Testing machine learning algorithms for the prediction of depositional fluxes of the radionuclides 7Be, 210Pb and 40K.

Pedro De la Torre Luque¹, Concepcion Duenas², Elisa Gordo³, and Sergio Cañete³

¹Istituto Nazionale di Fisica Nucleare Sezione Bari ²University of Malaga ³University of Málaga

November 23, 2022

Abstract

The monthly depositional fluxes of three natural radionuclides ($^{1}{Be}, ^{210}$ Be, 40 K) were measured at a Mediterranean coastal station (Malaga, Southern Spain) over a 14-year period from 2005 to 2018, corresponding to 168 monthly samples. The study of these radionuclides provides valuable information on the atmospheric air circulation, transportation and erosion processes as well as a control of the environmental radioactivity. In this work, the depositional fluxes of these radionuclides are investigated and their relations with several atmospheric variables, such as air temperature, pressure or precipitations, have been studied by applying two popular machine learning methods: Random Forest and Neural Network algorithms. We extensively test different configurations of these algorithms and demonstrate their predictive ability for reproducing depositional fluxes of $^{7}{7}$ Be, 210 Pb and 40 K. We use the Pearson-R correlation coefficient and the mean average error to evaluate the predictions of the developed models, revealing that the models derived with Neural Networks achieve slightly better results, in average, although similar, having into account the uncertainties. The mean Pearson-R coefficients, evaluated with a k-fold cross-validation method, are around 0.85 for the three radionuclides using Neural Network models, while they go down to 0.83, 0.79 and 0.8 for $^{7}{3}$ Be, 210 Pb and 40 SK, respectively, for the Random Forest models. Additionally, applying the Recursive Feature Elimination technique we determine the variables more correlated with the depositional fluxes of these radionuclides, which elucidates the main dependencies of their temporal variability.

Testing machine learning algorithms for the prediction of depositional fluxes of the radionuclides ⁷Be, ²¹⁰Pb and ⁴⁰K.

P. De La Torre Luque¹, C. Dueñas², E. Gordo³, S. Cañete³

¹Istituto Nazionale di Fisica Nucleare, Bari, via Orabona 4, I-70126 Bari, Italy
²Department of Applied Physics I, Faculty of Sciences, University of Malaga, Spain
³Central Research Services (SCAI), University of Malaga, 29071-Malaga, Spain

Key Points:

9	• Machine learning methods allow us to improve predictions on environmental ra-
10	dioactivity.
11	• Correlation found between Solar cycle and depositional fluxes of crustal radionu-
12	clides 210 Pb and 40 K.
13	- Machine learning models adapted for studying depositional fluxes of $^7\mathrm{Be},^{210}\mathrm{Pb}$

 14 and 40 K.

4

5 6 7

8

Corresponding author: P. De La Torre Luque, pedro.delatorreluque@fysik.su.se

15 Abstract

The monthly depositional fluxes of three natural radionuclides (⁷Be, 210 Pb and 40 K) were 16 measured at a Mediterranean coastal station (Malaga, Southern Spain) over a 14-year 17 period from 2005 to 2018, corresponding to 168 monthly samples. The study of these 18 radionuclides provides valuable information on the atmospheric air circulation, transporta-19 tion and erosion processes as well as a control of the environmental radioactivity. In this 20 work, the depositional fluxes of these radionuclides are investigated and their relations 21 with several atmospheric variables, such as air temperature, pressure or precipitations, 22 have been studied by applying two popular machine learning methods: Random Forest 23 and Neural Network algorithms. We extensively test different configurations of these al-24 gorithms and demonstrate their predictive ability for reproducing depositional fluxes of 25 ⁷Be, ²¹⁰Pb and ⁴⁰K. We use the Pearson-R correlation coefficient and the mean aver-26 age error to evaluate the predictions of the developed models, revealing that the mod-27 els derived with Neural Networks achieve slightly better results, in average, although sim-28 ilar, having into account the uncertainties. The mean Pearson-R coefficients, evaluated 29 with a k-fold cross-validation method, are around 0.85 for the three radionuclides using 30 Neural Network models, while they go down to 0.83, 0.79 and 0.8 for ${}^{7}\text{Be}$, ${}^{210}\text{Pb}$ and ${}^{40}\text{K}$, 31 respectively, for the Random Forest models. Additionally, applying the Recursive Fea-32 ture Elimination technique we determine the variables more correlated with the depo-33 sitional fluxes of these radionuclides, which elucidates the main dependencies of their tem-34 poral variability. 35

³⁶ 1 Introduction

The use of natural radionuclides as markers for studying the atmospheric circula-37 tion provides valuable information about the complex mechanisms involved. It is com-38 mon to employ different natural radionuclides as tracers and chronometers in aquatic and 39 atmospheric systems (Wogman et al., 1968; Martell, 1970; Schuler et al., 1991) and they 40 have demonstrated to be very useful in studies dedicated to understand the mechanisms 41 and rates of removal of aerosols (Baskaran et al., 1993). In this work, we aim at the study 42 of a predictive model for the depositions of fallout radionuclides ⁷Be, ²¹⁰Pb, and ⁴⁰K, 43 whose different origins allow us to infer important features of the atmospheric circula-44 tion, erosion processes, transportation and deposition of soils and sediments from episodic 45 to long-term timescales. 46

⁴⁷ ⁷Be is a cosmogenic radionuclide originated by spallation reactions of cosmic rays ⁴⁸ with light atmospheric nuclei, such as nitrogen and oxygen (Lal et al., 1958) that has ⁴⁹ a decay half-live of $T_{1/2} = 53$ day. Thus, this nuclide is mostly produced in the strato-⁵⁰ sphere and reach the troposphere in periods of air exchange between these two layers. ⁵¹ This is why the production of ⁷Be is dependent on altitude, latitude and solar cycle but ⁵² has negligible dependence on longitude (Baskaran et al., 1993; Dueñas et al., 2017).

In contrast, ²¹⁰Pb, with a decay half-live of $T_{1/2} = 22.3$ yr, is produced from the radioactive decay of ²²²Rn, the only gaseous decay product of ²³⁸U series. Therefore, ²¹⁰Pb is found in larger concentrations near the ground and with important dependence on the distribution of land and seas (Moore et al., 1973; Wilkening et al., 1975; Preiss et al., 1996; Garcia-Orellana et al., 2006)

The atmospheric ⁴⁰K ($T_{1/2} = 1.3 \cdot 10^9$ yr) is related to a crustal origin, from most kinds of soil, which is usually found in association with other re-suspended materials, as PM10 (particulate matter with diameter 10 μm) from the African continent (Karlsson et al., 2008; Dueñas et al., 2011).

Several works in the past have been dedicated to study the relations between the
 concentrations or depositional fluxes of these radionuclides with different environmen tal variables for different latitudes and longitudes. In this work, we employ a large dataset



Figure 1. Physical map showing the location of the study area. The zoomed window shows the exact position of study area, in Málaga.

(168 monthly measurements, from January 2005 to December 2018) of environmental
variables and the fluxes of ⁷Be, ²¹⁰Pb, and ⁴⁰K radioactivity in the Mediterranean coastal
region of Málaga (Southern Spain). Similar studies were carried out in the same zone
in the past and reported some important results, such as correlations with particulate
material (PM10 levels) or with other environmental variables included in this work (Dueñas
et al., 2004, 2009, 2011, 2017).

Here, we are exploring new methods of studying the complex relations between the 71 depositional flux of these radionuclides and atmospheric variables, using machine learn-72 ing algorithms. Machine learning (ML) techniques (Carbonell et al., 1983) provide a promis-73 ing tool in the prediction of any magnitude which depends on a large number of vari-74 ables and exhibits complex relations with them. Particularly, we are focused here on the 75 implementation of these methods for the prediction of depositional fluxes of the men-76 tioned radionuclides. These models allow us to identify subtle long-term relationships 77 between the temporal variability of the depositional fluxes and other environmental cy-78 cles, like the Solar cycle or atmospheric cycles. Additionally, reproducing these fluxes 79 allow us discern the real agents affecting the depositions of these radionuclides and could 80 provide another tracer of anomalous (artificial) radiation episodes. In addition, we ar-81 gue that these kind of models could be extended to different zones, always that measure-82 ments are available, to study relations with other variables not yet taken into account. 83

⁸⁴ 2 Materials and measurements

2.1 Study area

85

Málaga (4° 28' 8" W; 36° 43' 40"N), is the major coastal city in the Andalusian 86 region situated in the south-east of Spain (see Figure 1), on the Mediterranean coast and, 87 therefore, has a climate influenced by continental and maritime air masses. The predom-88 inant winds are easterly (SE) and westerly (NW). The climate is temperate, with con-89 trasting wet (approximately October–April) and dry (approximately May–September) 90 periods (Dueñas et al., 2012). The city is almost surrounded by mountains, which cause 91 a special wind regime. Due to its geographical proximity to the African continent, our 92 study area is frequently affected by intrusions of air masses with high concentrations of 93 atmospheric particulate matter (Escudero et al., 2005). The sampling point is located on the flat roof of the Central Research Services (SCAI) building at the University of 95 Málaga, at a height of 10 m above the ground and approximately at 5 km from the coast-96 line, near the airport and surrounded by roads with traffic exhaust. 97

98 2.2 Data extraction

Bulk deposition samples were collected from January 2005 to December 2018. Sam-99 ples were collected monthly using a collector that it is slightly tilted stainless steel tray 100 1 m^2 in area and a polyethylene vessel of 60 L capacity for rainwater sample reservoir. 101 A volume of 6 L of the bulk deposition (the sum of wet deposition flux and the gravi-102 tational sedimentation fraction of the dry deposition) was reduced via evaporation to ap-103 proximately 1 L and transferred to a Marinelli geometry container for gamma counting. 104 The method and processing procedures were described previously (Dueñas et al., 2011). 105 106 The atmospheric fluxes were calculated using the expression:

$$F = A/St \text{ (Bq m}^{-2} \text{ month}^{-1}), \tag{1}$$

where A is the activity in the sample obtained from the gamma spectra, S is the surface area of the collector and t is the duration of sampling time. Additionally, aerosol samples were collected weekly in cellulose filters of $0.8 \,\mu\text{m}$ pore size and 47 mm diameter with an air sampler (Radeco, mod AVS-28A) at a flow rate of 40 l/min. A monthly composite sample containing 4 or 5 filters (depending on the number of weeks each month) was formed for the gamma analysis.

Radiometric measurements were performed by low-level gamma spectrometry with 113 a coaxial-type germanium detector (Canberra Industries Inc., USA), with a relative ef-114 ficiency of 20% and it was calibrated using certified reference gamma ray cocktail. Each 115 sample was measured for 172,0000 s. Gamma spectra analyses were performed with the 116 Genie2K spectrometry software version 2.0 (Canberra Industries Inc., USA). The char-117 acteristic gamma peaks selected for the determination of the different radionuclides were: 118 477.6 keV for 7 Be, 1460.81 keV for 40 K and 46.5 keV for 210 Pb. To validate the meth-119 ods, our lab routinely participates in interlaboratory comparisons to measure gamma-120 emitting radionuclides, in different types of samples, organized by the International Atomic 121 Energy Agency (IAEA), the Joint Research Centre (JRC), and the Spanish Nuclear Safety 122 Council (CSN). Further details of the low-background gamma-ray detection system have 123 been previously described by refs. Dueñas et al. (1999, 2004). 124

The meteorological data (temperature, relative humidity, distance travelled monthly by the wind and precipitation) used in this study were obtained from the nearest station network of the Spanish Meteorological Agency (AEMET) (500 m away from the sampling site). Days affected by African dust outbreaks have been obtained from CALIMA project (www.calima.es). The monthly sunspots number were obtained from NOAA's Space Weather Prediction Center (SWPC).

Additionally, data of daily concentrations of particulate matter fraction PM10 were
 obtained from Carranque (36^o 43' 40" N; 4^o 28' 4" W), a monitoring station belonging
 to the regional Atmospheric Pollution Monitoring network managed by the Environmen tal Health Service of the Andalusian Government.

3 Methods: description of the algorithms applied and cross-validation framework

ML techniques have demonstrated their predictive power in a variety of fields, from 137 medicine (e.g. (Lapedes et al., 1988)) to astrophysics (e.g. (Schaefer, C. et al., 2018), (Graff 138 et al., 2014)), used for both classification (as in (Williams et al., 2006)) and numerical 139 forecasting (see, for example, refs. Sarkar et al. (2009); vStencl and Stastny (2011)). Gen-140 erally, ML methods are used to find the relation between a set of input variables and an 141 output variable one is interested in. These variables are usually called features and la-142 bels, respectively. In the present study, the labels are the monthly depositional fluxes 143 collected from 2005 to 2018 and the features are the atmospheric variables gathered in 144 the same period. Earlier studies have demonstrated that it is possible to find linear re-145 lations between atmospheric variables and the depositional fluxes of these radionuclides, 146

although the uncertainties related to this determination become too large to have accurate predictions. Using these methods we aim at obtaining more precise predictions
on the depositional fluxes that could be used, e.g., to reliably detect the emission of artificial radiation or other non-expected radiation sources.

The relation between features and labels is progressively adjusted by iterating over 151 the amount of data samples given to the algorithm, therefore the larger the amount of 152 samples used to feed (or train) the algorithm the better the predictions become. The data 153 sample used to adjust the algorithm is called training dataset and this adjustment pro-154 155 cess is known as the training phase, which basically consists on tuning some training parameters in order to predict the correct labels given. The algorithm adjusts itself in each 156 iteration by comparing its predicted label with the correct label. Then, in order to eval-157 uate the performance of the model one must provide it with new input data (i.e. these 158 features must be different from the training data to ensure unbiased or over-fitted eval-159 uations of the algorithm effectiveness). In this way, we can "grade" or "score" the model 160 performance by comparing the predicted outputs with the real labels in what is called 161 the test phase. The new set of data used in this phase is called test data. 162

Two different supervised algorithms have been implemented in this study; Neural Networks and Random Forest techniques, and their ability to predict depositional fluxes has been extensively tested for different configurations and for the depositional fluxes of the 7Be , ^{210}Pb and ^{40}K radionuclides. Very few works have been published using ML techniques to predict depositional fluxes and none of them systematically analyzing their performance. An example of these studies can be found in ref. Chham et al. (2018), but a deeper research on the efficiency of these techniques is necessary.

The most popular ML algorithm is the *Artificial Neural Network* (ANN) model. Neural networks can learn complex patterns using layers of neurons which mathematically transform the data. The layers between the input and output are referred to as "hidden layers". A Neural Network can learn relationships between the features that other algorithms cannot easily discover, including also complex non-linear relations.

Moreover, we used an alternative and less demanding (in terms of resources) technique, the *Random Forest* algorithm ¹, which, in turn, is not able to consider non-linear features in the relations between the features. This algorithm relies in an ensemble of decision trees which are combined to get averaged predictions. Each tree uses a sub-sample of the full data set, randomly selected, and progressively divides it into different nodes (or leaves) depending on certain quantitative (or qualitative, in case the tree is applied for a classification problem) criteria decided by the algorithm.

We have divided our collected data set into a training set, containing the 80-85%182 of the full data set, and a test set that allows us to quantify the performance of out pre-183 dictions. The list of features (meteorological or atmospheric variables) employed is based 184 on monthly averages (or monthly accumulated) and it consists of: Air temperature (in 185 0 C), relative humidity level (%) number of days affected by African dust outbreaks (in-186 trusions), distance travelled monthly by the wind (in km), pressure (hPa), sunspot num-187 ber, amount of rainfall (dm^3) , PM10 level $(\mu g/m^3)$, seasonal factor (from 1, for winter, 188 to 4, for spring), monthly factor (from 1, for January, to 12, for December), total rain-189 fall duration (min), humid days, dry days and time between rains (in days). For both 190 algorithms, the labels (depositional fluxes) are normalized, since this allows a better per-191 formance of the algorithm. 192

A Neural Network in which the input features first result into 8 units (1st hidden layer) and then into 4 units (second hidden layer) have been found to be the most adequate, as it is depicted in the Appendix A. The implementation of the Neural Network

 $^{^1\,{\}rm Specifically}$ the method RandomForestRegressor given by the package of sklearn.ensemble

	$^{7}\mathrm{Be}$	210 Pb	$^{40}\mathbf{K}$
Learning rate	2.1e-3	2.1e-3	2.2e-3
Decay rate	5.e-6	5.e-5	2.4e-6

Table 1.Main hyperparameters (i.e. the values needed to control the learning process in MLalgorithms) used in the Adam optimizer, adjusted for each of the radionuclides studied.

has been achieved by using the *Python Keras* (Chollet, 2015) library. The connections
between the input features and the first hidden layer, as well as between the first and
second hidden layers use the Rectified Linear Unit (ReLU) as activation function and
the connections from the second hidden layer and the output units are calculated with
a linear activation function.

The model performance was optimized including a step of batch normalization and dropout (finding the best results adjusting it to the 10% of the sample) after each of the hidden layers. In addition, the adaptive moment estimation optimizer, or *Adam* optimizer², was found to get the best performance for every one of the radionuclides. On top of this, the best results were found when taking the natural logarithm of the values for the features, as expected, and setting the mean absolute error metrics as the loss function.

Different configurations of the neural networks models and the hyperparameters 208 involved (i.e. the values needed to control the learning process in ML algorithms) were 209 refined by applying a simple random search method (i.e. probing different hyperparam-210 eters in an equally spaced grid of values) (Bergstra & Bengio, 2012). The optimization 211 of the combination of these hyperparameters is left for a next work. In table 3, we show 212 the main hyperparameters tuned for the Adam optimizer for each radionuclide. The rest 213 of hyperparameters needed by the optimizer were set to their default values given by the 214 keras method. 215

For the Random Forest algorithm, it was found that using the features values normalized, instead of their natural logarithm, gave better results. Then, the main hyperparameters were adjusted for each of the nuclides, setting the mean absolute error (MAE) as criterion for splitting the nodes and a minimum number of samples required to split an internal node (*min_samples_split*) to 3. The number of decision trees (also known as number of estimators) used in the model was set to be 680 for ⁷Be and ²¹⁰Pb and 280 for ⁴⁰K.

The results from both algorithms and for the three radionuclides are shown and compared in the next section, in which we fully demonstrate their ability for reproducing the data and systematically explore the statistical errors around these predictions as well as the main features involved.

4 Results: predictive power of the algorithms

As a first step before running our models, we randomly shuffle the features and labels and, then, they are divided into a training and a test sets. Once the model is trained, we rate its performance by comparing the predictions with the test labels, corresponding to a 15-20% of the full data sample, using the mean percentage error and the Pearson-

² https://keras.io/api/optimizers/

R index value. While the former is an indicator of the quantitative differences between
test labels and predictions, the latter is a good indicator of the trend similarities between
the two sets.

In order to compare these results with a reference model, we applied the same kind of evaluation as applied for the ML algorithms to the model found in the linear regression analysis presented in ref. Dueñas et al. (2017) for the ⁷Be radionuclide. This analysis yields a linear relation between the depositional flux of ⁷Be and the amount of rainfall (the variable which shows the largest correlation with the depositional flux of every radionuclide) of:

$$Flux_{Be} = 6.33 + 2.6 \times \text{rainfall} \tag{2}$$

Then, the evaluation is carried out by using a portion of 25 randomly selected mea-241 surements (similar to the amount of samples in the test sets used for the ML algorithms 242 applied) of ⁷Be and amount of rainfall (corresponding to the same date) and measured 243 the Pearson-R index and mean error of the predictions obtained with this reference model. 244 In order to have a robust idea on the value of these metrics, we repeated this for 100 times 245 (analogous to what is done in section 4.1), with different randomly selected samples of 246 25 measurements, and computed the average value. These metrics result in a mean R 247 index of $\sim 0.45 \pm 0.4$ and a maximum R index of 0.95, while the mean percentage er-248 rors were of 103±150%. Having these reference metric values is necessary to compare 249 to the quantitative results of the Random Forest and Neural Network algorithms stud-250 ied here. In Figure B1 (Appendix B), we display the comparison between the predictions 251 from the reference model and the depositional flux measurements for one of these sam-252 ples. 253

In comparison, in Figure 2, we show some of the best results acquired from the Neu-254 ral Network and Random Forest algorithms for all the studied radionuclides, which demon-255 strates that these algorithms can allow us to significantly improve our predictions on de-256 positional fluxes with respect to traditional methods. Here, we highlight that these are 257 predictions obtained from their corresponding atmospheric variables, and remark the im-258 portance of evaluating these predictions with data not used for the training phase, since 259 this highly biases our evaluation. As we can see by the Pearson-R value, these predic-260 tions are able to suitably reproduce the labels trend with respect to the atmospheric vari-261 ables. In addition, we find mean absolute errors of the order 50% usually, which are well 262 below the error levels found using linear regressions (as shown above) and are similar to 263 the experimental uncertainties in the determination of these fluxes, which can be $\mathcal{O}(10\%)$, 264 as shown in refs. Herranz et al. (2008); Heydorn (2004). In this case, it has been observed 265 that high-flux values are difficult to be matched, which may be related to periods of anoma-266 lous radiation doses. Nevertheless, this requires a dedicated study of those points and 267 their temporal behaviour, which is beyond the scope of this paper. Further sources of 268 uncertainty in these comparisons mainly come from the statistical uncertainties related 269 to the measurement of the atmospheric variables and variables not included in the model. 270

Surprisingly, the models make good predictions also for the ⁴⁰K nuclide, even with 271 a considerably smaller number of samples available for it. On top of this, we found that 272 the absolute percentage errors follow a similar distribution for each radionuclide and both 273 algorithms. They are well described with a Gamma probability distribution, which ex-274 hibits a slightly negative mode and a slightly positive median. This is likely due to the 275 fact that the distribution of depositional fluxes is also be very well reproduced with a 276 Gamma function. A representative example of these distributions for the Neural Net-277 work and Random Forest algorithms is shown in Figure 3 for the ⁷Be radionuclide af-278 ter gathering several repetitions for different test sets used. The fact that these errors 279 follow such distribution can be used to statistically diagnose anomalous episodes of ra-280 diation doses. We noticed that the Random Forest models produce slightly larger me-281

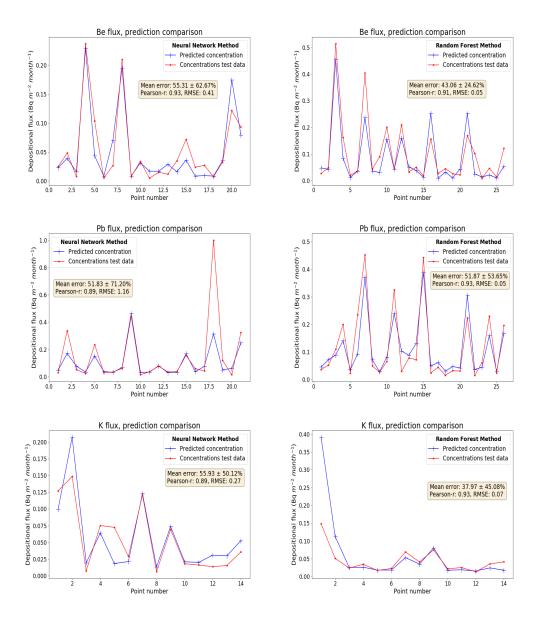


Figure 2. Example of the results of the predictions found from the Neural Network (left panels) and Random Forest (right panels) models. These predictions are limited to the test sample, which is chosen to be around a 20% of the full data set. We also include the values of the metrics used to evaluate the predictive ability of these methods, which are the Pearson-R correlation coefficient and the mean absolute error and its standard deviation. The root mean square error (RMSE), in units of Bq m⁻² month⁻¹, is also included for completeness.

Nevertheless, the evaluation of the models is highly dependent on the data set used.
From one side, the larger the test set, the more reliable is the model performance evaluation, but at the cost of reducing the number of samples used in the training set. On
the other side, if the test set is too short, the model performance evaluation will be very
uncertain. In this case, we observed that using around 20% of the full data set allowed

dian values and mode values more deviated from 0, but no significant differences between same algorithms for different nuclides was detected.

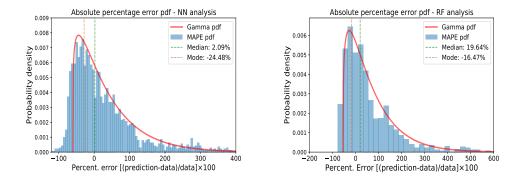


Figure 3. Probability distribution for the percentage errors found for various evaluations with (around 20) different tests sets. The left plot shows the results of these evaluations for the Neural Network algorithm and the right plot those for the Random Forest algorithm.

us to make consistent evaluations. Even though, they are still short enough to make our
evaluation very dependent on the data test used. This issue is well known by the ML
community and there are many possible strategies to deal with it to have an unbiased
evaluation of our model (Raschka, 2018) and its predictions uncertainties, as it is explored
in the next section.

4.1 Statistical evaluation

To prevent from biasing our model evaluation by the small amount of test data used and have into account the full uncertainty involved, we evaluate the algorithms by means of a k-fold procedure. In this process the data set is divided into k subsets. Each time, one of the k subsets is used as test set and the other k-1 subsets form the training set. Then, we statistically combine the results to get solid conclusions.

At this point, another difference between the Neural Network and the Random For-300 est algorithms should be taken into account to correctly manage the full uncertainties 301 involved: while the training process exactly results in the same model for the Random 302 Forest algorithm, this is subject to further fluctuations in the Neural Network algorithm. 303 This is due to the optimization procedure necessary for finding the minimum error or 304 loss when evaluating the examples in the training dataset. The main problems usually 305 faced are: getting stacked in local minimal or local optima (i.e. regions where the loss 306 is relatively low but it is not the lowest), saddle or flat points (regions where adjustments 307 of the training weights do not lead to an appreciable change in the loss) and other is-308 sues more related with the loss function, gradients and the dimensionality involved. More 309 precise information about these problems can be found, e.g., in ref. Bengio (2012). There-310 fore, each time the Neural Network is trained, specially when the number of samples is 311 not large enough, it is subject to small variations in the model predictions. For this rea-312 son, a good evaluation of the uncertainties involved in the predictions of the Neural Net-313 work model requires to add these fluctuations. 314

In particular, we repeated the training and test phase for 5 times with the same test and training datasets. Then, we perform the evaluations with 20 different randomlyselected test sub-datasets following the k-fold procedure. This means that we carry out a total of 100 training and evaluation steps to determine the Pearson-R value and the mean percentage error of our predictions with respect to the experimental data, as well as the uncertainties related to these determinations for the Neural Network model. In turn, as the Random Forest algorithm does not suffer from those training fluctuations, we performed 60 evaluations of the model, employing a different test and training subsets, accordingly, in each evaluation.

These results are shown in Figure 4, where we represent the mean Pearson-R in-324 dex values and the 1σ uncertainty related to its determination for both, the Neural Net-325 works and Random Forest algorithms and for the three nuclides with respect to the num-326 ber of iterations employed in the training phase. In general, we observe that the mean 327 Pearson-R index values are larger for the ⁷Be and 210 Pb radionuclides, while 40 K shows 328 the opposite, due to the smaller number of samples available. In addition, the uncertain-329 ties related to the determination of the R index value from the Random Forest algorithm 330 is slightly larger than that from the NN algorithm. The mean Pearson-R index values 331 obtained are between 0.75-0.88 for ⁷Be and ²¹⁰Pb, but around 0.7-0.8 for ⁴⁰K, although 332 the errors are still high for every radionuclide. In particular, the determination of ⁷Be 333 seems to be the most accurate in general, showing a 1σ uncertainty in the determina-334 tion of the R index value around ± 0.065 for the NN algorithm and ± 0.08 for the RF al-335 gorithm. A maximum mean R index value of around 0.87 and 0.88 are found for ⁷Be and 336 ²¹⁰Pb, respectively, at 1400 and 1300 iterations. The maximum mean R index value ob-337 tained for 40 K is slightly above 0.8, found with the RF algorithm. 338

As expected, the performance of these methods in reproducing depositional fluxes improves when having more samples, obtaining larger Pearson-R index values and lower uncertainties related. Nevertheless, we observed that the NN algorithm seems to accuse more the smaller number of samples with respect to the RF technique.

4.2 Selecting the main variables

To fully exploit the capability of ML techniques in improving our predictions in the 344 depositional fluxes, we determined which are the most important features using the re-345 cursive feature elimination algorithm (RFE), which allows us to reduce the complexity 346 and needed cpu time of the Neural Network and Random Forest algorithms and prevents 347 from over-fitting our results. In addition, we compared the results obtained with these 348 features with those obtained when using all the features. Specifically, we used the RFECV349 method from the *sklearn.feature_selection* python package. The RFE algorithm is a fea-350 ture selection method that allows a model to progressively eliminate the weakest features 351 and find the best scoring combination of features. 352

In Figure 5 we show the optimal important features found by the RFE algorithm, along with their relative importance. As expected, the rainfall duration and rainfall volume are selected by the three radionuclides. Then, we observe that other atmospheric variables are present, as the number of humid or dry days, the average monthly pressure or the mean air temperature. On the other hand, the PM10 level and sunspot number are selected as important for the ⁴⁰K nuclide.

The fact that the sunspot number arises as one of the most important variables de-359 scribing the depositional fluxes of ²¹⁰Pb and ⁴⁰K is unexpected. In principle, this vari-360 able is expected to be relevant for the production of 7 Be since it is related with the so-361 lar activity (this is, the Sun's magnetic field), which plays an important role on the flux 362 of cosmic rays reaching the atmosphere (Yoshimori et al., 2003). This fact is probably 363 due to the mild correlations between sunspot number and other atmospheric variables, but more data samples are needed to get a solid conclusion, since the sunspot number 365 follows cycles of 11 and 22 years, following the solar magnetic cycles (E.W., 2015). This 366 could be explained by the fact that there are other correlations found between the so-367 368 lar cycle and other atmospheric variables, as the atmospheric temperature (Qu et al., 2012) and correlations with the cosmic-ray intensity at Earth, which is known to be re-369 lated to climate and involved in processes of cloud formation (Veretenenko et al., 2018; 370 Svensmark et al., 2013; Marsh & Svensmark, 2000). 371

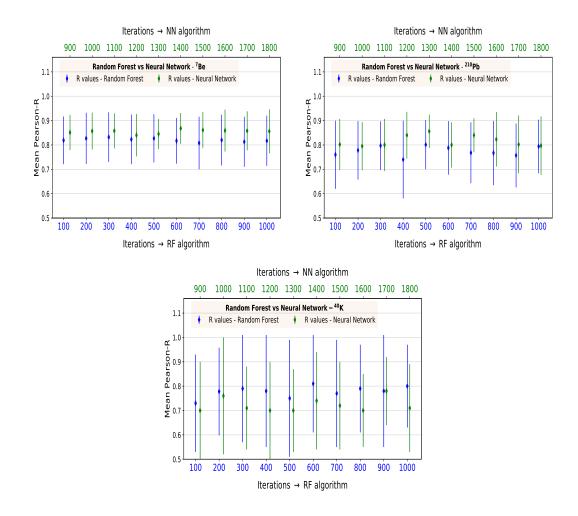


Figure 4. Results from the k-fold evaluation of the Pearson-R correlation coefficient for the Neural Network and Random Forest algorithms for the depositional fluxes of ⁷Be (upper panel), ^{210}Pb (middle panel) and ^{40}K (lower panel). The results obtained from the NN algorithm are shown in green while the results from the RF algorithm are shown in blue.

Once these features have been selected, we proceed to compare the NN and RF al-372 gorithms explored in this work using all the features and using just the important fea-373 tures, as displayed in Figure 6. From this figure, we can see that the NN models for 40 K 374 have significantly improved, restricting our features to be just the important ones. This 375 means that some of the eliminated features were over-fitting the model. This can be re-376 lated to the fact that this radionuclide actually comes from African zones and reach coastal 377 zones of Southern Spain after it is transported by winds in the correct direction. There-378 fore, some of the atmospheric variables measured in the zone of Malaga could not be suit-379 able to describe its amount and depositions in Malaga. Even though, the amount of rain-380 fall should still be crucial to make the African dust to definitely fall in the study region. 381 Furthermore, the presence of the sunspot number as an important feature have not been 382 pointed out in the past, which may mean that there are other atmospheric variables with 383 a considerable role in the amount and depositional flux of 40 K found in the Mediterranean 384 coastal zone of the Southern Spain. 385

On the other hand, we see that for ⁷Be and ²¹⁰Pb the results remain very similar to the case with all the features, which is quite remarkable given the number of variables

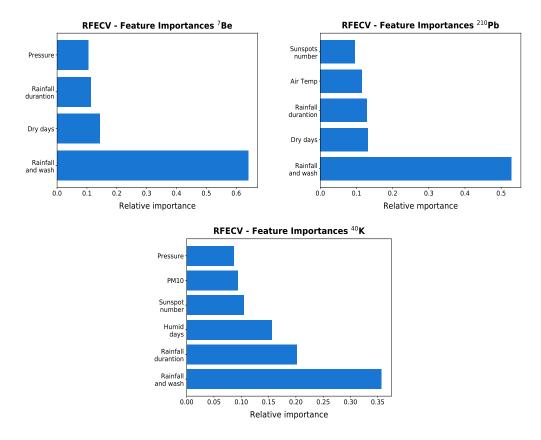


Figure 5. Histograms with the important features found with the implemented recursive feature elimination algorithm for the depositional fluxes of ⁷Be (upper panel), ²¹⁰Pb (middle panel) and ^{40}K (lower panel) with their relative importance.

needed. In addition, the uncertainties related to the determination of the Pearson-R correlation coefficient have been considerably reduced in the NN models for ⁴⁰K, while they seem to be almost identical for all other cases.

In general, these results are consistent with other previously found, but the use of these ML methods allow our predictions to be more complex and better adapt to the variability related to the depositional fluxes of different radionuclides.

³⁹⁴ 5 Conclusions

Modern computer algorithms allow us to refine our measurements and model pre-395 dictions via new statistical tools or artificial intelligence. In this work, we have made use 396 of two common machine learning algorithms, Neural Networks and Random Forests, in 397 order to predict and analyse the depositional fluxes of ⁷Be and ²¹⁰Pb and ⁴⁰K. This work 398 has shown, first, that these methods can be successfully applied to study the depositional 399 fluxes of different radionuclides from atmospheric variables as the amount of rainfall, pres-400 sure or air temperatures. Second, we have evaluated the performance of these models 401 using a k-fold method and the Pearson-R coefficient and mean absolute error as metrics 402 finding that these techniques can significantly improve old predictions made from mul-403 tivariate linear regression analyses. 404

A05 As expected, the performance of these methods in reproducing depositional fluxes improves when having more samples, obtaining larger Pearson-R index values and lower

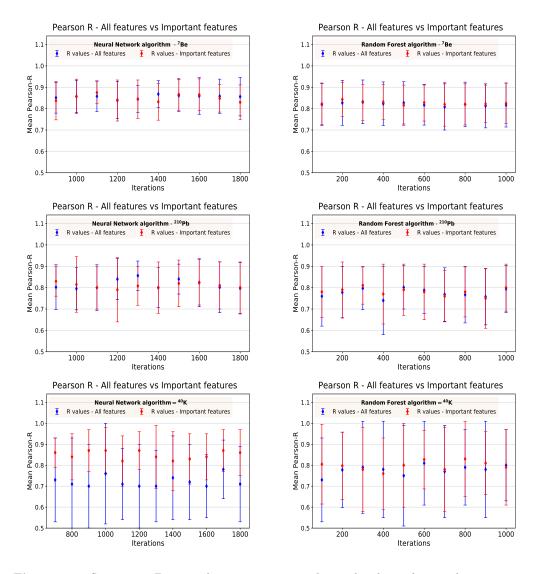


Figure 6. Same as in Figure 4, but comparing now the results obtained using the main variables obtained from the RFE algorithm and those obtained from the models trained with all the available variables in the data set.

uncertainties related. This, in fact, confirms the prospects on future models, with a larger 407 number of samples measured. This is mainly related to the long times involved in the 408 natural cycles of atmospheric variables, as, for example, the sunspot number, which is 409 known to follow 11 or 22-years periods (solar magnetic cycles). Nonetheless, we have demon-410 strated that the algorithms employed here are able to reproduce the experimental de-411 positional fluxes using monthly-averaged variables and that these predictions can help 412 identifying periods of anomalous radiation doses. Interestingly, we found that both, the 413 depositional fluxes of ²¹⁰Pb and ⁴⁰K, seem to be correlated with the Sunspot number. 414

The Neural Network models seem to reach higher mean Pearson-R index values, calculated using a k-fold cross-validation treatment, almost reaching 0.9, although the uncertainties are still quite high. Furthermore, the use of a Recursive Feature Elimination algorithm has been used to find the variables that perform the best predictions and allow us to reduce to 4, 5 and 6 the number of variables used for predicting the depositional fluxes of ⁷Be, ²¹⁰Pb and ⁴⁰K, respectively. The training of the Neural Network

and Random Forest models with these variables resulted into a negligible difference in 421 the Pearson-R index values and the uncertainties related to its determination except for 422 the 40 K nuclide in the Neural Network model, which showed a significant improvement. 423 Even with this reduced number of variables used for training our methods, we were able to obtain mean values for the Pearson-R index value above 0.80 for all the three nuclides 425 and both algorithms. A maximum mean R index value around 0.87 is found for ^{7}Be , ^{210}Pb 426 and ⁴⁰K, respectively, at 1400, 1300 and 1200 iterations for the Neural Network method. 427 For the Random Forest method, the maximum mean R index value of sim 0.81 is found 428 around 500 and 600 iterations for 210 Pb and 40 K and of almost 0.85 for the 7 Be radionu-429 clide. 430

In conclusion, we demonstrate that Random Forest and Neural Networks methods 431 are able to improve our current knowledge and predictions on the depositional fluxes of 432 radionuclides in the Mediterranean coastal zone of Malaga and these models can be ex-433 tended to other zones too, in order to build a more complex ensemble that could refine 434 the existent knowledge on deposition of different radionuclides. Thus, this work consti-435 tutes the first step into the study of a large-scale (in terms of geographical areas) model 436 able to make predictions on depositional fluxes for different geographical zones thanks 437 to the adaptability of these algorithms. The implementation of a recurrent neural net-438 work applied to the prediction of depositional fluxes can improve these models and will 439 be also investigated in a next work. 440

441 Acknowledgments

We would like to express our gratitude to the Consejo de Seguridad Nuclear, Spain, for their financial support to the Environmental Radioactivity Laboratory of the University of Málaga.

445 **References**

446	Baskaran, M., Coleman, C. H., & Santschi, P. H. (1993). Atmospheric depositional
447	fluxes of 7be and 210pb at galveston and college station, texas. Journal of
448	Geophysical Research: Atmospheres, 98(D11), 20555-20571. Retrieved from
449	https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/93JD02182
450	doi: https://doi.org/10.1029/93JD02182
	Bangia V (2012) Practical recommandations for gradient based training of deep

- Bengio, Y. (2012). Practical recommendations for gradient-based training of deep architectures. In G. Montavon, G. B. Orr, & K. Müller (Eds.), Neural networks: Tricks of the trade - second edition (Vol. 7700, pp. 437–478). Springer. Retrieved from https://doi.org/10.1007/978-3-642-35289-8_26 doi: 10.1007/978-3-642-35289-8_26
- Bergstra, J., & Bengio, Y. (2012, February). Random search for hyper-parameter optimization. J. Mach. Learn. Res., 13(null), 281–305.
- Carbonell, J. G., Michalski, R. S., & Mitchell, T. M. (1983). Machine learning: a
 historical and methodological analysis. AI Magazine, 4(3), 69–69.
- Chham, E., Piñero-García, F., Brattich, E., El Bardouni, T., & Ferro-García, M.
 (2018). 7be spatial and temporal pattern in southwest of europe (spain): Evaluation of a predictive model. *Chemosphere*, 205, 194 202. Retrieved from http://www.sciencedirect.com/science/article/pii/S0045653518307483
 doi: https://doi.org/10.1016/j.chemosphere.2018.04.099
- ⁴⁶⁵ Chollet, F. (2015). *Keras.* https://github.com/fchollet/keras. GitHub.
- ⁴⁶⁶ Dueñas, C., Fernández, M., Carretero, J., Liger, E., & Cañete, S. (2004). Long ⁴⁶⁷ term variation of the concentrations of long-lived rn descendants and cosmo ⁴⁶⁸ genic 7be and determination of the mrt of aerosols. Atmospheric Environ ⁴⁶⁹ ment, 38(9), 1291 1301. Retrieved from http://www.sciencedirect.com/
 ⁴⁷⁰ science/article/pii/S1352231003010549 doi: https://doi.org/10.1016/

471	j.atmosenv.2003.11.029
472	Dueñas, C., Fernández, M., Cañete, S., & Pérez, M. (2009). 7be to 210pb con-
473	centration ratio in ground level air in málaga (36.7 ^o n, 4.5 ^o w). Atmospheric
474	Research, 92(1), 49 - 57. Retrieved from http://www.sciencedirect.com/
475	science/article/pii/S0169809508002342 doi: https://doi.org/10.1016/
476	j.atmosres.2008.08.012
477	Dueñas, C., Fernández, M., Gordo, E., Cañete, S., & Pérez, M. (2011). Gross
	alpha, gross beta activities and gamma emitting radionuclides composition
478	of rainwater samples and deposition to ground. <i>Atmospheric Environment</i> ,
479	45(4), 1015 - 1024. Retrieved from http://www.sciencedirect.com/
480	science/article/pii/S1352231010009313 doi: https://doi.org/10.1016/
481	- ,, -, ,,
482	j.atmosenv.2010.10.045 Γ
483	Dueñas, C., Fernández, M., Gordo, E., Cañete, S., & Pérez, M. (2012). Chemical
484	and radioactive composition of bulk deposition in málaga (spain). Atmospheric
485	Environment, 62, 1-8. Retrieved from https://www.sciencedirect.com/
486	science/article/pii/S1352231012007595 doi: https://doi.org/10.1016/
487	j.atmosenv.2012.07.073
488	Dueñas, C., Fernández, M., Liger, E., & Carretero, J. (1999). Gross alpha, gross
489	beta activities and 7be concentrations in surface air: analysis of their varia-
490	tions and prediction model. Atmospheric Environment, $33(22)$, $3705 - 3715$.
491	Retrieved from http://www.sciencedirect.com/science/article/pii/
492	S1352231099001727 doi: https://doi.org/10.1016/S1352-2310(99)00172-7
493	Dueñas, C., Gordo, E., Liger, E., Cabello, M., Ca, S., Pérez, M., & de la
494	Torre Luque, P. (2017, 11). 7 be, 210 pb and 40 k depositions over 11 years
495	in m alaga. Journal of Environmental Radioactivity, 178-179, 325-334. doi:
496	10.1016/j.jenvrad.2017.09.010
497	Escudero, M., Castillo, S., Querol, X., Avila, A., Alarcón, M., Viana, M., Ro-
498	driguez, S. (2005, 09). Wet and dry african dust episodes over eastern spain.
499	Journal of Geophysical Research, 110, 18-8. doi: 10.1029/2004JD004731
500	E.W., C. (2015). The extended cycle of solar activity and the sun's 22-year magnetic
501	cycle. Space Sciences Series of ISSI, 53. doi: https://doi.org/10.1007/978-1
502	-4939-2584-1_6
502	Garcia-Orellana, J., Sanchez-Cabeza, J. A., Masqué, P., Àvila, A., Costa, E., Loÿe-
	Pilot, M. D., & Bruach-Menchén, J. M. (2006). Atmospheric fluxes of 210pb
504 505	to the western mediterranean sea and the saharan dust influence. Journal
	of Geophysical Research: Atmospheres, 111(D15). Retrieved from https://
506	agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/2005JD006660 doi:
507	https://doi.org/10.1029/2005JD006660
508	
509	
510	ficient and robust neural network training tool for machine learning in as-
511	tronomy. Monthly Notices of the Royal Astronomical Society, 441(2), 1741-
512	1759. Retrieved from https://doi.org/10.1093/mnras/stu642 doi:
513	10.1093/mnras/stu642
514	Herranz, M., Idoeta, R., & Legarda, F. (2008, 10). Evaluation of uncertainty and
515	detection limits in radioactivity measurements. Nuclear Instruments and Meth-
516	ods in Physics Research Section A: Accelerators, Spectrometers, Detectors and
517	Associated Equipment, 595, 526-534. doi: 10.1016/j.nima.2008.07.105
518	Heydorn, K. (2004, 10). Evaluation of the uncertainty of environmental measure-
519	ments of radioactivity. Journal of Radioanalytical and Nuclear Chemistry, 262,
520	249-253. doi: 10.1023/B:JRNC.0000040882.22365.a7
521	Karlsson, L., Hernandez, F., Rodriguez, S., Lopez-Perez, M., Hernandez-Armas, J.,
522	Alonso-Perez, S., & Cuevas, E. (2008). Using 137cs and 40k to identify natural
523	saharan dust contributions to pm10 concentrations and air quality impairment
524	in the canary islands. Atmospheric Environment, $42(30)$, 7034–7042.
525	Lal, D., Malhotra, P. K., & Peters, B. (1958, January). On the production of ra-

526	dioisotopes in the atmosphere by cosmic radiation and their application to
527	meteorology. Journal of Atmospheric and Terrestrial Physics, 12(4), 306-328.
528	doi: 10.1016/0021-9169(58)90062-X Lapedes, A., Barnes, C., Burks, C., Farber, R., & Sirotkin, K. (1988). Application
529	of neural networks and other machine learning algorithms to dna sequence
530 531	analysis (Tech. Rep.). Los Alamos National Lab., NM (USA).
532	Marsh, N., & Svensmark, H. (2000, 11). Cosmic rays, clouds, and climate. Space Sci-
532	ence Reviews, 94, 215-230. doi: 10.1023/A:1026723423896
534	Martell, E. (1970). Transport patterns and residence times for atmospheric trace
535	constituents vs. altitude. In (Vol. 93). ACS Publications. doi: 10.1021/ba-1970
536	-0093.ch009
537	Moore, H. E., Poet, S. E., & Martell, E. A. (1973). 222rn, 210pb, 210bi, and
538	210po profiles and aerosol residence times versus altitude. Journal of Geo-
539	physical Research (1896-1977), 78(30), 7065-7075. Retrieved from https://
540	agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/JC078i030p07065
541	doi: https://doi.org/10.1029/JC078i030p07065
542	Preiss, N., MéLièRes, MA., & Pourchet, M. (1996, December). A compilation
543	of data on lead 210 concentration in surface air and fluxes at the air-surface
544	and water-sediment interfaces. Journal of Geophysical Research, 101(D22),
545	28,847-28,862. doi: 10.1029/96JD01836
546	Qu, W., Zhao, J., Huang, F., & Deng, S. (2012, jun). CORRELATION BETWEEN
547	THE 22-YEAR SOLAR MAGNETIC CYCLE AND THE 22-YEAR QUA-
548	SICYCLE IN THE EARTH'S ATMOSPHERIC TEMPERATURE. The
549	Astronomical Journal, 144(1), 6. Retrieved from https://doi.org/10.1088/
550	0004-6256/144/1/6 doi: 10.1088/0004-6256/144/1/6
551	Raschka, S. (2018, 11). Model evaluation, model selection, and algorithm selection in
552	machine learning. ArXiv preprint https://arxiv.org/abs/1811.12808v3.
553	Retrieved from https://arxiv.org/abs/1811.12808v3
554	Sarkar, K., Ghalia, M. B., Wu, Z., & Bose, S. C. (2009). A neural network model
555	for the numerical prediction of the diameter of electro-spun polyethylene oxide
556	nanofibers. Journal of Materials Processing Technology, 209(7), 3156 - 3165. Retrieved from http://www.sciencedirect.com/science/article/pii/
557	S0924013608005827 doi: https://doi.org/10.1016/j.jmatprotec.2008.07.032
558	Schaefer, C., Geiger, M., Kuntzer, T., & Kneib, JP. (2018). Deep convolutional
559 560	neural networks as strong gravitational lens detectors. A&A, 611, A2. Re-
561	trieved from https://doi.org/10.1051/0004-6361/201731201 doi: 10.1051/
562	0004-6361/201731201
563	Schuler, C., Wieland, E., Santschi, P. H., Sturm, M., Lueck, A., Bollhalder, S.,
564	Wolfli, W. (1991). A multitracer study of radionuclides in lake zurich,
565	switzerland: 1. comparison of atmospheric and sedimentary fluxes of 7be, 10be,
566	210pb, 210po, and 137cs. Journal of Geophysical Research: Oceans, 96(C9),
567	17051-17065. Retrieved from https://agupubs.onlinelibrary.wiley.com/
568	doi/abs/10.1029/91JC01765 doi: https://doi.org/10.1029/91JC01765
569	Svensmark, H., Enghoff, M. B., & Pedersen, J. O. P. (2013). Response of cloud
570	condensation nuclei (¿50 nm) to changes in ion-nucleation. Physics Letters
571	A, 377(37), 2343 - 2347. Retrieved from http://www.sciencedirect.com/
572	science/article/pii/S0375960113006294 doi: https://doi.org/10.1016/
573	j.physleta.2013.07.004
574	Veretenenko, S., Ogurtsov, M., Lindholm, M., & Jalkanen, R. (2018). Galac-
575	tic cosmic rays and low clouds: Possible reasons for correlation rever-
576	sal. In Z. Szadkowski (Ed.), <i>Cosmic rays</i> (chap. 5). Rijeka: IntechOpen.
577	Retrieved from https://doi.org/10.5772/intechopen.75428 doi:
578	10.5772/intechopen.75428
579	vStencl, M., & Stastny, J. (2011, 04). Artificial neural networks numerical forecast- ing of aconomia time series. In (p. 15+). IntechOpen
580	ing of economic time series. In (p. $15+$). IntechOpen.

- Wilkening, M., Clements, W., & Stanley, D. (1975). Radon 222 flux measurements 581 in widely separated regions. Natural radiation environment II, 717-730. 582 Williams, N., Zander, S., & Armitage, G. (2006, October). A preliminary per-583 formance comparison of five machine learning algorithms for practical ip 584 traffic flow classification. SIGCOMM Comput. Commun. Rev., 36(5), 585 5 - 16.Retrieved from https://doi.org/10.1145/1163593.1163596 doi: 586 10.1145/1163593.1163596 587 Wogman, N. A., Thomas, C. W., Cooper, J. A., Engelmann, R. J., & Perkins, 588 R. W. (1968).Cosmic ray-produced radionuclides as tracers of atmo-589 spheric precipitation processes. Science, 159(3811), 189–192. Retrieved 590 from https://science.sciencemag.org/content/159/3811/189 doi: 591 10.1126/science.159.3811.189 592 Yoshimori, M., Hirayama, H., Mori, S., Sasaki, K., & Sakurai, H. (2003). Be-7 nuclei 593 produced by galactic cosmic rays and solar energetic particles in the earth's 594 atmosphere. Advances in Space Research, 32(12), 2691 - 2696. Retrieved from 595 http://www.sciencedirect.com/science/article/pii/S0273117703800856 596

597

doi: https://doi.org/10.1016/j.asr.2003.07.006

Appendix A Sketches of Neural Network and Random Forest structures

In this appendix, we show a sketch of the general structure of the Neural Network model employed and an example of a branch of a decision tree from the Random Forest algorithm investigated in this work.

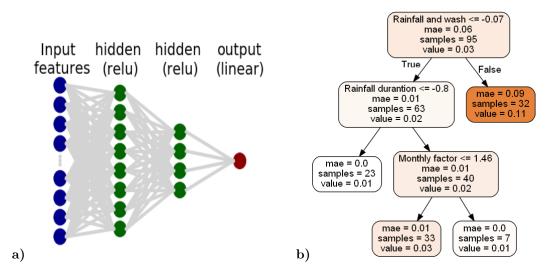


Figure A1. a): Sketch of the Neural Network model used, where there are two hidden layers that use the ReLU activation function and an output unit that linearly combines the nodes of the last hidden layer. b): Example of a decision tree used as part of a Random Forest model.

Appendix B Sketches of Neural Network and Random Forest structures

This appendix shows a comparison between the predictions from the reference model and the depositional flux measurements for one of these samples. It is crucial to have a reference model evaluated in the same way as for the ML algorithms studied in the paper, since this kind of evaluation is rather peculiar from ML algorithms. As we see, traditional models, based in linear regressions, are unable to reproduce the depositional fluxes behaviour, because of the complex relationships between variables.

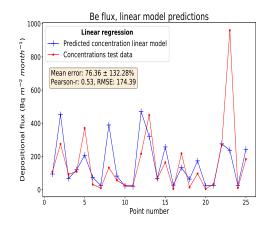


Figure B1. Predictions found from the reference linear model on one of the 25-length data samples, using the same evaluation as for the Random Forest and Neural Network algorithms studied in this work. Units of RMSE are of $Bq \ m^{-2} \ month^{-1}$.