Characterization of internal tide non-stationarity : Eulerian versus Lagrangian perspectives

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Abstract

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ABSTRACT

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1. Introduction

The disentangling of internal tides and balanced flow is a key issue for the Surface 26 Water and Ocean Topography (SWOT) mission (Morrow et al. 2019). This wide-swath 27 altimeter mission will provide instantaneous 2D sea level maps, with an expected horizontal 28 resolution of the order of 15-30 km (Morrow et al. 2019; Fu et al. 2012). With this 29 resolution, internal tides and mesoscale balanced flow will be captured, providing a unique 30 opportunity to study both motions and their interaction. While both motions have distinct 31 time scales, they can have similar length scales (order of hundreds of kilometers) which 32 makes their separation by spatial filtering difficult. SWOT coarse temporal resolution 33 (20 day repeat time approximately) will also prevent a separation by temporal filtering. Due to these two issues, internal tides and balanced flow will be entangled in SWOT data, 35 which is problematic for estimates of surface currents via geostrophy. 36

Internal tides (or baroclinic tides) are internal waves generated by the barotropic tide 37 when it passes over a topography (Garrett and Kunze 2007). Internal tides that are phase-38 locked with the tidal forcing are referred to as stationary internal tides. Stationary internal 39 tides are predictable as they are phase-locked and of known amplitude. A fraction of the 40 internal tides energy (mainly high modes) dissipates close to their generation's location (Whalen et al. 2020) but a significant part travels in the open ocean over potentially 42 great distances - up to thousands of kilometers - with a low-mode vertical structure. As 43 they travel, internal tides encounter variations of the background stratification as well as 44 obstacles (continental shelf, islands, etc) which can cause energy dissipation or scattering 45 toward higher modes (Whalen et al. 2020; Savva and Vanneste 2018; Savage et al. 2020). Internal tides can also become nonstationary. In particular, loss of stationarity has been 47 observed when internal tides travel in a background stratification that varies in time (Buijsman et al. 2017), or pass through a turbulent jet (Ponte and Klein 2015; Dunphy 49

et al. 2017; Savage et al. 2020). They also become nonstationary when they reach the continental shelf (Nash et al. 2012b,a).

Several works used altimeter observations to study baroclinic tide including its non-52 stationary component (Ray and Zaron 2016; Zaron 2017, 2019; Nelson et al. 2019). To 53 overcome limitations of altimeter data, the use of the global drifter program (GDP) dataset has been investigated in addition to the altimeter data (Zaron 2017, 2019; Nelson et al. 55 2019). The GDP complete dataset, combining Argos and GPS localisation system, is available interpolated over a temporal grid of 1h (Elipot et al. 2016). Lately, the number 57 of drifters localised via GPS has increased, with only GPS tracked drifters recently de-58 ployed, thus decreasing the localisation error in the data (Elipot et al. 2016). This dataset 59 has already been used to investigate high frequency near-surface near-inertial oscillations 60 (NIOs) (Elipot et al. 2010). 61

There are however challenges and questions raised by the use of Lagrangian data. One 62 of the challenges associated with the analysis and interpretation of Lagrangian data is that 63 data collected by a drifter as it moves along with the flow may entangle Eulerian spatial 64 and temporal variability and, give a deformed perspective of the Eulerian field. (LaCasce 65 2008) reviewed conceptual frameworks that have been developed in order to tackle this issue (Lumpkin et al. 2002; Middleton 1985; Davis 1983, 1985). Two regimes are typically 67 identified : fixed float and frozen turbulence. The prevalence of one regime over the other is determined by the parameter $\alpha = T_E/T_a$, where T_E is the Eulerian evolution timescale 69 of the flow and T_a is the time required for a drifter to travel the Eulerian characteristic 70 spatial scale of the observed fluctuation. T_a is given by L/U, with U corresponding to an 71 advection velocity and L the fluctuation's spatial scale. If $\alpha \ll 1$, the time required for 72 the drifter to travel the length L is greater than the timescale of the fluctuation, T_E . In 73 this case, one can expect an agreement between the Lagrangian and Eulerian timescales. 74 Conversely, if $\alpha \gg 1$, it takes a drifter a time smaller than T_E to travel a distance L, causing 75

a more rapid fluctuation in the Lagrangian perspective. In this paper we apply these ideas
to fit the case of internal tides interacting with a turbulent jet.

(Zaron and Elipot 2020) found a spectral broadening of the barotropic tide peaks in 78 Lagrangian data compared to the Eulerian one, due to flow and/or tides spatial non-79 stationarity. It is therefore not trivial that Lagrangian drifter data may be used to study 80 internal tides and their loss of stationarity depending on the regions of the ocean and the 81 associated dynamical regime. It is also necessary to evaluate the use of Lagrangian data in the context of SWOT. Can Lagrangian data be used to filter and identify internal tides? 83 In this study, we compare the non-stationarity timescales and amplitudes of an internal 84 tide field crossing a turbulent jet, in Eulerian and Lagrangian frameworks. This allows us 85 to test analytical models describing balanced flow and internal tides in an idealized set-up 86 and the relationship of Eulerian and Lagrangian timescales.

We first present the numerical set-up used in this study as well as the method and model used to estimate signal amplitudes and decorrelation timescales. The results are shown in the second part for one simulation at first and several simulations varying the jet's strength later, . Lastly, we develop a theoretical model to predict Lagrangian autocovariance from Eulerian one and confront our results to it.

2. Numerical simulations and Lagrangian data

⁹⁴ a. Numerical simulations

Idealized numerical simulations of an internal tide crossing a turbulent jet are considered. The numerical model is the Coastal and Regional Ocean COmunity model (CROCO) solving the hydrostatic equations. Its configuration follows Ponte et al. (2017) with a rectangular numerical domain (1024 km x 3072 km), zonally periodic. The Coriolis frequency follows the beta-plane approximation and is defined so that the domain is representative of mid-latitudes. A turbulent zonal jet crosses the domain at its center

along the meridional direction. It results from initialising the numerical simulation with a 101 baroclinically unstable jet and is maintained by relaxing zonally averaged fields (velocities, 102 temperature, sea level) toward the initial jet. Simulations with different turbulent jet 103 strength are obtained by modulating the strength of the initial jet or equivalently the 104 latitudinal thermal gradient. After 500 days, relaxation of the zonal mean fields toward 105 the initial jet is ceased. The jet has a mean velocity amplitude maximum around 1450km 106 in the center of the jet (Fig.1a, red line). The jet's amplitude decays over the observed period of time with a maximum around 0.6 m/s at the beginning and around 0.4 m/s at 108 the end. The jet's velocity is computed by averaging each velocity component (u and v) 109 over 2 days. 110

A mode-1 internal wave is generated at y = 400 km with a semi-diurnal frequency (2 cpd). Its signature at the surface contributes significantly to the total velocity amplitude in the northern and southern areas (Fig.1a, green line compared to red line). Sponge layers are put on the top and the bottom of the domain (y < 300 km and y > 2700 km). About 8000 simulated near surface drifters (referred to as drifters in the rest of this study) are also initialised at day 500 on a regular grid extending from 600 km to 2400 km and their advection is performed online (Fig.1b).

118 b. Lagrangian outputs overview

In the central part of the domain, the jet turbulence dominates the drifter net motion with a displacement of about 300 km in the *x*-direction and 160 km in the *y*-direction, i.e. an internal tide wavelength, over a 40 day time window (Fig.2c). When averaged over all drifters in this area, the net displacement over such time window is about 200 km. Away from the jet (Fig.2a and e), the net distance travelled by selected drifters, in the y-direction, is of about 20–30 km which is a fraction of an internal tide wavelength. Internal tides produce on the other hand periodical displacements of the order of 2–3 km. Eulerian and Lagrangian zonal velocity time series exhibit significant differences in the jet at both low and internal tide frequency (modulation of the envelope and phase) over the 40 day temporal window (Fig2d). To quantify these differences, we will estimate amplitude and decorrelation/nonstationary time scales associated with the jet and internal tides and compare the results in different parts of the domain.

¹³¹ Zonal velocity time series outside the jet (Fig2b and f) exhibit weaker differences ¹³² between both perspectives, as seen from their envelopes. Modulations of internal tide ¹³³ fluctuations in the south and north parts of the domain differ markedly from each other ¹³⁴ with faster fluctuations of their envelope in the north compared to the south. This feature ¹³⁵ occurs in both Eulerian and Lagrangian time series and reflects the loss of stationarity of ¹³⁶ the internal tide as it propagates northward and interacts with the turbulent jet.

¹³⁷ c. Methods

138 1) ESTIMATION OF EULERIAN AND LAGRANGIAN AMPLITUDES AND TIMESCALES

¹³⁹ We estimate the amplitudes and decorrelation timescales of slow and fast motions by ¹⁴⁰ computing the velocity autocovariance in Eulerian and Lagrangian model outputs and ¹⁴¹ fitting it to a theoretical model for a superposition of two motions (fast and slow) keeping ¹⁴² timescales and amplitudes as variables. We assume a velocity component *v* may be written ¹⁴³ as the sum of an internal tide part \tilde{v} and a turbulent (or jet) part \bar{v} :

$$v = \widetilde{v} + \overline{v} \tag{1}$$

Assuming internal tide velocities and jet velocities are uncorrelated, the total autocovariance, C, equals to the sum of the autocovariances of \tilde{v} and \bar{v} :

$$C(\tau) = \langle v(t)v(t+\tau) \rangle = \overline{C}(\tau) + \overline{C}(\tau), \qquad (2)$$

where $\langle \cdot \rangle$ is a time averaging operator.

¹⁴⁷ The high frequency signal for mode-1 internal tides is written as

$$\widetilde{v} = \Re \left[\widetilde{v}_e(t) e^{i\omega t} \right]$$
 with \Re the real part (3)

where \tilde{v}_e is the envelop of the oscillations of the tides and depends on the time, thus capturing the non-stationarity of the waves. $\omega/2\pi$ is the frequency of the waves, 2 cycles per day.

The internal tide signal can be decomposed into stationary and nonstationary contributions. The stationary part is defined with a coherent temporal averaging operator :

$$\widetilde{\nu}_s = \langle \widetilde{\nu} \rangle_c, \tag{4}$$

$$= \Re \left[\langle \tilde{v}_e \rangle_c e^{i\omega t} \right] \tag{5}$$

¹⁵³ Hence the nonstationary part, defined as the total velocity minus the stationary part :

$$\widetilde{\nu}_{ns} = \widetilde{\nu} - \langle \widetilde{\nu} \rangle, \tag{6}$$

$$= \Re \left[(\tilde{v}_e - \langle \tilde{v}_e \rangle_c) e^{i\omega t} \right]$$
⁽⁷⁾

Assuming the envelope of the nonstationary signal is exponential with a decay timescale \widetilde{T} , the fast autocovariance is expressed as:

$$\widetilde{C}(\tau) = \left[\widetilde{V}_s^2 + \widetilde{V}_{ns}^2 e^{-\tau/\widetilde{T}}\right] \times \cos(\omega\tau)$$
(8)

¹⁵⁶ Following Veneziani et al. (2004) the turbulent velocity autocovariance is assumed to ¹⁵⁷ have the form :

$$\overline{C}(\tau) = \overline{V}^2 e^{-\tau/\overline{T}} \cos(\Omega \tau) \tag{9}$$

- where \overline{T} is the decorrelation timescale and Ω accounts for eddies and meanders.
- ¹⁵⁹ The total autocovariance is then given by:

$$C(\tau) = \widetilde{C}(\tau) + \overline{C}(\tau) = \left[\widetilde{V}_s^2 + \widetilde{V}_{ns}^2 e^{-\tau/\widetilde{T}}\right] \times \cos(\omega\tau) + \overline{V}^2 e^{-\tau/\overline{T}} \cos(\Omega\tau)$$
(10)

This model is then fitted with the autocovariance obtained from our data and averaged in bins in the y-direction. The variables \widetilde{T} , \widetilde{V}_s , \widetilde{V}_{ns} , \overline{T} , \overline{V} and Ω are estimated to find the best fit. The fit is done using a non linear least square regression (Jones et al. 2001–). Lower bounds are fixed to zero for all parameters except timescales, for which they are fixed at one day. The weight is uniform with no dependence on the timelag.

Because of drifters displacements across the domain, the time window used for the 165 computation of Lagrangian autocovariances has to be short enough for the result to be 166 typical of a specific area, while being long enough to capture the long decorrelation 167 timescale. For each drifter's trajectory the velocity time series is split into segments of length T_w , overlapping each other by 50%. A time window of 40 day is chosen. Eulerian 169 mean velocities, averaged in time and zonal direction is interpolated on drifters trajectories 170 and removed. No significant impacts of this removal were observed on the results for the 171 fast signal. The autocovariance is computed over each segment and averaged within 50 km 172 wide meridional bins. Each segment is attributed to a bin depending on the mean position 173 over the period T. We did not find a significant sensitivity of our results to the length 174 of the window. The Eulerian autocovariance is computed at each grid point using the 175 same time windows and bin-averaged meridionally as for the Lagrangian autocovariance. 176 Autocovariances are then divided by the autocovariance at timelag zero to obtain the autocorrelation. 178

3. Signatures of internal tides and turbulent jet in Eulerian and Lagrangian perspec-

180 tive

¹⁸¹ a. Autocorrelation functions

Three main regimes stand out on Lagrangian and Eulerian 2D autocorrelation (function of timelag and y) (Fig.3 first and second column respectively), corresponding to typical drifters' trajectories shown in Fig2a, c and e. Autocorrelation at these latitudes of interest are further shown in Fig. 4. Fast oscillations, corresponding to the internal tides, are seen in the northern and southern parts of the domain for all autocorrelation functions. In these

areas, the signal seems to be dominated by internal tides, with a negligible influence of 187 the slow advection of the drifters by the balanced flow. No decay of the amplitude of the 188 oscillations with the time lag can be visually detected south of the domain, especially in 189 the Eulerian perspective (see Fig. 4, panels e and f) indicating that internal tides are nearly 190 stationary there. On the contrary, the amplitude of the fast oscillations decays mildly 191 in the northern area, indicating internal tides non-stationarity. There are no significant 192 visual differences between Lagrangian and Eulerian autocorrelations in the northern and southern areas. Conversely, the central area ($y \in [1000; 2000]$ km) exhibits a fast decay – 194 especially in the Lagrangian perspective – of fast oscillations combined to a slower general 195 decay associated with the slower jet turbulence. As observed in drifters trajectories and 196 velocity time series (Fig. 2, panels c and d), Lagrangian diagnostics diverge from Eulerian 197 ones in this area. The damping of fast oscillations in the central area occurs over a much shorter timescale in Lagrangian autocorrelation compared to Eulerian one : fast 199 oscillations disappear after only a few days ($\tau \leq 5$ days) in Lagrangian perspective while 200 they are still visible in the Eulerian one for the largest τ (20 days). The decorrelation of the 201 slower motion is also faster in Lagrangian autocorrelation compared to Eulerian one, and 202 exhibits a negative lobe around $\tau \sim 4$ days in the Lagrangian autocorrelation of v, which 203 we attribute to the slow evolution of jet meanders. We also note that the autocorrelation 204 of u does not decrease toward zero in the jet. Although this aspect is not the focus of our 205 study, we attribute this to the fact that drifters oversample the meandering jet, resulting in 206 a larger mean for the zonal velocity compared to the Eulerian mean. The fact that Eulerian 207 autocorrelation decays more slowly than Lagrangian one is consistent with previous studies 208 comparing Eulerian and Lagrangian decorrelation timescales (Lumpkin et al. 2002). 209

To estimate Eulerian and Lagrangian decorrelation timescales and velocity amplitudes for both slow and fast motions, we fit the autocovariance from our data (Eulerian and Lagrangian) to the model given by eq. (10).

Eulerian and Lagrangian diagnostics (blue and red lines Fig. 5) show a loss of the 214 stationarity of internal tides during the crossing of the jet. Eulerian stationary amplitude 215 (Fig. 5c) exhibits high values relatively to the non-stationary one (Fig. 5d) ($\sim 0.06 \,\mathrm{m\,s^{-1}}$ 216 versus $\sim 0.01 \text{ m s}^{-1}$) in the south with Eulerian envelope (Fig. 4 e and f) near constant indi-217 cating internal tides nearly stationary in this area. During the crossing of the jet, internal 218 tides loss of stationarity is captured in Eulerian diagnostics as the nonstationary amplitude 219 becomes larger than the stationary one, increasing up to $\sim 0.08 \,\mathrm{m\,s^{-1}}$ in the northern part 220 (y larger than ~ 1700 km), while the stationary part decreases slightly (~ 0.05 ms⁻¹). A 22 bump is observed in nonstationary amplitude due to the combined effects of a bump in 222 stratification in the jet, Coriolis and the loss of stationarity. In this dominantly nonstation-223 ary area, nonstationarity timescales are between 10 and 20 days. We note that the Eulerian 224 envelope in the north (blue lines Fig. 4a and b) does not seem to reach zero but a plateau, 225 consistent with a remaining stationary component. 226

Lagrangian diagnostics present patterns significantly different from Eulerian ones as 227 expected from drifters trajectories (Fig. 2 a, c and e) and autocorrelations (Fig. 3). In 228 the south, the Lagrangian envelope (red lines Fig. 4 e and f) decays faster than Eulerian 229 one. Lagrangian nonstationary and stationary amplitudes (red lines Fig. 5c and d) present 230 similar values ($\sim 0.04 \,\mathrm{m\,s^{-1}}$ for the stationary part and $\sim 0.03 \,\mathrm{m\,s^{-1}}$ for the nonstationary 23 part). Nonstationary timescales (Fig. 5a) remain between 10 and 20 days. In the jet area, 232 the nonstationary amplitude largely dominates as it reaches values around $\sim 0.08 \, \mathrm{m \, s^{-1}}$ 233 while the stationary part decreases toward values smaller than $\sim 0.01 \text{ m s}^{-1}$. Nonstationary 234 timescale decreases rapidly in this region and reaches 1 day in its center. This apparent non-235

stationarity in Lagrangian perspective is likely due to a dominant slow motion strongly
advecting the drifters in this area. In the north, we observe a decay of Lagrangian
autocorrelation envelope close to the Eulerian one. Stationary and nonstationary amplitude
are similar to Eulerian ones with non-stationarity dominant. Timescales also have similar
values, between 10 and 20 days. The non-stationarity in this area is captured similarly in
both Lagrangian and Eulerian diagnostics.

The slow motion decorrelation timescales (Fig. 5b) reach their lowest values in the 242 central area, ~ 20 day in Eulerian data and ~ 10 day in Lagrangian outputs. It corresponds 243 to the area of high amplitude (Fig. 5e). It also coincides with the area of low Lagrangian 244 non-stationarity timescales which supports an apparent non-stationarity in Lagrangian 245 diagnostics dominant in this part. As seen in Lagrangian autocorrelation (Fig. 3), oscilla-246 tions due to slow motion meanders are observed in the autocorrelation of v but well fitted 247 by our model (Veneziani et al. 2004). The fit for the autocorrelation of u seems however to 248 fail in the center of the jet with a decorrelation timescale of the slow motion overestimated 249 in that bin as a tendency towards a value superior to zero is not expected by our model for 250 slow motion. 251

The sensitivity of internal tide nonstationarity timescales to the jet EKE is investigated with five numerical simulations of increasing jet's strength as shown by the meridional distributions of velocity amplitude (Figure 6b).

The jet strength, as measured by the velocity amplitude maximum value, varies by a factor of about 2 across five simulations (Fig.6b). The internal tides' total velocity amplitude, defined by $\sqrt{\tilde{V}_s^2 + \tilde{V}_{ns}^2}$, increases in the northern area with the jet's strength (Fig.6e). It increases with latitude similarly in both Eulerian and Lagrangian perspectives, regardless of the region of the domain or the jet's strength.

For the 2 most energetic simulations, S₃ and S₄, both Eulerian and Lagrangian diag-26 nostics show a loss of stationarity of internal tides occurs when internal tides cross the 262 jet. As in the previously studied simulation, the internal tides are nearly stationary in 263 the southern area in Eulerian amplitudes (continuous lines in Fig. 6c and d) while La-264 grangian stationary and nonstationary ones (circular markers in Fig. 6c and d) remain of 265 similar order. In the jet, Eulerian and Lagrangian nonstationary amplitudes increase while 266 Lagrangian stationary amplitude drops to zero. Lagrangian timescales (Fig. 6a) reach 267 minimal values (\leq 5days) while Eulerian ones remain around or above 10 days in every 268 simulation. The width of this area increases with the jet's strength. As the drifters move 269 along with the flow they might capture spatial variations as internal tides' non-stationarity 270 This may explain Lagrangian timescales shorter than Eulerian ones in the area where the 271 slow motion dominates. In the northern area, the same pattern is found for the two cases : 272 non-stationarity amplitudes and timescales are similar in both framework, the Lagrangian 273 non-stationarity is not/weakly biased by the drifters' perspective. Interestingly, the inter-274 mediate case, S_2 exhibits features similar to S-3 and S_4 in the south and jet but distinct 275 stationary amplitude and nonstationary timescale. The autocorrelation does not reach a 276 plateau (stationary amplitude) in the chosen time window. This case is then considered 277

completely nonstationary in the north by our method, with small stationary amplitude and 278 overestimated timescale. Contrary to previous cases, the two least energetic simulations, 279 S_0 and S_1 , show weak loss of stationarity in Eulerian perspective as the stationary am-280 plitude remains dominant for all bins and the nonstationary timescales near the highest 281 values allowed in our fitting procedure. Lagrangian stationary amplitude drops to zero in 282 the jet as the nonstationary one exhibits a bump in the same area and the timescale drop to 283 1 day supporting Lagrangian apparent non-stationarity due to advection even for weakly 284 energetic simulations. 285

4. Lagrangian model for autocovariance and comparison to fitted autocovariance

²⁸⁷ a. Theoretical expectation for the fast Lagrangian autocorrelation

We assume that the fast signal is a modulated monochromatic wave propagating in a single direction (say *x*) and characterized by a frequency ω and wavenumber *k*:

$$\tilde{v}(t) = \Re\left\{\tilde{v}_e(x,t)e^{i(\omega t - kx)}\right\},\tag{11}$$

where \tilde{v}_e is the slowly varying envelope. Let's consider a parcel traveling with the flow with trajectory X(t). The autocovariance of \tilde{v} as measured along the parcel trajectory is given by:

$$\tilde{C}_L(\tau) = \langle \tilde{v}(t+\tau)\tilde{v}(t) \rangle, \tag{12}$$

$$= \frac{1}{2} \Re \left\{ \left\langle \widetilde{v}_e \left[X(t+\tau), t+\tau \right] \widetilde{v}_e^* \left[X(t), t \right] e^{i \left[\omega \tau - k(X(t+\tau) - X(t)) \right]} \right\rangle \right\},\tag{13}$$

where we assume that fast oscillation terms ($\propto e^{\pm 2i\omega t}$) are smoothed out by the averaging procedure. Neglecting the contribution of the wave velocity to the displacement: $X(t) = \int^{t} \bar{v}(s) ds$, as well as the spatial dependency of the wave envelope (this point is discussed in the discussion, sect. b, below), and further assuming that the envelope of the wave and the slow flow are not correlated, we obtain:

$$\tilde{C}_L(\tau) = \Re \left\{ \tilde{C}_E(\tau) \times \left\{ e^{i \left[\omega \tau - k(X(t+\tau) - X(t)) \right]} \right\} \right\},\tag{14}$$

where $\tilde{C}_E(\tau)$ is the autocovariance of the envelope. For the latter, we will re-use the model previously introduced in eq. (8), now denoting the Eulerian decorrelation term as \tilde{T}_E :

$$\tilde{C}_E(\tau) = \tilde{C}(\tau)/\cos(\omega\tau) = \tilde{V}_s^2 + \tilde{V}_{ns}^2 \exp(-\tau/\tilde{T}_E).$$

²⁹⁸ We assume now that the slow flow is a stationary Gaussian process, with a rms amplitude ²⁹⁹ \overline{V} (over one direction) and an exponential decorrelation in time with a typical time scale ³⁰⁰ \overline{T} – somewhat consistent with the model proposed in section 1, eq. (9) dropping the ³⁰¹ $\cos(\Omega \tau)$ term for simplicity. Such model – sometimes referred as an unbiased correlated ³⁰² velocity model in the literature (Gurarie et al.) – corresponds to the time-homogeneous ³⁰³ Ornstein-Uhlenbeck process.

The displacement $\delta X = X(t + \tau) - X(t)$ is also a Gaussian process with null mean and variance given by (Pope 2015, Chap. 12):

$$\langle \delta X(t)^2 \rangle \equiv \sigma_X^2 = 2\bar{T}^2 \bar{V}^2 \left[\tau / \bar{T} - \left(1 - e^{-\tau / \bar{T}} \right) \right]. \tag{15}$$

It is worth noting that the variance of the displacement admits two asymptotic regimes: $\sigma_X^2 \rightarrow \bar{V}^2 \tau^2$ for $\tau \ll \bar{T}$, and $\sigma_x^2 \rightarrow 2\bar{V}^2\bar{T}\tau$ for $\tau \gg \bar{T}$. From this variance for the displacement, one obtains the final expression for the autocovariance of the fast motion in the Lagrangian frame:

$$\widetilde{C}_{L}(\tau) = \widetilde{C}_{E} \int_{-\infty}^{\infty} \cos(\omega\tau - k\delta X) p(\delta X) d\delta X$$
(16)

$$= \left(\tilde{V}_s^2 + \tilde{V}_{ns}^2 \exp(-\tau/\tilde{T}_E)\right) \cos(\omega\tau) \int_{-\infty}^{\infty} \cos(k\delta X) \frac{e^{-\delta X^2/(2\sigma_X'^2)}}{\sigma_X'\sqrt{2\pi}} d\delta X$$
(17)

$$= \tilde{C}e^{-\sigma_X^2 k^2/2} = \tilde{C}\exp\left(-k^2 \bar{V}^2 \bar{T}^2 \left[\tau/\bar{T} - (1 - e^{-\tau/\bar{T}})\right]\right)$$
(18)

The resulting Lagrangian autocorrelation has no stationary part and decays faster than the Eulerian autocorrelation, as follows from the exponentially decay due to drifter transports. A non-dimensional parameter $k\bar{V}\bar{T}$ readily appears in this exponential, which compares the time taken by the a drifter to travel a wave length compare to the typical decorrelation time of the slow flow. Several regimes are identified, depending on the effect of the advection, $k^2\sigma^2$, at a fixed time, typical of the decorrelation of the envelope :

• Weak advection $k^2 \sigma^2 \ll 1$: The signature of the wave signal in the Lagrangian frame of reference matches the Eulerian one: $\tilde{C}_L(\tau) \sim \tilde{C}_E(\tau)$. Lagrangian stationary and non-stationary contributions directly reflect Eulerian ones. If $\tilde{V}_{ns} \gg \tilde{V}_s$, Eulerian is nonstationary. Amplitudes and non-stationarity timescales are expected to be similar in both perspectives. If $\tilde{V}_{ns} \ll \tilde{V}_s$, Eulerian signal is stationary.

• Strong advection, $k^2 \sigma^2 \gg 1$: the Lagrangian perspective will deform the Eulerian one. $k\bar{V}\bar{T}$ will control the form of the Lagrangian envelope. If $k\bar{V}\bar{T} \gg 1$ the exponential decay of the Lagrangian autocorrelation scales is quadratic in τ with decay time scale $1/k\bar{V}$: $\tilde{C}_L(\tau) \sim \tilde{V}_s^2 \cos(\omega \tau) \times e^{-k^2 \bar{V}^2 \tau^2}$. If $k\bar{V}\bar{T} \ll 1$ the slow flow decorrelation induces a linear exponential decay with decay time scale $1/k^2 \bar{V}^2 \bar{T}$: $\tilde{V}_s^2 \cos(\omega \tau) \times e^{-k^2 \bar{V}^2 \bar{T} \tau}$

This predicted form of the autocorrelation in the Lagrangian framework is in qualitative agreement with the Lagrangian autocorrelations shown on Fig. 5 and 6. In particular, the stationary part of the fitted Lagrangian autocorrelation becomes negligible in a large region around the jet latitude. The fact that it is not zero at the very south and very north of the domain is due to the fact that the parameter $k\bar{V}\bar{T}$ is very small in this region, associated with a very slow decay of the Lagrangian envelop (see also Fig. 3), which implies a non-vanishing stationary fraction as recovered by the model fit from eq. (8).

³³⁴ b. Comparison of fitted autocovariances to predicted Lagrangian ones

The predicted Lagrangian autocovariance's envelope (Fig. 7 right column) is computed from the fitted Eulerian autocovariance envelope (Fig. 7 left column) following Eq. (18) so that it may be compared to the fitted Lagrangian one (Fig. 7 middle column). For all simulations (S_0 , S_2 and S_4 shown in Fig. 7, first, second and third line respectively),

the Eulerian autocorrelation envelope's decay rate increases from south to north and with 339 the jet's strength as commented previously. This is expected of a loss of stationarity 340 as the wave propagates northward. For the fitted Lagrangian autocorrelation however, 341 low values of the envelop are reached for small time lags in the jet due to the advection 342 in the area. In the northern part the Lagrangian envelope exhibits slower decay rates 343 that tend to Eulerian ones. The predicted Lagrangian autocorrelation envelope reaches 344 values inferior to 0.1 for time lags smaller than 5 days in the center of the domain which 345 is consistent with what is observed in the fitted Lagrangian autocorrelation in a strong 346 advection area. In northern and southern area, the advection impact represented in the 347 predicted autocorrelation also fits what is observed in fitted Lagrangian autocorrelation 348 with decay rates qualitatively close to the Eulerian ones. This decay seems however 349 slightly overestimated (values smaller than 0.2 for $\tau > 30$ days when similar values are 350 not reached in Eulerian or Lagrangian fits). Overall, following Eq. (18), the differences 351 between Lagrangian and Eulerian can be qualitatively explained as the effect of drifters 352 being advected by the slow motion. 353

³⁵⁴ Depending on the significance of Eulerian non-stationarity and the relative importance ³⁵⁵ of \tilde{T}_L/\tilde{T}_E and $k^2\sigma^2(\tilde{T}_L)$ (Fig. 8b and c, respectively), the Lagrangian non-stationarity ³⁶⁶ timescale is likely due to Eulerian non-stationarity or drifters strong advection by slow ³⁵⁷ motion. For simplification sake, both terms will be referred to as $r_E(\tilde{T}_L)$ and $r_{adv}(\tilde{T}_L)$ ³⁵⁸ respectively.

In the southern area, the ratio of Eulerian nonstationary and stationary amplitudes (Fig. 8a) is smaller than one for all simulations, internal tides area stationary. The advection is weak, $r_{adv}(\tilde{T}_L) \ll 1$. The Eulerian nonstationary timescale has no significant impact on the form of envelope in this area and its estimation is expected not to be significant. r_E is found larger than one is this area, however a Lagrangian timescale larger than Eulerian one cannot be explained by our model. As stipulated previously, this can be due to an underestimation of the Eulerian timescale. It also could be caused by an estimation error of Lagrangian nonstationary timescale and/or the result of an inappropriate form of decay employed for the fit (e.g. linear exponential decay vs quadratic) for Lagrangian autocovariance as the slow motion decorrelation timescale is of the order of the time window size (40 days).

Northern and central areas illustrate the two different regimes for internal tides non-370 stationarity in Lagrangian data, described in the previous subsection. In the north, ratio 37 of nonstationary and stationary amplitudes is larger than one for S_2 , S_3 and S_4 , where 372 nonstationary amplitudes dominates. For least energetic slow motion (S_0 and S_1) stationary 373 tide still dominates but with significant nonstationary contribution. The advection term, 374 r_{adv} is small in the north and r_E values are around 1 : the Lagrangian autocovariance is 375 close to the Eulerian one, following eq.(18). The non-stationarity captured in Lagrangian 376 perspective corresponds to the one in Eulerian perspective. In the jet, the ratio of Eulerian 377 nonstationary and stationary amplitudes increases with nonstationary component not yet 378 dominant. In the same area, r_E is small as the Lagrangian timescale is small in front of 379 the Eulerian one, the decorrelation is faster in Lagrangian perspective than in Eulerian 380 one. This coincides with r_{adv} around 1, the internal tides' non-stationarity captured in the 381 Lagrangian perspective is due to advection by slow motion. We call this non-stationarity, 382 apparent non-stationarity. 383

5. Discussion

a. Lagrangian apparent non-stationarity

We now discuss the reported alteration of the internal tide surface signature in the Lagrangian perspective in the more general context of observation of ocean high-frequency dynamics. Low mode internal tides have by definition large vertical scales – similar to that of the background flow. Advection by the slow flow is of particular importance for discussing the Eulerian/Lagrangian distorsion, even though it does not fully capture the interaction between the slow flow and the internal tide (Dunphy et al. 2017; Savage et al. 2020). A vertical mode expansion of equations of motions linearized around the slow background flow shows that advection of the internal tide mode is driven by a non-trivial weighted average of the background flow.

This effective advection is expressed as $H^{-1} \int_{-H}^{0} \phi_n^2 \mathbf{U} dz$ (Kelly and Lermusiaux 2016), where ϕ_n is the standard pressure mode for an internal wave with vertical mode number *n* (see also Duda et al., for a more technical approach).

Thus, for a surface intensified background flow, the flow advecting the drifter (at the surface) and the one advecting the internal tide mode is different, explaining why the Lagrangian observer renders a distorted view of the internal wave signal.

For the simulation with moderate jet intensity S2, for instance, the mode 1 effective advection velocity (computed, but not shown) is of order $0.2 \,\mathrm{m\,s^{-1}}$ at its maximum, while the surface velocity is typically greater than $1 \,\mathrm{m\,s^{-1}}$: the Eulerian distortion, driven by the effective advection velocity, is therefore smaller than the Lagrangian distortion, driven by the difference between this effective advection and the surface velocity transporting the drifter.

This somewhat justifies a central approximation of the theoretical model derived in section a, where we neglected the advection of the wave by the slow flow.

For small scale internal waves on the other hand, ray theory can be used to describe their propagation through the background flow (Broutman et al. 2004). This approach shows that wave packets are advected by the local flow, which is associated with a Doppler shifting of the Eulerian frequency: $\omega = \hat{\omega} + \mathbf{k} \cdot \mathbf{U}$, where ω and $\hat{\omega}$ are respectively the wave absolute (or Eulerian) and intrinsic (as measured in a frame of reference moving with the slow flow) frequencies, **k** is the wave vector, and **U** is the slow flow.

19

Ignoring advection of the drifter by the wave current, the signal measured by the drifter coincides with the wave field in the frame comoving with the mean flow with least distortion in the Lagrangian frame of reference.

This situation is opposite to the configuration investigated here, as Lagrangian autocorrelation exhibits faster decrease with time lag compared to Eulerian auto-correlation, and the theoretical model proposed here would obviously not be relevant.

⁴²¹ In a realistic configuration, the range of validity of each of these two regimes (e.g. small ⁴²² vs large scale waves) remains to be quantified.

423 b. On internal tide non-stationary spatial envelope

Another assumption of the theoretical model is that the spatial envelope of the wave is 424 spatially uniform but temporally variable. In reality the envelope of the wave propagates 425 with the internal tide group speed resulting in spatial variability if a temporal one is 426 admitted. We assume the typical size of the envelope should however scale as the product 427 of the group velocity and the Eulerian non-stationary time scale and that this will in general be larger than several wavelengths. The apparent non-stationarity relies on the 429 conversion of the spatial variability at the scale of a wavelength and thus do not expect 430 the spatial variability of the envelope would affect the form (10) at first order. Synthetic 431 experiments could help verify this point. If necessary, the spatial inhomogeneities of 432 the wave envelop could be included in the model, at the cost of adding complexity. This, 433 however, requires additional to characterize these spatial inhomogeneities (through spatio-434 temporal autocovariance), which has not been reported in the context of internal tides - to 435 our knowledge. 436

437 c. Autocorrelations models and sationary/non-stationary decomposition

⁴³⁸ Several ad-hoc choices have been made regarding the shape of internal tide and slow ⁴³⁹ motion autocorrelation. Limits to these choices are visible on Figure 4c for slow motions

and are speculated to affect estimates of internal tide nonstationary time scales in the 440 southern part of the domain. The internal tide envelope autocorrelation initially chosen 441 included a single exponential decaying term instead of the sum of stationary/nonstationary 442 contributions. We abandoned eventually this choice as it does not naturally lead to the 443 decomposition of the signal into stationary/non-stationary contributions as well as was 444 required to consider time scales overly large in stationary cases (>1000 days). One 445 may also have fitted the more general form (Eq.(18)) to Lagrangian autocorrelations, for 446 example, and evaluated its relevance compared to the single linear term exponential form. 447 This would add one more parameter to estimate however and would require to determine 448 whether this more general form leads to an improvement which we felt was a study on its 449 own. We did not attempt to do this eventually in favor of a more qualitative assessment 450 of the theory. Determining what form is more appropriate to describe the Eulerian and 451 Lagrangian internal envelope autocorrelation is a study on its own that would best reserved 452 to realistic configurations. 453

6. Conclusion

In order to investigate the use of Lagrangian data to characterize internal tides propagating through a turbulent jet, we compare Eulerian and Lagrangian internal tides characteristics in an idealized simulation. Characteristics of mode-1 internal tides were estimated via fitting data in both perspectives and an internal tide autocovariance envelope including the sum of a constant stationary contribution and an exponentially decaying nonstationary one.

⁴⁶¹ Near their generation site and far from the jet, internal tides are found to be nearly ⁴⁶² stationary in Eulerian perspective. As internal tides propagate through the jet (i.e. ⁴⁶³ energetic area), the drifters are strongly advected by slow motions which causes ⁴⁶⁴ Lagrangian non-stationary timescales to be lower than Eulerian ones. We call this

phenomenon, found in Lagrangian perspective, apparent non-stationarity. After crossing 465 the jet, internal tides propagate in an area of weaker energy but have lost stationarity. 466 Internal tides non-stationary amplitude is significant there and even dominates in the most 467 energetic cases. In this area, the intrinsic non-stationarity captured in Eulerian diagnostics 468 is present similarly in Lagrangian amplitudes and timescales. Regardless of slow motion 469 amplitude or Eulerian stationarity of internal tides the total amplitude (stationary and 470 nonstationary components) is successfully recovered in Lagrangian perspective. The 471 deformation of internal tides characteristics in Lagrangian perspective was qualitatively 472 predicted by a theoretical model for Lagrangian autocovariances (eq.(18)). This model 473 modifies Eulerian autocovariances to take into account the advection of drifters. 474

475

One of the main result of this study is the estimation of the total amplitude in Eulerian 476 and Lagrangian perspectives using autocovariances and parametric fit instead of non-477 parametric spectral analysis. This method has the advantage to avoid the dependence 478 on an arbitrary choice of frequency band. Indeed, in spectral analysis the chosen band 479 has a strong impact on evaluating the energy contribution of motions identified by their 480 frequency (Yu et al. 2019). This choice is further complicated by the variation of the width 481 of tidal peaks depending on the internal tides' stationarity and drifters advection by slow 482 motion, in the case of Lagrangian spectra. Parametric spectral analysis could be explored 483 as an alternative to free from this constraint. The efficiency of our method should also be 484 investigated in a more realistic set-up. 485

Lagrangian diagnostics and their comparison to Eulerian ones should help to assess how drifters data could be used, notably in the context of SWOT. As Lagrangian and Eulerian total amplitudes were found to be similar, Lagrangian estimates could help to identify where the internal tides contribution to ocean surface energy would be significant. This leading to where balanced flow and internal tides could be entangled in SWOT

data. It involves however a transition from internal tides kinetic energy amplitude to 491 pressure, which should be investigated. In areas of weak advection, we might be able 492 to separate stationary and non-stationary amplitudes in Lagrangian data. The theoretical 493 model developed in this study should help to define regimes for which this separation is 494 possible. Areas of high non-stationary internal tides contribution and therefore of difficult 495 prediction of internal tides may thus be flagged. Note that these specific regions are also 496 the ones where we found a correct estimation of non-stationarity timescales. Again these preliminary results should be investigated in a realistic set-up, numerical model or in situ 498 data. 499

This potential information brought by Lagrangian data could also be complementary material to map internal tides. Mappings of internal tides have been studied from altimetry (Zaron 2017, 2019). Drifters data, via an estimation of total amplitude, might contribute to this mapping, setting at least an upper limit to stationary and nonstationary amplitudes. A direct decomposition of total amplitude in stationary and nonstationary components may also be derived, in some cases, from drifters data, as noted previously.

As anticipated in other studies (Zaron and Elipot 2020), apparent non-stationarity was found in our Lagrangian data and predicted by our theoretical model, in areas of strong advection. Areas and regimes for which Lagrangian perspective would deform internal tides characteristics may be flagged. It could have consequences on the use of Lagrangian filtering to filter out tidal signal from surface data. Indeed, low decorrelation timescales translate in the frequency domain by a broadening of tidal peaks which is expected to complicate the separation of motions by their frequency.

The estimation of internal tides contribution in ocean surface energy through the total – and, in some cases, of stationary and nonstationary – amplitudes from drifters data might also help to validate realistic models resolving high frequency variability.

23

In the shorter term, a natural extension of our study is to apply our findings to the 516 analysis of actual drifter trajectories or numerically predicted ones in realistic simulations. 517 Eulerian outputs from a realistic simulation (LLC4320) have already been compared to 518 drifters data (Yu et al. 2019). A future study will aim at applying our findings to simulated 519 drifter trajectories from the same numerical simulation. Similar methods applied to 520 energetic regions of the globe would allow us to test our method and analytical models to 521 more realistic set-ups in particular to compare Lagrangian and Eulerian amplitudes and 522 non-stationarity timescales. 523

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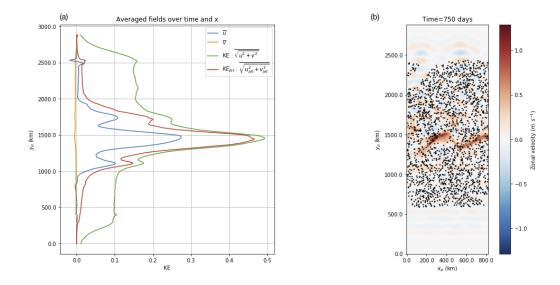
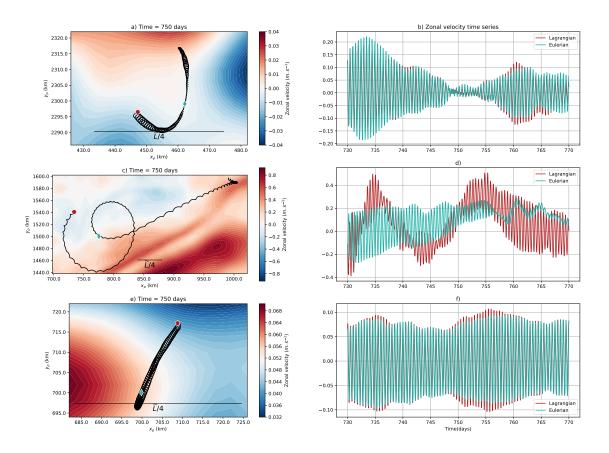


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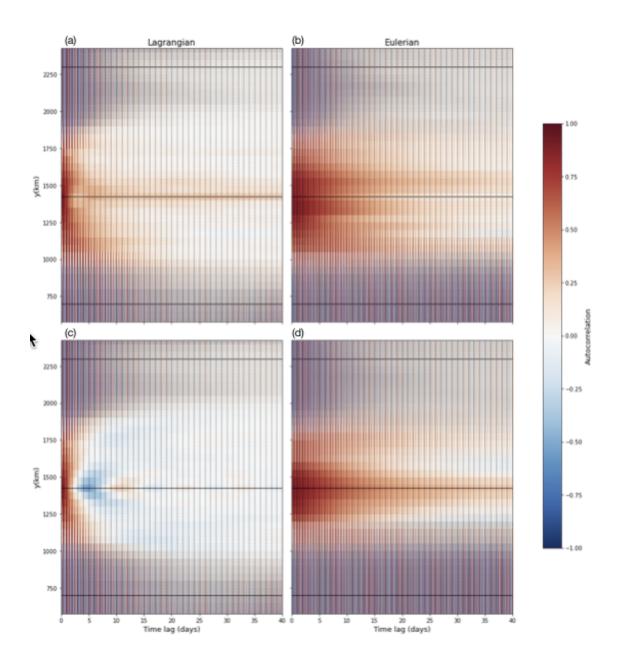
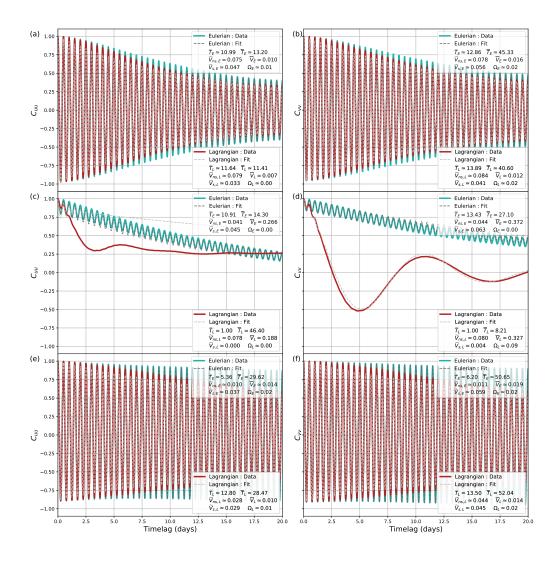


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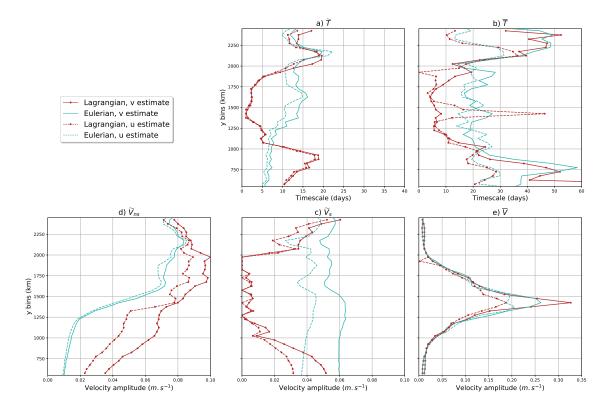


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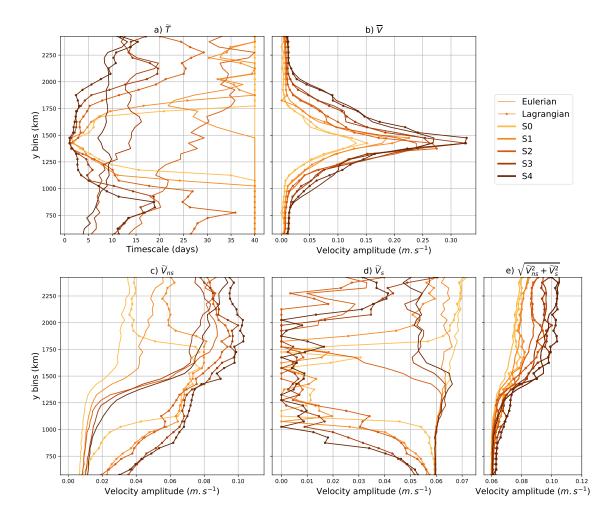


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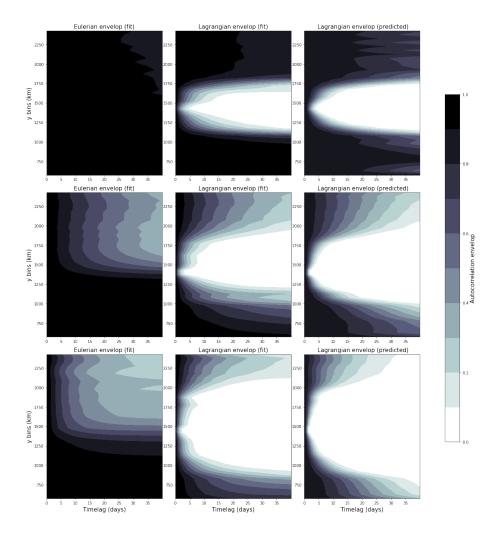


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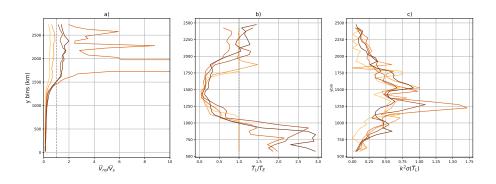


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