Characterizing Charge Structure in Central Argentina Thunderstorms During RELAMPAGO Utilizing a New Charge Layer Polarity Identification Method

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Abstract

A new automated method to retrieve charge layer polarity from flashes, named Chargepol, is presented in this paper. Using data from the NASA Lightning Mapping Array (LMA) deployed during the RELAMPAGO field campaign in Cordoba, Argentina, from November 2018 to April 2019, this method estimates the polarity of vertical charge distributions and their altitudes and thicknesses (or vertical depth) using the very-high frequency (VHF) source emissions detected by LMAs. When this method is applied to LMA data for extended periods of time, it is capable of inferring a storm's bulk electrical charge structure throughout its life cycle. This method reliably predicted the polarity of charge within which lightning flashes propagated and was validated in comparison to methods that require manual assignment of polarities via visual inspection of VHF lightning sources. Examples of normal and anomalous charge structures retrieved using Chargepol for storms in Central Argentina during RELAMPAGO are presented for the first time. Application of Chargepol to five months of LMA data in Central Argentina and several locations in the United States allowed for the characterization of the charge structure in these regions and for a reliable comparison using the same methodology. About 13.3% of Cordoba thunderstorms were defined by an anomalous charge structure, slightly higher than in Oklahoma (12.5%) and West Texas (11.1%), higher than Alabama (7.3%), and considerably lower than in Colorado (82.6%). Some of the Cordoba anomalous thunderstorms presented enhanced low-level positive charge, a feature rarely if ever observed in Colorado thunderstorms.

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12 Key Points:

- A new automated method to estimate the polarity, altitude, and vertical depth of charge layers from flashes is presented.
- Thunderstorm charge structures in Central Argentina are characterized from months of
 LMA data, with 13.3% anomalous charge structure.
- Cordoba anomalously charged thunderstorms have a distinct charge altitude distribution
 when compared to those in Colorado.

19 Abstract

A new automated method to retrieve charge layer polarity from flashes, named Chargepol, is 20 presented in this paper. Using data from the NASA Lightning Mapping Array (LMA) deployed 21 during the RELAMPAGO field campaign in Cordoba, Argentina, from November 2018 to April 22 2019, this method estimates the polarity of vertical charge distributions and their altitudes and 23 24 thicknesses (or vertical depth) using the very-high frequency (VHF) source emissions detected by LMAs. When this method is applied to LMA data for extended periods of time, it is capable 25 of inferring a storm's bulk electrical charge structure throughout its life cycle. This method 26 reliably predicted the polarity of charge within which lightning flashes propagated and was 27 validated in comparison to methods that require manual assignment of polarities via visual 28 inspection of VHF lightning sources. Examples of normal and anomalous charge structures 29 30 retrieved using Chargepol for storms in Central Argentina during RELAMPAGO are presented for the first time. Application of Chargepol to five months of LMA data in Central Argentina and 31 several locations in the United States allowed for the characterization of the charge structure in 32 these regions and for a reliable comparison using the same methodology. About 13.3% of 33 Cordoba thunderstorms were defined by an anomalous charge structure, slightly higher than in 34 Oklahoma (12.5%) and West Texas (11.1%), higher than Alabama (7.3%), and considerably 35 lower than in Colorado (82.6%). Some of the Cordoba anomalous thunderstorms presented 36 37 enhanced low-level positive charge, a feature rarely if ever observed in Colorado thunderstorms.

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38 **1 Introduction**

39 Past studies have associated the severity of thunderstorms with patterns in charge distribution (Wiens et al., 2005; Fuchs et al., 2018). The dominant meteorological environment 40 provides initial conditions that would influence the kinematics and microphysics within 41 thunderstorms, which in turn affects its charge structure and dominant cloud-to-ground lightning 42 (CG) polarity. Relatively few studies have documented the charge structure over continents other 43 than North America (Lopez et al., 2019; Pawar & Kamra, 2004; Pineda et al., 2016; Qie et al., 44 2005). Furthermore, documenting charge structure in regions such as Argentina, which has 45 perhaps some of the highest flash rate thunderstorms in the world (Zipser et al., 2006), is crucial 46

47 for understanding such understudied thunderstorms.

Due to the nature of lightning processes and their characteristic emission in the VHF 48 spectrum, thunderstorm charge structures associated with flashes can be inferred from Lightning 49 Mapping Array (LMA) observations (Lang & Rutledge, 2011; Rust et al., 2005; Wiens et al., 50 2005). Based on knowledge of radiation propagation by lightning, VHF-based sensors primarily 51 detect radiation from negative breakdown of lightning that propagates through regions of 52 positive charge (Mazur & Ruhnke, 1993; Rison et al., 1999). Then, mapping of VHF sources is 53 used to manually determine the location of positive and negative charge layers (Bruning et al., 54 2007; Lang and Rutledge, 2008; Rust et al., 2005; Wiens et al., 2005). An intra-cloud lightning 55 (IC) flash initiates in a region with strong electric field, in between regions of charge with 56 opposite polarities. Upon initiation, bi-direction leaders are formed and move into opposite 57 regions of charge: a positive leader moves to a region of net negative charge, and a negative 58 leader moves to a region of positive charge in the cloud (Coleman et al., 2003; Kasemir, 1960). 59 When a leader reaches a charge layer, it propagates horizontally through the charge layer away 60 from the flash initiation location (Shao & Krehbiel, 1996). Flashes propagating through charge 61 regions that constitute a vertical dipole with positive charge located above negative charge, are 62

referred to as positive cloud flashes (+IC), while flashes that propagate through negative over 63 positive dipoles are defined as negative cloud flashes (-IC, Bruning et al., 2014). K-processes 64 may also occur, transporting charge to the base of the initial channel (Shao & Krehbiel, 1996). 65 Observations of the VHF altitude source distributions for long periods of time are generally used 66 to infer the location of charge regions, as the altitude with most sources are often associated with 67 positive charge layers (Fuchs et al., 2018; Fuchs & Rutledge 2018; Lang et al., 2020; Lang & 68 Rutledge, 2011). Tessendorf et al. (2007b) infer charge layer polarity automatically by using the 69 first LMA source altitude for a flash, and the number of sources above and below that altitude. 70 Stough and Carey (2020) utilized the DBSCAN (Density-Based Spatial Clustering of 71 Applications with Noise, Ester et al., 1996) algorithm to identify regions of dense sources and 72 infer charge region polarity. Electric field soundings have been deployed to infer polarity of 73 charge regions within and nearby thunderstorms (Marshall et al., 1995; Rust & MacGorman, 74 2002; Stolzenburg et al., 1998), and have been compared to LMA-inferred charge regions (Rust 75 et al., 2005). In order to interpret an electric field dataset with altitude, the Gauss' Law 76 approximation is assumed, where the charge density is proportional to the electric field variation 77

78 with height (Stolzenburg et al., 1998).

In order for clouds to build regions of net charge polarity and become electrified, the non-79 80 inductive charging (NIC) mechanism is thought to dominate, which does not require a preexisting electric field to polarize the cloud and precipitation size particles. In the NIC 81 mechanism, the polarity that graupel particles acquire when colliding with ice crystals in the 82 presence of supercooled liquid water (Saunders et al., 1991; Takahashi, 1978) depends on the 83 temperature, and the effective liquid water content (EWC, the accreted fraction of the liquid 84 water content). High (low) temperature and large (small) EWC are associated with graupel 85 charging positively (negatively), and ice crystals charging negatively (positively) (Berdeklis & 86 List, 2001; Perevra et al., 2000; Saunders et al., 1991, 2001; Saunders & Peck, 1998; Takahashi, 87 1978). As the rimer particle (e.g., graupel) accretes supercooled liquid water, it is heated by 88 latent heat, which sublimates the ice surface and reduces the diffusional growth (Williams et al., 89 90 1991). According to the relative diffusional growth theory (Baker et al., 1987; Emersic & Saunders, 2010), the ice particle growing faster by diffusion acquires positive charge. Particle 91 differential fall speeds and updrafts lead to storm-scale charge separation, with ice crystals being 92 93 transported upward to cloud tops, and graupel residing in the mixed-phase region in the midlevels forming the two largest charge regions during the developing-to-mature stage of 94

95 thunderstorms (Williams, 1985).

Thunderstorms with upper-level negative and mid-level positive charge layers define an 96 97 anomalous charge structure, as observed in thunderstorms during the STEPS field campaign 98 conducted in Kansas, Colorado, and Nebraska (MacGorman et al., 2005; Rust et al., 2005; Rust 99 & MacGorman, 2002; Tessendorf et al., 2007a; Tessendorf et al., 2007b; Weiss et al., 2008; Wiens et al., 2005). They have also been observed in thunderstorms in Oklahoma by Marshall et 100 101 al. (1995) and Emersic et al. (2011), during the TELEX field campaign (MacGorman et al., 2008), in Texas (Chmielewski et al., 2018), Alabama (Stough & Carey, 2020), and Spain (Pineda 102 103 et al., 2016). Storms with a normal charge structure would have a dominant net negative charge in the mixed-phase layer, and net positive above, as demonstrated in early foundational studies 104 reviewed by Williams (1985), in the in-situ aircraft studies by Dye et al. (1986, 1988, 1989), 105 during TELEX (Bruning et al., 2007) and STEPS (Weiss et al., 2008) field campaigns, among 106 others. A low-level charge layer with opposite polarity to the nearest charge region is 107 occasionally present (Lopez et al., 2019; Pawar & Kamra, 2004; Williams, 1989) and, if positive 108

and abnormally large, may also be termed anomalous (Bruning et al., 2014; Fuchs et al., 2015;

110 Qie et al., 2005). Some events can have multiple charge regions, such as mesoscale convective

systems (MCSs) (Lang & Rutledge, 2008; Lund et al., 2009; Stolzenburg et al., 1998), multicell
storms (Bruning et al., 2007), and supercells (Bruning et al., 2010; Calhoun et al., 2013; Wiens et

112 storms (Br113 al., 2005).

114 Fuchs and Rutledge (2018) analyzed a large lightning flash dataset for isolated cells in four different regions in the United States, and found that Colorado storms have a prevalence of 115 anomalous charge structures compared to other regions. Colorado's highest flash rate mode was 116 observed at lower levels (warmer temperatures and higher radar reflectivity values) than in other 117 regions. In addition, they suggested that Colorado is followed by Oklahoma in terms of 118 anomalous storm frequency, followed by Alabama and Washington D.C. with rare anomalous 119 120 observations. A large occurrence of positive cloud-to-ground lightning (+CG) is often associated with anomalous charge structure storms, as a main net positive charge region is at the middle or 121 low levels of a storm instead of near its top, facilitating the propagation of positive leaders 122 toward the ground, especially if a small opposite (negative) charge region is present at lower 123 levels. Orville and Huffines (2001) found that the percentage of +CGs in the United States varies 124 from 2% in Florida to 10-20% in a region extending from the High Plains of Eastern Colorado to 125 the Upper Midwest. In the central and north Great Plains, a high percentage (>50%) of severe 126 127 storm reports were found to be associated with predominantly +CG lightning (>50% of CGs being positive), when compared to southern Great Plains and eastern United States (Carey et al., 128 2003). 129

This study aims to characterize the charge structure in the Central Argentina region for 130 the first time, utilizing a large dataset, as it is a key science goal of the RELAMPAGO (Remote 131 sensing of Electrification, Lightning, And Mesoscale/microscale Processes with Adaptive 132 Ground Observations) field campaign (Nesbitt et al., 2021). This novel research is achieved by 133 first developing and testing a new automated method to retrieve thunderstorm charge layer 134 135 polarity using Lightning Mapping Array (LMA) source and flash data, which is described in this paper. Southeast South America has among the most severe thunderstorms in the world in terms 136 of high flash rate (Zipser et al., 2006), hail size (Cecil & Blankenship, 2012), heavy 137 precipitation, and flash floods (Rasmussen et al., 2014). Lightning characteristics have only been 138 139 documented using LMA data recently in this region (Lang et al., 2020), and the distribution of charge within Argentina thunderstorms is explored for the first time in great detail in this study. 140 141 The general charge structure is estimated for a large dataset with a new algorithm, allowing for the inference of the likelihood of normal and anomalous charge structure. Similar to Tessendorf 142 et al. (2007b) and Stough and Carey (2020), this method automatically infers charge polarity 143 144 from flashes, more closely resembling Tessendorf et al. (2007b) method but with improved procedures, better emulating the steps that a human expert would perform when assigning 145 polarity to LMA sources for a flash by detecting the negative leader in a bi-directional model and 146 assigning polarity to sources of a flash (e.g., Rust et al. 2005). In this study, if a given lightning 147 flash passes a series of conditions, an algorithm analyzes its source location and time in order to 148 produce a prediction of charge layer polarity for that flash. This method has the capability to be 149 quickly applied to a large number of lightning flashes in a large LMA dataset (e.g., a few 150 minutes to process 24 hours of LMA flash level data within 100 km of the network center), 151 which allows for the inference of the general charge structure and its evolution in time for a 152 thunderstorm or for a large area of interest, as demonstrated by examples in this paper. The new 153 algorithm infers three-dimensional charge distribution on the flash level but its output is 154

simplified to vertical charge layer profiles for the science applications highlighted in this study. 155 Hence, output of this method is similar to manual assignment of polarity, providing positive and 156 negative layer altitude and vertical depth, but it is much less labor intensive. This algorithm 157 provides a detailed inference of the charge layer distribution in the vertical, including altitude 158 and vertical depth of negative charge layers, which is often not possible to be analyzed from the 159 VHF source distribution analysis, a method in which positive charge altitude is inferred from its 160 peak distribution. Lastly, this paper will present a detailed application of the new charge layer 161 polarity algorithm by characterizing the charge structure of Central Argentinian thunderstorms 162 by processing a large multi-month sample of LMA observations for the first time. The algorithm 163 performance is then further demonstrated through its application to multi-month LMA datasets 164 from several locations in the United States in which charge structure has already been 165 documented using the LMA-based charge layer retrieval techniques discussed above. The 166 additional application herein allows the charge structure of Central Argentinian thunderstorms to 167 be compared for the first time to several well-studied locations in the United States such as 168 Colorado, Oklahoma, West Texas and Alabama using the same algorithm. Consistency with 169 prior studies of charge structure in well observed regions of the United States ensures that this 170

171 method is applicable for future work.

172 2 Lightning Networks Deployed During RELAMPAGO and DC3

The Lightning Mapping Array (LMA) is a GPS-based network (Goodman et al., 2005; 173 Koshak et al., 2004; Krehbiel et al., 2000; Rison et al., 1999) that operates in the VHF 174 electromagnetic spectrum (Krehbiel et al., 2000), in which radiation events detected are often 175 associated with lightning breakdown processes (Rison et al., 1999). LMAs locate and report the 176 time of VHF sources emitted during lightning breakdown processes using a time-of-arrival 177 technique (Koshak et al., 2004; Koshak & Solakiewicz, 1996; Lhermitte & Krehbiel, 1979; 178 Thomas et al., 2004), where a χ^2 goodness-of-fit function with a threshold of lower than 5 is 179 utilized to minimize location errors, and minimum of 6 operating network sensors are required to 180 ensure location accuracy (Chmielewski and Bruning, 2016). The Imatools Python package 181 (Bruning et al., 2015) was used to process LMA source data into lightning flash datasets. This 182 package is based on the DBSCAN (Ester et al., 1996) algorithm, a machine learning algorithm 183 used to cluster VHF sources to reconstruct the shapes (structure) of entire lightning flashes. 184 DBSCAN randomly searches for clusters of VHF sources in space and time, and groups each 185 cluster individually. These groups are used to define individual flashes using the following 186 criteria: source-to-source minimum distance and time thresholds of 3000 meters and 150 ms, 187 respectively, and a maximum flash duration of 3 seconds (Fuchs et al., 2016). 188

As part of the RELAMPAGO field campaign (Nesbitt et al., 2021), an LMA of 11 189 sensors was deployed by NASA Marshall Space Flight Center to the eastern side of the Sierra de 190 Cordoba mountains in the province of Cordoba, Central Argentina, from mid-November 2018 to 191 mid-April 2019 (Lang et al., 2020). RELAMPAGO LMA data was used in this study for 192 development and validation of the charge layer inference method, and characterization of the 193 charge structure climatology in the Cordoba warm season. In addition, LMA datasets from the 194 DC3 (Deep Convective Clouds and Chemistry, Barth et al., 2015) field campaign are used to 195 independently estimate the charge structure in a variety of climatological regimes of the United 196 197 States to compare results of the presented algorithm with those of other studies, and with storms in the Cordoba region of Argentina examined during RELAMPAGO. During the DC3 field 198 campaign, LMA networks were deployed simultaneously in Alabama, West Texas, Oklahoma, 199

and Colorado in May and June 2012 (Barth et al., 2015; DiGangi et al., 2016; Mecikalski et al.,

201 2015). For each dataset, only flashes with centroid location within the 100 km range distance

from the LMA network center are being considered in this study, as altitude errors are expected

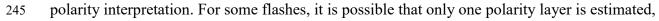
to be smaller and the flash detection efficiency to be higher (Chmielewski & Bruning 2016;

Koshak et al. 2004; Lang et al., 2020; Thomas et al. 2004) within this range. In addition, flashes with less than 20 detected sources were not considered in this study (see Section 3 for details).

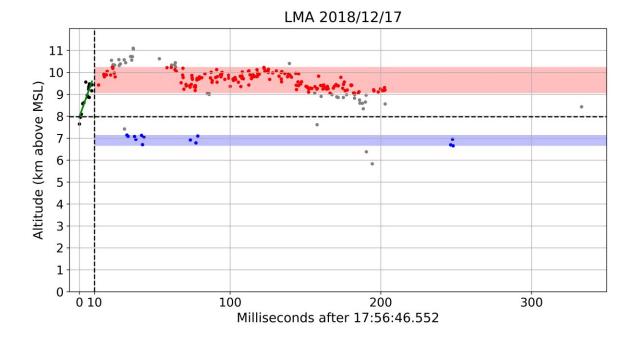
206 **3 Description of the Charge Layer Polarity Identification Method**

The charge layer polarity identification method (hereafter Chargepol) consists of an 207 automated algorithm that applies a series of procedures to each lightning flash retrieved by the 208 lmatools, in order to infer charge layer polarity from a flash (link in the Acknowledgments). For 209 reference, Figure 1 shows a flash example with the procedures illustrated. First, flashes with less 210 than 20 sources are disregarded because those flashes would not allow a sufficient number of 211 212 sources to characterize the initial negative leader breakdown, negative leader propagation through a positive charge region, and sources associated with a negative charge region. Then, all 213 sources contained in the first 10 ms of a flash, referred to here as the Preliminary Breakdown 214 sources (PB sources), are analyzed. A minimum of 4 PB sources is required, and the time 215 interval between the first and last PB source has to be at least 2 ms, in order to better characterize 216 the initial vertical motion of the negative leader. Typical duration periods are between 4 and 10 217 ms for PB (Zheng et al., 2019). We make the assumption that PB sources are associated with 218 negative breakdown having a predominant vertical motion toward a region of positive charge 219 220 (Shao & Krehbiel, 1996). Hence, linear regression is applied to the PB sources time-height dimension. The linear regression slope is used as a proxy for the vertical speed of the leader, and 221 has to be greater than a threshold of absolute value of 0.05 (0.5 km height variation in 10 ms), 222 which is equivalent to a vertical speed of 5×10^4 ms⁻¹, or half the typical order of magnitude speed 223 of a negative leader (Behnke et al., 2005; Shao & Krehbiel, 1996; van der Velde & Montanya, 224 2013). By applying that slope threshold, flashes with no clear initial vertical motion are 225 discarded, facilitating further a correct depiction of charge region polarity. In addition, the linear 226 regression fit to the PB sources required a mean squared error (MSE) of less than 0.25 to prevent 227 228 fitting a regression to noisy sources.

Only flashes that satisfy all the aforementioned conditions are used for charge layer 229 depiction, which is typically about 16% of all flashes (more on this in Section 5). The fact that 230 not all flashes are analyzed does not interfere with the objective of this study, because estimating 231 charge polarity for some flashes is sufficient to determine the charge structure evolution over 232 long periods of many hours, as demonstrated in the next section. Next, non-PB sources (sources 233 after 10 ms from flash initiation) are used to infer charge layer polarity, altitude and vertical 234 depth. The PB linear regression intercept altitude is used as a threshold, referred to here as the 235 Charge Height Threshold (CHT), in order to separate positive and negative charge layers 236 candidate sources. For a positive PB linear regression slope (i.e., a flash with initial negative 237 leader moving upward), all non-PB sources above (below) the CHT are candidate sources to 238 define a positive (negative) charge layer. A flash with initial downward motion (negative PB 239 linear regression slope) would have all non-PB sources below (above) the CHT as candidate 240 sources for positive (negative) charge layer. Then, among the candidate sources for each layer 241 polarity, the interval between the 10th and the 90th percentile source heights is used to define a 242 charge layer, which provided a better distinction between positive and negative vertical 243 histograms when compared to other intervals and also neglected sources with possibly doubtful 244



- which leads to the total number of estimated positive layers from flashes for a large period of
- time being slightly larger than the number of estimated negative layers from flashes.



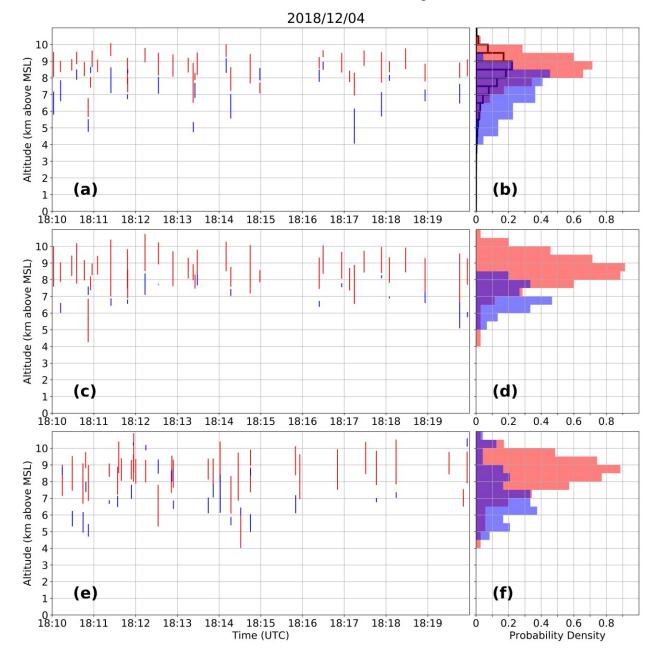
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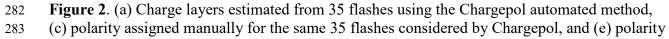
Figure 1. A time-height plot for a positive intracloud flash. The vertical dashed line marks the 10 249 250 ms time limit that defines the PB sources (black dots). The green line is the linear regression fit on the PB sources. The horizontal dashed line is the CHT (Charge Height Threshold), that 251 separates candidate sources for positive and negative charge layers. Red and blue dots (and 252 shaded areas) define the positive and negative charge layers altitudes and width for this flash. 253 found by applying the interval between the 10th and 90th percentile source altitudes for each 254 polarity candidate sources. Grav dots are candidate sources outside the 10th-90th percentile 255 interval, which were not used to define charge layers. 256

257 3.1 Validation using Manual Analysis of LMA

In order to validate the automated Chargepol identification method, manual polarity 258 inference (Rust et al., 2005; Wiens et al., 2005) was performed on some lightning flashes, and 259 compared with the Chargepol algorithm output. First, a 10-minute period (4 December from 260 1810 to 1820 UTC) with a predominance of normal charge structure (i.e., normal dipole with 261 positive charge over negative charge) was chosen from the RELAMPAGO LMA dataset. Among 262 the 168 flashes that occurred in this period with a normal charge structure, the algorithm 263 estimated charge layers from 35 of them (21%) (Figure 2a). Figures 2b shows a histogram 264 density with the altitude where each charge layer polarity was detected (a peak of positive 265 polarity of 0.7 between 8.5 and 9 km height means that 70% of all positive charge occurred at 266 that level). Source polarities were manually assigned for the same 35 flashes, shown in Figures 267 2c and 2d. The positive charge altitude was estimated to be between about 8 km and 9.5 km from 268 both Chargepol (Figure 2b) and the manual method (Figure 2d). Manual assignment of negative 269 charge (Figure 2d) proved to be challenging, as it could not be estimated from all lightning 270 flashes. Even so, it is notable that the negative charge layer is located at altitudes generally below 271

- the altitude of positive charge, with peak occurrence between 6.5 and 7 km height (Figure 2d).
- Additional validation was performed by assigning polarity for another 35 randomly chosen
- flashes among the 133 flashes during the same 10-minute period that were not considered by C_{1}
- Chargepol (Figure 2e). Most of these flashes did not have a clear vertical trend of the initial
 leader (not shown). However, as shown in Figure 2f, most positive charge layer detections from
- flashes were estimated to be between 8 and 9.5 km, consistent with the automated method
- (Figure 2b), while negative charge is located at lower altitudes. The analysis of an independent
- subset of flashes from Figures 2e-f demonstrates that Chargepol analysis on a fraction of total
- 280 flashes is sufficient for the determination of thunderstorm charge structure.





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assigned manually for 35 other flashes not considered by Chargepol during the same time period.
Each red (blue) vertical line represents a positive (negative) charge layer estimated from a flash.
(b) (d) and (f) shows histograms (0.5 km his size) with the probability density of retrieved

- (b), (d), and (f) shows histograms (0.5 km bin size) with the probability density of retrieved
- 287 positive and negative charge layers with height for (a), (c) and (e), respectively (overlap of histograms in purple). Black histogram in (b) shows the source height distribution. The 10-
- 289 minute period chosen had a predominance of normal charge structure as clearly shown by the
- 290 Chargepol algorithm, manual analysis, and source distribution.

This procedure was repeated for a 10-minute period (5 December from 1800 to 1810 291 UTC) with a predominance of anomalous charge structure (dipole with positive charge located 292 below negative charge), shown in Figure 3. During this period, a high flash rate storm produced 293 mostly negative ICs propagating through a lower positive charge. Another storm with low flash 294 295 rate and upper positive charge layer was active at the same time. A total of 107 flashes occurred during this period, in which Chargepol estimated charge layers for 36 of them (Figure 3a-b). 296 Manual depiction of charge polarity for these same 36 flashes (Figure 3c) show that the altitudes 297 of positive and negative charge layers (Figure 3d) are in agreement with Chargepol, although 298 manual inference of negative charge is at a slightly higher altitude. From Figure 3d, more than 299 50% of the negative charge layers occurred at altitudes from 6 to 7.5 km, while Chargepol 300 estimated negative charge layers from 5.5 to 7 km height (Figure 3b). The small differences in 301 302 charge layer altitudes between the manual and automated method demonstrate the small uncertainty of the method. Manual inference for a different set of 36 flashes during the same 303 time period that was not considered by Chargepol is shown in Figures 3e and 3f, and it is 304 consistent with other flashes (Figures 3a and 3c) in locating lower positive charge and mid-level 305 negative charge. The altitude distance between positive and negative charge layers centers (from 306 histogram plots) for all methods is about 2 km. The few upper positive charge layers located 307 above 8.5 km by all methods are from the normal charge structure storm aforementioned. 308

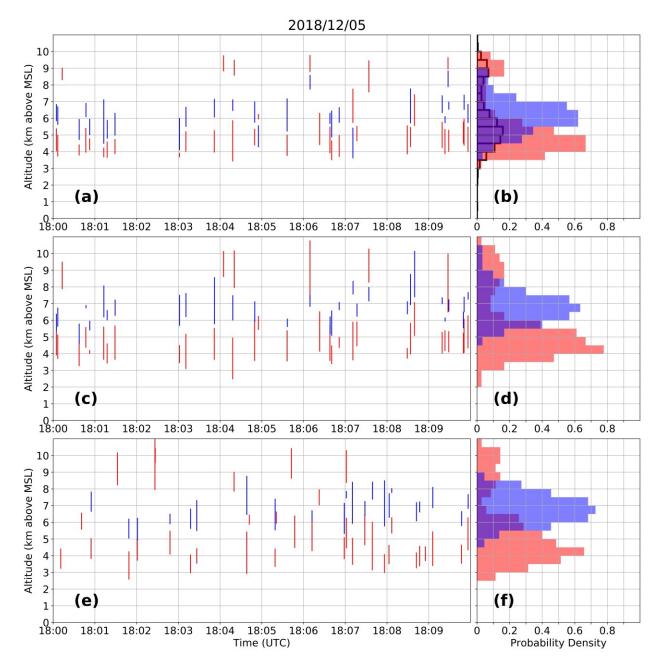


Figure 3. Same as in Figure 2, but for a 10-minute period with predominance of anomalous charge structure.

The manual depiction of charge layers polarity agrees qualitatively well with the automated depiction. The vertical distance between each polarity's vertical source distribution maxima were sufficiently large by more than 1 km (Fig. 2 and 3 histograms), leading to charge layers being well identified in the vertical dimension.

316 3.2 Validation using Vertical Distribution of VHF Sources

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An additional method to validate the Chargepol algorithm is the estimate of the positive charge layer altitude from the peak in the VHF source histogram (Fuchs et al., 2018; Fuchs & Rutledge 2018; Lang et al., 2020; Lang & Rutledge, 2011). Figures 2b and 3b show an additional

- histogram of the vertical source density. The histogram for the normal case (Figure 2b) presents
- the peak at the same altitude the Chargepol method shows a peak with the most occurrences of
- positive charge. A comparison of these two methods shows that the Chargepol method has the advantage of inferring negative charge layer altitude, which is not possible to estimate from the
- advantage of inferring negative charge layer altitude, which is not possible to estimate from the LMA VHE source distribution. For the anomalous area (Figure 2b), the main law layer pack
- LMA VHF source distribution. For the anomalous case (Figure 3b), the main low-level peak from the anomalous charge structure storm and the secondary peak from the normal storm are
- depicted. The peak from the source histogram is at a slightly higher altitude, 5 to 5.5 km,
- compared to Chargepol's positive inference at 4 to 5 km. However, both methods generally agree
- and the depiction of the negative layer by Chargepol is notable.

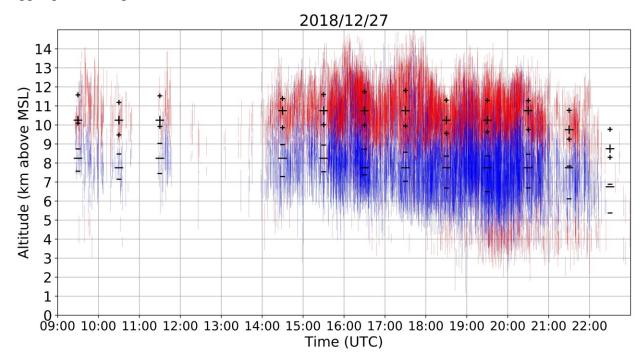
329 4 Chargepol Method Applied to RELAMPAGO Thunderstorms

How charge is structured in Argentinian thunderstorms is generally unknown, and so we 330 331 make use of the Chargepol method to examine them. During the five-month period the RELAMPAGO LMA network was operating in Cordoba, Argentina. Different storm modes 332 were observed and included isolated convection, multicellular storms, supercells, and mesoscale 333 334 convective systems (Nesbitt et al. 2021). In order to demonstrate the capability of the algorithm to depict charge structures, examples of distinct Cordoba cases and their evolution in time are 335 presented. Examples of thunderstorms with different charge structures in Cordoba are shown and 336 included normal, anomalous, a case with an enhanced lower positive charge layer, and one that 337 demonstrated a change from one archetype to another through its lifetime. The altitudes at which 338 either positive or negative polarities were classified most frequently, and the mean altitude of 339 tops and bottoms of each charge layer polarity were examined every hour to show how charge 340 layer altitude varied with time for all presented cases. The variation of a dipole's altitude with 341 time depicts a storm's charge structure evolution and it is shown in this study in order to 342 demonstrate a possible application a user can generate with this dataset. For a lower charge layer 343 polarity from a given dipole, the mean altitude of tops and bottoms of charge layers estimated 344 from flashes were only calculated for charge layers in which its top was at a lower altitude than 345 the mean upper dipole polarity altitude. Similarly, for the upper charge layer polarity from a 346 347 dipole, mean altitude of its top and bottom was obtained from charge layers with its bottom above the altitude of the mean altitude of the lower charge layer polarity. These restrictions were 348 put in place to focus analysis on the top and bottom altitudes of the dominant positive and 349 negative charge layers in the main dipole. To further demonstrate the algorithm's capabilities 350 over regions of the United States that have been studied and well characterized with other charge 351 retrieval methods (e.g., Bruning et al., 2010; MacGorman et al., 2005; Mecikalski et al., 2015; 352 353 Wiens et al., 2005), an example from each of the LMA networks deployed during DC3 are shown in the supporting information and included a normal tripole case in Alabama, anomalous 354 storms in Colorado, a case with a transition from anomalous to normal charge structure in 355 Oklahoma, and a normal dipole in West Texas at typical altitudes (negative in mid-levels, 356 positive in the upper levels) but with a very high altitude negative charge layer observed above 357 the upper positive. 358

4.1 27 December 2018 Case: Normal Charge Structure

Figure 4 shows the estimate of charge layer polarity for all convective storms (most of them multicellular) that occurred in the RELAMPAGO LMA domain for a 14-hour period on 27 December 2018. Most thunderstorms that occurred on this day presented an upper-level positive

charge layer above 9-10 km height, and a mid-level negative charge layer between about 5 and 9 363 km height. Altitude variation in charge layers is speculated to be due to different thunderstorms 364 having varying updraft strength and cloud-top heights. As the number of charge layers vary 365 within a storm, where more charge layers are found where flash rates are highest (Brothers et al., 366 2018; DiGangi et al., 2016, for example), we inferred flash rates from periods when charge 367 layers were estimated frequently in short periods of time. For example, more frequently 368 estimated charge layer polarities in shorter timespans were indicative of higher storm flash rates 369 than those less frequently estimated. For most of the period between 1500 and 2100 UTC, the 370 total flash rates of all storms was higher than 50 flashes/minute, considering flashes with more 371 than 10 sources and all active thunderstorms in the domain. The total flash rate of storms in the 372 domain peaked at 195 flashes/minute at 1609 UTC. The dominance of positive over negative 373 charge structure means that most flashes depicted by Chargepol were +ICs, with a typical initial 374 upward motion of a negative leader and further propagation through the upper positive charge 375 layer. This general dipole structure characterizes a typical normal charge structure, as it is 376 common in many regions of the United States as shown in similar LMA-based charge retrieval 377 studies, such as in Alabama (Mecikalski et al., 2015) and Oklahoma (Bruning et al., 2007). Some 378 flashes propagated through a lower positive charge layer below 5 km height, principally after 379 1920 UTC. That was caused by -IC flashes with initial negative leaders with downward motion 380 and further propagation through the low-level positive charge region. Hence, from 1900 to 2200 381 UTC, a typical tripole charge structure (Williams, 1989) was present, though the upper positive 382 region is considerably more active than the lower positive due to more flashes contributing to the 383 upper positive depiction. 384



385

Figure 4. Charge layers estimated from flashes using the Chargepol automated method for all
 RELAMPAGO thunderstorms on 27 December 2018 from 0900 to 2300 UTC. Each red (blue)

vertical line represents a positive (negative) charge layer estimated from a flash. Large black

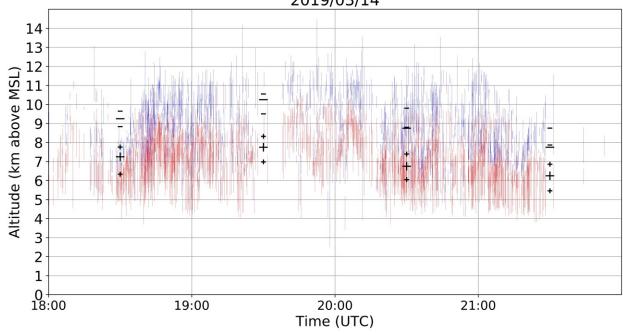
389 symbols represent the altitudes in which most charge layers of a certain polarity were estimated

390 for each hour period, as long as more than 30 layers with that polarity were present in that hour.

- Small black symbols represent the mean altitudes of the top and bottom of charge layers for eachpolarity and hour.
- 393

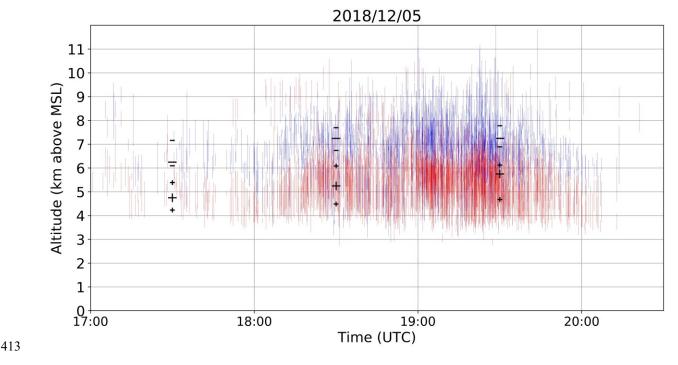
4.2 14 March 2019 and 5 December 2018 Cases: Anomalous Charge Structure

A cluster of RELAMPAGO storms on 14 March 2019, all with anomalous dipole charge 394 structures, are shown in Figure 5. These storms had a dominant mid-level positive charge layer 395 and upper-level negative charge layer, similar to some anomalous storms over Colorado (Fuchs 396 et al. 2015). As multiple storms are shown in Figure 5, a large altitude variation is noticeable for 397 the charge layers, which is possibly dependent on individual storm intensity. Storms with 398 stronger updrafts are thought to initiate flashes between charge layers residing at higher altitudes 399 (Stolzenburg et al., 1998). Most flashes in these storms presented -IC lightning, which means 400 that negative breakdown had an initial downward propagation, hence negative charge is 401 402 estimated at higher levels than positive charge. A similar anomalous dipole case occurred in an isolated thunderstorm on 5 December 2018 (Figure 6). This storm had a flash rate higher than 30 403 flashes/minute for most of the period between 1815 and 1945 UTC, with a peak flash rate of 128 404 405 flashes/minute at 1902 UTC. This anomalous case is different from the 14 March 2019 anomalous case because estimated charge layers are located at lower levels: negative charge is 406 located in the mid-levels, while positive charge is in the low-levels. Also, this was a relatively 407 shallow storm system exhibiting a radar echo top at about 10-km height (not shown), hence no 408 upper positive charge layer had developed. Upper positive charge at about 9 km height from 409 1800 to 1900 UTC was from another storm in the domain (see discussion in Section 3.1). 410

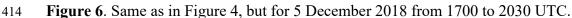


2019/03/14

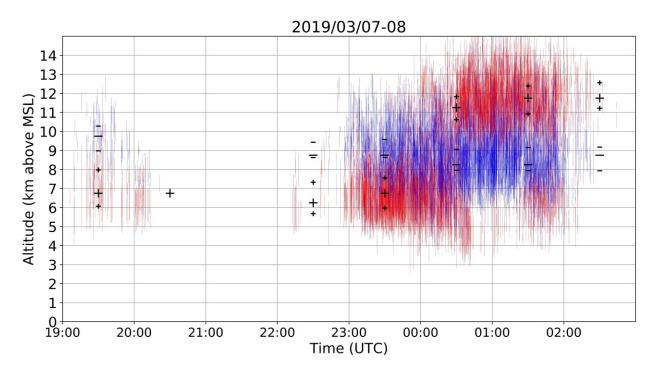
411



412 **Figure 5**. Same as in figure 4, but for 14 March 2019 from 1800 to 2200 UTC.



4.3 7-8 March 2019 Case: Transition from Anomalous to Normal Charge Structure 415 Thunderstorms on 7 March 2019 (Figure 7) during RELAMPAGO presented an 416 anomalous charge structure with mid-level positive charge and upper-level negative charge. 417 From 1900 to 2300 UTC, storms that occurred in the LMA domain had a low flash rate (less than 418 30 flashes/minute considering all thunderstorms in the domain), then few charge layers were 419 depicted by Chargepol, but an anomalous dipole is clearly present, similar to the storm studied 420 by Fuchs et al. (2018) over Colorado. After 2300 UTC, a MCS formed with a dominant 421 anomalous charge structure, with its flash rate rapidly increasing to more than 100 flashes/minute 422 in the LMA domain. On the following UTC day, high flash rates remained, reaching a peak of 423 496 flashes/minute at 0124 UTC in the domain, and an upper positive charge layer formed above 424 425 10 km height. This upper positive layer became visible because flashes started propagating through that layer. After 0045 UTC, fewer flashes propagated through the lower positive charge 426 layer. Hence, this case characterizes a transition from anomalous to normal charge structure. This 427 case demonstrates how complex charge structure evolution can be estimated by the Chargepol 428 method, such as the presence of anomalous and normal main dipoles, tripoles, and their 429 evolution in time. 430







433 **5** Frequency of Anomalous Charge Structure in Central Argentina Compared to the U.S.

As the described Chargepol method allows for a relatively fast processing time for large datasets of months of LMA data, one can obtain the general charge structure evolution in time for a domain area, as shown in the previous section. Hence, in order to characterize the likelihood of normal and anomalous charge structure for the five months in which the LMA was deployed in the Cordoba, Argentina region for the first time, the Chargepol layer polarity output was summarized for a better understanding and interpretation of the general dominant charge structure.

In order to achieve a summary of the general charge structure typically occurring in 441 Argentinian storms for long periods of time, the charge polarity information was initially 442 subdivided into time periods of one hour to obtain the dominant dipole for every hour period. 443 Then, the number of charge layers of a given polarity were counted for every altitude in 0.5-km 444 bins for every hour. The altitude with the most positive charge layers estimated from flashes, and 445 the altitude with most negative charge layers, define a single altitude bin for each layer polarity, 446 which characterizes the dominant dipole for that hour, as long as both maximum polarities occur 447 at different heights. A minimum threshold of 30 charge layers from each polarity occurring in 448 one hour was applied, in order to remove the influence of thunderstorms with low flash rates 449 contributing to the charge structure estimation. The large black symbols present in Figures 4-7 450 451 represent the altitude with most occurrences of a charge layer polarity for each hour, and the resultant main dominant dipole for an hour period. 452

In this study, an estimated dipole structure for a one hour period is referred to as a "sample". Samples in which dipoles had positive located at a higher altitude than negative are referred in this study as normal charge structures (Dye et al., 1986; Williams, 1985). A normal charge structure sample could have few flashes that estimated the presence of a low-level

positive charge layer, however if more flashes contributed to the maximum height occurrence of 457 positive being at high levels, it would be considered a normal charge structure sample. Figure 4 458 shows an example of a normal tripole charge structure (Williams, 1989) with more positive 459 layers estimated at high levels, leading to a normal dipole estimation. Samples with negative 460 charge over positive charge are considered to have a dominant anomalous charge structure. The 461 most common type of anomalous dipole sample is the type with positive charge at the mid-levels 462 or mixed-phase layer, and negative in the upper levels of a storm (Figure 5). Another structure 463 that could lead to a negative-over-positive dipole is when enhanced positive charge is at low 464 levels of a storm, while negative charge is at the mid-levels (Figure 6, Bruning et al., 2014; 465 Fuchs et al., 2015). In this scenario, an upper positive charge could be present, which could lead 466 to an interpretation of a normal tripole charge structure, an uncommon characteristic during 467 RELAMPAGO as the enhanced low-level positive charge layer is not typically accompanied by 468 an upper-level positive charge layer (Figure 6). However, in this study and others (e.g., Fuchs et 469 al. 2015), a normal tripole scenario with more flashes propagating through the lower positive 470 charge layer than through the upper positive would imply the characterization of an anomalous 471 dipole. Both anomalous scenarios (positive in the mid-levels, and in the low levels) imply that 472 most flashes consisted of -ICs with negative leaders having an initial downward motion, rather 473 than +ICs that would initially move upward. Hence, in this study, scenarios with a dominance of 474 -ICs, or negative-over-positive dipoles, are considered anomalous. 475

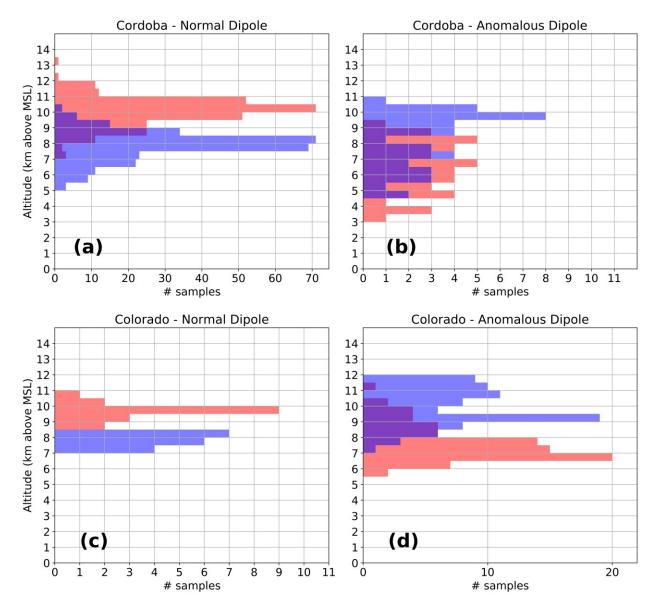
During the five months that the LMA was operating in Cordoba, Argentina, 306 samples 476 were observed, which means 306 hours with lightning activity in which the aforementioned 477 methodology estimated a dominant dipole structure. Among the 306 Cordoba samples, 265 478 consisted of normal dipole charge structure, while the other 41 were anomalous (Table 1). That 479 means that 13.3% of samples had a dominant anomalous charge structure, which can be 480 interpreted as an approximate frequency of occurrence of anomalous storms in Cordoba, 481 Argentina. Table 1 shows the number of normal and anomalous samples for the Cordoba LMA 482 deployed during RELAMPAGO, as well as for the four LMA networks deployed during DC3 in 483 484 several locations across the United States (e.g., Colorado, West Texas, Oklahoma and Alabama) for comparison, all sampled in the warm season (May and June 2012). Table 1 also shows the 485 total number of flashes with more than 20 sources, and the fraction of flashes that were 486 considered by the algorithm, being 16.7% for all LMA networks. The comparison of Cordoba 487 dipole samples with DC3 networks is shown to demonstrate the usefulness of the application and 488 capabilities of Chargepol, as it made estimates of charge structure climatologies similar to those 489 observed in other studies. Even though the sample numbers vary for the different locations, 490 consistent with Carey et al. (2003) and Fuchs and Rutledge (2018), Alabama showed the lowest 491 percentage of anomalous storms (7.3%), and Colorado anomalous frequency was much higher 492 than any other region (82.6%). Oklahoma and West Texas fell in between these two regions, and 493 with similar anomalous frequencies to Cordoba (12.9% for Oklahoma and 11.1% for West 494 Texas). From the flash centroid altitude distribution for the entire RELAMPAGO LMA dataset, 495 Lang et al. (2020), observed a peak at 10 km height and a secondary peak at 6 km height, the 496 latter attributed to anomalous storms and stratiform lightning. For the normal and anomalous 497 Cordoba events shown in Lang et al. (2020), normal and anomalous samples were consistently 498 depicted by the Chargepol algorithm. 499

500	Table 1. Number of normal	l and anomalous samples	for Cordoba, Alabama,	West Texas,
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501 Oklahoma, and Colorado.

	Cordoba	Alabama	West Texas	Oklahoma	Colorado
Number of days	157	32	48	41	61
Number of flashes (>20 sources)	808416	39046	261713	497139	545005
Number of flashes considered by Chargepol	165767 (20.5%)	7653 (19.5%)	65309 (24.9%)	58900 (11.8%)	62556 (11.4%)
Total number of samples	306	41	99	80	98
Normal samples	265	38	88	70	17
Anomalous samples	41	3	11	10	81
% Anomalous	13.3	7.3	11.1	12.5	82.6

502 The distribution of normal samples with altitude demonstrated that most normal dipoles were present in the mid-to-upper levels; i.e., with mid-level negative and upper-level positive 503 charge. Figure 8a shows the distribution of normal dipoles for Cordoba, Argentina. The altitude 504 distribution of anomalous samples (i.e., negative over positive dipoles) in Argentina shows that 505 there were cases in which negative charge was present in the upper levels with positive in the 506 mid-levels, and cases of negative in the mid-levels, with enhanced positive in the low levels 507 508 (Figure 8b). In Colorado, few normal samples were observed, but their altitude distribution is similar to Cordoba (Figure 8c). The distribution of anomalous samples with altitude in Colorado 509 showed that most dipoles had upper level negative and mid-level positive. No apparent presence 510 of an anomalous dipole located in the low-mid-levels occurred in the Colorado DC3 LMA 511 dataset. Therefore, the Cordoba 5 December 2018 case (Figure 6) demonstrates a singular 512 thunderstorm charge structure that is either rare or completely absent in Colorado. The normal 513 514 sample distributions in height for the other 3 U.S. locations (not shown) were similar to Cordoba and Colorado, while the anomalous sample distribution for these 3 locations (not shown) proved 515 inconclusive due to the low sample number. 516



517

Figure 8. Distribution of normal and anomalous samples with altitude for Cordoba (a, b) andColorado (c, d).

520 6 Summary and Discussion

This paper presented charge structures for the warm season thunderstorms in Cordoba, 521 Argentina for the first time through thunderstorm examples and long-term statistics utilizing a 522 new method that identifies charge layer polarity at a flash level from LMA VHF data. This 523 method is able to estimate general charge structures such as normal and anomalous dipoles, 524 tripoles, altitude and vertical depth of charge layers. Chargepol was applied to months of LMA 525 data, allowing for the inference of the frequency of anomalous and normal charge structure 526 thunderstorms in Cordoba, Argentina, and comparison to four well-studied U.S. regions using 527 the same methodology. 528

This method was developed from a meteorological standpoint, which means that the 529 530 objective was to obtain the general charge structure evolution through the entire thunderstorm life cycle, or for many hours of data. In order to achieve that, there was no need to retrieve 531 charge polarity from every flash as demonstrated in the comparison of Chargepol relative to 532 manual charge structure analysis and the VHF source distribution peak. Instead, only flashes 533 with less doubtful characteristics were used to provide an accurate charge polarity retrieval. 534 Hence, when considering such long periods of time, the frequency of anomalous and normal 535 charge structures can be estimated. Also, we found that it is sufficient to summarize the data into 536 the main dominant dipoles for every hour in order to characterize the charge structure for a 537 region. It is important to emphasize that, once charge layers are retrieved from individual 538 lightning flashes, one can organize this same dataset in any other manner depending on the user's 539 purpose. Examples include considering the algorithm output as a database to be organized into 540 shorter or longer time periods, obtaining the density of charge layers polarity over the time-541 altitude domain, calculating statistics for comparison with observations from other 542 instrumentation such as radar, etc. 543

The complexities of a three-dimensional charge structure that may be present at sub-544 storm scale, with charge layers extending through different altitudes depending on distance to an 545 updraft core, are not being fully accounted for in this study. For a flash analysis, we consider the 546 charge distribution over the vertical dimension only, which proved to be sufficient for this 547 study's objectives. For a given flash, the Chargepol method can estimate no more than two 548 charge layers with opposite polarities. However, when observing charge layers output for 549 numerous flashes, it is possible to infer the presence of dipoles, their altitude and time evolution, 550 the presence of tripoles and even multiple charge layers if flashes propagate through it. Only 551 charge layers that had flashes moving through them can be inferred. In the case of a positive 552 charge layer without a lightning flash moving through it, the charge layer cannot be visualized as 553 a product of the algorithm, which is a fundamental limitation of all LMA-based charge retrieval 554 methods (Rust et al., 2005). The fact that Chargepol neglects small flashes for charge layer 555 556 estimation, as it discards flashes with less than 20 sources, makes it hard to locate small pockets of charge within thunderstorms. Even if these charge regions were located, it could be hard to 557 visualize and interpret their evolution over minutes. Also, differentiating charge structure of 558 small flashes from noise would be challenging, an issue to be addressed in a future study. In 559 order for small flashes to be included in the analysis for identification of finer charge structures, 560 threshold of parameters have to be relaxed prior to running Chargepol. However, estimating the 561 general dipole and tripole charge structures is feasible with the conditions used in this study, 562 satisfying our purpose. 563

The Chargepol method proved capable for analyzing large LMA datasets in a reasonable 564 processing time of minutes, allowing for efficient interpretation of charge structures over 565 Cordoba, Argentina during the recent RELAMPAGO field campaign and a consistent 566 comparison of these novel results with thunderstorms from different regions of the United States 567 whose charge structures have been sampled with LMA and are more well understood. A high 568 frequency of anomalous storms were found for Colorado consistent with other studies (Fuchs et 569 al., 2018). Examples of Cordoba anomalous thunderstorms with altitude distributions of positive 570 charge layers that are uncommon in Colorado were presented. Interestingly, Cordoba showed 571 slightly higher anomalous charge structure frequency compared to Oklahoma and West Texas, 572 while Alabama presented the lowest anomalous frequency among all studied regions consistent 573 with prior work (Fuchs et al., 2018). Reasonings for these results were not explored in this study. 574

- 575 The meteorological, environmental, kinematic and microphysical conditions in Central
- 576 Argentina are speculated to be important contributors to the observed charge structures
- documented herein during RELAMPAGO, and they will be explored in future studies and
- compared to past work from other regions throughout the world. The charge polarity outputs
- 579 presented in this study have the potential to be useful for numerous applications in lightning
- research, and Chargepol has been made available as an open-source algorithm with options to choose parameter thresholds.

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- 589 https://doi.org/10.5067/RELAMPAGO/LMA/DATA101. NSF DC3 LMA data are available on
- 590 https://data.eol.ucar.edu/master_lists/generated/dc3/. Chargepol algorithm is available at
- 591 https://github.com/brmedin/chargepol.

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