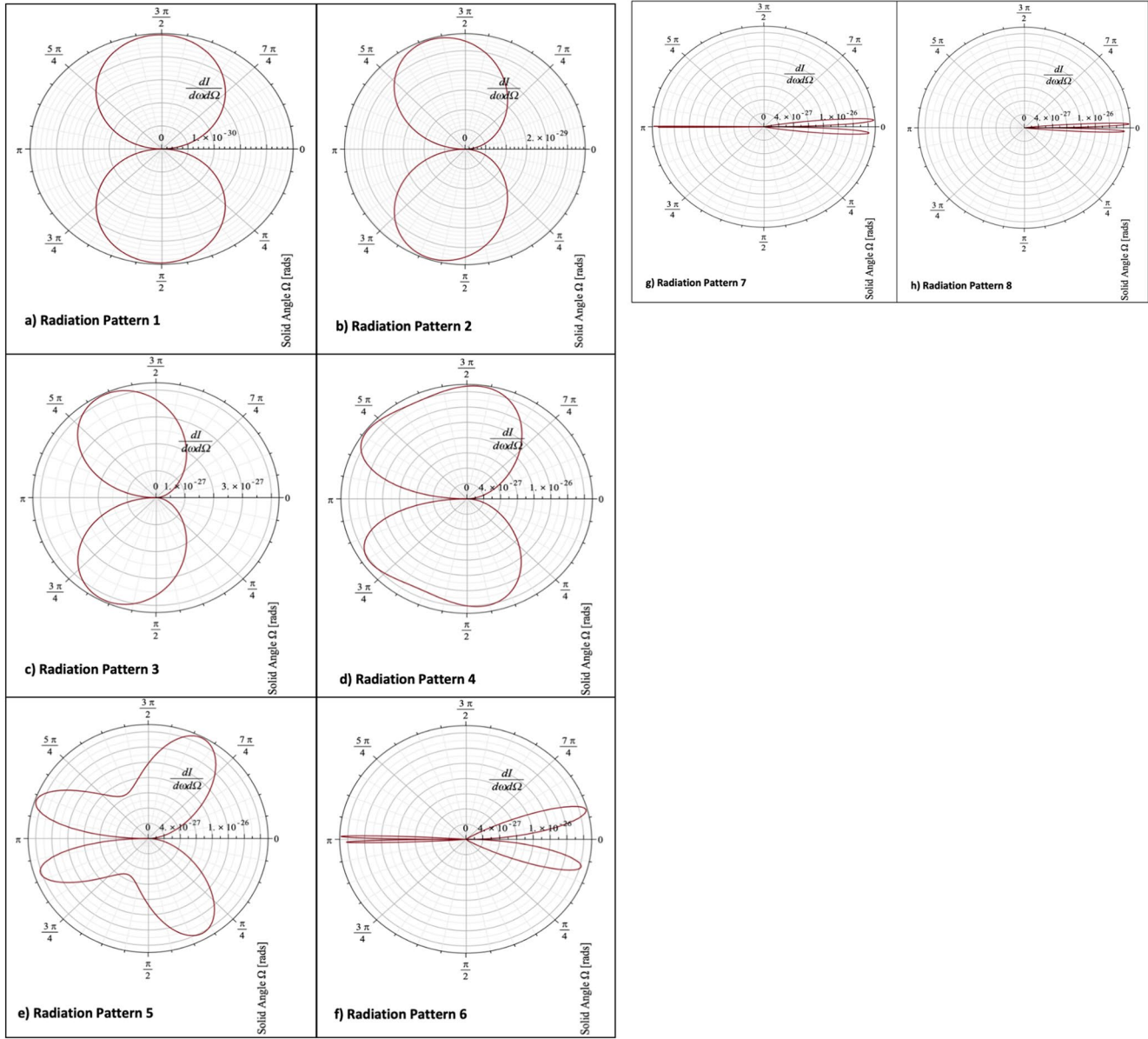


Figure 7. The Radiation patterns emitted by anti-clockwise rotating charged particle - bremsstrahlung process. Plot is in Polar co-ordinates. Horizontal axis gives the radiation intensity per Solid angle, Ω , per emitted angular radiation frequency, ω . In addition, angle of the Polar plot is the Solid angle, Ω . ((a)–(f)) The radiation pattern connecting dipole with a forward-backward peaking radiation pattern is called the transition pattern. The radiation patterns and formula explaining the radiation patterns is not complete if it does not demonstrate the transition patterns. The transition patterns explain how the particle attains relativistic forward-backward peaking lobes. The radiation pattern of a particle during the bremsstrahlung, starting from a low velocity up until a relativistic velocity starts from the dipole radiation pattern and exhibits the forward and backward peaking radiation pattern. The transition is demonstrated in radiation patterns 4 and 5, where the dipole collapses to form four maxima. The values used for plotting are: mean free time $\tau = 30 \mu\text{s}$, number of charges $z = 1$, $a = 100 \mu\text{m}$, $b = 1 \text{ nm}$ (a and b are related to mean free path), $s_{fl} = 1$, $s_f = 1$, $S_{\text{Spectral}R} = 1$, velocity-time scaling factor $s_{fv} = 1 \times 10^9$ and velocity scaling factor $s_{fv} = 8.19 \times 10^{-11}$. Finally, the bremsstrahlung asymmetry is $R = 1/8$. In addition, $\frac{1}{9} \leq R \leq \frac{1}{3}$. (g) A particle reaching relativistic speed attains a focused beam in both forward and backward direction. (h) A particle reaching to ultra-relativistic speed experiences the take over of the Doppler effect that minimizes the low-frequency radiation in the backward direction and emits almost fully in the forward direction.

Figure 8 displays the effects of the bremsstrahlung radiation asymmetry control quantity R . The bremsstrahlung asymmetry R depends on the radius of the curvature of the incoming particle trajectory undergoing the bremsstrahlung process. In addition, the bremsstrahlung asymmetry R increases as the radius of curvature of the bremsstrahlung trajectory decreases.

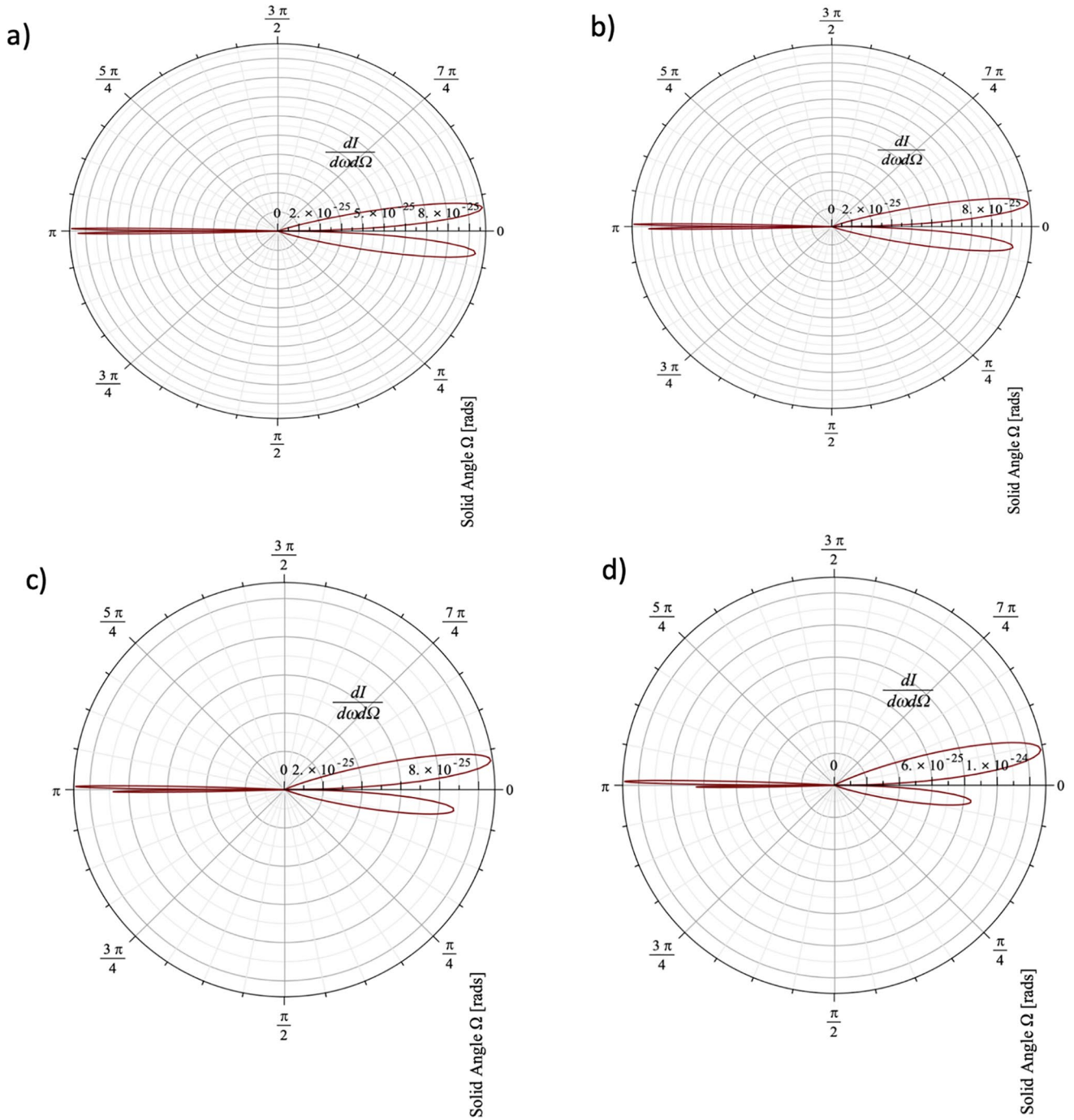


Figure 8. Bremsstrahlung Asymmetry Quantity R with different values. (a) $R = \frac{1}{9}$, (b) $R = \frac{1}{6}$, (c) $R = \frac{1}{4}$, (d) $R = \frac{1}{3}$.

9. Discussion

First, the created model describing radiation patterns specifically for the bremsstrahlung process is modeled to be generic and applicable to all bremsstrahlung process on earth, outer space, or other planets. As mentioned in the third paragraph of the introduction, TGFs are believed to originate as a result of the bremsstrahlung process of RREA driven by external leader tip electric field. Hence, the theoretical model assumes a single particle accelerated under an external leader tip electric field interacting with a target

particle or ion via Coulomb field. In addition, the incoming particle is assumed to follow a spiral particle trajectory. This assumption is based on the in-balance of centripetal (Coulomb field between two particles + Earth's external magnetic field) and centrifugal (external leader tip electric field) forces acting on the incoming particle when it interacts with a target particle via Coulomb field. Any in-balance in centripetal and centrifugal force acting on an incoming particle that follows a curved trajectory would result in a spiral particle trajectory. Hence, the choice of spiral trajectory considers the effects of the Earth's external magnetic field, B on an incoming particle trajectory in an indirect way.

Particle speed defined in Equation 9 relies on the ansatz of a sharp rise and slow decay form of leader tip electric field. This ansatz originates and is backed up by the asymmetric γ -ray bursts (GRBs) measurement that reveals the internal charge structure of the leader tip described in the introduction section, in the second paragraph.

Although acceleration is modeled from the electric field that assumes the same behavior as the leader tip electric field observed in the atmosphere, it only provides a magnitude component of the electric field acceleration. Independent input parameters such as particle trajectory, the bremsstrahlung asymmetry, R , and the Doppler effect are fundamental physics where the particle trajectory, and the bremsstrahlung asymmetry, R are specific to the bremsstrahlung process. These input parameters enable the created model to apply to any bremsstrahlung processes.

The bremsstrahlung process has a continuous electromagnetic spectrum meaning it is capable of explaining both ionizing and low-frequency radiations. The bremsstrahlung radiation patterns presented in Figure 7 are unique to a bremsstrahlung process. They are always the same in geometry with forward-backward radiation patterns at relativistic particle speeds and all the asymmetries. However, when the model is applied to any other bremsstrahlung processes other than TGFs, only radiation intensities are subject to change but not the radiation shapes.

Overall, the eight different radiation patterns plotted in Figure 7 are not the complete radiation patterns, especially for high-frequency radiation. As mentioned in the previous section, a high-frequency radiation pattern is sensitive to the interaction process described by scattering cross-sections that an incoming bremsstrahlung electron experiences against target particles. Moreover, plotted radiation patterns are the second-order derivative term $\frac{d^2 I}{d\omega d\Omega_{rad}}$ of Equation 12, and it misses the information of interaction described by Coulomb scattering cross-section. In future work, a complete radiation pattern will be calculated using the complete Equations 12 and 36, and Equation 1 provided in the supplemental information, S2 with appropriate form factors, and a dimensional analysis to describe high-frequency emissions. Furthermore, the first part of the derived equation demonstrates a success as Figure 7a (patterns 1) and 7h (pattern 8), display the common expected radiation patterns such as low-frequency dipole radiation pattern and relativistic forward peaking radiation patterns (Jackson, 1999, p. 669, Figure 14.4).

The dipole radiation pattern 1 in Figure 7a displays perfect symmetry about both the velocity axis and the axis perpendicular to the velocity vector. In the absence of a Doppler and bremsstrahlung asymmetric radiation effect, the four maxima of the relativistic radiation pattern are symmetric about both the velocity axis and the axis perpendicular to the velocity. Therefore, following the statement of Noether's theorem, conservation of energy arising from the symmetry in the radiation lobes about the two axes mentioned above, causes the dipole radiation pattern 4 and 5 (in Figures 7d and 7e) to collapse and form 4 symmetric maxima. Initially, there were symmetries about two axes, and as the particle gains up speed, the number of symmetric axes should still be preserved. Hence, if there is only a forward peaking radiation with none in the backward direction, one of the symmetry is broken (about axes perpendicular to the velocity). The idea of preserving the initial number of symmetric axes is quite powerful in understanding why there are four maxima (two forward peakings and two backward peakings) and why the dipole collapses to form 4 maxima of the radiation pattern. All the radiation patterns in Figure 7 incorporate the Doppler effect and the bremsstrahlung trajectory asymmetry. These two effects cause the radiation pattern to be asymmetric. However, the emitted total radiation energy is still conserved compared between the presence and absence of radiation pattern asymmetry causing physical effects (The bremsstrahlung trajectory and the Doppler effect) at any given particle velocity. In other words, the total energy radiated at a given velocity is the same regardless of the Doppler and the bremsstrahlung trajectory effect. While these two effects increase the

radiation beam intensity in one direction, they also decrease the radiation intensity in other directions by the same amount such that the overall energy radiated by the particle at a given speed remains the same.

On the other hand, the reported radiation asymmetry of the bremsstrahlung (R -parameter) about the velocity vector due to the particle's curved trajectory was found to be existing at relativistic speeds when forward and backward peaking radiation patterns are present. This bremsstrahlung radiation asymmetry (R -parameter) we get is a novel finding. At the non-relativistic speeds, when radiation has a dipole pattern, asymmetry of the bremsstrahlung was found to be absent. This absence indicates that asymmetry is associated with being a physical effect rather than a mathematical artifact. The absence of the bremsstrahlung radiation asymmetry at a non-relativistic dipole radiation pattern of a particle following a curved trajectory can again be compared to a car having its headlights on each side of the door radiating perpendicular with respect to the velocity vector of the car. In this case, one can see that two emitted radiation beams would not be asymmetric as two beams are in the opposite direction to each other and are independent of the curvature of the trajectory as headlights radiate perpendicular to the trajectory.

Asymmetries are important because they alter the radiation intensity of the emitted radiation. Asymmetries increase the radiation intensity in one direction and decrease it in the other direction. The bremsstrahlung radiation pattern of a single particle with involved asymmetries can be used with a Monte Carlo simulation for multiple particles to predict radiation intensity map across an area where lightning discharge takes place. The radiation intensity map is the expected radiation power at each different observation location around the area of lightning discharge location. The predicted radiation intensity map allows improving the lightning detection and location systems. It helps to find the location of peak radiation intensity, hence contributes to the development of better radiation protection systems.

10. Summary

In summary, the radiation patterns were found to be peaking in a backward direction as well as an already known forward direction. In addition, four maxima (2 in the forward direction and 2 in the backward direction) were found to be due to the conservation of symmetry axes arising from the initial dipole radiation pattern. With the introduction of the Doppler effect, peaking lobes of radiation in forward and backward directions were found to be asymmetric about an axis perpendicular to the particle's velocity vector. Moreover, the novel second asymmetry of peaking lobes with respect to particle's velocity vector was found to be unique to the bremsstrahlung due to the particle following a curved trajectory. Finally, it is also found that low-frequency radiation peaks in the backward direction, whereas high-frequency peaks in the forward direction.

Data Availability Statement

The Maple worksheets used to simulate the particle trajectory, external lightning leader tip electric field, particle velocity, and the radiation patterns are openly available from the University of Bath Research Data Archive at <https://doi.org/10.15125/BATH-00810>.

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Acknowledgments

M. Yücemöz would like to thank to his supervisor M. Füllekrug for all the opportunities, suggestions, guidance, reviews, and support throughout his PhD study. The Engineering and Physical Sciences Research Council (EPSRC) and MetOffice sponsor M. Yücemöz's PhD project under contract numbers EG-EE1239 and EG-EE1077. M. Füllekrug acknowledges support from the Natural Environment Research Council (NERC) for under grants NE/L012669/1 and NE/H024921/1. M. Yücemöz wishes to thank Dr. Adrian Hill for mathematical support in helping to solve integral with a divergence problem and his family for their support and good wishes.

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