Quantifying Intrinsic and Extrinsic Contributions to Seismic Anisotropy in Tomographic Models

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Abstract

Seismic anisotropy in the Earth's mantle inferred from seismic observations is usually interpreted either in terms of intrinsic anisotropy due to Crystallographic Preferred Orientation (CPO) of minerals, or extrinsic anisotropy due to rock-scale Shape Preferred Orientation (SPO). The coexistence of both contributions misconstrues the origins of seismic anisotropy observed in seismic tomography models. It is thus essential to discriminate CPO from SPO. Homogenization/upscaling theory provides means to achieve this goal. This theory enables to compute the effective elastic properties of a heterogeneous medium, as seen by long-period waves. In this work, we investigate the effects of upscaling an intrinsically anisotropic and highly heterogeneous Earth's mantle. We show analytically in 1-D that the full effective radial anisotropy ξ^* is approximately the product of the effective intrinsic radial anisotropy ξ^* CPO and the extrinsic radial anisotropy ξ^* SPO : ξ^* [?] ξ^* CPO x ξ^* SPO. This law is verified numerically in the case of a 2-D marble cake model of the mantle with a binary composition, and in the presence of CPO obtained from a micro-mechanical model of olivine deformation. We compute the long-wavelength effective equivalent of this mantle model using the 3-D non-periodic elastic homogenization technique. Our numerical findings predict that for wavelenghts smaller than the scale of deformation patterns, tomography may overestimate the true anisotropy (i.e. intrinsic anisotropy due to CPO) due to significant SPO-induced extrinsic anisotropy. However, at wavelenghts larger than deformation patterns, intrinsic anisotropy is always underestimated in tomographic models due to the spatial averaging of the preferred orientation of anisotropic minerals. Thus, we show that it is imperative to homogenize a CPO evolution model first before drawing comparisons with tomographic models. As a demonstration, we use our composite law with a homogenized CPO model of a plate-driven flow underneath a mid-ocean ridge, to estimate the SPO contibution to an existing tomographic model of radial anisotropy.

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8 Key Points:

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9	•	We propose a theoretical expression that relates the observed seismic anisotropy
.0		to its intrinsic and extrinsic contributions.

- For wavelengths longer than the scale of deformation patterns in the mantle,
 we show that observed anisotropy in tomographic models underestimates in trinsic anisotropy, due to the spatial averaging of intrinsic anisotropy over long
 wavelengths.
- For wavelengths shorter than the scale of deformation patterns, observed anisotropy
 overestimates intrinsic anisotropy due to the presence of extrinsic anisotropy.

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17 Abstract

Seismic anisotropy in the Earth's mantle inferred from seismic observations is usually 18 interpreted either in terms of intrinsic anisotropy due to Crystallographic Preferred 19 Orientation (CPO) of minerals, or extrinsic anisotropy due to rock-scale Shape Pre-20 ferred Orientation (SPO). The coexistence of both contributions misconstrues the ori-21 gins of seismic anisotropy observed in seismic tomography models. It is thus essential 22 to discriminate CPO from SPO. Homogenization/upscaling theory provides means to 23 achieve this goal. This theory enables to compute the effective elastic properties of 24 a heterogeneous medium, as seen by long-period waves. In this work, we investigate 25 the effects of upscaling an intrinsically anisotropic and highly heterogeneous Earth's 26 mantle. We show analytically in 1-D that the full effective radial anisotropy ξ^* is 27 approximately the product of the effective intrinsic radial anisotropy $\xi^*_{\rm CPO}$ and the 28 extrinsic radial anisotropy ξ_{SPO}^* : 29

$$\xi^* \approx \xi^*_{\rm CPO} \times \xi^*_{\rm SPO}.$$

This law is verified numerically in the case of a 2-D marble cake model of the man-31 tle with a binary composition, and in the presence of CPO obtained from a micro-32 mechanical model of olivine deformation. We compute the long-wavelength effective 33 equivalent of this mantle model using the 3-D non-periodic elastic homogenization tech-34 nique. Our numerical findings predict that for wavelenghts smaller than the scale of 35 deformation patterns, tomography may over-estimate the true anisotropy (*i.e.* intrinsic 36 anisotropy due to CPO) due to significant SPO-induced extrinsic anisotropy. How-37 ever, at wavelenghts larger than deformation patterns, intrinsic anisotropy is always 38 underestimated in tomographic models due to the spatial averaging of the preferred 30 orientation of anisotropic minerals. Thus, we show that it is imperative to homogenize 40 a CPO evolution model first before drawing comparisons with tomographic models. 41 As a demonstration, we use our composite law with a homogenized CPO model of a 42 plate-driven flow underneath a mid-ocean ridge, to estimate the SPO contibution to 43 an existing tomographic model of radial anisotropy. 44

45 **1** Introduction

Seismic anisotropy in the Earth's mantle originates from various processes and
can be observed at different spatial scales (Hansen et al., 2021). At the mineral
scale, crystallographic preferred orientation (CPO) of anisotropic mantle minerals due

to progressive shearing over time produces large-scale intrinsic anisotropy (Nicolas
& Christensen, 1987; Maupin & Park, 2015). On the other hand, rock-scale shape
preferred orientation (SPO) such as layered heterogeneous materials, seismic discontinuities, preferentially-oriented cracks or conduits containing fluid intrusions unresolved
by long period seismic waves are mapped as large-scale extrinsic anisotropy (Backus,
1962; Crampin & Booth, 1985).

Although these two mechanisms are completely different, a medium may be ei-55 ther (or both) intrinsically anisotropic and extrinsically anisotropic at a given scale, 56 depending on the minimum wavelength of the observed wavefield used (Maupin et al., 57 2007; Wang et al., 2013; Fichtner et al., 2013; Bodin et al., 2015). Backus (1962) 58 showed that a horizontally-layered isotropic medium is equivalent to a homogeneous 59 radially anisotropic medium with a vertical axis of symmetry when sampled by seismic 60 waves whose wavelength is much longer than the thickness of layers. This urged seis-61 mologists to interpret tomographic models separately depending on the type of data 62 used (*i.e.*, different data-types sample different length scales). Scattering studies use 63 high frequency body waves and interpret small-scale isotropic heterogeneities in terms of phase changes (e.g. Tauzin & Ricard, 2014) or chemical stratification (e.g. Tauzin 65 et al., 2016). On the other hand, long period surface waves with typical wavelengths of 66 the order 10^2 km retrieve a smooth anisotropic mantle with scales consistent with con-67 vective flow (e.g. Debayle & Ricard, 2013; Bodin et al., 2015; Maupin & Park, 2015). 68 Surface waves however lack the resolving power to recover sharp seismic discontinuities 69 and instead, map these as long wavelength radial anisotropy (Backus, 1962; Capdev-70 ille et al., 2013). Anisotropic structures retrieved from tomography may therefore be 71 a combination of apparent extrinsic anisotropy due to SPO and deformation-induced 72 intrinsic anisotropy. The ambiguity on the origin of observed anisotropy (i.e. whether 73 a material is intrinsically anisotropic or strongly heterogeneous) may mislead seismolo-74 gists in interpreting the structural origin of seismic anisotropy observed in tomographic 75 images. 76

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1.1 Intrinsic anisotropy due to Crystallographic Preferred Orientation

Intrinsic anisotropy results from the preferred alignment of anisotropic crystals
 in an aggregate when subjected to a macroscopic deformation. In the mantle, single
 crystal olivine exhibits orthorhombicity, and hence suffers variations in fast and slow

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P- and S-wave velocities up to 20 % (Kumazawa & Anderson, 1969). When olivine and pyroxene form a polycrystalline aggregate and are subsequently deformed in the mantle flow, their CPO can be described in terms of a hexagonally symmetric medium (e.g. J. Montagner & Nataf, 1988).

Observations of large-scale anisotropy in tomographic models appear to be ubiq-85 uitous in regions associated with strong deformation, and have often been interpreted 86 in terms of convective flow (McKenzie, 1979). For instance, tomographic imaging 87 has revealed the presence of positive radial anisotropy (*i.e.*, horizontally propagating 88 SH-waves traveling faster than SV-waves) of about 4% in the upper ~ 250 km of 89 the mantle and has been interpreted as lateral flow (refer to Long and Becker (2010) 90 for a comprehensive review). Long wavelength seismic anisotropy is also prevalent 91 in the transition zone (e.g. Trampert & van Heijst, 2002; Wookey & Kendall, 2004) 92 although its origin is still highly debated (Chen & Brudzinski, 2003; Chang & Fer-93 reira, 2019; Sturgeon et al., 2019). Probing deeper depths, the lower mantle appears to be isotropic (e.g. Meade et al., 1995) barring the D" layer where enough evidence 95 have shown it to be anisotropic (e.g. Kendall & Silver, 1998; McNamara et al., 2002; 96 Panning & Romanowicz, 2006). 97

Since CPO maps the deformation patterns, CPO may deviate from the flow direction. This is because the deformation patterns relate not to the velocity field itself, but to the velocity gradient. Moreover, CPO is not instantaneous, but depends on the history of the deformation. As a result, regions with short deformation trajectories such as beneath mid-ocean ridges appear to have under-developed CPO, and would lag behind the direction of shear deformation (É. Kaminski & Ribe, 2002).

Based on laboratory experiments of simple shear, the fast axis of olivine tends 104 to align parallel to the long axis of the finite strain ellipsoid (FSE) at low strains due 105 to plastic deformation (Zhang & Karato, 1995). At larger strains, dynamic recrys-106 tallization facilitates the alignment of the olivine fast axis towards the direction of 107 shear (Zhang & Karato, 1995; Bystricky et al., 2000). Mechanical models of CPO 108 evolution, coupled with geodynamic flow modeling have been developed to replicate 109 these results and have been extrapolated at scales consistent with mantle deformation 110 patterns. Among these is the viscoplastic self-consistent (VPSC) model which is used 111 to explain the mechanical response of polycrystals to plastic deformation (Tommasi et 112

al., 2000). Such tools however are computationally expensive, especially when applied
to 3-D and non-steady state flows (Lev & Hager, 2008). Another well-received method
is the D-Rex model, that utilizes a simple kinematic approach (E. Kaminski et al.,
2004). The predicted CPO is then converted to an elastic medium in which seismic
waves can propagate, and may explain anisotropic signatures observed in seismic data
recorded at the surface.

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1.2 Extrinsic anisotropy due to Shape Preferred Orientation

Extrinsic anisotropy is observed under two conditions: (1) when the scale of the heterogeneities is much smaller than the minimum wavelength of the observed wavefield, and (2) when the contrast between seismic wave velocities (*i.e.* the amplitude of heterogneities) is large.

One of the known configurations at which extrinsic anisotropy is produced is rock-124 scale shape preferred orientation (SPO). In the Earth's mantle, rock-scale SPO can 125 be the result of igneous differentiation, or more generally of the stirring of chemical 126 heterogeneities by tectonic or convective deformation (e.g. Faccenda et al., 2019). 127 Since magmatically differentiated oceanic lithosphere is composed of a basaltic crustal 128 layer blanketed by a depleted harzburgitic mantle (Allègre & Turcotte, 1986), mantle 129 structure is often modeled in terms of a mechanical mixture of these two end-member 130 compositions (e.g. Hofmann, 1988; Xu et al., 2008; Ballmer et al., 2015). 131

Large-scale thermal convection in the mantle triggers the constant injection of 132 oceanic lithosphere into the mantle (Coltice & Ricard, 1999). It then mechanically 133 stirs with the surrounding mantle and experiences a series of stretching and thinning 134 due to the normal and shear strains associated with convection (Allègre & Turcotte, 135 1986). This led Allègre and Turcotte (1986) to develop a geodynamic model of the 136 mantle that would depict marble cake-like patterns. In their model, the layering may 137 be erased either by dissolution processes when the stripes become thin enough that 138 chemical diffusion becomes efficient, or by mantle reprocessing at mid-ocean ridges. 139 Assuming that the mixing preserves the physical properties of the two-end members 140 with depth and over geological time scales, such processes may explain rock-scale seis-141 mic heterogeneities observed in the mantle in agreement with the spectrum of isotropic 142

anomalies obesrved along ridges (Agranier et al., 2005; Xu et al., 2008; Stixrude &
Jeanloz, 2015).

In this paper, we extend the work of Alder et al. (2017) by estimating the long-145 wavelength effective equivalent of a marble cake mantle as hypothesized by Allègre 146 and Turcotte (1986), but in the presence of intrinsic anisotropy. Our aim is to quan-147 tify the level of effective anisotropy resulting from elastic homogenization, that is, the 148 relegated version of the true Earth as seen by long-wavelength seismic tomography. 149 Section 2 is a brief overview of the homogenization theory and provides a definition of 150 some terms and notations to guide the reader throughout the paper. Section 3 shows 151 1-D analytical expressions for homogenization and highlights a composite law that 152 separates intrinsic and extrinsic anisotropy for a layered and anisotropic media. Here, 153 we demonstrate that the effective anisotropy varies with the square of isotropic het-154 erogeneities, as well as with the square of anisotropic heterogneities, plus a cross term 155 related to their coupling. In section 4, we build a 2-D media analogous to the marble 156 cake model where we consider a mechanical mixture of two end-member compositions. 157 We follow this by introducing intrinsic anisotropy due to mantle deformation associ-158 ated with convection patterns consistent with the marble cake model. We compute 159 the long-wavelength effective equivalent of the 2-D models using the Fast-Fourier Ho-160 mogenization algorithm (Capdeville et al., 2015). Section 5 presents the results of the 161 previous section: one of the major findings is that in the absence of isotropic het-162 erogeneities, intrinsic anisotropy is always underestimated upon homogenization due 163 to the spatial averaging of the preferred orientation of the anisotropic minerals. We 164 also verify numerically that the composite law derived in section 3 can be extended 165 to 2-D media. Finally in section 6, we apply the composite law to infer the extrinsic 166 component of anisotropy from a tomographic model of the upper-mantle beneath a 167 mid-ocean ridge with the help of a homogenized CPO model. 168

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2 Elastic homogenization

Seismic tomography is only able to recover a smooth representation of the real Earth due to the limited frequency band of seismic data. This smooth average, however, is not just a simple spatial average but is produced from highly non-linear upscaling relations. In the context of wave propagation, such upscaling relations, also known as elastic homogenization, remove seismic heterogeneities whose scales are much

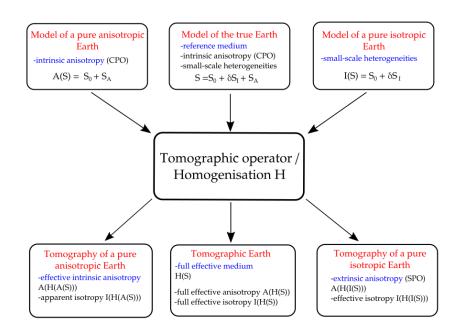


Figure 1. Homogenization of different Earth models and their respective outputs. The true Earth mantle (top middle box) is described by an average isotropic model \mathbf{S}_0 , isotropic heterogeneities, $\delta \mathbf{S}_{\mathbf{I}}$ and intrinsic anisotropy $\mathbf{S}_{\mathbf{A}}$, the sum of which being the elastic model \mathbf{S} that tomography tries to recover. However, tomographic methods have only access to a homogenized model $\mathcal{H}(\mathbf{S})$ (or full effective medium). This model has both isotropic components symbolized by $\mathcal{I}(\mathcal{H}(\mathbf{S}))$ and anisotropic components, $\mathcal{A}(\mathcal{H}(\mathbf{S}))$. The goal of this paper is to quantify the differences between $\mathcal{A}(\mathcal{H}(\mathbf{S}))$ and $\mathcal{A}(\mathbf{S})$, $\mathcal{I}(\mathcal{H}(\mathbf{S}))$ and $\mathcal{I}(\mathbf{S})$. Numerically we can also discuss how an anisotropic model without isotropic heterogeneities (boxes on the left) can be recovered and if the tomographic inversion can lead to apparent isotropic model is recovered by the tomographic inversion and what is the level of extrinsic anisotropy (SPO) that can be estimated.

smaller than the minimum wavelength of the observed wavefield, and instead replace
them with effective properties.

Hereafter, what we refer to as the *reference medium* $\mathbf{S}(\mathbf{r})$ is an elastic model of the real Earth varying in space \mathbf{r} that accounts for both intrinsic anisotropy due to CPO and small-scale isotropic heterogeneities that resemble marble cake-like patterns. This reference medium can be treated as a sum of several decompositions resulting from a cascade of orthogonal projections (Browaeys & Chevrot, 2004). One can then express $\mathbf{S}(\mathbf{r})$ in terms of an isotropic tensor $\mathbf{S}_{\mathbf{I}}(\mathbf{r})$ plus an intrinsically anisotropic component $\mathbf{S}_{\mathbf{A}}(\mathbf{r})$ related to CPO:

$$\mathbf{S}(\mathbf{r}) = \mathbf{S}_{\mathbf{I}}(\mathbf{r}) + \mathbf{S}_{\mathbf{A}}(\mathbf{r}), \tag{1}$$

where $\mathbf{S}_{\mathbf{I}}(\mathbf{r})$ can be decomposed further into:

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$$\mathbf{S}_{\mathbf{I}}(\mathbf{r}) = \mathbf{S}_{\mathbf{0}} + \delta \mathbf{S}_{\mathbf{I}}(\mathbf{r}). \tag{2}$$

Here, S_0 is an isotropic tensor uniform in space, and $\delta S_I(\mathbf{r})$ is a deviation from S_0 related to small-scale isotropic heterogeneities. The reference medium becomes:

$$\mathbf{S}(\mathbf{r}) = \mathbf{S}_0 + \delta \mathbf{S}_{\mathbf{I}}(\mathbf{r}) + \mathbf{S}_{\mathbf{A}}(\mathbf{r}). \tag{3}$$

For convenience, let us introduce an operator \mathcal{I} that extracts the isotropic component from \mathbf{S} , and an operator \mathcal{A} that extracts the anisotropic component from \mathbf{S} :

$$\mathcal{I}(\mathbf{S}(\mathbf{r})) = \mathbf{S}_{\mathbf{I}}(\mathbf{r})$$

$$\mathcal{A}(\mathbf{S}(\mathbf{r})) = \mathbf{S}_{\mathbf{0}} + \mathbf{S}_{\mathbf{A}}(\mathbf{r}).$$
(4)

These notations will be used heavily in the rest of the text to denote the isotropic 193 and anisotropic components of an elastic medium. Radial anisotropy, in particular, 194 quantifies the level of anisotropy when the medium is averaged azimuthally (J.-P. Mon-195 tagner, 2007; Maupin et al., 2007). In such a vertically transverse isotropic medium 196 (VTI), the level of radial anisotropy is given by $(V_{SH}/V_{SV})^2$, where V_{SV} is the ve-197 locity of vertically traveling S waves or horizontally traveling S waves with vertical 198 polarization, and V_{SH} is the velocity of horizontally traveling S waves with horizontal 199 polarization. The *intrinsic radial anisotropy* associated with $\mathcal{A}(\mathbf{S})$ (*i.e.* due to the 200 component $\mathbf{S}_{\mathbf{A}}$) will be denoted by ξ_{CPO} . 201

In the event where long-period waves sample this reference medium, small-scale heterogeneities are seen only through their effective properties. Computing these effective properties is designated by a mathematical operator \mathcal{H} called *upscaling* or

homogenization. Setting aside the imperfections of inversion algorithms and data cov-205 erage, performing seismic tomography can be viewed as applying the operator \mathcal{H} that 206 homogenizes \mathbf{S} . The seismic tomography model/long-wavelength effective medium of 207 \mathbf{S} is then $\mathcal{H}(\mathbf{S}) = \mathcal{H}(\mathbf{S}_0 + \delta \mathbf{S}_I + \mathbf{S}_A)$ which we now refer to as the *full effective medium*. 208 The anisotropic component of the full effective medium given by $\mathcal{A}(\mathcal{H}(\mathbf{S}))$ will be re-209 ferred hereafter as the *full effective anisotropy* and its isotropic component $\mathcal{I}(\mathcal{H}(\mathbf{S}))$ is 210 the full effective isotropy. We will symbolize the full effective radial anisotropy corre-211 sponding to $\mathcal{A}(\mathcal{H}(\mathbf{S}))$ with ξ^* . 212

On the other hand, the homogenized counterpart of a pure anisotropic Earth (*i.e.*, a model where only the anisotropic component varies spatially) is $\mathcal{H}(\mathcal{A}(\mathbf{S})) =$ $\mathcal{H}(\mathbf{S_0} + \mathbf{S_A})$ where $\mathcal{A}(\mathcal{H}(\mathcal{A}(\mathbf{S})))$ is the *effective intrinsic anisotropy*. The *effective intrinsic radial anisotropy* corresponding to $\mathcal{A}(\mathcal{H}(\mathcal{A}(\mathbf{S})))$ will then be designated as ξ^*_{CPO} . Note that due to the non-linearity of \mathcal{H} , homogenization creates apparent isotropic heterogeneities in the elastic tensor $\mathcal{I}(\mathcal{H}(\mathcal{A}(\mathbf{S})))$ as a byproduct, albeit of low amplitude.

Finally, the tomographic counterpart of a pure isotropic Earth (*i.e.*, a model where the isotropic component varies spatially, and the anisotropic component is zero) is $\mathcal{H}(\mathcal{I}(\mathbf{S})) = \mathcal{H}(\mathbf{S_0} + \delta \mathbf{S_I})$ where the non-negligible apparent anisotropic component due to SPO $\mathcal{A}(\mathcal{H}(\mathcal{I}(\mathbf{S})))$ is called *extrinsic anisotropy*. Here, *extrinsic radial anisotropy* will be denoted by ξ_{SPO}^* (Refer to Figure 1 for a comprehensive summary).

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3 Analytical expressions in the 1-D case

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3.1 Backus homogenization

A vertically transverse isotropic (VTI) medium is an elastic medium with hexagonal symmetry and vertical symmetry axis. It can be described by five elastic parameters A, C, F, L, and N, also known as the Love parameters (Love, 1906). Supposing that axis 3 is the symmetry axis, the local **S** for a VTI solid can be expressed in Mandel 231 notation as:

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$$\mathbf{S} = \begin{pmatrix} A & A-2N & F & 0 & 0 & 0 \\ A-2N & A & F & 0 & 0 & 0 \\ F & F & C & 0 & 0 & 0 \\ 0 & 0 & 0 & 2L & 0 & 0 \\ 0 & 0 & 0 & 0 & 2L & 0 \\ 0 & 0 & 0 & 0 & 0 & 2N \end{pmatrix},$$
(5)

and the level of shear wave radial anisotropy can be written as:

$$\xi = \frac{N}{L}.\tag{6}$$

Backus (1962) explicitly showed analytical upscaling relations for seismic waves propagating in a 1-D stratified medium. For a 1-D layered medium where each layer is a VTI medium, the long-wavelength effective medium is also a VTI medium. The effective equivalent of the elastic constants, for instance, N and L concerning the shear wave velocities are given by an arithmetic mean and a harmonic mean, respectively:

$$N^* = \langle N \rangle, \tag{7}$$

$$L^* = \langle 1/L \rangle^{-1} \,, \tag{8}$$

where $\langle . \rangle$ refers to the spatial average over a wavelength of any periodic function (in this case, N and 1/L), and * denotes a long wavelength effective property. The effective density ρ^* is simply the arithmetic mean of the local density ρ :

 $\rho^* = \langle \rho \rangle \,. \tag{9}$

The effective shear wave radial anisotropy ξ^* is essentially the ratio between the effec-

tive equivalents of N and L:

$$\xi^* = \frac{N^*}{L^*} = \langle N \rangle \langle 1/L \rangle \,. \tag{10}$$

In this way, for a 1-D fine-scale medium where each layer is isotropic (N = L), the long-wavelength effective medium is transversely isotropic, and the level of extrinic radial anisotropy is given by $\langle N \rangle \langle 1/N \rangle$ (Alder et al., 2017).

3.2 An analytical expression to quantify CPO and SPO in a 1-D lay ered media

Let us consider an intrinsically anisotropic (CPO component) and finely-layered (SPO component) VTI medium. Similar as to how we defined our reference medium in section 2, we regard the elastic parameters N and L as the sum of an isotropic component defined by the shear moduli μ , and local anisotropic perturbations N_A and L_A , respectively:

$$N(z) = \mu(z) + N_A(z),$$
 (11)

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$$L(z) = \mu(z) + L_A(z).$$
 (12)

The isotropic equivalent of μ can be computed as (J.-P. Montagner, 2007; Maupin et al., 2007):

$$\mu = \frac{1}{15}(C + A - 2F + 6L + 5N), \tag{13}$$

where A and C are elastic parameters concerning P-waves, and F relates to the so-called 'ellipticity'(*i.e.*, the velocity along the direction interposing fast and slow velocities). Assuming no P-wave anisotropy and setting F to unity, one can simplify equation (13) to:

$$\mu = \frac{1}{3}(2L + N). \tag{14}$$

(16)

Knowing equations (6) and (14), one can re-write N and L in terms of μ and $\xi_{CPO} = N/L$ giving:

$$N = \xi_{\rm CPO} \frac{3\mu}{2 + \xi_{\rm CPO}},\tag{15}$$

$$L = \frac{3\mu}{2 + \xi_{\rm CPO}}.$$

It is now straightforward to determine the anisotropic components N_A and L_A by equating equations (11) and (15). For the sole purpose of segregating the isotropic and anisotropic components, the forms of N and L consistent with equation (1) is therefore:

$$N(z) = \mu(z) + \left(\xi_{\rm CPO}(z) \frac{3\mu(z)}{2 + \xi_{\rm CPO}(z)} - \mu(z)\right),\tag{17}$$

$$L(z) = \mu(z) + \left(\frac{3\mu(z)}{2 + \xi_{\rm CPO}(z)} - \mu(z)\right),\tag{18}$$

where any variable as a function of z implies variations in space.

To calculate the long-wavelength effective equivalent of such a medium, let us first write the parameters μ and ξ_{CPO} as:

 $\mu(z) = \langle \mu \rangle + \delta \mu(z), \tag{19}$

$$\xi_{\rm CPO}(z) = \langle \xi_{\rm CPO} \rangle + \delta \xi_{\rm CPO}(z), \tag{20}$$

where $\langle \mu \rangle$ and $\langle \xi_{\rm CPO} \rangle$ are the spatially-averaged counterparts. $\delta \mu$ and $\delta \xi_{\rm CPO}$ are smallscale heterogeneities in the shear modulus and intrinsic radial anisotropy, respectively. They verify $\langle \delta \mu \rangle$ and $\langle \delta \xi_{\rm CPO} \rangle = 0$. The long-wavelength effective equivalents N^* and $1/L^*$ are:

$$N^* = \langle N \rangle = \left\langle \xi_{\rm CPO} \frac{3\mu}{2 + \xi_{\rm CPO}} \right\rangle = \left\langle (\langle \xi_{\rm CPO} \rangle + \delta \xi_{\rm CPO}) \frac{3(\langle \mu \rangle + \delta \mu)}{2 + \langle \xi_{\rm CPO} \rangle + \delta \xi_{\rm CPO}} \right\rangle, \tag{21}$$

$$\frac{1}{L^*} = \langle 1/L \rangle = \left\langle \frac{2 + \xi_{\rm CPO}}{3\mu} \right\rangle = \left\langle \frac{2 + \langle \xi_{\rm CPO} \rangle + \delta \xi_{\rm CPO}}{3(\langle \mu \rangle + \delta \mu)} \right\rangle.$$
(22)

We can simplify equations (21) and (22) by assuming a weak contrast in the shear modulus $\delta \mu \ll \langle \mu \rangle$. Using a second-order Taylor expansion, we get:

$$N^* \approx \frac{3\langle\mu\rangle}{2 + \langle\xi_{\rm CPO}\rangle} \bigg(\langle\xi_{\rm CPO}\rangle - \frac{2}{(2 + \langle\xi_{\rm CPO}\rangle)^2} \langle\delta\xi_{\rm CPO}^2\rangle + \frac{2}{\langle\mu\rangle(2 + \langle\xi_{\rm CPO}\rangle)} \langle\delta\mu \cdot \delta\xi_{\rm CPO}\rangle\bigg),\tag{23}$$

$$^{299}_{300} \qquad 1/L^* \approx \frac{2 + \langle \xi_{\rm CPO} \rangle}{3\langle \mu \rangle} \left(1 + \frac{1}{\langle \mu \rangle^2} \langle \delta \mu^2 \rangle - \frac{1}{\langle \mu \rangle (2 + \langle \xi_{\rm CPO} \rangle)} \langle \delta \mu \cdot \delta \xi_{\rm CPO} \rangle \right). \tag{24}$$

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Using equation (10), we multiply equations (23) and (24) and neglect terms higher than order two to obtain the full effective radial anisotropy ξ^* due to both fine-layering and intrinsic anisotropy:

$$\xi^* \approx \langle \xi_{\rm CPO} \rangle - \frac{2}{(2 + \langle \xi_{\rm CPO} \rangle)^2} \langle \delta \xi_{\rm CPO}^2 \rangle + \frac{\langle \xi_{\rm CPO} \rangle}{\langle \mu \rangle^2} \langle \delta \mu^2 \rangle + \frac{2 - \langle \xi_{\rm CPO} \rangle}{\langle \mu \rangle (2 + \langle \xi_{\rm CPO} \rangle)} \langle \delta \mu \cdot \delta \xi_{\rm CPO} \rangle. \tag{25}$$

Equation (25) explicitly shows the separate effects of the small-scales in the 305 isotropic component and in the intrinsically anisotropic component onto the effective 306 anisotropy as 'seen' by long-period seismic waves. Assuming the medium to be de-307 void of intrinsic anisotropy (*i.e.*, $\xi_{CPO} = 1$ and $\delta \xi_{CPO} = 0$), the full effective radial 308 anisotropy ξ^* directly relates to the level of heterogeneities in the shear moduli. Here, 309 $\langle \delta \mu^2 \rangle$ refers to the variance of small-scale heterogeneities of the shear modulus $\delta \mu$. 310 It can be interpreted as the extrinsic radial anisotropy ξ_{SPO}^* due to the seismically 311 unresolved small-scale isotropic heterogeneities. It varies as the square of the hetero-312 geneities, in agreement with the result of Alder et al. (2017). 313

On the other hand, when the isotropic component is uniform (*i.e.*, $\delta \mu = 0$), ξ^* also varies with the square of heterogeneities, but now in intrinsic anisotropy. This can be interpreted as the effective intrinsic radial anisotropy ξ^*_{CPO} , i.e. the intrinsic anisotropy that gets smoothed out as a result of upscaling. Interestingly, its overall effect is to reduce the level of intrinsic anisotropy as indicated by the minus sign in front of the second term. In the absence of small-scale isotropic heterogeneities, we anticipate anisotropy to be always underestimated by tomography.

Finally, equation (25) suggests the existence of a cross-term $\langle \delta \mu \cdot \delta \xi_{CPO} \rangle$ due to the spatial correlation between intrinsic anisotropy and shear modulus. Supposing spatial variations in both components are significant such as at major seismic discontinuities, the correlation term should influence the anisotropy mapped in tomographic models. Nevertheless, it is propounded that this correlation term is usually negligible and hence may need not be accounted for (Bakulin, 2003).

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Similarly, the effective Voigt-averaged shear modulus μ^* is given by:

$$\mu^* = \frac{2L^* + N^*}{3}.$$
(26)

Plugging equations (23) and (24) into equation (26), we get:

$$\mu^* = \langle \mu \rangle - \frac{2}{\langle \mu \rangle (2 + \langle \xi_{\rm CPO} \rangle)} \langle \delta \mu^2 \rangle - \frac{2 \langle \mu \rangle}{(2 + \langle \xi_{\rm CPO} \rangle)^3} \langle \delta \xi_{\rm CPO}^2 \rangle + \frac{4}{(2 + \langle \xi_{\rm CPO} \rangle)^2} \langle \delta \mu \cdot \delta \xi_{\rm CPO} \rangle. \tag{27}$$

Ignoring intrinsic anisotropy (*i.e.*, $\xi_{CPO} = 1$ and $\delta \xi_{CPO} = 0$), the effective shear 331 modulus μ^* is always smaller than its spatially-averaged version $\langle \mu \rangle$. Such a result 332 is logical in the 1-D case. Here, radial anisotropy induced by fine-layering is always 333 positive (equation (25)) thereby having $N^* > L^*$. Since L 'counts' twice and N once 334 in its isotropic average, its long-wavelength effective equivalent μ^* is always slower 335 than $\langle \mu \rangle$. Contrastingly, if one neglects isotropic heterogeneities and only consider 336 variations in intrinsic anisotropy, homogenization also results in the underestimation 337 of the shear modulus. One would predict that homogenization leads to the creation 338 of apparent isotropic heterogeneities due to small-scale variations in CPO. Lastly and 339 as expected, the cross term recurs due to the spatial correlation between the shear 340 modulus and intrinsic anisotropy. 341

We acknowledge that the homogenized expressions given by equations (25) and (27) in terms of the isotropic shear modulus μ may not be particularly convenient for seismologists. In practice, we observe spatial distributions in V_S and not in μ . If one assumes that density is uniform, then $\delta \mu / \mu$ can be simply replaced by $2\delta V_S / V_S$. On the other hand, if one assumes that density increases with V_S , one could also establish long-wavelength effective expressions for V_S in the same manner as μ using simple empirical relations for density such as that of Tkalčić et al. (2006).

In the Earth's asthenosphere, we do not expect velocity variations (*i.e.* $\Delta V_S/(V_{S1}+$ 349 V_{S2})) oftentimes to be larger than 5% (e.g. Xu et al., 2008; Stixrude & Jeanloz, 2015). 350 This roughly translates to $\sim 10\%$ heterogeneities in the shear modulus assuming con-351 stant density ρ . To perform a numerical estimate, let us examine a stack of planar 352 layers with alternating shear moduli values determined by $\pm \Delta \mu/2$ (Figure 2a middle 353 panel) corresponding to shear wave velocity variations $\Delta V_S/(V_{S1}+V_{S2})$ of about 10%. 354 The 1-D depth profiles depict periodic variations with layers of equal thicknesses of 355 20 km. Positive intrinsic radial anisotropy ($\xi = 1.2$) is prescribed in the even layers, 356 whereas the odd layers are isotropic ($\xi = 1$) (Figure 2a right panel). After upscaling 357 over a wavelength much larger than 20 km, the resulting profiles for N^* and L^* are 358 homogeneous, and simply given by their arithmetic and harmonic means, respectively 359 (Figure 2a left panel). Once the long-wavelength effective N^* and L^* are acquired, we 360 can compute the full effective radial anisotropy ξ^* through equation (10) (solid red line 361 in Figure 2a right panel), and the effective shear modulus μ^* through equation (26) 362 (solid red line in Figure 2a middle panel). Figure 2b illustrates a different scenario 363 where ξ only exists in the odd layers (Figure 2b right panel). In essence when the 364 shear modulus and intrinsic anisotropy are uncorrelated, the homogenized parameters 365 μ^* and ξ^* should be the same regardless. However, a slight offset in μ^* and ξ^* of 366 Figure 2b with respect to Figure 2a can be observed which is exclusively attributed to 367 this cross term as hinted by equations (25) and (27). Strictly speaking, the reduction 368 in the amplitude of the effective properties arises from the switch in signs in the cross 369 term from positive to negative $\langle \delta \mu \cdot \delta \xi \rangle$, implying that in the second scenario, the shear 370 modulus and intrinsic anisotropy are now anti-correlated. 371

To validate the second-order approximation, we also show ξ^* and μ^* using equa-372 tions (25) and (27) respectively (dashed blue lines in Figures 2a and 2b middle and 373 right panels). Clearly by applying equation (25), the intrinsic component (first term) 374 contributes the most to the effective anisotropy with $1 - \langle \xi_{\rm CPO} \rangle = 0.1$ wherein its spa-375 tial variations' overall effect is to tone-down the amplitude of anisotropy by an amount 376 of $\sim -10^{-3}$ (second term). This is followed by the SPO component (third term) which 377 is responsible for the amplification of anisotropy (~ $+10^{-2}).$ Lastly, the cross term 378 provides the least contribution (~ $\pm 10^{-3})$ and therefore can reasonably be ignored in 379 most cases. The \pm sign denotes that it may increase or decrease anisotropy depending 380 on the coupling pattern between the shear modulus and intrinsic anisotropy. 381

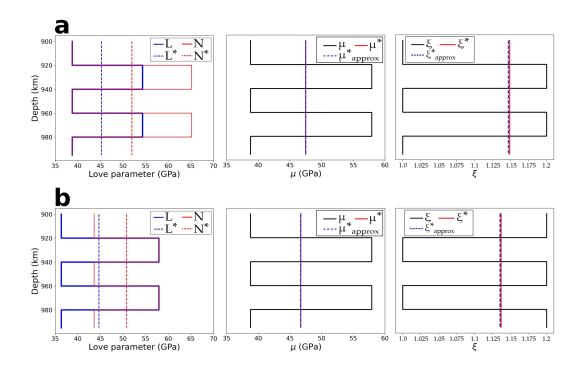


Figure 2. 1-D binary and periodic media with 10% isotropic heterogeneities in terms of $\Delta V_S/(V_{S1} + V_{S2})$ prescribed across: (a) even layers, and (b) odd layers. Upon homogenization, the resulting profiles are homogeneous (variables denoted by (*)). The dashed blue lines at the middle (μ_{approx}^*) and right panels (ξ_{approx}^*) correspond to the predicted long-wavelength effective equivalents using the second order approximations from equations (27) and (25), respectively. The difference in the homogenized shear moduli and radial anisotropy between (a) and (b) is attributed to the cross term as implied by equation (25). Since the medium is periodic, it is enough to only display a portion of the medium.

3.3 Composite law for radial anisotropy

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In this section, we show how the total effective anisotropy can be expressed in terms of its intrinsic and extrinsic contributions. For that, we investigate two special cases: (1) a purely isotropic 1-D layered medium, (2) a purely anisotropic 1-D medium (*i.e.*, no spatial variations in isotropic component), and find equivalent expressions for extrinsic radial anisotropy ξ_{SPO}^* and effective intrinsic radial anisotropy ξ_{CPO}^* . By doing so, we elicit a simple composite law related to equation (25) that can be extrapolated to 2-D and 3-D media.

In the case of an isotropic medium with spatially-varying shear modulus, the radial anisotropy is entirely due to SPO. Equation (25) reforms into:

$$\xi_{\rm SPO}^* \approx 1 + \frac{1}{\langle \mu \rangle^2} \langle \delta \mu^2 \rangle.$$
 (28)

On the other hand, a purely anisotropic medium without spatial variations in the shear modulus leads to an effective intrinsic anisotropy:

$$\xi_{\rm CPO}^* \approx \langle \xi_{\rm CPO} \rangle - \frac{2}{(2 + \langle \xi_{\rm CPO} \rangle)^2} \langle \delta \xi_{\rm CPO}^2 \rangle.$$
⁽²⁹⁾

By taking the product between equations (28) and (29), neglecting terms higher than order two, one has simply:

$$\xi_{\rm CPO}^* \times \xi_{\rm SPO}^* \approx \langle \xi_{\rm CPO} \rangle - \frac{2}{(2 + \langle \xi_{\rm CPO} \rangle)^2} \langle \delta \xi_{\rm CPO}^2 \rangle + \frac{\langle \xi_{\rm CPO} \rangle}{\langle \mu \rangle^2} \langle \delta \mu^2 \rangle, \tag{30}$$

which is approximately equal to ξ^* in equation (25) but without the cross term. Therefore, ignoring spatial correlations between intrinsic anisotropy and shear modulus, the full effective radial anisotropy can be quantified through the following composite law:

$$\xi^* \approx \xi^*_{\rm CPO} \times \xi^*_{\rm SPO}.$$
 (31)

In practice, ξ^* can be estimated from a tomographic inversion (Debayle & Ken-403 nett, 2000; Plomerová et al., 2002; Gung et al., 2003; Nettles & Dziewoński, 2008a; 404 Fichtner et al., 2010). Seismologists often compare ξ^* with the intrinsic radial anisotropy 405 $\xi_{\rm CPO}$ computed from a geodynamically-based CPO model (Becker et al., 2003, 2006; 406 Ferreira et al., 2019; Sturgeon et al., 2019). The comparison should be done instead 407 with an effective model ξ^*_{CPO} , which is difficult to estimate without access to any 408 elastic homogenization tools. Furthermore, equation (25) suggests that there is a non-409 negligible extrinsic component of anisotropy due to the unresolved small-scale isotropic 410

heterogeneities. While it is difficult to rigorously establish analytical solutions in the case of a 2-D/3-D complex media, following the logic above, we hypothesize that the mismatch often observed between homogenized CPO models and tomographic models is the extrinsic radial anisotropy ξ_{SPO}^* .

415 4 Methods for 2-D media

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4.1 Homogenization in 2-D and in 3-D media

The classic homogenization method of Backus is only applicable in 1-D to media 417 exhibiting spatial periodicity. The true Earth, however, is a complex 3-D and multi-418 scale medium. To alleviate this problem and quantify effective elastic properties in 419 a mantle-like medium, we rely on the non-periodic elastic homogenization technique 420 developed by Capdeville and Marigo (2007); Capdeville et al. (2010); Guillot et al. 421 (2010); Capdeville et al. (2015). Originally, this method has been developed as a pre-422 processing step enabling one to solve the elastostatic wave equation using a simple 423 mesh, speeding up the computations for wave propagation in complex media. It has 424 also been used to improve the convergence and computational cost of full waveform 425 inversion (Capdeville & Métivier, 2018; Hedjazian et al., 2021). 426

In the non-periodic case, the homogenization is not performed with respect to the 427 periodicity of the medium, but with respect to the minimum wavelength present in the 428 wavefield. The assumption that this minimum wavelength λ_{\min} exists is required for 429 non-periodic medium with no scale separation such as the true Earth. Scales smaller 430 than λ_{\min} are seen by the wavefield only through their effective properties. To sep-431 arate the small and the large scales, we need to define a threshold wavelength λ_h , 432 called the homogenization wavelength. λ_h is a user-defined parameter, and all scales 433 smaller than λ_h are homogenized. Numerical examples suggest that, for all natural 434 media, homogenization with a value $\lambda_h = 0.5 \lambda_{\min}$ is sufficient to accurately reproduce 435 the complete wavefield (Capdeville et al., 2010). Hence, this value is chosen in the 436 rest of the present study. Computing the effective properties of an elastic medium 437 with homogenization wavelength λ_h requires to solve an elastostatic problem numer-438 ically. To do this, we use the 3-D Fast-Fourier Homogenization algorithm developed 439 by Capdeville et al. (2015). 440

Assuming a good data coverage, Capdeville and Métivier (2018) numerically verified that for complex elastic media, a seismic tomography model and the homogenized model are in agreement at spatial wavelengths higher than λ_h . Hence, homogenization can be viewed as a first-order tomographic operator. We will consider the homogenized model as the best image one could get from seismic tomography. This can be translated to:

$$\mathbf{S}^* = \mathcal{H}(\mathbf{S}) \tag{32}$$

where \mathcal{H} is the tomographic operator, **S** is the reference medium, and the homogenized model **S**^{*} is the full effective medium (*i.e.*, the best recovered image as seen by a wavefield of a given minimum wavelength λ_{\min} and assuming perfect data coverage). In this paper, we apply this 'tomographic operator' to a 2-D composite medium by upscaling the marble cake model in the presence of deformation-induced anisotropy.

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4.2 Isotropic heterogeneities in a 2-D mechanically-mixed mantle

To define our 2-D incompressible flow model imitating mantle convection, we use a stream function similar to that of Alder et al. (2017):

$$\Psi(x, z, t) = \sin(a\pi z) \left[\sin(b\pi x) + \alpha(t) \sin((b+1)\pi x) + \beta(t) \sin((b+2)\pi x) \right]$$
(33)

where $\alpha(t)$ and $\beta(t)$ are sinusoidal functions of time that introduces chaotic mixing. The variables *a* and *b* relate to the number of distinguishable convection cells and are chosen arbitrarily. The form of the function Ψ ensures free-slip boundary conditions. Finally, the resulting velocity field is scaled using a reference value of 1 cm·yr⁻¹.

We replicate the marble cake patterns by deforming a circular anomaly at the center of the box using our prescribed flow field. To do this, control points that define the contour of the anomaly are advected using fourth-order Runge Kutta methods with variable time-stepping (Press et al., 1992). To achieve a final configuration for the anomaly, we define an advection mixing time $T_{\rm SPO}$. Figure 3 illustrates the evolution of the pattern when subjected to the flow field defined in equation (33). Setting $a = 1, b = 2, \text{ and } T_{\rm SPO} = 75$ My, we have a mechanically-mixed medium with two characteristic convection cells.

Using the last panel of Figure 3, the binary system is defined by assigning a reference S-wave velocity value $V_{S_2} = 4.52 \text{ km} \cdot \text{s}^{-1}$ to the yellow region, and $V_{S_1} =$

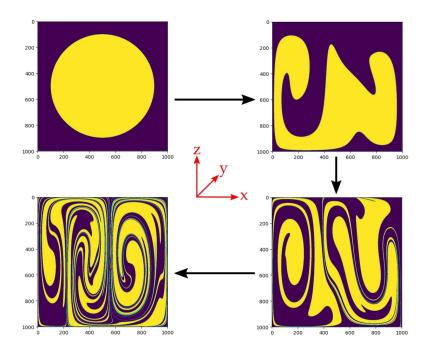


Figure 3. Initially a circle, the anomaly is deformed progressively until the medium reaches a stage resembling marble cake-like patterns.

3.7 km·s⁻¹ to the purple region so that the level of isotropic heterogeneities is given by $100\% \times (V_{S_1} - V_{S_2})/(V_{S_1} + V_{S_2})) = 10\%$. *P*-wave velocities are computed by imposing a constant ratio $V_P/V_S = 1.7$ (Obrebski et al., 2010). Following the work of Tkalčić et al. (2006), we compute the density ρ using the empirical relation $\rho =$ $2.35 + 0.036(V_P - 3)^2$. These values are used to define the local isotropic tensor **S**_I in equation (1).

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4.3 Modeling of Crystallographic Preferred Orientation

Crystallographic preferred orientation of mantle minerals forms due to strain 478 accumulation over time. The velocity gradient tensor field of the marble cake model 479 $\nabla \mathbf{u}$ is derived from the stream function Ψ described previously. We then model CPO 480 evolution of olivine aggregates using D-Rex, a program which calculates strain-induced 481 CPO by plastic deformation, and dynamic recrystallization (E. Kaminski et al., 2004). 482 The activities of olivine slip systems are chosen to correspond to dry mantle conditions, 483 while other parameters are taken as in the reference D-Rex model. A time scale for 484 CPO evolution $T_{\rm CPO}$ is defined to control the level of intrinsic anisotropy. 485

In our numerical experiments, we compute CPO everywhere irrespective of the actual mineralogical phase. We scale the elastic tensor derived from D-Rex so that its isotropic component is identical to the binary system derived in Section 4.2. The reference medium can be constructed from equation (1) where S_{I} now relates to the small-scale isotropic heterogeneities in the mechanically-mixed mantle, and S_{A} is the intrinsically anisotropic component computed with D-Rex.

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4.4 Quantifying the level of anisotropy

In this section, we define two ways to quantify the level of seismic anisotropy for any given elastic tensor **S**. The first one is radial anisotropy. We project the elastic tensor in terms of an azimuthally-averaged vertically transverse isotropic (VTI) medium to obtain a tensor described as in equation (5). Here, the parameters L and N can be computed from **S** as follows (J.-P. Montagner & Nataf, 1986):

$$L = \frac{1}{2}(S_{44} + S_{55}) \tag{34}$$

$$N = \frac{1}{8}(S_{11} + S_{22}) - \frac{1}{4}S_{12} + \frac{1}{2}S_{66}.$$
(35)

The level of radial anisotropy is then given by equation (6).

Another convenient way to quantify anisotropy is to compute the percentage of total anisotropy by taking the L2-norm fraction of the anisotropic part of the elastic tensor with respect to the isotropic part. This quantity is called the anisotropy index and is given by:

anisotropy index =
$$\frac{||\mathbf{S} - \mathbf{S}_{\mathbf{I}}||}{||\mathbf{S}_{\mathbf{I}}||}$$
. (36)

507 5 Elastic homogenization of a 2-D mechanically-mixed mantle in the 508 presence of CPO

Figure 4 displays some seismic properties of the reference medium \mathbf{S} before and 509 after homogenization. The left panels are the true structures, whereas the middle and 510 right panels are the structures equating to the full effective medium $\mathcal{H}(\mathbf{S})$ at homog-511 enization wavelengths λ_h of 200 km and 500 km, respectively. The first row depicts 512 the shear wave velocities. The anisotropy characterized by its radial component, and 513 by its norm fraction are depicted in the second row and in the third row, respectively. 514 Each pixel initially contains an isotropic part derived from the marble cake model with 515 a mixing time for advection $T_{\rm SPO} \sim 75 {\rm ~My}$, and an anisotropic part computed from a 516

⁵¹⁷ CPO model with a time scale for CPO evolution of $T_{\rm CPO} \sim 40$ My corresponding to a ⁵¹⁸ moderately developed crystal fabric. Hereafter, we run our simultaion in a box of size ⁵¹⁹ 1000 × 1000 km, define $\lambda_{\rm max} = 1000$ km as the length scale of our simulations.

Several glaring features can be observed such as the presence of positive radial anisotropy ($\xi > 1$) at the top and bottom boundaries where flow is sub-horizontal, and likewise negative ($\xi < 1$) at regions where the flow is sub-vertical. As expected, homogenization results in the smoothing of the structures with the level of smoothing modulated by λ_h . However, homogenization is not just a simple spatial average but a product of highly non-linear upscaling relations. With increasing homogenization wavelengths, the full effective medium becomes devoid of anisotropy in some areas.

After decomposing S into an isotropic tensor $\mathcal{I}(S)$ and an anisotropic tensor $\mathcal{A}(S)$ 527 through equations (2) and (4), one can also homogenize and analyze each component 528 separately, i.e. $\mathcal{H}(\mathcal{I}(\mathbf{S}))$ and $\mathcal{H}(\mathcal{A}(\mathbf{S}))$. Figure 5 shows the level of effective radial 529 anisotropy of these two separate components after homogenization. The top panels 530 recreate the results of Alder et al. (2017). Indeed, homogenizing the fine-layered 531 isotropic medium produces extrinsic radial anisotropy ξ_{SPO}^* (*i.e.*, radial anisotropy of 532 model $\mathcal{H}(\mathcal{I}(\mathbf{S})))$. Notice that the patterns of effective intrinsic radial anisotropy and 533 extrinsic radial anisotropy maps are roughly similar. For example, they both induce a 534 positive radial anisotropy $\xi > 1$ in the horizontal layers: the stretched heterogeneities 535 that induce SPO become elongated along the direction of the maximum principal strain 536 rate that also controls the CPO. 537

Figure 6 depicts the apparent isotropic heterogeneities created upon homogenization of $\mathcal{A}(\mathbf{S})$. It produces maximum velocity perturbations of about 0.25 % at $\lambda_h = 200$ km and 0.2 % at $\lambda_h = 500$ km. It appears to be a small effect, especially considering the large and sharp variations of intrinsic anisotropy in our CPO model.

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To better illustrate the behaviour of different contributions to anisotropy, we plot in Figure 7 the amplitude of radial anisotropy (in terms of its standard deviation over the entire 2-D model domain) against the wavelength of homogenization λ_h . In the following cases, the intrinsic anisotropy component of the reference medium **S** is computed for a CPO developing over increasing duration T_{CPO} of 5, 40, or 75 Myr. Several point can be noted in Figure 7:

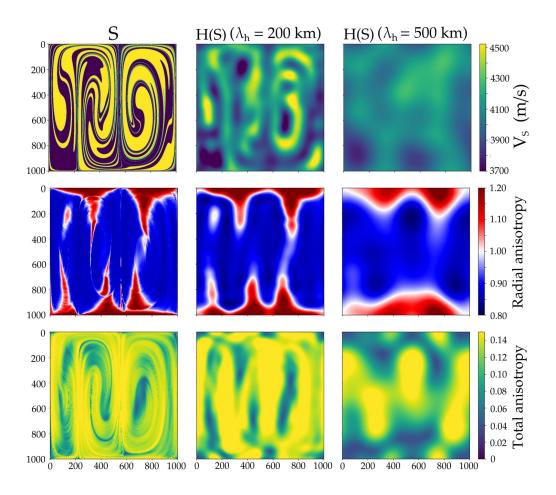


Figure 4. Seismic properties of the reference medium **S** before and after homogenization. The model dimensions are 1000 km × 1000 km × 1000 km, with cell sizes of 5 km × 5 km × 5 km. Here, each pixel contains an **S** which consists of small-scale isotropic heterogeneities and an intrinsically anisotropic perturbation computed with D-Rex (E. Kaminski et al., 2004). The present-day marble cake patterns correspond to a mixing time for advection $T_{\text{SPO}} \sim 75$ My, whereas the time scale for CPO evolution is $T_{\text{CPO}} \sim 40$ My. We homogenized **S** using the Fast-Fourier homogenization algorithm of Capdeville et al. (2015). (From left to right) First row: V_s models derived from **S**, $\mathcal{H}(\mathbf{S})$ at $\lambda_h = 200$ km, and $\mathcal{H}(\mathbf{S})$ at $\lambda_h = 500$ km. Second row: ξ_{CPO} , ξ^* at $\lambda_h = 200$ km, and ξ^* at $\lambda_h = 500$ km. Last row: Total anisotropy in terms of the norm fraction of **S**, $\mathcal{H}(\mathbf{S})$ at $\lambda_h = 200$ km, and $\mathcal{H}(\mathbf{S})$ at $\lambda_h = 500$ km. Elastic homogenization can be viewed as the best possible model reconstructed by seismic tomography assuming perfect ray-path coverage.

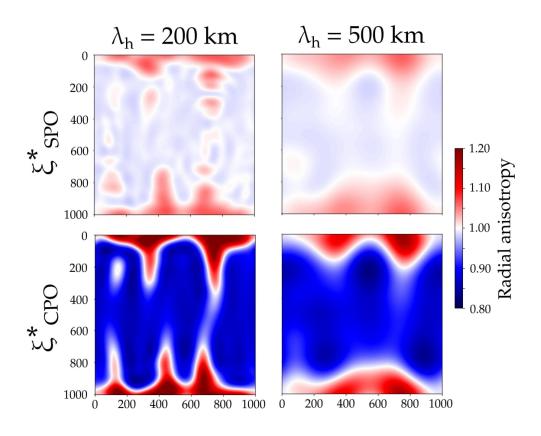


Figure 5. Extrinsic radial anisotropy ξ_{SPO}^* (*i.e.*, radial anisotropy of model $\mathcal{H}(\mathcal{I}(\mathbf{S}))$) (top panels) at two different wavelengths of homogenization λ_h . Here, $\xi_{\text{SPO}}^* > 1$ is now interpreted as horizontal layering whereas < 1 as vertical layering. The bottom panels show the effective intrinsic radial anisotropy ξ_{CPO}^* (*i.e.*, radial anisotropy of model $\mathcal{H}(\mathcal{A}(\mathbf{S}))$).

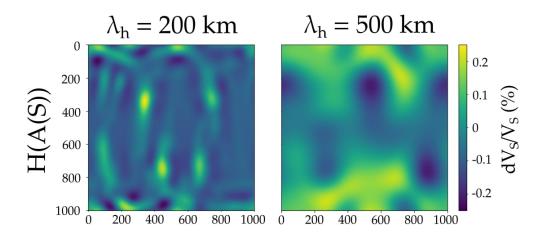


Figure 6. Apparent isotropic velocity perturbations with respect to a mean velocity V_S at two different wavelengths of homogenization λ_h . $\mathcal{H}(\mathcal{A}(\mathbf{S}))$ pertains to the homogenized model of a purely anisotropic medium. Even when placed in a very favorable scenario for intrinsic anisotropy, homogenizing a purely anisotropic medium produces a meager 0.25% artificial heterogeneities at $\lambda_h = 200$ km and 0.2% at $\lambda_h = 500$ km.

(i) The resulting intrinsic radial anisotropy $\xi_{\rm CPO}$ in terms of its standard deviation over the entire region (dashed lines) increases with $T_{\rm CPO}$, although some saturation is observed (*i.e.*, the orientation of crystals depends mostly on their recent deformation, and lose the memory of the deformation they underwent too long ago).

(ii) The level of intrinsic anisotropy is diminished upon homogenization. ξ^*_{CPO} 552 (hollow squares) is always lower than the reference value ξ_{CPO} (dashed lines), and 553 diminishes with λ_h . This effect can be easily understood. For small λ_h , the wavelength 554 of homogenization is small compared to the scale of deformation patterns (of order 100 555 km). At each point of the 2-D map, the direction of CPO is therefore locally constant 556 over λ_h , which yields $\xi^*_{CPO} \approx \xi_{CPO}$. At larger scales, when λ_h increases compared to 557 the scale of convection, this direction becomes likely random and CPO heterogeneities 558 averaged over λ_h have different orientations: there is less of a preferential direction 559 and the averaged level of CPO anisotropy is diminished. 560

(iii) On the contrary, the full effective radial anisotropy ξ^* at short wavelengths of homogenization λ_h is larger than ξ_{CPO} . This is in agreement with the analytical expression given by equation (25). This additional anisotropy is of course due to the existence of SPO (black circles) which reinforces the total level of effective anisotropy.

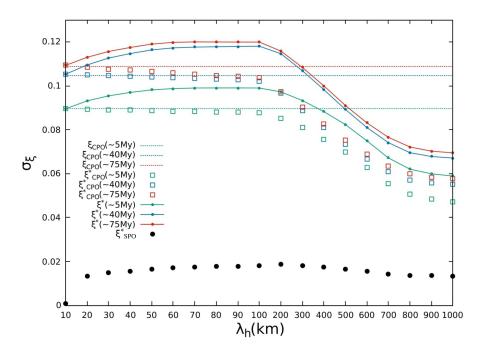


Figure 7. Effective radial anisotropy in terms of its standard deviation σ_{ξ} over the entire 2-D image, plotted as a function of homogenization length. The time scales indicated in million years pertain to the evolution history of CPO (a larger time scale leads to stronger CPO). Dashed lines represent the standard deviation of ξ_{CPO} in model **S** and serve as reference values. In this experiment, ξ_{SPO}^* of model $\mathcal{H}(\mathcal{I}(\mathbf{S}))$ (black circles) deemed to be five times smaller than ξ_{CPO}^* of model $\mathcal{H}(\mathcal{A}(\mathbf{S}))$ (hollow squares). Since SPO is mostly in-phase with CPO, the two anisotropic components add constructively giving the full effective radial anisotropy ξ^* (solid line-dots).

(iv) Both ξ^*_{CPO} and ξ^* converge toward ξ_{CPO} at infinitely short homogenization wavelengths. Only in this unrealistic case (*i.e.*, the perfect recording of the seismic wavefield up to infinitely short periods), would seismic tomography be able to map the true intrinsic anisotropy.

(v) Extrinsic radial anisotropy ξ_{SPO}^* here has an amplitude that is five times smaller than ξ_{CPO}^* . Such a result, of course, is specific to this numerical experiment, and that CPO is indeed stronger than SPO might not be always true. For instance, a longer mixing time would have resulted in a thinner and more complex layering that would have increased the SPO. We are unfortunately limited by the number of tracers necessary to describe the phase stirring which is exponentially increasing with time.

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5.1 Verifying the composite law $\xi^* = \xi^*_{CPO} \times \xi^*_{SPO}$ in 2-D

In this section, we aim to verify numerically equation (31) in 2-D by plotting ξ_{SPO}^* $\times \xi_{\text{CPO}}^*$ against ξ^* for each pixel in our 2-D maps of radial anisotropy. Here again, the three quantities ξ_{SPO}^* , ξ_{CPO}^* , and ξ^* are respectively computed from $\mathcal{H}(\mathcal{I}(\mathbf{S}))$, $\mathcal{H}(\mathcal{A}(\mathbf{S}))$, and $\mathcal{H}(\mathbf{S})$. In this case, there is no spatial correlation between CPO and isotropic heterogeneities in the marble cake, and the effect of the cross-term should be mitigated. Figure 8b shows this for two different homogenization wavelengths λ_h . We can see that the relation holds exceptionally well even for large λ_h .

In practice however, tomographic models of ξ^* are interpreted in terms of intrinsic 583 anisotropy, and directly compared with ξ_{CPO} computed from CPO models (Becker et 584 al., 2003, 2006; Ferreira et al., 2019). We mimic this scenario by comparing $\xi_{\rm SPO}^*$ 585 $\times \xi_{\rm CPO}$ instead with ξ^* (Figure 8a). As it turns out, the relation only holds for 586 small values of λ_h . At larger values of λ_h , the trend appears to be more dispersed as a 587 consequence of the averaging process, losing its viability to some extent. In the absence 588 of a homogenized CPO model, we project that this composite law would remain true 589 in general under the condition that the minimum wavelength used in tomography is 590 sufficiently small. 591

To test the effect of the rigidity-anisotropy cross-term, we consider another mantle model where CPO is only present in one of the two phases of the 2-D marble cake illustrated in Figure 3. In addition, we increase the percentage of isotropic heterogeneities in V_S to 15%. This increases the correlation between the shear modulus and

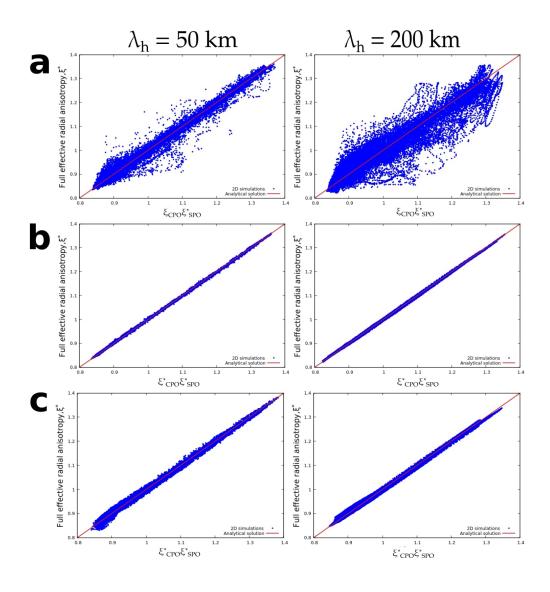


Figure 8. For each pixel *i* on our 2-D maps of ξ^* , ξ_{CPO} , ξ^*_{CPO} , and ξ^*_{SPO} Figure 8a: , we plot a cloud of points with coordinates: (a) $[\xi_{\text{CPO}} \times \xi^*_{\text{SPO}}(i), \xi^*(i)]$, (b) $[\xi^*_{\text{CPO}} \times \xi^*_{\text{SPO}}(i), \xi^*(i)]$ overlaying the analytical solution equation (31) (solid red lines). Our numerical simulations suggest that CPO models should be homogenized first before comparing with tomographic models. Figure 8c displays the effect of the cross term by increasing isotropic heterogeneities to 15% and by prescribing CPO only in the yellow phases of the marble cake model in Figure 3.

⁵⁹⁶ intrinsic anisotropy. Figure 8c displays the numerical solution at $\lambda_h = 50$ km and ⁵⁹⁷ 200 km when CPO is computed in the yellow phases alone. Based on our analytical ⁵⁹⁸ results, the points are much more spread-out than that of Figure 8b. In this scenario, ⁵⁹⁹ CPO now varies sharply and in the same places as isotropic discontinuities (*i.e.*, $\delta\xi_{CPO}$ ⁶⁰⁰ terms in equation (25) are much larger), and as expected the cross-term is much more ⁶⁰¹ apparent. Nonetheless, this only produces small departures from the composite law ⁶⁰² (red line), implying that the predictions carried out by the composite law are robust.

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5.2 Discussion

We investigated the effects of elastic homogenization to a specific class of finescale, marble cake-like models of the mantle in the presence of deformation-induced anisotropy. The homogenization procedure can be viewed as a tomographic operator applied to a reference elastic model (Capdeville et al., 2013).

We showed that the extrinsic radial anisotropy produced by fine-layering could 608 reach up to 2% (see Figure 7) assuming 10% of isotropic heterogeneities. This anisotropy 609 is much lower than the one induced by CPO where the effective intrinsic radial 610 anisotropy could peak at nearly 11%. This result is however modulated by some 611 parameters that regulate the level of effective anisotropy. For example, the layered fil-612 aments contrived from our marble cake models are of the order 10-100 km whereas of 613 those proposed by Allègre and Turcotte (1986) are much thinner and can stretch even 614 further down to the centimeter scale, which would induce larger extrinsic anisotropy. 615 Next is more of on the intrinsic aspect. Since CPO results from finite strain accumu-616 lation over time, the amplitude of intrinsic anisotropy increases with the time scale for 617 CPO evolution $T_{\rm CPO}$. Such presumptions may only be valid in regions where rock de-618 formation varies over extended periods of time, although recrystallization and damage 619 would limit the CPO that can be eventually accumulated (Ricard & Bercovici, 2009). 620 Furthermore, we considered olivine of type-A crystal fabric as the solitary anisotropic 621 mineral in our mantle models. Because of this, the intrinsic anisotropy produced from 622 finite deformation should be seen as an upper bound. Inclusion of other anisotropic 623 minerals such as pyroxene which make up a fraction in mantle periodotite (Maupin 624 & Park, 2015) would change the net anisotropy. For instance, we anticipate that in-625 cluding a substantial amount of enstatite would dilute the amount of anisotropy (e.g. 626 E. Kaminski et al., 2004). Therefore, whether CPO accounts for most of the bulk 627

anisotropy observed in tomographic images remains inconclusive and needs furtherverification.

In light of the simulations conducted, we expect large-scale anisotropy to be 630 only overestimated when CPO coexists with significant SPO as exemplified in our 631 simulations. In the absence of SPO, homogenization can only decrease the strength of 632 anisotropy. By accounting for both contributions, we showed that $\xi > 1$ is attributed 633 to a combination of lateral flow and horizontal layering, and $\xi < 1$ is a combination 634 of flow ascent and vertical layering. Indeed, the direction of shear not only dictates 635 the preferred orientation of the anisotropic minerals, but also of the orientation of the 636 folded strips that gives rise to fine-layering. 637

The repercussion of homogenizing intrinsic anisotropy alone amounts to the 638 convection-scale averaging of the CPO as evidenced by our simulations. By defini-639 tion, intrinsic anisotropy develops from the preferred alignment of the minerals at the 640 crystal level. This explains why we can map-out the fast axis of anisotropy at an 641 arbitrary location. When long period observations sample an intrinsically anisotropic 642 medium, the wavefield spatially-averages these orientations. As a result, preferential 643 orientations that are products of imbricated convection tend to appear more het-644 erogeneous, thereby ostensibly losing intrinsic anisotropy upon homogenization. In 645 contrast, spatially-coherent preferential orientations produced by simpler convection 646 patterns are less susceptible to the dilution of intrinsic anisotropy. 647

The applicability of equation (31) in a 2-D complex media may be of interest 648 to geodynamicists and tomographers alike. Not only does it permit one to directly 649 quantify the discrepancy between the full effective radial anisotropy inferred from 650 a tomographic model and the effective intrinsic radial anisotropy computed from a 651 homogenized CPO model, it further solidifies the supposition that the mismatch is in-652 deed a result of extrinsic radial anisotropy due to the seismically-unresolved small-scale 653 isotropic heterogeneities. We have conducted several numerical experiments to prove 654 that the composite law still holds exceptionally well even when the rigidity-intrinsic 655 anisotropy cross term is amplified. Erroneous interpretation of a CPO model to explain 656 tomographic observations may therefore be avoided with the help of homogenization. 657

The conclusions reached in this section are based on two assumptions: (1) We held the isotropic velocity contrast at a fixed value and assumed it to be representative

of the entire mantle. In reality however, V_S variations between two-end member com-660 positions generally decrease with depth (Xu et al., 2008; Stixrude & Jeanloz, 2015). 661 This is not to mention the local presence of melt and water that contributes to the vari-662 ations in wave velocities, and hence the strength of heterogeneities which completely 663 alters the level of apparent anisotropy. (2) We disregarded the dependency of the elas-664 tic constants built from our mantle models on pressure P and temperature T. Future 665 avenues one could take would be to incorporate P - T dependence using empirical 666 relations constrained from laboratory experiments. For instance, one may compute 667 P-T dependence using first-order corrections around a reference elastic tensor at 668 ambient P-T conditions (Estey & Douglas, 1986). The availability of self-consistent 669 thermodynamic models based on free-energy minimization schemes (J. A. Connolly, 670 2005; J. Connolly, 2009) can also be employed in lieu of the simpler relations for more 671 accurate predictions of seismic wave velocities in any given bulk composition (Stixrude 672 & Lithgow-Bertelloni, 2011). 673

674 6 Separating SPO from CPO in tomographic models: Application to radial anisotropy beneath oceanic plates

676

6.1 Radial anisotropy beneath oceanic plates

Within the context of seismic tomography, surface waves offer the capability to 677 image upper-mantle structure providing an in-depth view of large-scale anisotropy. 678 Surface wave tomography images positive radial anisotropy underneath oceanic basins 679 $(V_{SH} > V_{SV})$, characterized by a layer of strong signatures lying in between $\sim 80 - 200$ 680 km depth, corresponding to the asthenosphere (e.g. J.-P. Montagner, 1985; Ekström & 681 Dziewonski, 1998; Panning & Romanowicz, 2006; Nettles & Dziewoński, 2008b). The 682 maximum positive vertical gradient of ξ^* , at ~ 80 km depth, independent of plate age, 683 is a recurrent feature in these tomographic models. This has raised questions about 684 the potential use of radial anisotropy as a marker of the lithosphere-asthenosphere 685 boundary (LAB), which is expected on the contrary to deepen with plate age (Burgos 686 et al., 2014; Beghein et al., 2019). The strong radial anisotropy in the asthenosphere 687 is usually explained by geodynamic models including CPO evolution (Becker et al., 688 2006, 2008). 689

Across the oceanic lithosphere, plate-averaged radial anisotropy (*i.e.*, all points in the radial anisotropy models with the same plate age are averaged) displays modest

-30-

levels of about 1-3%. Several models have been proposed to explain these observa-692 tions. Hansen et al. (2016) and Hedjazian et al. (2017) suggest that CPO-related radial 693 anisotropy developed below the ridge is subsequently frozen in the lithosphere, lead-694 ing to an age-independent signature. It has also been proposed quasi-laminated melt 695 structures, preserved during lithospheric thickening, can also explain this frozen-in 696 signature of anisotropy (e.g. Auer et al., 2015; Debayle et al., 2020). Hence SPO may 697 also be a potential explanatory mechanism, and a substantial fraction of the observed 698 lithospheric anisotropy may be due to small-scale isotropic heterogeneities (Wang et 699 al., 2013; Kennett & Furumura, 2015). 700

The corroboration of the composite law in a 2-D complex medium prompted us to assess the discrepancy between a tomographic model and a CPO model of uppermantle radial anisotropy underneath a mid-ocean ridge. In our hypothesis, this should commensurate to the extrinsic radial anisotropy due to the unresolved small-scales in seismic velocities.

706

6.2 The tomographic model

In conjunction with the pre-existing global V_{SV} model of the upper-mantle constrained from Rayleigh wave data DR2012 (Debayle & Ricard, 2012), we adopt the recent global V_{SH} model CAM2016SH of Ho et al. (2016) to acquire a plate-averaged 2-D profile of radial anisotropy associated with slow-spreading oceanic ridges.

The V_S models were reconstructed by independently inverting Love (for V_{SH} models) and Rayleigh (for V_{SV} models) waveforms up to the fifth overtone between the period range 50 – 250 s using an extension of the automated waveform inversion approach of Debayle (1999). We refer the reader to Debayle and Ricard (2012) and Ho et al. (2016) for a more detailed description of the inversion procedure.

From the V_{SV} and V_{SH} models of the upper-mantle, we compute the tomographic counterpart of radial anisotropy using $\xi^* = (V_{SH}/V_{SV})^2$. Here, we acknowledge that ξ^* is not directly inferred from simultaneous inversions of Love and Rayleigh data but is a rudimentary estimate from the two S-wave velocity models that may conceivably have different qualities. We view the following exercise as only a proof-of-concept and therefore the results should be interpreted with caution.

The depth distribution of ξ^* spanning from 35 - 400 km is shown in Figure 9 722 (top panel). Positive radial anisotropy values ($\xi^* > 1$) are confined in the upper ~ 723 200 km of the model domain which is in close agreement with previous studies (e.g. 724 J.-P. Montagner, 1985; Ekström & Dziewonski, 1998; Panning & Romanowicz, 2006). 725 Although the origin of anisotropy imaged in the asthenosphere is well-understood 726 purely in terms of CPO, anisotropy observed in the lithopshere may be a combination 727 of CPO and SPO (Wang et al., 2013). Here our task is to invoke the composite law 728 to isolate SPO from CPO in this tomographic model with the help of a homogenized 729 CPO model. 730

731

6.3 The CPO model

In this section, we re-interpret the results of Hedjazian et al. (2017) where they 732 examined radial anisotropy profiles predicted from CPO models produced by plate-733 driven flows underneath a mid-ocean ridge. From their work, we borrowed two CPO 734 models that correspond to a fast-developing CPO and a slow-developing CPO. The 735 rate is dictated by the dimensionless grain boundary mobility parameter M which con-736 trols the kinetics of grain growth (and hence, the degree of dynamic recrystallization) 737 (E. Kaminski et al., 2004). In the first case, a value of M = 125 constrained from 738 laboratory experiments (Nicolas et al., 1973; Zhang & Karato, 1995) corresponding 739 to CPO produced from uniform deformation and initially-random CPO was imposed 740 (E. Kaminski et al., 2004). Subsequently, the second case considers a case where M741 = 10 (*i.e.*, slower CPO evolution) which also reproduces experimental results but in 742 the case of an initially developed CPO (Boneh et al., 2015). Hedjazian et al. (2017) 743 compared the CPO anisotropy directly with tomographic models. They concluded 744 that the patterns of radial anisotropy predicted with the slow CPO evolution were 745 in better agreement with tomographic models. We homogenize the two CPO models 746 and obtain their long-wavelength effective equivalent, and again appraise the resulting 747 profiles in comparison with tomographic observations. 748

749

6.3.1 The intrinsic CPO mineralogical model

2-D surface-driven mantle flows were acquired using the code Fluidity (Davies
 et al., 2011). In both models, upper-mantle deformation is governed by a composite
 dislocation and diffusion creep rheology following the implementation of Garel et al.

(2014). D-Rex was used to model CPO evolution. A complete description of the
methodology can be found in Hedjazian et al. (2017).

Figure 9 displays the intrinsic radial anisotropy profiles ξ_{CPO} belonging to the 755 fast-evolving CPO with reference D-Rex values $M = 125 \pmod{A}$ and the slow-756 evolving CPO with $M = 10 \pmod{B}$. Model A predicts a layer with strong levels 757 of intrinsic radial anisotropy of about 10% ($\xi_{\rm CPO} \approx 1.1$) at a depth of ~ 80 km 758 starting at approximately 20 My. At about the same depth, tomographic models 759 yield approximately 5% radial anisotropy (e.g. Panning & Romanowicz, 2006; Nettles 760 & Dziewoński, 2008b; Burgos et al., 2014). Hence, it has been argued that model A 761 overpredicts the observed level of large-scale anisotropy in the upper-mantle (Hedjazian 762 et al., 2017). On the contrary, model B predicts modest levels of intrinsic radial 763 anisotropy, about 5% ($\xi_{\rm CPO} \approx 1.05$) across the oceanic lithosphere which is more 764 consistent with tomographic observations. 765

766

6.3.2 The homogenized CPO model

Figure 9 now shows the effective intrinsic radial anisotropy profiles ξ_{CPO}^* of model 767 A^* and model B^* . In both cases, the ensuing patterns of radial anisotropy are smoothed 768 out as a result of homogenization. For instance, the apparent two-layered distribution 769 of intrinsic radial anisotropy with depth (down to ~ 250 km) in model A vanishes 770 after homogenization. The depth profile of effective intrinsic radial anisotropy as a 771 result contain one layer of radial anisotropy centered at ~ 100 km depth, making it 772 now compatible with tomographic models of the asthenosphere. Furthermore, it was 773 implied that radial anisotropy predicted with typical laboratory-derived parameters 774 exceeds tomographic observations. Here, we argue that it may also be the other way 775 around. Due to finite-frequency effects and eventually limitations in resolution power, 776 seismic tomography may underestimate the strength of intrinsic anisotropy, at least in 777 the absence of small-scale isotropic heterogeneities. The level of radial anisotropy in 778 models A and B are therefore larger than their homogenized/tomographic counterparts 779 (models A^* and B^*). As opposed to common practice, the physical parameters used 780 in CPO models of which are initially constrained by experimental data may need not 781 be manually tuned, and perhaps that the action of varying such parameters to conform 782 with tomographic observations deems unnecessary. We therefore conclude that direct 783 visual comparison between a CPO model and a tomographic model could lead to wrong 784

interpretations, and that homogenization is necessary to have correct interpretationsof the CPO models.

787

792

6.4 Deriving an SPO model

The SPO models of Figure 9 (models C and D) can be estimated by using our composite law in equation (31). The extrinsic radial anisotropy is obtained by simply dividing the tomographic model of radial anisotropy by that of the homogenized CPO model:

$$\xi_{\rm SPO}^* = \frac{\xi^*}{\xi_{\rm CPO}^*}.\tag{37}$$

In this way, models C and D are obtained from models A^{*} and B^{*}, respectively.

Strong levels of positive extrinsic radial anisotropy near the ridge axis may be 794 due to the inability of surface waves to register vertical flow because of its limited 795 lateral resolution. Model D, associated with the slow-evolving CPO model B, is almost 796 devoid of SPO. This is expected since model B was tailored to fit seismic tomography 797 observations from CPO only. Based on our results, one should favor SPO model C 798 that corresponds to a fast-evolving CPO model. It displays positive extrinsic radial 799 anisotropy above 200 km depth. This is more consistent with the existence of lateral 800 fine-scale structures at the base of the lithosphere (e.g. Auer et al., 2015; Kennett & 801 Furumura, 2015). 802

7 Conclusion

81

Differentiating the relative contributions of CPO and SPO to the full effective 804 medium is not a simple, straightforward process. The tomographic operator (here ap-805 proximated by \mathcal{H}) acts as a smoothing operator, and its inverse is highly non-unique. 806 It is therefore clearly impossible to separate the CPO and SPO contributions in a 807 tomographic model. One of the most logical courses of action is to compare tomo-808 graphic models of anisotropy with existing micro-mechanical models of CPO evolution 809 (e.g. Becker et al., 2003, 2006; Ferreira et al., 2019). Here, we proposed a very simple 810 composite law that directly relates the separate contributions of CPO and SPO to the 811 full effective radial anisotropy ξ^* inferred from tomographic models: 812

$$\xi^* = \xi^*_{\text{SPO}} \times \xi^*_{\text{CPO}},$$

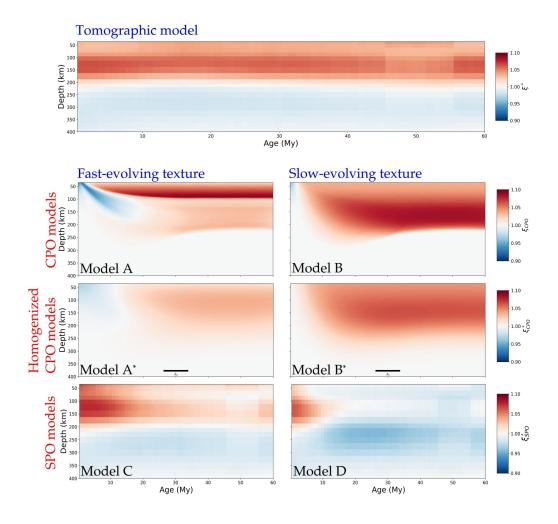


Figure 9. Plate-averaged radial anisotropy across the upper-mantle beneath oceanic basins with ages ranging between 0 and 80 Myrs obtained from a tomographic model (top panel), reference CPO models corresponding to fast and slow-evolving textures (models A and B), homogenized versions of model A (model A^{*}) and of model B (model B^{*}). Models C and D, respectively, are the extrinsic radial anisotropy profiles computed by dividing ξ^* of the tomographic model, by ξ^*_{CPO} of model A^{*} and B^{*}, using the composite law. Positive lithospheric radial anisotropy in model C implies the existence of horizontally-laminated structures. This is absent in model D which is expected since model B^{*} is designed to fit observations.

which we have numerically verified using simple 2-D toy models of an intrinsically anisotropic and a heterogeneous mantle. Although our numerical experiments were mainly a proof-of-concept, comparing a CPO model directly to an existing tomographic model is unwarranted and we highly recommend homogenizing a CPO model as an intermediate step.

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