

An analysis of recorded and simulated SH wave reverberations in the upper mantle beneath the USArray

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November 30, 2022

Abstract

Long-period ($T > 10$ s) shear-wave reverberations between the surface and reflecting boundaries below seismic stations are useful for studying the mantle transition zone (MTZ) but finite-frequency effects may complicate the interpretation of waveform stacks. Using waveform data from the USArray and spectral-element method synthetics for 3-D seismic models, we illustrate that a common-reflection point (CRP) modeling of layering in the upper mantle must be based on 3-D reference structures and accurate calculations of reverberation traveltimes. Our CRP mapping of recorded waveforms places the 410-km and 660-km phase boundaries about 15 km deeper beneath the western US than beneath the central-eastern US if it is based on the 1-D PREM model. The apparent east-to-west deepening of the MTZ disappears in the CRP image if we account for shear-wave velocity variations in the mantle. We also find that ray theory overpredicts the traveltime delays of the reverberations if 3-D velocity variations in the mantle are prescribed by global models S40RTS, SEMUCB-WM1, and TX2015. Undulations of the 410-km and 660-km are underestimated in the analysis when their wavelengths are smaller than the Fresnel zones of the wave reverberations in the MTZ.

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31 1. Introduction

32
33 Recordings of long-period ($T > 10$ s) shear waves are useful data to map seismic discontinuities and velocity
34 gradients in the mantle transition zone (MTZ) (e.g., [Shearer, 1990](#)). The mineral-phase transitions near
35 depths of 410 and 660 km produce the highest amplitude shear-wave reflections after the S wave arrival
36 (e.g., [Shearer and Buehler, 2019](#)), before the SS arrival (e.g., [Flanagan and Shearer, 1998](#)), and between
37 multiple ScS reflections (e.g., [Revenaugh and Jordan, 1991](#)) in stacks of transverse-component
38 seismograms. We call these boundaries the “410” and “660” in this paper and define the MTZ as the layer
39 of the mantle between the 410 and 660. Constraints on the depths of the 410 and 660 and the thickness of
40 the MTZ constrain the temperature and composition of the mantle (e.g., [Bina and Hellfrich, 1994](#); [Xu et al.](#)
41 [2009](#)) and heat and mass transfer between the upper and lower mantle.

42
43 Most seismological studies of hundreds to thousands of waveforms are based on 1-D seismic reference
44 profiles and ray theory to facilitate the analysis and computations. However, long-period shear waves are
45 sensitive to seismic structure in the mantle beyond the geometric ray so ray-theoretical calculations of
46 traveltimes and waveform shifts may be inaccurate (e.g., [Tromp et al., 2005](#)). Modeling inaccuracies have
47 been discussed thoroughly for the SS wave and its precursors (e.g., [Neele et al., 1997](#); [Zhao and Chevrot,](#)
48 [2003](#); [Bai et al., 2012](#); [Guo and Zhou, 2020](#); [Koroni and Trampert, 2016, 2021](#)), but they apply to all long-
49 period seismic wave reflections and conversions in the MTZ, including the multiple ScS reverberations
50 (e.g., [Haugland et al., 2020](#)) and receiver functions (e.g., [Deng and Zhou, 2015](#)).

51
52 The receiver-side shear-wave reverberation in the upper mantle is the phase of interest in this paper. It has
53 been introduced by [Shearer and Buehler \(2019\)](#), a study we abbreviate as SB19 from hereon, as a new wave
54 type for probing the upper mantle and the MTZ. Using USArray waveforms and a common-reflection-point
55 (CRP) imaging method, SB19 estimated the depths of the 410 and 660 to be 40–50 km deeper beneath the
56 western US than beneath the central and eastern US. This is an important study outcome as it implies that
57 the seismic contrast in the uppermost mantle beneath the tectonically active western US and tectonically
58 stable central and eastern US extends into the MTZ.

59
60 SB19 used ray theory and the 1-D iasp91 velocity model to relate traveltimes to reflector depth. They
61 acknowledged that 3-D seismic velocity heterogeneity may have a significant effect on the amplitude,
62 coherence, and depths of the 410 and the 660 in the CRP images. In this paper, we follow up on their
63 recommendation to investigate how 3-D velocity structure changes the interpretation of CRP imaging
64 results and to test the hypothesis that the 410 and 660 beneath the US are unperturbed. Using a stacking
65 method instead of an inversion method we confirm that reverberation traveltimes are longer and therefore
66 that the 410 and 660 are 40–50 km deeper beneath the western US than the central-eastern US if the analysis
67 is based on a 1-D reference structure (section 2). We explore how strongly 3-D shear-velocity heterogeneity,
68 as constrained by global tomography, perturb reverberation traveltimes and how ray-theoretical traveltime
69 corrections change the CRP images (section 3). Using spectral-element method seismograms we evaluate
70 the accuracy of ray theory in predicting the reverberation traveltimes and whether undulations on the 410
71 and 660 are resolvable by long-period shear wave reflections (section 4). In section 5, we discuss our key
72 findings.

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74

75 2. Mapping of the 410 and 660 by 1-D common reflection point imaging

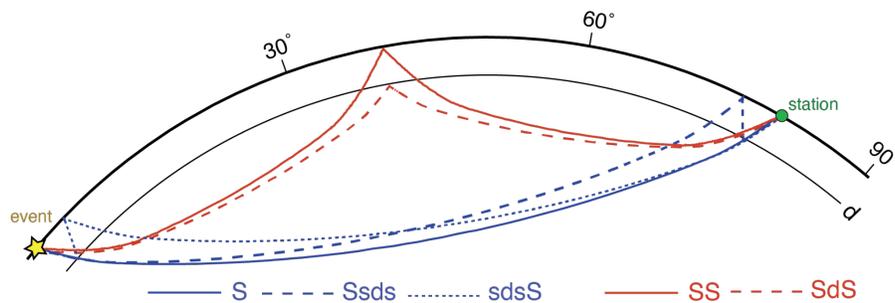
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77 2.1 The Ssds phase

78

79 A shear-wave reverberation beneath the receiver is abbreviated here as “Ssds” following the notation of
 80 SB19. Ssds is a shear-wave that follows a similar path in the mantle as the direct S wave and reflects off
 81 the free surface and off the top of a reflector at depth d before it is recorded by a seismometer on the surface
 82 (Figure 1). The arrival time of Ssds after S depends primarily on the reflector depth d and the shear-wave
 83 speed above the reflecting layer. For PREM, an earthquake at the surface, and an epicentral distance of 80° ,
 84 Ss410s and Ss660s arrive 159.6 s and 242.2 s after S, respectively. Ssds can interfere with SS precursors
 85 but the two phases have different slownesses and are distinguishable in waveforms recorded over a wide
 86 epicentral distance range. The top-side reflection sdsS near the source has the same traveltimes as Ssds at
 87 any distance for a 1-D velocity structure. For stations at similar azimuths, source-side reflections points are
 88 virtually identical whereas the Ssds reflection points are separated beneath the arrays of stations. Therefore,
 89 variations in the Ssds traveltimes are primarily due to seismic structure in the upper mantle beneath the
 90 seismic stations. There is no source-side and receiver side ambiguity if the analysis is limited to earthquakes
 91 deeper than the reflecting boundaries of interest (Liu and Shearer, 2021) but the data set would be
 92 significantly smaller.

93



94

95 **Figure 1.** Ray diagram of the phases S (solid blue line), Ssds (dashed blue line), sdsS (dotted blue line), SS (solid red
 96 line) and the SS precursor SdS (dashed red line) for an epicentral distance of 80 degrees.

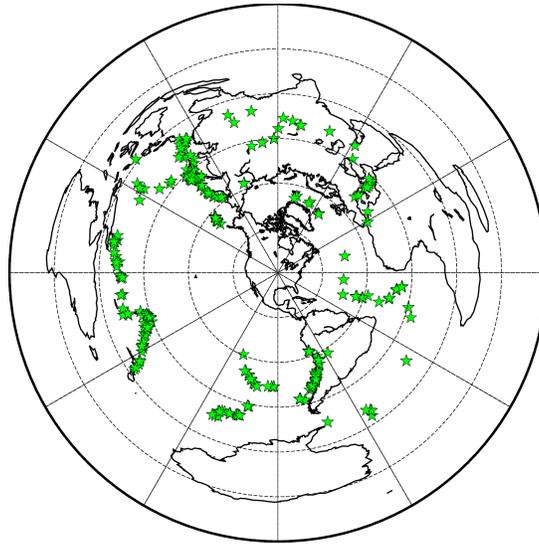
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98 2.2 USArray waveforms

99

100 Our data set includes 59,517 transverse-component displacement waveforms from 337 global earthquakes
 101 (Figure 2) recorded by stations from the USArray and other regional networks in the forty-eight
 102 conterminous United States. The earthquakes are shallower than 35 km, so the direct and depth phases form
 103 a single pulse at long periods. The earthquakes have moment magnitudes smaller than 7.0 so rupture
 104 complexity does not affect the long-period waveforms strongly. The epicentral distances are between 60°
 105 and 110° and waveforms have been filtered using a bandpass Butterworth filter with corner frequencies of
 106 20 mHz and 80 mHz. We align the waveforms on the peak S-wave displacement and normalize them, so
 107 the S waves have the same polarities and maximum displacements of +1. In all waveforms, the S-wave
 108 displacement is at least six times larger than the signal in the 100-s long window prior to the S wave onset.
 109 The maximum and the root-mean-square displacement in the window [30 s, 220 s] after the S-wave arrival
 110 time are six times and three times smaller than the peak S-wave displacement, respectively. We remove
 111 earthquakes with fewer than 20 seismograms left after these quality control steps.

112



113

114 **Figure 2.** (stars) Epicenters of earthquakes used in this study. The dashed circles have a common center of [40N, -
 115 95E] and radii of 30, 60, 90, 120, and 150 degrees.

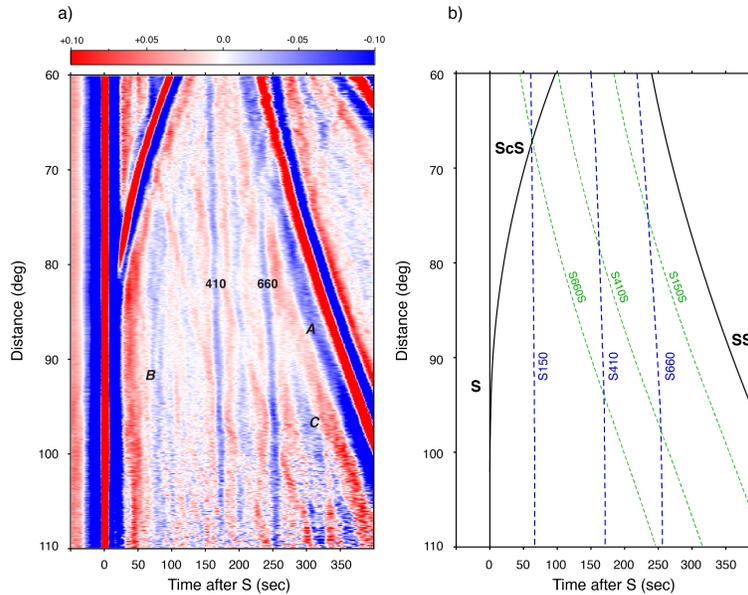
116

117 A record section of the sum of these waveforms bring out Ss410s and Ss660s as the strongest mantle
 118 reflections (Figure 3). The Ss410s and Ss660s have mean amplitudes of about 0.05 and are recorded without
 119 interference with ScS and SS at distances larger than 60° and 75°, respectively. The SS precursors S410S
 120 and S660S are weaker than Ss410s and Ss660s at distances smaller than 110° (e.g., Shearer, 1991). The
 121 signal labeled ‘A’ in Figure 3a, which arrives about 50 seconds before SS, appears to be an SS-precursor
 122 reflection at a depth of about 125 km. The signals near the label ‘B’ in Figure 3a between about 45 and 70
 123 seconds may be Ssds reflections from the lithosphere-asthenosphere boundary (LAB), a boundary that has
 124 also been studied with P-wave and S-wave receiver functions (Rychert et al., 2007; Abt et al., 2010; Hopper
 125 and Fischer, 2018) and multiple S-wave reflection (Liu and Shearer, 2021). We suspect that the signal 45 s
 126 after S with a positive polarity is a side lobe due to the applied butterworth filter because a signal with
 127 similar strength is present about 45 s before S.

128

129 Ssds reflections from the uppermost lower mantle below the 660 arrive more than 250 s after S and do not
 130 interfere with SS and the S410S precursor at distances larger than 95° in region ‘C’ of Figure 3a. However,
 131 it is difficult to differentiate reflections below the 660 from shallower SS precursors because S waves are
 132 attenuated by diffraction around the core and slowness resolution is poor. The high-amplitude signal about
 133 330 s after S has a slowness of roughly 1.0 sec/degree which is smaller than the slowness of any SS
 134 precursor. Its traveltimes is similar to that of the phases Ss410s410s (i.e., the shear-wave reverberation with
 135 two up-and-down shear-wave segments between the surface and the 410) and the phase PSs660s (i.e. the
 136 PS phase with an additional top-side reflection off the 660). However, it is unlikely that these phases can
 137 be recorded with high amplitudes on transverse component records.

138



139
 140 **Figure 3.** (a) Record section of transverse component seismograms used in this study. Shown is the amplitude of
 141 ground displacement in red and blue for positive and negative polarities, respectively, with a color intensity
 142 proportional to the absolute value. The seismograms have been aligned to the S wave at time 0. (b) The arrival times
 143 of S, ScS, and SS (black lines), S150, S410 and S660 (blue dashed lines), and the SS-precursors S150S, S410S and
 144 S660S (green dashed lines) have been computed for PREM for a source depth of 20 km.

145

146 2.3 Common reflection point imaging

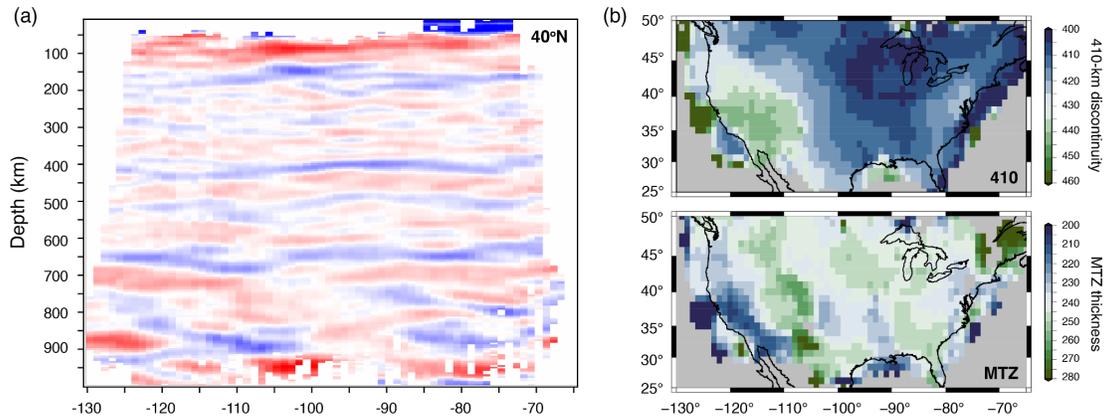
147

148 By 1-D common reflection point (CRP) mapping we convert the Ssds-S difference times to the locations
 149 of the Ssds reflecting points in the upper 800 km of the mantle beneath the USArray. We use the TauP
 150 software (Crotwell et al., 1999) and the PREM velocity structure to calculate Ssds reflection points and
 151 traveltimes. At a depth d , 1,716 reflection points are uniformly distributed on a $1^\circ \times 1^\circ$ horizontal grid
 152 between 25°N and 50°N and between -130°E and -65°E . The horizontal grids are separated by 5 km from
 153 10 to 1,000 km, for a total of 199 depths. For a gridpoint X , we select waveforms for which the Ssds
 154 reflection points are within the $1^\circ \times 1^\circ$ bin around X and for which the theoretical Ssds arrival time differs
 155 more than 15 s from the theoretical arrival times of sS, ScS, and sScS, and more than 50 s from the arrival
 156 time of SS to avoid wave interference. If fewer than five waveforms are available, we deem the mean
 157 displacement of Ssds to be inaccurately determined.

158

159 Since we use shallow focus earthquakes, the source-side and the receiver-side reflections have identical
 160 traveltimes. From synthetic seismograms for PREM, we have verified that they are equally strong so we
 161 attribute half of the mean Ssds amplitude to a source-side reflection. To construct the CRP images, we
 162 estimate source-side and receiver-side reflections sequentially for each of the 337 earthquakes following
 163 three steps. First, we determine the mean of the Ssds displacement at the theoretical arrival time of Ssds for
 164 all seismograms. Second, we subtract this mean value from the Ssds displacement of each waveform. Third,
 165 we assume the residual displacement to be due to reflections beneath the USArray. After mapping the Ssds
 166 signals onto the grid ray theoretically, we average the receiver-side reflection amplitudes within $1^\circ \times 1^\circ$
 167 bins, which are narrower than the Fresnel zones of 10-s period Ssds reflections in the mantle transition
 168 zone, as shown by SB19. We have also implemented the approach by SB19, who estimate the source-side

169 and receiver-side contributions to Ssds in one step using a sparse-matrix inversion solver. This approach
 170 yields smaller amplitudes of the Ssds reflections but the overall character of the CRP image, including the
 171 depths of the 410 and 660, are similar (Supplementary Figure 1).
 172



173
 174 **Figure 4.** (a) Vertical section of the CRP image along 40°N. Red and blue indicate positive and negative polarities,
 175 respectively. The color scale used is the same as in Figure 3. (b) Depth of the 410 (top) and the thickness of the mantle
 176 transition zone (bottom).
 177

178 Figure 4a shows a vertical section of the CRP image along the 40°N parallel. As expected from Figure 3,
 179 the 410 and 660 are the clearest reflectors. Variations of the Ss410s-S and Ss660s-S difference times project
 180 as spatial variations in the depth of 410 and 660. The 410 and 660 are deeper and more complex beneath
 181 the western US (west of -100°E) than beneath the central and eastern US. This is also apparent in other
 182 sections through the CRP images, not shown here. The 410 is strongest between longitudes -100°E and -
 183 75°E. The 520-km discontinuity may be responsible for a relatively weak Ssds reflection between the 410
 184 and 660. The CRP images near the 410 and 660 west of -100°E are complex, which was also noted by SB19.
 185 Strong reflectors corresponding to the Ssds signals in region B of the record section of Figure 3 are mapped
 186 at about 100 km and 150 km depth, but their depths and strengths vary. The incoherent structures at depths
 187 larger than 800 km are most likely imaging artifacts because these structures correspond to the amplified
 188 signals in region C of Figure 3, where S are diffracting waves and slowness resolution is relatively poor.
 189

190 Figure 4b shows maps of the depth of the 410 and the thickness of the MTZ. These are estimated from the
 191 absolute minimum values of the mean displacements in the CRP image in the depth ranges of 350–470 km
 192 (for the 410) and 620–730 km (for the 660) by cubic spline interpolation. We do not estimate the depth of
 193 the 410 and 660 where a secondary absolute minimum is stronger than 40% of the absolute minimum in
 194 these depth ranges. The depth of the 410 varies by 40–50 km. The 410 is deepest beneath the southern Basin
 195 and Range and the Colorado Plateau and shallowest beneath the central plains and the Atlantic coast. The
 196 thickness of the MTZ varies less than 10 km because the 410 and 660 depth variations are similar. The
 197 MTZ is thinnest beneath California and thickest beneath the Southern Rocky Mountains and the Colorado
 198 plateau. The MTZ thickness is anomalous in small regions near the margins of our model domain. This
 199 includes the extremely thin (210 km) MTZ beneath the west coast of central California which was also
 200 resolved by SB19. However, the CRP images have low resolution here because the data coverage is poor.
 201

202 SB19 resolved similar maps as Figure 4b, indicating that our and SB19’s data sets contain consistent
 203 variations of the Ss410s-S and Ss660s-S difference times and that estimates of the depths of the 410 and
 204 660 do not strongly depend on the applied mapping method.

205
 206

207 3. Influence of 3-D seismic heterogeneity on the CRP images

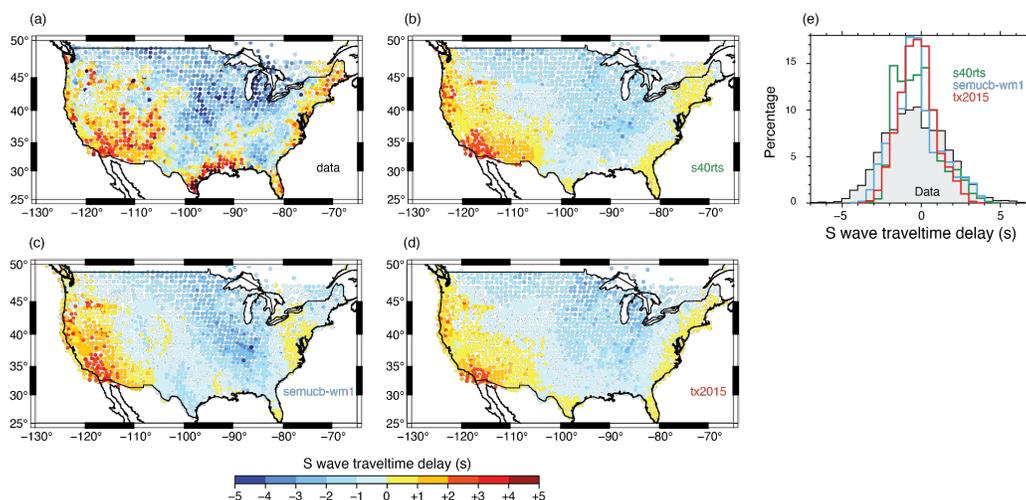
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209 The map of the depth of the 410 shown in Figure 4b is reminiscent of the estimated shear-wave velocity
 210 variations in the upper mantle beneath the US from the modeling of regional S-waves (e.g., Grand and
 211 Helmberger, 1984), surface-waves (e.g., Van der Lee and Nolet, 1997), and, more recently, ambient noise
 212 (e.g., Bensen et al., 2008), P-wave polarization (Park et al., 2019) and surface-wave amplification (Eddy &
 213 Ekstrom, 2014; Bowden & Tsai, 2017). This indicates strongly that shear-velocity variations in the mantle
 214 affect Ss410s-S and Ss660s-S difference times and that a mapping method based on a 1-D velocity structure
 215 would overestimate undulations of the 410 and 660.

216

217 3.1 S-wave traveltimes variations

218



219
 220 **Figure 5.** The recorded (a) and predicted (b, c, d) traveltimes delays of S waves by tomographic models S40RTS (in
 221 b), SEMUCB-WM1 (in c), and TX2015 (in d). Each circle indicates the location of a seismic station. Its color indicates
 222 the mean of the S-wave traveltimes delays with respect to the PREM model for at least five S waves. (e) Histograms
 223 of the S wave traveltimes delay in the data (grey fill) and predictions by S40RTS (green line), SEMUCB-WM1 (blue
 224 line), and TX2015 (red line) for the stations in panels a–d.

225

226 Figure 5a shows shear-wave velocity variations in the upper mantle affect the traveltimes of S waves.
 227 Plotted are the average S-wave delay with respect to the PREM velocity model of at least five S-waves
 228 recorded at seismic stations from the USArray. The delay times have been corrected for ‘source terms’,
 229 representing the effects of a potential mislocation of the earthquake location and origin time on the absolute
 230 S wave traveltimes. S waves recorded by USArray stations in the western US (the tectonically active region)
 231 arrive on average 5–6 seconds later than at stations in the central and eastern US (the stable platform). The
 232 global-scale mantle models S40RTS (Ritsema et al., 2011), SEMUCB-WM1 (French & Romanowicz,
 233 2014), and TX2015 (Lu & Grand, 2016) predict a similar traveltimes pattern (Figures 5b–d) but the range is
 234 slightly smaller than in the data (Figures 5e). Our calculations, not included in Figure 5, indicate that the

235 crustal structure from CRUST1.0 (Laske et al., 2013) enhances the east-west contrast only slightly, so wave
 236 speed variations in the mantle are primarily responsible for the S-wave traveltime differences.

237
 238 The imperfect match between the recorded and the predicted S wave traveltime is expected because
 239 tomographic models do not perfectly explain the recorded traveltime variation of any shear-wave (e.g.,
 240 Ritsema et al., 2004). Nevertheless, it is obvious that shear-velocity heterogeneity affects teleseismic S
 241 wave traveltimes across the USArray. Since Ssds has two additional propagation legs through the upper
 242 mantle, the Ss410s-S and Ss660s-S difference times are likely to be double the variation shown in Figure
 243 5a due to shear velocity heterogeneity only. If shear wave speed variations in the upper mantle beneath
 244 North America are ignored in the modeling, a variation of the Ss410s-S and Ss660s-S difference times of
 245 more than 10 s would imply that the depths of the 410 and 660 vary by about 18 and 20 km or more. This
 246 is of the same magnitude as resolved in Figure 4.

247

248 3.2 Ray-theoretical corrections

249

250 Since the tomographically predicted S-wave traveltime variation of 5–6 seconds across the USArray is a
 251 significant fraction of the recorded traveltime variation, we suspect that shear-velocity variations in the
 252 upper mantle influence the CRP imaging and our estimate of the 410 depth. To quantify this, we determine
 253 the CRP image for “corrected” Ssds-S difference times. From the measured Ssds-S difference time, we
 254 subtract the predicted difference time anomaly (positive or negative) by shifting a 5-s wide segment of the
 255 waveform around the theoretical Ssds arrival time. We predict the Ssds-S difference time by ray tracing
 256 through S40RTS, SEMUCB-WM1, or TX2015. In the calculations, we do not include a crust in the velocity
 257 models so Ssds-S difference times are artificially short by about 3.5 s.

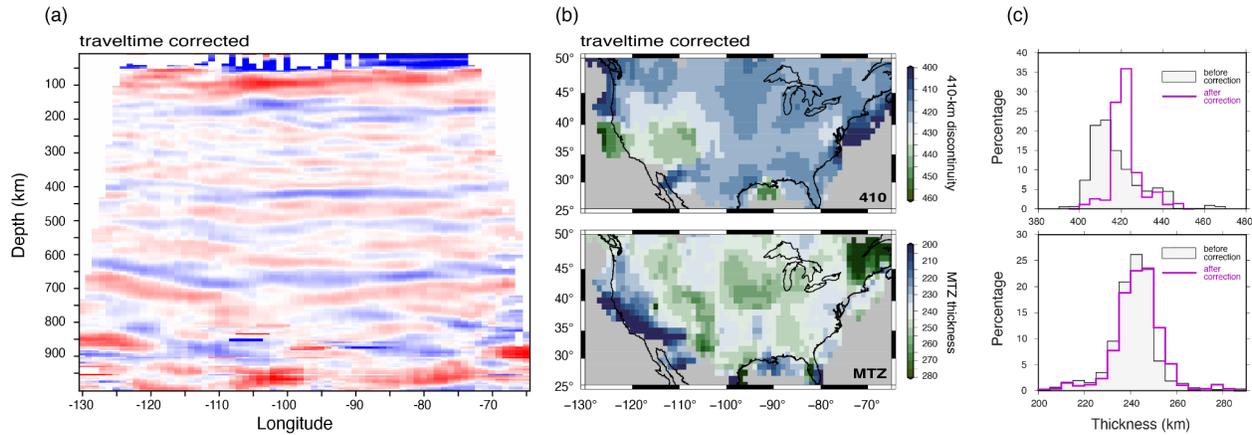
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259 Figure 6a shows the CRP image along the 40°N parallel after travel time corrections using S40RTS. The
 260 images for SEMUCB-WM1 and TX2015 are similar, as expected from Figure 5. Compared to the
 261 uncorrected profile shown in Figure 4a, The character of the corrected and uncorrected CRP images are the
 262 same but the 410 and 660 are flatter boundaries across the USarray. This is especially clear for the region
 263 between -100°E and -80°E where the 410 and 660 are relatively simple. Figure 6b emphasizes that the depth
 264 variations of the 410 is smaller when the CRP image is based on tomographically corrected Ssds-S
 265 difference times. The thickness of the MTZ in the corrected and uncorrected images are similar because
 266 shear velocity variations are relatively weak in the MTZ compared to the uppermost mantle. The histograms
 267 shown in Figures 6c illustrate that the depth variation of the 410 is about a factor of two smaller when
 268 traveltime corrections have been applied to the data and that the corrections do not change the range in
 269 MTZ thickness values. The travel time corrections change the mean depth of the 410 and 660 by about 10
 270 km, which is similar to the change obtained by Shearer & Buehler (2019) using ray-theoretical corrections
 271 computed for a regional 3-D velocity model. The 410 and 660 are imaged deeper than for the uncorrected
 272 waveforms because we underpredict Ssds-S traveltimes in our calculation by excluding the crust in the
 273 velocity model.

274

275 One cannot argue that the ray-theoretically corrected images reflect the actual depth variations of the 410.
 276 Since S40RTS and any other tomographic model does not explain perfectly the recorded S-wave traveltime
 277 variation (see Figure 3), it is unlikely that the traveltimes corrections have not completely removed the
 278 effects of the shear-velocity structure on the CRP image. Further, we show in the next section that ray-

279 theoretical predictions of long-period Ssds-S traveltimes are inaccurate and that corrections can project as
 280 spurious signals in the CRP images.
 281



282
 283 **Figure 6.** Depth of the 410 (in a) and the thickness of the transition zone (in b) estimated after ray-theoretical travel
 284 time corrections have been applied. Panels c and d show histograms of the resolved 410 depth and the thickness of the
 285 MTZ beneath the USArray with (black line) and without (purple line) traveltime corrections. Compare with Figure
 286 4b.
 287
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289 4. Resolution tests using spectral-element-method waveforms

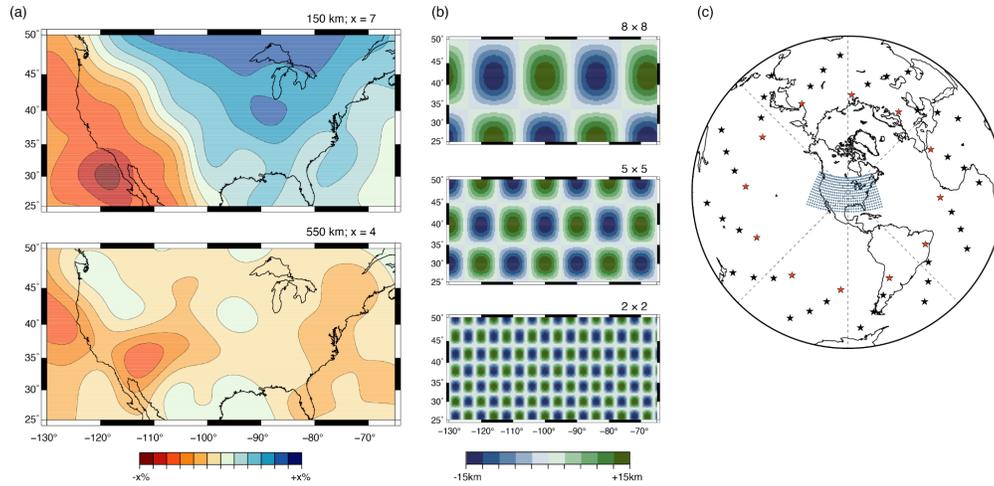
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291 We analyze synthetic waveforms to test the robustness of our CRP imaging approach, the resolution of 410
 292 and 660 undulations from long-period Ssds waveform data, and the accuracy of ray-theoretical corrections.
 293 The waveforms are computed using SPECFEM3D-Globe software (e.g., Komatitsch and Tromp, 2002;
 294 Komatitsch et al., 2016) modified by us to allow for undulations of the 410 and 660. The seven test
 295 structures are PREM, S40RTS (Figure 7a), SEMUCB-WM1, TX2015, and structures T2, T5 and T8 (Figure
 296 7b). In each structure, the density and velocities in the uppermost mantle extend to the surface. We remove
 297 the crust from seismic models to suppress reverberations in the crust that complicate the waveforms
 298 following the S wave.
 299

299

300 The one-dimensional PREM structure with discontinuities at 220 km, 400 km, 670 km depth serves as a
 301 baseline test for determining artefacts in the CRP images unrelated to 3-D structure in the upper mantle. In
 302 our calculations, S40RTS, SEMUCB-WM1, and TX2015 represent models of the 3-D shear velocity
 303 structure in the mantle. We do not include the crustal structure, adopt PREM as the reference structure for
 304 each of the three models, and assume the Voigt average shear-velocity variations in the anisotropic
 305 SEMUCB-WM1 model. The 220, 410 and 660 are horizontal boundaries at the same depths as in the PREM
 306 model. Models T2, T5, and T8 have the same layered velocity structure as PREM but the 410 and the 660
 307 are sinusoidal boundaries with amplitudes of 15 km and wavelengths of 2°, 5°, and 8°, respectively. The
 308 undulations of the 410 and 660 are in the opposite sense so the thickness of the mantle transition zone
 309 between the 410 and the 660 varies up to 30 km with respect to the average of 270 km.
 310

310



311
 312 **Figure 7.** (a) Maps of the shear-velocity variation at (top) 150 km and (bottom) 500 km depth according to S40RTS.
 313 The east-west contrast across the US is similar for SEMUCB-WM1 and TX2015. (b) Harmonic undulations of the
 314 410 and 660 for (from top to bottom) models T8, T5, and T2 with spatial wavelengths of 8° , 5° , and 2° . (c) Distribution
 315 of hypothetical earthquakes (stars) and stations (circles). For models PREM, S40RTS, T8 and T5 we compute
 316 waveforms for the twelve earthquakes indicated by red stars. For T2, we compute waveforms for these earthquakes
 317 and the additional 36 earthquakes indicated by black stars.

318

319 For each of the seven structures, we compute waveforms at periods longer than 10 s for 462 stations in a
 320 rectangular $2^\circ \times 2^\circ$ grid between longitudes 130° – 65° W and latitudes 25° – 50° N (Figure 7c). We calculate
 321 waveforms for 12 earthquakes uniformly distributed at a distance of 75° from $[-100^\circ\text{E}, 40^\circ\text{N}]$. We use 48
 322 earthquakes distributed in a spiral for structure T2. The uniform data coverage is sufficient to investigate
 323 the effects of velocity heterogeneity on Ssds-S traveltimes and the resolution of 410 and 660 undulations
 324 using long-period Ssds reflections. Because of the high computational cost, we cannot afford to reproduce
 325 the source-station combinations in the data and, therefore, we cannot estimate CRP mapping artefacts due
 326 to inhomogeneous slowness and azimuthal sampling.

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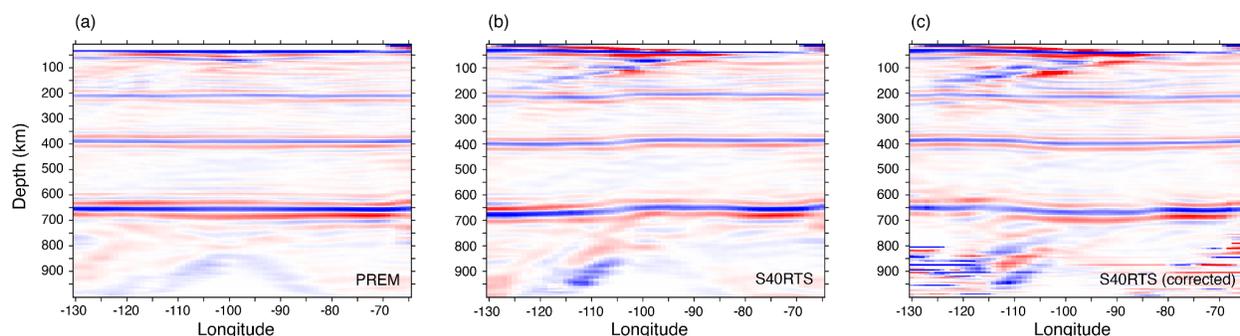
328 4.1 Testing ray-theoretical traveltimes corrections

329

330 Figure 8 shows the CRP images along the 35° N parallel in the central region of the model domain
 331 determined for the PREM and S40RTS models. Supplementary Figure 2 shows similar CRP images for
 332 SEMUCB-WM1 and TX2015. The CRP image for PREM in Figure 8a is the ideal case as the assumed
 333 velocity structure of the mantle is identical to the structure used to calculate traveltimes and ray paths.
 334 Artifacts are entirely due to the implementation of the CRP mapping procedure, the limited frequency band
 335 of the waveforms, and wave interference. PREM's velocity discontinuities at 220, 400, and 670 km depth
 336 are resolved about 10 km shallower in the mantle because the crust is not included in the waveform
 337 computations. Since the waveforms are computed for periods longer than 10 s and since shear wave speed
 338 increases with depth, reflectors at larger depths are more stretched than at shallower depths. The imaged
 339 660 is therefore about 60% stronger than the 410 even though the impedance contrast at the 660 is a factor
 340 of two stronger than at the 410. Overall, the CRP image derived from PREM waveforms is free of artificial
 341 layering between 150 km and 750 km depth. The side lobes of the 660 near -65° E are artifacts near the
 342 boundaries of the station grid. Layering near 100 km depth, which is especially strong near the center of
 343 the CRP image, shows that the Ssds reverberation is not an ideal wave type for imaging the uppermost

344 mantle when the analysis is based on shallow earthquakes. The arch-shaped structure below 750 km depth
 345 is likely the projection of shallow SS precursors misinterpreted as Ssds reflections deep in the transition
 346 zone as discussed in section 2.1.

347



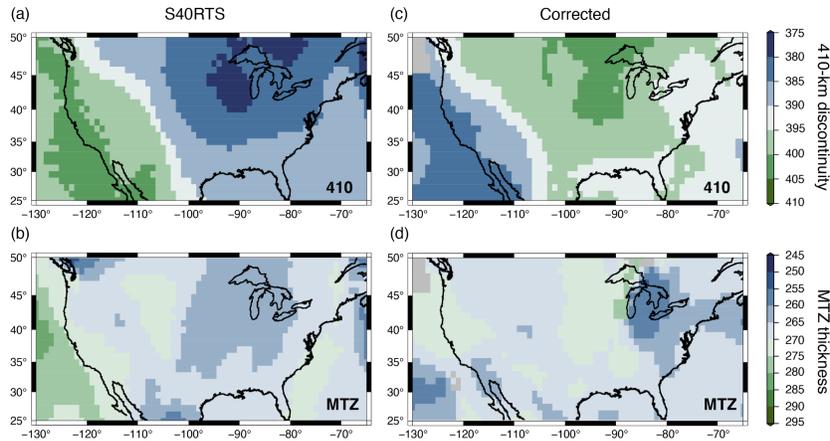
348
 349 **Figure 8.** CRP image along the 35° parallel determined for (a) PREM synthetics, (b) S40RTS synthetics, and (c)
 350 S40RTS synthetics after ray-theoretical corrections have been applied.

351
 352 The CRP image derived from S40RTS waveforms is more complex (**Figure 8b**). The 410 and 660 deepen
 353 from east to west because S40RTS predicts that Ssds travel times through the upper mantle are shorter
 354 beneath the central and eastern US than beneath the western US and we use the PREM velocity structure
 355 to convert traveltimes to reflector depths. The velocity heterogeneity in S40RTS causes misalignments of
 356 Ssds signals and therefore fluctuations in the strength of 410 and 660 from west to east by up to a factor of
 357 two. For example, the 660 appears as a relatively weak reflecting boundary between longitudes -120°E and
 358 -110°E, near the transition between the low-velocity upper mantle of the western US and the high-velocity
 359 upper mantle beneath the central US. In addition, spurious reflectors are particularly strong between -120°E
 360 and -100°E, where horizontal gradients in the uppermost mantle are strongest. It is difficult to identify how
 361 complex wave propagation produced the complexity in the CRP image but we note that the CRP image
 362 based on USArray waveforms is also most complex for the western US and a tilted reflective structure in
 363 the upper mantle has been observed by SB19 in their data image, albeit with an eastward dip and a greater
 364 depth extent.

365
 366 **Figure 8c** shows the CRP image based on the S40RTS synthetics after applying ray-theoretical
 367 traveltimes corrections following the procedure outlined in section 4.3. The traveltimes corrections do not remove,
 368 and may even amplify the CRP image artifacts for depths shallower than 100 km and deeper than 750 km. More
 369 significantly, the ray-theoretical calculations appear to overpredict the contribution of shear-velocity
 370 heterogeneity to the Ss410s-S and Ss660s-S difference times. After traveltimes correction, the 410 and 660
 371 are projected shallower beneath the western US than the central US, opposite to the imaged depths of the
 372 410 and 660 prior to corrections.

373
 374 The inaccuracy of ray theory in predicting the shear-wave traveltimes perturbations is illustrated further in
 375 **Figure 9**. It shows the estimated depths of the 410 and the 660 and the thickness of the MTZ based on the
 376 1-D CRP method applied to synthetic waveforms computed for S40RTS. **Supplementary Figure 3** shows
 377 that we obtain similar results for SEMUCB-WM1 and TX2015. The total variation in the depths of the 410
 378 and 660 is about 15–20 km. As expected, the depth of the 410 (**Figure 9a**) mimics the shear-velocity
 379 variations in the upper mantle of S40RTS (**Figure 7a**) and the S-wave traveltimes delay map shown in **Figure**
 380 **5**. The depth variation of the 660 (**Figure 9b**) is slightly different because shear velocity variations in the

381 MTZ also influence the traveltimes of Ss660s. Variations in the thickness of the MTZ (Figure 9c) of about
 382 10 km are small compared to the depth variations of the 410 and 660 because shear velocity variations in
 383 the MTZ are much weaker than in the uppermost mantle.
 384

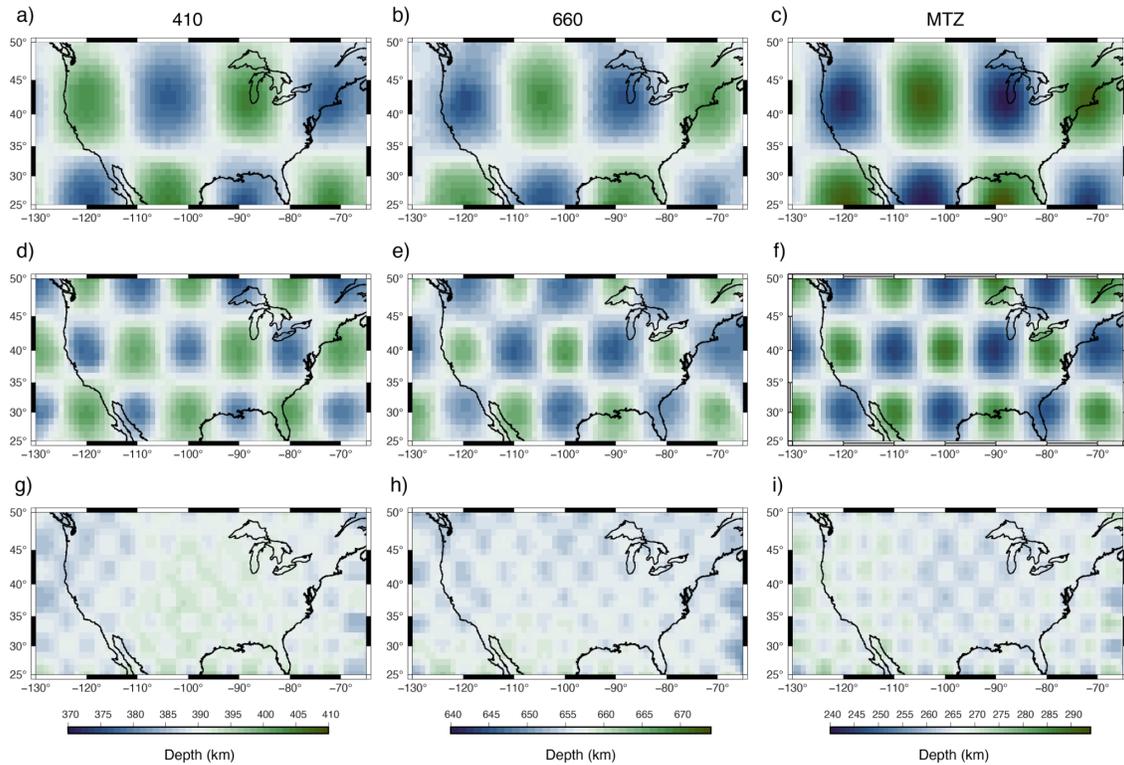


385
 386 **Figure 9.** Depths of the 410 (top row) and the thickness of the transition zone (bottom row). Panels (a) and (b) are
 387 estimated from spectral-element-method seismograms calculated for model S40RTS. Panels (c) and (d) show the same
 388 estimates after ray-theoretical time shifts have been applied to the waveforms.
 389

390 If ray-theoretical traveltimes are precise, we must expect that the CRP images of the ray-
 391 theoretical corrected S40RTS, SEMUCB-WM1 and TX2015 waveforms are similar to the CRP image for
 392 the PREM model because the 410 and 660 are horizontal boundaries in all models. However, we find this
 393 not to be the case. While the elevation of the 410 and 660 beneath the western US (by 10 and 11 km,
 394 respectively) and their depressions beneath the central-eastern US (by 11 and 12 km, respectively) have the
 395 expected trends, the corrections are larger than expected ray-theoretically. In the corrected image, the 410
 396 and 660 are shallower in the western US than in the eastern US (Figure 9d and 9e) opposite to the
 397 uncorrected CRP image (Figure 9g and 9h). The corrections are least accurate for the central-eastern US.
 398 Here, the depth correction of the 410 is 10 km but expected to be 5 km. In the western US the inferred and
 399 predicted depth correction differ by a factor of 1.7.
 400

401 4.2 The resolution of 410 and 660 undulations

402
 403 Figure 10 shows the depths of the 410 and the 660 and the thickness of the MTZ resolved for models T2,
 404 T5, and T8. The checkerboard pattern of the undulations on the 410 and 660 are resolved for T5 and T8 but
 405 the amplitude of the undulations is underestimated. The resolved thickness of the MTZ varies, on average,
 406 12 km and 6 km less than in the original T8 and T5 models. The resolution of the undulations in T2 is poor
 407 despite using a larger set of waveforms for 48 earthquakes. From experiments we have found that the
 408 resolution does not improve if we densify the grid of stations to a 1-degree spacing (not shown in Figure
 409 10). Therefore, fluctuations of the depth of the 410 or 660 with a wavelength of about 200 km are
 410 intrinsically unresolvable from long-period Ssds waveforms because the Fresnel zone of Ss410s and Ss660s
 411 in the upper mantle at the dominant frequency of about 0.05 Hz are about 500 km, much wider than the
 412 undulations of the 410 and 660 in T2.
 413



414
 415 **Figure 10.** Depth maps of 410, 660, and MTZ thickness using the depth stack method for 8×8 (a, b, c), 5×5 (d, e,
 416 f), and 2×2 (g, h, i) input topography models.

417
 418

419 5. Discussion and conclusions

420

421 The receiver-side S wave reverberation, denoted as Ssds, is a useful data type to map the shear velocity
 422 structure in the upper mantle, including undulations of the 410 and 660 mineral phase transitions. Ssds
 423 complement SS precursor and P-to-S wave conversion (i.e., receiver function) imaging of the mantle
 424 because of its unique wave path geometry. In agreement with the analysis by SB19, we observe in record
 425 sections of waveform stacks that the Ss410s-S and Ss660s traveltime differences vary by up to 10 s across
 426 stations from the USArray. If the traveltime differences are attributed entirely to 410 and 660 undulations,
 427 it implies that the 410 and 660 are 40–50 km deeper beneath the western US than the central and eastern
 428 US. In turn, this would mean that the geological contrast between the tectonically active western US and
 429 the stable central and eastern US persists as a temperature or compositional contrast in the mantle
 430 zone and thus a link between uppermost mantle and mantle transition zone dynamics.

431

432 However, the correlation between the resolved depth of the 410 (and the 660) and tomographic maps of the
 433 shear-velocity structure in the upper mantle is high. This indicates that velocity heterogeneity in the
 434 uppermost mantle contributes significantly to the Ss410s-S and Ss660s-S traveltimes and the spatial
 435 variations of the 410 depth inferred from common-reflection-point (CRP) imaging. Ray-theoretical
 436 corrections of traveltimes for velocity heterogeneity by shifting segments of the waveforms containing
 437 Ss410s and Ss660s prior to CRP stacking reduces the variation in the 410 depth by a factor of two.

438

439 For at least two reasons we find ray-theoretical corrections imprecise. First, seismic tomography has
440 uncertainty. Global models S40RTS, SEMUCB-WM1, and TX2015 agree on the east–west contrast but
441 disagree on the magnitude of the traveltime perturbations (see [Figure 4](#)). Each model underestimates the S-
442 wave traveltime delay at USArray stations (see [Figure 5](#)) which is consistent with the fact that tomographic
443 models underestimate the magnitude of traveltime and waveform perturbations. Hence, the effect on the
444 estimated depths of the 410 and 660 depends on the chosen tomographic model. SB19 note that the
445 traveltime corrections may introduce incoherence in the CRP images and use that as a factor in determining
446 the value of traveltime corrections.

447
448 Second, our experiments with spectral-element method synthetics demonstrate that ray-theoretical
449 predictions of the Ss410s-S and Ss660s traveltime differences are inaccurate. CRP images derived from
450 waveforms computed for a mantle with 3-D velocity heterogeneity and horizontal phase boundaries show
451 the expected deepening of the 410 and 660 below the western US and shallowing beneath the central and
452 eastern US where the shear velocities are relatively low and high, respectively. After applying traveltime
453 corrections for the 3-D wave speed structure, the 410 and 660 remain undulating boundaries. In fact, the
454 410 and the 660 in the corrected CRP image are deeper beneath the central-eastern US than beneath the
455 western US, opposite to the uncorrected CRP image. This indicates that ray theory overpredicts the Ssds-S
456 difference time by about a factor of two. This is the case for S40RTS, SEMUCB-WM1 and TX2015 and
457 presumably also finer-scale regional tomographic models when finite-frequency effects are stronger. The
458 inaccuracy of ray-theoretical predictions of the traveltime perturbations of long-period waves has been
459 studied previously. For example, [Neele et al. \(1997\)](#) and [Zhao and Chevrot \(2003\)](#) have pointed out that the
460 broad SS sensitivity kernels at the reflection points on the surface or the mantle discontinuities. [Bai et al.](#)
461 [\(2011\)](#) and [Koroni and Trampert \(2013\)](#) illustrate how the finite wave effects affect CRP images built from
462 SS precursors similarly to the study here.

463
464 Finally, we note that the resolution of the depths of the 410 and 660 depends on spatial scales of the
465 undulations. Our experiments with spectral element method synthetics indicate that the Ssds-S traveltime
466 difference is sensitive to $5^\circ \times 5^\circ$ and $8^\circ \times 8^\circ$ sinusoidal variations of the 410 and 660 depths albeit that the
467 height of the undulations is underestimated. Spatial variations of the 410 and 660 on a $2^\circ \times 2^\circ$ scale are not
468 resolvable because such variations are smaller than the width of the Fresnel zone of Ssds at a period of 10
469 s.

470
471 Although it is beyond the scope of this work, it is better to simultaneously estimate 410 and 660 topography
472 and shear velocity heterogeneity in the mantle of multiple data sets (e.g., [Gu et al., 2003](#), [Moulik and](#)
473 [Ekström, 2014](#)) using finite-frequency kernels that relate waveform perturbations to velocity heterogeneity
474 and phase boundary topography (e.g., [Guo and Zhou, 2020](#)) or, preferably, using an adjoint tomography
475 approach ([Koroni and Trampert, 2021](#)). Based on our experiments, the evidence for large-scale variations
476 of the depth of the 410 beneath the USArray is weak. As is well established, estimates of the thickness of
477 the MTZ are not affected strongly by shear velocity heterogeneity. We find the thickness of the MTZ to
478 vary by about 10 km, which is consistent with the receiver-function study of USArray data by [Gao & Liu](#)
479 [\(2014\)](#) and much smaller than global variations of the MTZ observed in SS precursors studies (e.g.,
480 [Flanagan & Shearer, 1998](#); [Chambers et al., 2005](#)).

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483 Acknowledgments

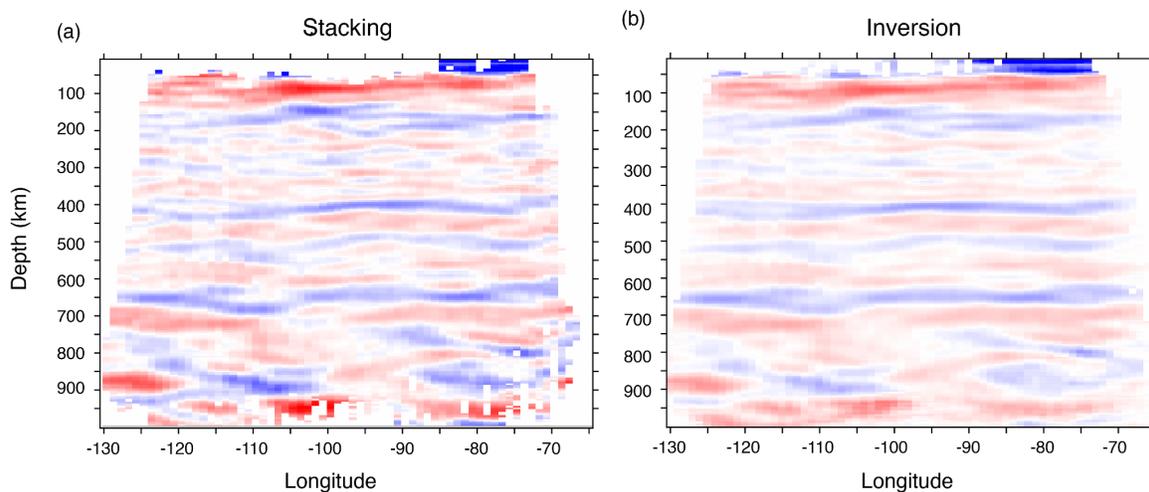
484
 485 This research was supported by the NSF (EAR-1644829). The facilities of IRIS Data Services, and
 486 specifically the IRIS Data Management Center, were used for access to waveforms, related metadata, and/or
 487 derived products used in this study. IRIS Data Services are funded through the Seismological Facilities for
 488 the Advancement of Geoscience (SAGE) Award of the National Science Foundation under Cooperative
 489 Support Agreement EAR-1851048. Data from the TA network were made freely available as part of the
 490 EarthScope USArray facility, operated by Incorporated Research Institutions for Seismology (IRIS) and
 491 supported by the National Science Foundation, under Cooperative Agreements EAR-1261681. The
 492 SPECMFEM3D_Globe software was downloaded from the Computational Infrastructure for Geodynamics
 493 (<https://geodynamics.org/>).

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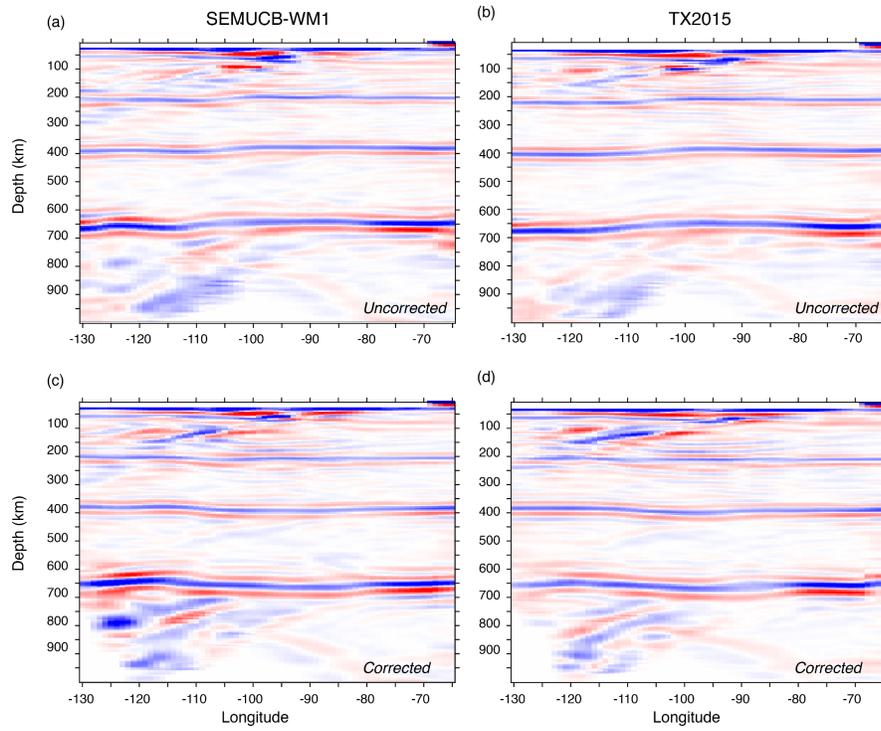
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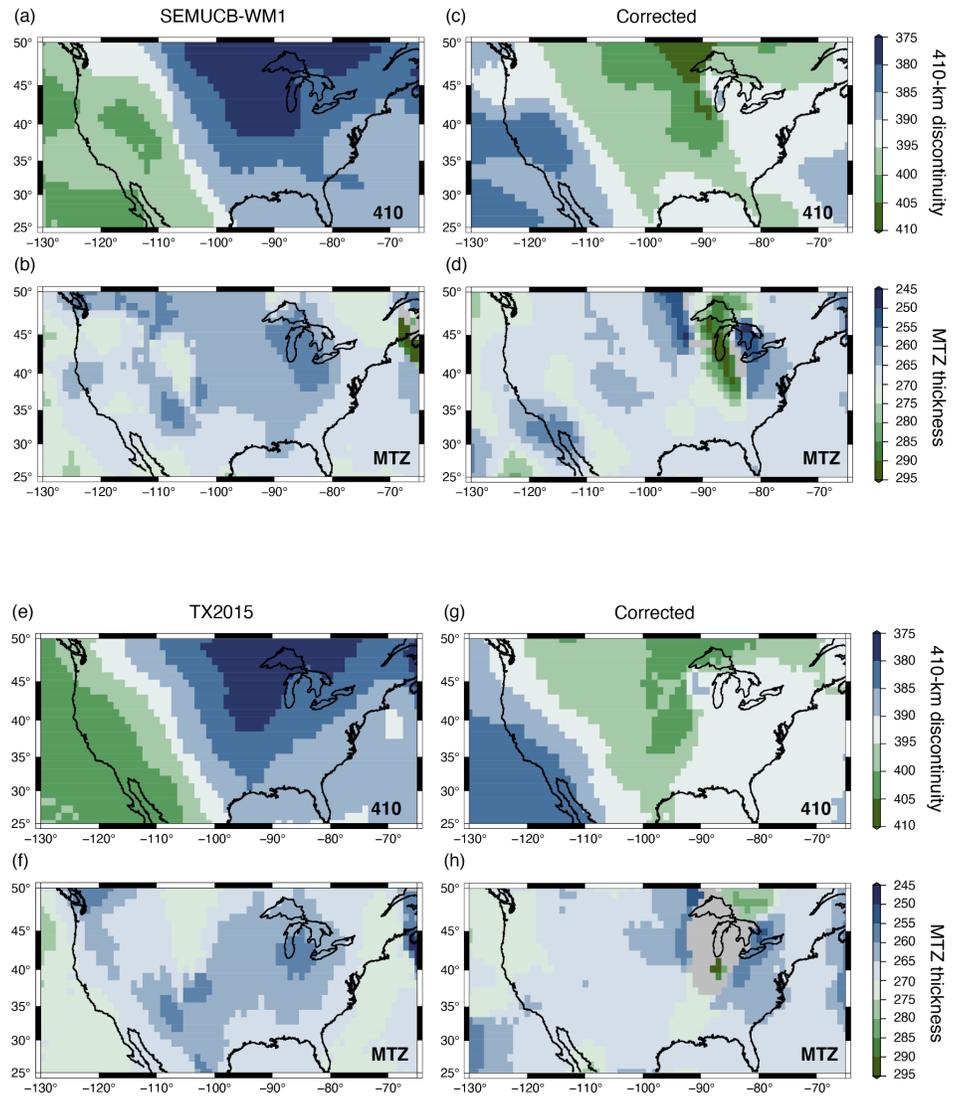
Supplementary Figure 1. Vertical section of the CRP image along 40°N determined by our stacking method (a) and SB19's inversion method (b). See also Figure 4. Because the CRP imaging is based on 1-D wave propagation, it is difficult to estimate the amplitude of Ssds produced by local reflecting boundaries and undulating global discontinuities. If a reflecting boundary exists only beneath the source, the source-side contribution to Ssds is underestimated because half of the amplitude of Ssds is attributed to a reflection on the receiver side. On the other hand, the source-side reflection is overestimated if a reflection boundary exists only beneath the USArray. We expect therefore that the impedance contrasts of reflecting boundaries are uncertain despite our large set of amplitudes from earthquakes at all azimuths from the USArray. Our implementation of the inversion approach results in a misfit reduction smaller than 10%, underscoring the difficulty to separate source-side and receiver-side contributions to Ssds waveforms and that the impedance contrasts are uncertain. The synthetic tests by SB19 also illustrate this.



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Supplementary Figure 2. CRP image along the 35° parallel determined for (a and c) SEMUCB-WM1 and (b and d) TX2015 synthetics without (in a and b) and with (in c and d) ray-theoretical corrections.

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Supplementary Figure 3. Depths of the 410 and the thickness of the MTZ obtained by CRP imaging spectral-element-method seismograms computed for model SEMUCB-WM1 (a–d) and TX2015 (e–h) with and without ray-theoretical traveltimes corrections. Compare to Figure 9.