

Grain size and transport biases in an Ediacaran detrital zircon record

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Abstract

Detrital zircon records of provenance are used to reconstruct paleogeography, sediment sources, and tectonic configuration. Recognition of the biases in detrital zircon records that result from hydraulic sorting of sediment and the initial characteristics of zircons in source regions (e.g., size and abundance) has added new complexity and caution in the interpretation of these records. In this study, we examine the role of transport process and sediment sorting in these records. We begin our analysis by investigating the influence of grain size and transport process in biasing detrital zircon provenance records in an idealized sedimentary system. Our modeling results show that settling and selective entrainment can leave distinct, process-dependent fingerprints in detrital zircon spectra if initial size variation between source zircon populations exists. We then consider a case study: a detrital zircon record from Ediacaran to Terreneuvian Death Valley. We focus on the Rainstorm Member, which is geochemically, mineralogically, and sedimentologically unusual. In addition to Earth's largest negative carbon isotope excursion (the Shuram excursion), the Rainstorm Member also contains anachronistic carbonate structures and a detrital mineral suite enriched in heavy minerals. We evaluate the detrital zircon provenance record of the Rainstorm Member, and find that, despite its unusual character, the provenance of the unit is similar to other units in the succession, with substantial input from Yavapai-Matzatzal provinces. Size and density measurements of heavy and light density components of the deposit suggest that its enriched heavy mineral suite is best explained through concentration by selective entrainment and winnowing. We find that our detrital zircon dataset is susceptible to hydrodynamic fractionation, so that grain size exerts influence on its provenance record, in particular for large Grenville-aged (1.0-1.2 Ga) grains.

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ABSTRACT

Detrital zircon records of provenance are used to reconstruct paleogeography, sediment sources, and tectonic configuration. Recognition of the biases in detrital zircon records that result from hydraulic sorting of sediment and the initial characteristics of zircons in source regions (e.g., size and abundance) has added new complexity and caution in the interpretation of these records. In this study, we examine the role of transport process and sediment sorting in these records. We begin our analysis by investigating the influence of grain size and transport process in biasing detrital zircon provenance records in an idealized sedimentary system. Our modeling results show that settling and selective entrainment can leave distinct, process-dependent fingerprints in detrital zircon spectra if initial size variation between source zircon populations exists. We then consider a case study: a detrital zircon record from Ediacaran to Terreneuvian Death Valley. We focus on the Rainstorm Member, which is geochemically, mineralogically, and sedimentologically unusual. In addition to Earth's largest negative carbon isotope excursion (the Shuram excursion), the Rainstorm Member also contains anachronistic carbonate structures and a detrital mineral suite enriched in heavy minerals. We evaluate the detrital zircon provenance record of the Rainstorm Member, and find that, despite its unusual character, the provenance of the unit is similar to other units in the succession, with substantial input from Yavapai-Matzatzal provinces. Size and density measurements of heavy and light density components of the deposit suggest that its enriched heavy mineral suite is best explained through concentration by selective entrainment and winnowing. We find that our detrital zircon dataset is susceptible to hydrodynamic fractionation, so that grain size exerts influence on its provenance record, in particular for large Grenville-aged (1.0-1.2 Ga) grains.

INTRODUCTION

Zircon's durability and utility as a geochronometer make it a valuable tool in provenance studies. After zircon grains are eroded, transported, and redeposited, an interpretation of their original source regions can be made by matching their measured U/Pb ages to source regions of known age. High-throughput dating by laser ablation has enabled a rich literature in which detrital zircon grains constrain paleogeography, sediment provenance, and tectonics (e.g., (Stevens et al., 2010; Cawood et al., 2012, 2013; Mackey et al., 2012; Gehrels, 2014; Blum et al., 2017). Fundamental to the use of detrital zircon as a provenance tracer is the assumption that, to some substantial degree, the detrital zircon grains analyzed within a sample are representative of the rest

32 of the bulk sediment. However, geological phenomena may lead to dissimilarities between the
33 detrital zircon population and bulk sediment of a sample.

34 Zircon is a robust mineral (Smithson, 1950; Carroll, 1953; Morton and Hallsworth, 2007).
35 This hardness allows zircon grains to survive not only transport over long distances, but also
36 multiple sedimentary cycles of erosion, transport, and deposition (Fedo et al., 2003; Hawkesworth
37 et al., 2009). This recalcitrance can lead to difficulty in interpreting the record of provenance
38 captured by detrital zircons: the measured ages of zircon grains may correspond either with igneous
39 or metamorphic terranes actively eroded at the time of deposition of the sediment (primary cycle
40 grains) or during previous sedimentary cycles (polycyclic grains). Careful analysis of grain cores
41 and rims and comparison with other minerals can reveal the polycyclic character of some zircons
42 (Hietpas et al., 2011a; Flowerdew et al., 2019; Moecher et al., 2019).

43 Other deviations can arise from source area variation in zircon fertility, which can span
44 several orders of magnitude even over regional spatial scales (Dickinson, 2008; Malusà et al.,
45 2016). Zircon-rich and zircon-poor terranes are not represented in the detrital zircon record in
46 proportion to the bulk sediment they contribute, but in proportion to their zircon fertility (Moecher
47 and Samson, 2006; Dickinson, 2008; Hietpas et al., 2011b; Malusà et al., 2016; Spencer et al.,
48 2018; Chew et al., 2020; Guo et al., 2020). In Laurentia, the Mesoproterozoic Grenville province
49 (1.0-1.2 Ga) is particularly zircon-rich (Moecher and Samson, 2006; Samson et al., 2018); these
50 and other igneous rocks associated with the assembly of Rodinia have high Zr concentrations (Liu
51 et al., 2017).

52 The co-variation of detrital zircon age, or source region, and grain size is the source of
53 deviation most relevant to the present study. Some evidence suggests that, in general, older zircons
54 are finer and less variable in size than younger grains (Lawrence et al., 2011; Yang et al., 2012),
55 although not all studies support this conclusion (Muhlbauer et al., 2017). Terrane-specific size
56 differences may exist, too. In Laurentia, Grenville sources may produce large (long axis on the
57 order of 1000 μm) zircons (Moecher and Samson, 2006; Samson et al., 2018). By comparison, the
58 nearby Neoproterozoic Carolina terrane produces much smaller (50 micron long axis) zircons.
59 These smaller zircons are, volumetrically, hundreds to thousands of times smaller than the *tips* of
60 large Grenville zircons (Moecher and Samson, 2006, see their Fig. 5).

61 If detrital zircons from a particular source region are characteristically large or small, then
62 sediment sorting will systematically emphasize or deemphasize the apparent contribution of that

63 source, independent of sediment provenance. Studies of some modern (Hietpas et al., 2011b;
64 Lawrence et al., 2011; Slama and Kosler, 2012; Ibañez-Mejia et al., 2018; Malkowski et al., 2019)
65 and ancient systems (Augustsson et al., 2017, 2019) show that grain size does exert control on
66 some records of provenance. Studies of modern sediments even capture statistically significant
67 variation in detrital zircon provenance over a single sand dune (Lawrence et al., 2011) or river
68 profile (Ibañez-Mejia et al., 2018) due to hydrodynamic sorting of sediment, despite these samples'
69 common transport history and provenance. In other successions, grain size does not appear to be a
70 major control on detrital zircon records of provenance (Muhlbauer et al., 2017; Leary et al., 2020).

71 Analytical procedures may enhance some size-dependent effects. Selection of large grains
72 ($>100\ \mu\text{m}$) for analysis biases selection for coarse-grained granitoid sources (Gehrels, 2000).
73 Mineral separation and hand-picking of grains may contribute additional size-dependent analytical
74 bias that minimizes the impact of smaller grains (Gehrels, 2000; Moecher and Samson, 2006;
75 Slama and Kosler, 2012). In studies of Laurentia, Grenville-aged sources are reported to have
76 particularly large zircon grains that may be systematically selected for analysis (Moecher and
77 Samson, 2006; Samson et al., 2018), contributing analytical bias to geological exaggeration of this
78 fertile source region in provenance records. Analysis of a consistent size class (63-125 μm) is
79 suggested as a strategy to mitigate the effects of hydrodynamic fractionation (Morton and
80 Hallsworth, 1994; Morton et al., 1996). However, this strategy enhances, rather than minimizes,
81 size-dependent effects (Garzanti et al., 2009) because this narrow size window will contain heavy
82 mineral grains for one bulk sediment size, but light mineral grains for a finer sediment. The current
83 best practice to mitigate these size-dependent biases is to measure both grain size and grain age
84 and determine what, if any, relationship exists between these variables (Malusà and Garzanti,
85 2019).

86 In this study, we explore how sedimentary size sorting, as well as size-density relationships
87 induced by transport process, may affect detrital zircon provenance records. First, we use size-
88 density relationships from both meta-analyses of settling experiments (Dietrich, 1982; Bagheri and
89 Bonadonna, 2016; Raffaele et al., 2020) and observations of natural sands (Hand, 1967;
90 Steidtmann and Haywood, 1982; Garzanti et al., 2008) to model the process-dependent
91 depositional size difference between the light mineral quartz (density or $\rho = 2.65\ \text{g/cm}^3$) and the
92 heavy mineral zircon, ($\rho = 4.65\ \text{g/cm}^3$) in an idealized sedimentary system. We apply these insights
93 to investigate the role of grain size in shaping a detrital zircon record from an Ediacaran to

94 Terreneuvian sedimentary succession from Death Valley, Nevada and California, USA. This
95 succession includes the Rainstorm Member of the Johnnie Formation, a unit that is
96 sedimentologically, mineralogically, and geochemically anomalous. We leverage insights from
97 our analysis of grain size and provenance to further constrain the depositional setting and transport
98 mechanisms related to Rainstorm Member's peculiar deposits.

99 *Size-Density Relationships in Clastic Deposits*

100 Grains with the same uniform settling velocity are hydraulically equivalent and will be
101 deposited together (Rubey, 1933). The settling velocity of a particle is a function of its size, shape,
102 roundness, and density, as well as of the properties of the fluid through which it settles. Grain size
103 and density are the most important variables in determining settling velocity for subspherical
104 grains (Dietrich, 1982), and increases in grain density are compensated by decreases in grain
105 diameter. For this reason, zircon ($\rho = 4.65 \text{ g/cm}^3$) grains are typically deposited with larger quartz
106 ($\rho = 2.65 \text{ g/cm}^3$) grains.

107 Empirical relationships derived from meta-analyses of settling experiments that span a
108 wide range of grain sizes, densities, shapes, and fluid properties (e.g., Clift and Gauvin, 1971;
109 Gibbs et al., 1971; Dietrich, 1982; Cheng, 1997; Ferguson and Church, 2004; Bagheri and
110 Bonadonna, 2016; Raffaele et al., 2020) capture the combined effects of turbulence and fluid
111 viscosity on settling velocity. **Figure 1** illustrates empirically derived relationships between grain
112 size and settling velocity for quartz and zircon in both water (using the relationship defined by
113 Dietrich, 1982; assuming a Corey Shape Factor of 0.7 and a Powers roundness value of 3.5, typical
114 values for natural sediments) and air (using the equation of Bagheri and Bonadonna, 2016; with
115 coefficients from Raffaele et al., 2020). These relationships predict that, relative to a constant
116 quartz grain size, zircon grains deposited in air will be larger than zircon grains deposited in water.
117 This difference is driven by the greater contrast between the submerged density of quartz and
118 zircon grains in water compared to air, and is observed in some natural sands (e.g., McIntyre, 1959;
119 Friedman, 1961; Steidtmann and Haywood, 1982).

120 However, many natural sediments exhibit size-density relationships inconsistent with these
121 predictions. Garzanti et al. (2008) find that, within aeolian and beach sands sourced from the
122 modern Po River delta, aeolian heavy mineral grains are smaller than beach heavy mineral grains
123 relative to the quartz sediment in each. In beach and aeolian sands from New Jersey, Hand (1967)
124 observes heavy minerals that are smaller than predicted by settling relationships. Steidtmann and

125 Haywood (1982) compare tourmaline and quartz grains within the aeolian Casper Formation, and
126 find that although grainfall laminae have size-density relationships consistent with settling through
127 air, grains from subcritical climbing ripples have size-density relationships in which tourmaline
128 grains are much smaller than predicted. Size-density relationships that deviate from equal settling
129 velocity, in which heavy minerals are smaller than predicted, are also found in other natural
130 sediments (Rittenhouse, 1943; McIntyre, 1959; Briggs, 1965; Grigg and Rathbun, 1965; Lowright
131 et al., 1972; Slingerland, 1977; Komar and Wang, 1984).

132 Various processes operating during sediment transport can lead to size-density
133 relationships other than that predicted by settling velocity (Lowright et al., 1972; Slingerland,
134 1977, 1984; Komar and Wang, 1984; Slingerland and Smith, 1986; Komar et al., 1989; Komar,
135 2007). Larger grains resting on a bed project further into the profile of the flow above them and
136 have a smaller pivot angle than smaller grains. This encourages selective entrainment of larger,
137 lighter grains rather than smaller, denser grains (McIntyre, 1959; Lowright et al., 1972;
138 Slingerland, 1977, 1984; Komar and Li, 1988; Komar et al., 1989). Shear sorting during transport
139 leads density-stratified deposits in which dispersive equivalence, rather than entrainment or
140 settling velocity, controls the size-density distribution of deposits (Emery and Stevenson, 1950;
141 Clifton, 1969; Sallenger, 1979). Variations in flow over bedforms like ripples and dunes can lead
142 to preferential deposition of heavy minerals near the crest of bedforms (McQuivey and Keefer,
143 1969; Brady et al., 1973). Localized, preferential deposition of heavy minerals can also occur in
144 response to flow changes around larger geomorphic features like bars or river confluences (Smith
145 and Beukes, 1983; Day and Fletcher, 1989).

146 We consider two end-member possibilities related to selective entrainment in our study of
147 detrital zircon provenance records. At one extreme, detrital zircon grains within a deposit are as
148 large as quartz grains. In this case, size and pivot angle control entrainment more strongly than
149 grain density; coarse light grains are selectively entrained; and the deposit becomes heavy mineral-
150 enriched and well-sorted, as observed in some placer deposits (Slingerland, 1977). At the other
151 extreme, detrital zircon grains are much smaller than predicted by settling, a relationship observed
152 in many natural sands (Rittenhouse, 1943; McIntyre, 1959; Briggs, 1965; Grigg and Rathbun,
153 1965; Hand, 1967; Lowright et al., 1972; Slingerland, 1977; Steidtmann and Haywood, 1982;
154 Komar and Wang, 1984; Komar, 2007; Garzanti et al., 2008). This scenario is thought to arise due
155 to the shielding of small heavy minerals by larger, lighter grains, leading to a difference in settling

156 velocity but equivalence in entrainment potential between larger light grains and smaller dense
157 grains (Hand, 1967; Steidtmann and Haywood, 1982; Komar and Wang, 1984; Komar and Li,
158 1988; Komar, 2007). Both extremes are the result of complex interactions between grain size and
159 shape, bed roughness, and fluid flow during transport near the bed. Size-density relationships for
160 sediments deposited by settling are intermediate between these two extremes. We hypothesize that,
161 for sediments in which detrital zircon populations exhibit significant size differences between
162 source areas, transport processes may impart different, size-dependent biases in their provenance
163 record even if bulk sediment size is constant.

164 GEOLOGICAL CONTEXT

165 We now consider the role of grain size in shaping a specific detrital zircon record. The
166 Johnnie Formation, Stirling Quartzite, and Wood Canyon Formation are Ediacaran to
167 Terreneuvian units within the mixed siliciclastic-carbonate, rift-to-passive margin succession of
168 Death Valley, USA, which thickens westward (**Figure 2**) (Stewart, 1970; Summa, 1993; Fedo and
169 Cooper, 2001; Corsetti and Kaufman, 2003; Clapham and Corsetti, 2005; Verdel et al., 2011;
170 Schoenborn et al., 2012). The underlying Noonday Dolomite is correlated with post-Marinoan cap
171 carbonate units (Prave, 1999; Halverson et al., 2005; Petterson et al., 2011), especially those also
172 containing tubestone (Cloud et al., 1969; Hegenberger, 1987; Hoffman and Schrag, 2002; Hoffman
173 et al., 2002; Rodrigues Nogueira et al., 2003; Corsetti and Grotzinger, 2005; Bold et al., 2016), so
174 is interpreted as earliest Ediacaran in age at its base (Prave, 1999; Petterson et al., 2011; Creveling
175 et al., 2016). The Wood Canyon Formation spans the Ediacaran-Cambrian boundary and contains
176 Ediacaran fossils in its lower member (Diehl, 1979; Corsetti and Hagadorn, 2000; Hagadorn and
177 Waggoner, 2000; Smith et al., 2017). Direct radiometric age constraints within the succession are
178 limited; a detrital zircon dated to 640.33 ± 0.09 from the Johnnie Formation is consistent with an
179 Ediacaran age for the unit (Verdel et al., 2011). Correlation of a carbon isotope excursion in the
180 Rainstorm Member, the uppermost unit of the Johnnie Formation, with dated intervals in northwest
181 Canada and Oman suggest that deposition of this unit took place after 574 ± 4.7 Ma (Rooney et
182 al., 2020).

183 These units record the development of the Laurentian passive margin including some
184 evidence for extensional tectonics (Stewart and Poole, 1974; Summa, 1993; Fedo and Cooper,
185 2001; Clapham and Corsetti, 2005). Braided channel facies and marine reworking of fluvial sand
186 are common in the Johnnie Formation, along with intertidal to shallow subtidal carbonates

187 containing stromatolites and ooids (Summa, 1993). Desiccation cracks, tepee structures, caliche,
188 and karst surfaces throughout the shallow environments of the Johnnie Formation provide
189 evidence for frequent, and sometimes long-lasting, exposure (Summa, 1993). Basinal
190 environments in the lower Johnnie Formation record a submarine channel complex with no
191 evidence for exposure (Williams et al., 1974; Troxel et al., 1982; Wright et al., 1984; Prave, 1999;
192 Petterson et al., 2011; Macdonald et al., 2013a; Creveling et al., 2016) The Stirling Quartzite
193 abruptly overlies the Johnnie Formation and is dominated by braided fluvial quartzite, though its
194 middle member includes interbedded intertidal shale, dolomite, and siltstone (Stewart, 1970; Fedo
195 and Cooper, 2001). Siltstone, locally arkosic quartzite, and minor carbonate make up the lower
196 and upper members of the Wood Canyon Formation (Stewart, 1970). These units record deposition
197 in shallow marine settings; its middle member, a locally conglomeratic, feldspathic quartzite, was
198 deposited in a braided fluvial system (Stewart, 1970; Fedo and Cooper, 1990, 2001; Corsetti and
199 Hagadorn, 2000; Hogan et al., 2011; Muhlbauer et al., 2017, 2020). Detrital zircon provenance
200 records through this succession show varying contributions from the Grenville, Yavapai-
201 Matzatzal, and mid-continent regions, and a smaller contribution from Archean sources, with the
202 vast majority of zircons older than 1 Ga (Stewart et al., 2001; Verdel et al., 2011; Schoenborn et
203 al., 2012; Muhlbauer et al., 2017). The Wood Canyon Formation features a sharp peak associated
204 with Grenville-aged sources (Stewart et al., 2001; Muhlbauer et al., 2017).

205 *The Rainstorm Member*

206 The Johnnie Formation is subdivided into six members recognized by Stewart (1970). The
207 uppermost of these, the Rainstorm Member, is a shallow marine, storm-dominated, mixed
208 siliciclastic-carbonate unit with unusual sedimentological, mineralogical, and geochemical
209 features (Stewart, 1970; Summa, 1993; Corsetti and Kaufman, 2003; Corsetti et al., 2004;
210 Kaufman et al., 2007; Pruss et al., 2008; Bergmann et al., 2011, 2013). The detrital zircon and
211 heavy mineral record of the Rainstorm Member, and what it reveals about provenance and process
212 during deposition of this peculiar unit, is the focus of this study. The provenance of the Rainstorm
213 Member is also investigated by Schoenborn et al. (2012). That study uses Nd isotopes in carbonate
214 to constrain the provenance of the Rainstorm Member and suggests that the source of sediment is
215 primarily the Yavapai-Matzatzal provinces. However, Schoenborn et al. (2012) note these data do
216 not definitively distinguish between this scenario and one in which the sources include the Mojave
217 and/or Yavapai-Matzatzal provinces and the Grenville province.

218 The regionally extensive Johnnie Oolite is a 2 m-thick, cross-stratified oolite bed near the base
219 of the Rainstorm Member, and is bounded above and below by siltstone and shale (Stewart, 1970;
220 Summa, 1993). Summa (1993) suggests a cryptic unconformity beneath the Johnnie Oolite. Above
221 the Johnnie Oolite, siltstone and sandstone are interbedded with pink and gray limestone
222 containing ooid grainstone, intraclast and edgewise conglomerate, and, locally, crystal fans
223 (Stewart, 1970; Summa, 1993; Corsetti et al., 2004; Pruss et al., 2008; Bergmann et al., 2013). The
224 siltstone and carbonate strata of the Rainstorm Member have been interpreted as a deposit
225 associated with highstand and maximum flooding; initial marine transgression is recorded by the
226 time-transgressive Johnnie Oolite (Summa, 1993; Bergmann et al., 2011). Hummocky cross-
227 stratification, amalgamated scours, and planar lamination are abundant throughout the Rainstorm
228 Member (Summa, 1993). Above these units, an incision surface cuts into the Rainstorm Member,
229 in some locations eroding it completely; these incised valleys are filled by thick conglomerate and
230 may represent either subaerial or submarine erosion (Stewart, 1970; Summa, 1993; Corsetti and
231 Kaufman, 2003; Clapham and Corsetti, 2005; Trower and Grotzinger, 2010; Verdel et al., 2011).
232 This incision surface is interpreted either as tectonic (Summa, 1993; Clapham and Corsetti, 2005)
233 or glacioeustatic (Christie-Blick and Levy, 1989; Abolins et al., 2000; Witkosky and Wernicke,
234 2018) in origin.

235 Pink-gray Rainstorm carbonate units are the only limestone strata in the Johnnie
236 Formation; other carbonate is dolostone (Summa, 1993). In the Southern Nopah range, Rainstorm
237 limestone beds also contain carbonate crystal fans neomorphosed to calcite from primary aragonite
238 (Summa, 1993; Corsetti et al., 2004; Pruss et al., 2008; Bergmann et al., 2013) (**Figure 3**). These
239 delicate crystal fans grew from the seafloor, nucleating on horizons of detrital, iron-rich material
240 (Summa, 1993; Pruss et al., 2008; Bergmann et al., 2013). These nucleation surfaces contain a
241 striking diversity of minerals, including calcite, dolomite, quartz, hematite, rutile, anatase,
242 ilmenite, muscovite, biotite, apatite, barite, monazite, and zircon (Bergmann et al., 2013). Crystal
243 fans are about a centimeter in height and individual blades have a horizontal hexagonal cross
244 section with a maximum diameter of 230 μm (Pruss et al., 2008). Sediment between crystal blades
245 includes ooids, detrital mineral grains, and intraclasts (Pruss et al., 2008; Bergmann et al., 2013)
246 (**Figure 3**).

247 Crystal fans like these are more typically associated with Archean or Paleoproterozoic
248 successions and Mesoproterozoic peritidal environments (Grotzinger and Read, 1983; Hofmann

249 and Jackson, 1987; Kah and Grotzinger, 1992; Kah and Knoll, 1996; Sumner and Grotzinger,
250 1996; Grotzinger and James, 2000; Allwood et al., 2009; Higgins et al., 2009; Bergmann et al.,
251 2013; Cantine et al., 2020). Where crystal fans occur in Neoproterozoic sediments, they are
252 typically associated with cap carbonate units (Grotzinger and James, 2000; Sumner and
253 Grotzinger, 2000; James et al., 2001; Hoffman and Schrag, 2002; Babinski et al., 2007;
254 MacDonald et al., 2009) and not with middle Ediacaran strata. Rainstorm crystal fans are thus
255 anachronistic.

256 Yet like other examples of crystal fans through Earth history (Bergmann et al., 2013),
257 Rainstorm crystal fans are interpreted to have developed during a period of maximum flooding
258 and low background sedimentation rate (Summa, 1993; Corsetti et al., 2004; Pruss et al., 2008;
259 Bergmann et al., 2013) The formation of Rainstorm crystal fans may be linked to minimal organic
260 delivery and limited aerobic respiration maintaining favorable conditions for carbonate growth at
261 the sediment-water interface (Bergmann et al., 2013). Most Rainstorm crystal fan layers remain *in*
262 *situ*, and the small size and delicate shape of the crystal fans suggests that they formed in relatively
263 quiet conditions (Summa, 1993; Pruss et al., 2008). Crystal fan intraclasts (Bergmann et al., 2013,
264 their figure 5) show that some crystal fan-bearing units were disrupted and reworked during high-
265 energy events.

266 In addition to crystal fans, the Rainstorm Member's pink and gray limestone hosts the nadir
267 of an extremely negative carbon isotope excursion, which contains values as low as $-12\text{‰ } \delta^{13}\text{C}$;
268 the onset of the excursion is within the Johnnie Oolite (Corsetti and Kaufman, 2003; Kaufman et
269 al., 2007; Bergmann et al., 2011; Verdel et al., 2011). The asymmetric excursion is characterized
270 by a rapid onset and gradual recovery (Corsetti and Kaufman, 2003; Kaufman et al., 2007;
271 Bergmann et al., 2011; Verdel et al., 2011). The excursion is synchronous across the Death Valley
272 area (Bergmann et al., 2011; Minguéz et al., 2015). An estimate for the duration of the Rainstorm
273 excursion, derived from magnetostratigraphic constraints and extrapolation of sedimentation rate,
274 yields 8.2 ± 1.2 million years (Minguéz et al., 2015). Another independent estimate derived from
275 subsidence modeling yields a duration of c. 6 million years (Witkosky and Wernicke, 2018).

276 The Rainstorm excursion is putatively correlated with Earth's most negative carbon isotope
277 excursion, the Shuram excursion (Corsetti and Kaufman, 2003; Kaufman et al., 2007; Bergmann
278 et al., 2011; Grotzinger et al., 2011; Verdel et al., 2011; Minguéz et al., 2015; Witkosky and
279 Wernicke, 2018). The Shuram excursion is observed on multiple continents in middle Ediacaran

280 strata. It is consistently characterized by a rapid onset to strongly negative $\delta^{13}\text{C}$ values followed
281 by gradual recovery (e.g., Burns and Matter, 1993; Fike et al., 2006; Kaufman et al., 2006;
282 Le Guerroué et al., 2006c; Bowring et al., 2007; Zhu et al., 2007; Jiang et al., 2007; Melezhik et
283 al., 2008; Le Guerroué, 2010; Macdonald et al., 2013b; Husson et al., 2015; Canfield et al., 2020).
284 Recent geochronological work indicates the Shuram excursion was a global and synchronous event
285 with a duration of less than 6.7 ± 5.6 million years (Canfield et al., 2020; Matthews et al., 2020;
286 Rooney et al., 2020). Because the Rainstorm excursion shares its magnitude; asymmetric onset
287 and recovery; position within middle Ediacaran strata; and a similar duration with the Shuram
288 excursion, like others (Corsetti and Kaufman, 2003; Kaufman et al., 2007; Bergmann et al., 2011;
289 Grotzinger et al., 2011; Verdel et al., 2011; Minguez et al., 2015; Witkosky and Wernicke, 2018),
290 we consider the Rainstorm excursion a regional expression of the global Shuram excursion and
291 refer to the negative carbon isotope excursion within the Rainstorm Member as the Shuram
292 excursion.

293 The magnitude and global extent of the Shuram excursion have made it a particular puzzle
294 for geochemists and geobiologists. Geochemical questions revolve around the primary (Rothman
295 et al., 2003; Fike et al., 2006; Bristow and Kennedy, 2008; McFadden et al., 2008; Ader et al.,
296 2009; Bjerrum and Canfield, 2011; Johnston et al., 2012; Lee et al., 2015; Miyazaki et al., 2018;
297 Shields et al., 2019) or post-depositional (Knauth and Kennedy, 2009; Derry, 2010; Schrag et al.,
298 2013; Oehlert and Swart, 2014) processes capable of driving such a large and apparently
299 synchronous global excursion globally. Determining the impact, if any, of the Shuram excursion
300 on the diversification, extinction, and habitats of early metazoans is also a critical geobiological
301 question (Macdonald et al., 2013b; Wood et al., 2015; Cui et al., 2017; Darroch et al., 2018;
302 Muscente et al., 2018; Zhang et al., 2019; Li et al., 2020; Rooney et al., 2020). The relationship
303 between the regional Gaskiers glaciation c. 580 Ma (Pu et al., 2016) and the Shuram excursion has
304 also provoked interest (Xiao et al., 2016), although recent age constraints (Canfield et al., 2020;
305 Matthews et al., 2020; Rooney et al., 2020) and the recognition of a possible Shuram excursion
306 equivalent in clastic-hosted carbonates in Newfoundland (Canfield et al., 2020) suggest that the
307 two events are temporally distinct.

308 Our understanding of the Shuram excursion globally relies on a sound sedimentological
309 understanding of the sediments that record it locally. The Rainstorm Member contains
310 anachronistic crystal fans and unusual detrital mineral assemblages, suggesting it records unusual

311 environmental conditions. It also contains Earth’s largest negative carbon isotope excursion,
312 suggests a tantalizing interplay between the unit’s unusual sedimentology and the processes
313 driving the excursion.

314 MATERIALS AND METHODS

315 *Grain mixing model*

316 To explore the influence of process-based size-density relationship on detrital zircon
317 records, we built a numerical model. The model generates synthetic quartz and zircon grains for
318 two source regions with defined age and size characteristics (**Fig. 4**). Quartz grains of all sizes are
319 generated; by contrast, zircon grains have a characteristic size distribution. No initial size-density
320 relationship between quartz and zircon grain sizes exists in the source regions.

321 Each individual grain is randomly assigned an age and grain size based on defined
322 parameters for its source region. Each grain is also assigned a 2σ age uncertainty of 5% of the
323 grain’s age, mirroring the uncertainty associated with laser ablation $^{206}\text{Pb}/^{207}\text{Pb}$ age measurements.
324 The source regions are mixed in a defined ratio. From the combined pool, 100 grains of quartz and
325 zircon are then randomly selected, representing the deposition of grains. For quartz grains, the
326 probability of being deposited are determined according to a defined mean grain size and standard
327 deviation representing the characteristics of the sedimentary deposit. The deposited zircon grains
328 are determined by a size-density relationship between the defined grain size of quartz to deposit
329 and the corresponding zircon size. The age spectrum of the resulting deposited zircon “sample” is
330 determined. Synthetic spectra are generated this way 10,000 times.

331 The model considers four possible different size-density relationships between quartz and
332 zircon. In one end-member relationship, the “equal sizes” scenario, the zircons deposited have the
333 same size as the quartz grains deposited—the greater density of the zircon grains has no effect on
334 the size of zircon grains deposited. This scenario is potentially applicable to well-sorted, heavy
335 mineral-rich deposits like placer sands that form under specific flow and bed roughness conditions
336 (Slingerland, 1977). The other end-member relationship, the “grain shielding” scenario, uses the
337 size-density relationship reported by Garzanti et al. (2008) from a modern aeolian sand, which we
338 interpret as the result of selective entrainment and grain shielding. Similar deposits, in which heavy
339 mineral grains have lower settling velocities than the quartz with which they are deposited, are
340 reported elsewhere (Hand, 1967; Steidtmann and Haywood, 1982). Intermediate to these end-
341 members are the size-density relationships empirically determined for quartz and zircon settling

342 in water (Dietrich, 1982) and air (Bagheri and Bonadonna, 2016; Raffaele et al., 2020). For each
343 case, these size-density relationships are shown in **Fig. 4, column II**.

344 In all cases described here, the two source regions are mixed in a 1:1 ratio, and the zircon
345 fertility, or ratio of quartz grains to zircon grains in each region, is held constant at 10:1, an
346 illustrative value. Geological measurements of zircon fertility are smaller (Malusà et al., 2016).
347 The parameters for all runs are shown in **Table 1**. See the Data Availability section for information
348 on the code used for this analysis.

349 *Base case*

350 This case mixes two source regions that are identical in every regard, including the grain
351 sizes of zircons, except for the ages associated with the source regions. Grains from Source 1 have
352 a population age of 1000 Myr with a standard deviation of 50 Myr; Source 2 grains have a
353 population age of 500 Myr with a standard deviation of 50 Myr. The sizes of quartz grains from
354 both source regions are uniformly distributed between 0 and 300 μm . Zircon from both regions
355 has a mean grain size of 150 μm with a standard deviation of 15 μm . The deposited quartz sediment
356 has a mean grain size of 200 μm with a standard deviation of 15 μm .

357 *Case 1*

358 This case is identical to the base case, with the single change that zircons from Source 2
359 are smaller than zircons from Source 1. Source 2 zircons have a mean grain size of 100 μm with a
360 standard deviation of 15 μm .

361 *Case 2*

362 This case explores behavior in finer-grained sediments than previous cases. In this case,
363 both Source 1 and 2 quartz grains have a mean grain size of 80 μm with a standard deviation of 15
364 μm . Source 1 zircons have a mean grain size of 40 μm with a standard deviation of 10 μm ; Source
365 2 zircons have a mean grain size of 30 μm with a standard deviation of 10 μm . Deposited quartz
366 has a mean grain size of 55 μm with a standard deviation of 5 μm . All other parameters are
367 unchanged from the base case.

368 *Rainstorm sample characterization*

369 Samples were collected in the southern Nopah Range, including a very fine-grained quartz
370 sandstone below the Johnnie Oolite and crystal fan-bearing strata within the Rainstorm Member.
371 Sample locations are available in the Supplement. Standard zircon separation procedures,
372 including magnetic and heavy liquid separation, were used for the sandstone sample. The crystal

373 fan-bearing samples were dissolved in 20% hydrochloric acid to remove carbonate. After
374 dissolution, the insoluble residue followed standard zircon separation procedures. Prior to analysis,
375 zircons were mounted, polished, and imaged by either cathodoluminescence or light microscope.

376 *Geochronology*

377 U-Pb ages of zircons were obtained via LA-ICP-MS using a Photon Machines 193 nm laser
378 ablation system coupled to a Thermo iCapQc™ mass spectrometer at Rutgers University. Laser
379 power was set at 50% power and a 10 Hz rep rate, resulting in a fluence of 4.25 J/cm² at the
380 surface of the zircon. The laser spot size was 20 μm. This small spot size was used to minimize
381 sample destruction during analysis, so that any potential young, concordant grains could be dated
382 with high-resolution U/Pb thermal ionization mass spectrometry (TIMS). Ablated material was
383 carried to the plasma via Helium gas at a total flow rate of 0.8 LPM. For each analysis, the
384 iCapQc™ was set up to acquire data for ~220 sweeps with a 10 millisecond per isotope dwell time.
385 Laser firing and ablation lasted for 30 seconds during each analysis while gas blanks were
386 measured for approximately 20 seconds immediately prior to each zircon analysis using the same
387 instrumental conditions (without firing the laser). Washout time between each analysis and the
388 start of the next background was approximately 1 minute. Data processing was performed
389 primarily using the Iolite software package with ²⁰⁶Pb/²⁰⁷Pb ages calculated using ISOPLOT
390 (Ludwig, 2003). ²⁰⁶Pb/²⁰⁷Pb ages are used in order to make direct comparison to the results of
391 Schoenborn et al. (2012). The use of ²⁰⁶Pb/²³⁸U ages rather than ²⁰⁶Pb/²³⁸U ages would yield the
392 same conclusions.

393 The primary standard used was the zircon standard 91500, with an age of 1065 Ma (Fisher et
394 al., 2014). 91500 was run 190 times, yielding a weighted mean ²⁰⁶Pb/²³⁸U age of 1061.7 ± 1.8 Ma
395 (2σ; MSWD = 2.1) and a ²⁰⁷Pb/²⁰⁶Pb age of 1085.7 ± 8.6 Ma (2σ; MSWD = 1.1). 91500 was run
396 after every 10-15 samples. The results from the 91500 standard were used to apply a final
397 fractionation correction to the unknowns via a bracketing approach. The zircon standards Plesovice
398 (published age: 337 Ma; Slama et al., 2008), R33 (419 Ma), and Mud Tank (732 Ma) were also
399 run during the session and processed the same way as the unknowns (Fisher et al., 2014). The
400 Plesovice standard was run a total of 161 times and yielded a weighted mean ²⁰⁶Pb/²³⁸U age of
401 337.80 ± 0.71 Ma (2σ; MSWD = 3.5) and a ²⁰⁷Pb/²⁰⁶Pb age of 341.9 ± 7.5 Ma (2σ; MSWD = 0.9).
402 The R33 standard was run a total of 84 times and yielded a weighted mean ²⁰⁶Pb/²³⁸U age of 419.6
403 ± 1.8 Ma (2σ; MSWD = 7.4) and a ²⁰⁷Pb/²⁰⁶Pb age of 468 ± 18 Ma (2σ; MSWD = 1.2). The Mud

404 Tank standard was run a total of 55 times and yielded a weighted mean $^{206}\text{Pb}/^{238}\text{U}$ age of $724.1 \pm$
405 3.4 Ma (2σ ; MSWD = 3.0) and a $^{207}\text{Pb}/^{206}\text{Pb}$ age of $737 \pm 33 \text{ Ma}$ (2σ ; MSWD = 0.9).

406 Due to the small spot size used and low Pb signal, the uncertainty of individual $^{207}\text{Pb}/^{206}\text{Pb}$
407 ages is substantial, especially in the Phanerozoic secondary standards. However, our samples do
408 not include Phanerozoic zircon grains and typically have significantly lower individual $^{207}\text{Pb}/^{206}\text{Pb}$
409 age uncertainties than those of the standards. The concordance of the Mesoproterozoic 91500
410 standard supports a good degree of confidence in the analyses; and our analysis focuses on distinct
411 Proterozoic detrital zircon age subpopulations that are resolvable within the error achieved. Grains
412 were screened for concordance between their $^{206}\text{Pb}/^{238}\text{U}$ and $^{207}\text{Pb}/^{206}\text{Pb}$ ages, and only those with
413 $\pm 10\%$ concordance were included in the analysis. Ages presented are $^{207}\text{Pb}/^{206}\text{Pb}$ ages.

414 Age data from the lower and middle Johnnie Formation and the Stirling Quartzite were
415 previously published in Schoenborn et al. (2012). Age data from the upper Stirling Quartzite and
416 Wood Canyon Formation were published in Muhlbauer et al. (2017).

417 **Zircon size determination**

418 The long and short axis of each grain from the Rainstorm Member was measured, either by
419 cathodoluminescence (CL) imaging or light microscope. Cathodoluminescence images of grains
420 from the lower and middle Johnnie Formation and lower and upper Stirling Quartzite (Schoenborn
421 et al., 2012) are available in the appendices of (Schoenborn, 2010). Grain long axes for upper
422 Stirling Quartzite and Wood Canyon Formation samples were previously published in Muhlbauer
423 et al. (2017). Only concordant grains for which a grain size could be determined are included in
424 this study.

425 Previous detrital zircon grain size studies have used either the equivalent spherical diameter
426 (ESD) of zircon grains (e.g., Lawrence et al., 2011) or the long axis of grains (e.g., Muhlbauer et
427 al., 2017). The ESD better accounts for the hydrodynamic characteristics of a zircon grain than a
428 measurement of its long axis alone (Garzanti et al., 2008; Lawrence et al., 2011), however, data
429 easily available for all grains was limited to measurements of long axes. This study uses the long
430 axis of detrital zircons to quantify their grain size, though the size trends described in this study
431 are consistent using ESD within the subpopulation of samples for which ESD can also be
432 calculated.

433 **Crystal fan size-density determination**

434 A sample of crystal fan-bearing carbonate from the Rainstorm Member was dissolved as
435 described above in 20% hydrochloric acid. The residue was rinsed with deionized water and dried.
436 This sample did not undergo magnetic separation and was separated into a heavy and light fraction
437 using the heavy liquid methylene iodide ($\rho = 3.32 \text{ g/cm}^3$). The majority of the detritus, both by
438 volume and by mass, separated into the heavy fraction (**Table 2**). The density of both fractions
439 was determined through volume displacement with a 10 mL volumetric flask, sensitive balance,
440 and deionized water. The grain sizes of the heavy and light fractions were determined using a laser
441 particle size analyzer at the University of Oklahoma.

442 *Modeling size-provenance relationships in the Ediacaran-Terreneuvian Death Valley succession*

443 To explore the signature of grain size in data from this succession, we pooled all measured
444 age and grain size data for zircon grains from these units. We defined 4 size distributions, each
445 corresponding to the size distributions of zircon grains from a particular unit: $69 \pm 16 \text{ }\mu\text{m}$ (1 sd;
446 Rainstorm crystal fans); $89 \pm 21 \text{ }\mu\text{m}$ (1 sd; Rainstorm sandstone); $124 \pm 32 \text{ }\mu\text{m}$ (1 sd; Lower
447 Johnnie Formation) and $154 \pm 50 \text{ }\mu\text{m}$ (1 sd; Wood Canyon Formation). Then, these size
448 distributions were used to randomly select 60 grains from the pool of all zircons, a process repeated
449 10,000 times. The resulting age spectra from each subset of 60 grains was determined. See the
450 Data Availability section for information on the code used for this analysis.

451 **RESULTS**

452 *Size mixing model*

453 Results from our exploration of size-density relationships within detrital zircon illustrate
454 that process-driven size-density relationships together with systematic initial size variation can
455 result in different measured spectra (**Fig. 4**). For each case, **Figure 4** shows the age and size
456 distribution of quartz and zircon grains from two source regions (column I). The two sources are
457 combined, and then filtered according to a size-density relationship corresponding to a set of
458 depositional conditions (column II). 100 zircon grains and 100 quartz grains are selected from each
459 pool according to the probability of deposition determined by the depositional filter. This selection
460 is repeated 10,000 times, and the median and 95% confidence interval of all results are determined
461 (column III). Across all runs, the proportion of quartz and zircon grains deposited that come from
462 each source region are shown in pie charts within column III.

463 **Base Case**

464 Results from the base case provide a point of comparison for other cases because all
465 parameters, except for the age of detrital zircon grains, are consistent between the two sources.
466 The measured age spectra for all size-density relationships accurately record the 1:1 sediment input
467 for both quartz and zircon grains with no effect from transport process.

468 **Case 1**

469 Case 1 is identical to the base case except that Source 2 zircons are smaller than Source 1
470 grains. The resulting “deposits” show a range of spectra with shifting Source 1 and Source 2 inputs
471 based on grain size. The “grain shielding” deposit shows an average of 85% Source 2 input,
472 whereas the “equal sizes” case shows nearly 100% Source 1 input, with <0.1% Source 2 input.
473 Deposits formed from the settling of grains through air and water both show a majority of Source
474 1 (larger) grains, but with differences in the proportion of the sources. The case shown for settling
475 through air has 6% Source 2 grains on average and the case shown for settling through water has
476 35% Source 2 grains on average. Across all transport processes, quartz grains in the deposit are
477 split evenly between the source regions.

478 **Case 2**

479 Case 2 explores process-dependent grain size shifts at smaller quartz and zircon grain sizes
480 than Case 1. Deposits show a range of spectra with variable contributions from Source 1 and
481 Source 2. At one extreme, the “grain shielding” case contains a slight majority of Source 2 grains.
482 The other three transport cases show the opposite trend, with more Source 1 grains than Source 2
483 grains. Both the air and water settling cases contain a majority of Source 1 grains, but fewer Source
484 1 grains than the “equal size” case, which contains more than 80% Source 1 grains.

485 *Rainstorm sample characterization*

486 **Geochronology**

487 Age spectra from combined new and previously published (Schoenborn et al., 2012;
488 Muhlbauer et al., 2017) zircon ages, as well as grain size measurements, are shown in **figures 5**
489 **and 6**. Four primary subpopulations (~1.0-1.2 Ga; ~1.3-1.5 Ga; ~1.6-1.9 Ga; and >2.0 Ga) are
490 present. These subpopulations are interpreted to have their origin in the Grenville orogen (~1.0-
491 1.2 Ga); the midcontinent (~1.3-1.5 Ga); the Yavapai-Mazatzal or Central Plains regions (~1.6-
492 1.9); and older cratonic sources (>2.0 Ga) (Dickinson and Gehrels, 2009). The relative contribution
493 of Grenville-aged material waxes and wanes through the section. This peak is absent or minimal
494 in the Lower Johnnie, Rainstorm crystal fans, and upper Stirling Quartzite. It is particularly

495 pronounced in the Wood Canyon Formation, as others note (Stewart et al., 2001; Muhlbauer et al.,
496 2017).

497 *Zircon size characterization*

498 We document that Grenville-aged grains are, on average, larger than their Paleoproterozoic
499 counterparts (**Fig. 6**). Zircon grains from the sub-Rainstorm sandstone and Rainstorm carbonate
500 are distinctly finer than other samples. To assess if the detrital zircon population is susceptible to
501 size-age fractionation by hydraulic effects, we split all measured detrital zircons into four size
502 groups based on the length of their major axis (silt, very fine sand, fine sand, and medium sand,
503 (Malusà et al., 2013; Malusà and Garzanti, 2019). The empirical cumulative distribution functions
504 of these four size groups are shown in **Figure 7**. We assessed the null hypothesis that these size
505 groups were drawn from the same age distribution using the Kolmogorov-Smirnov method (**Table**
506 **4**). Our data do not reject the null hypothesis for two pairs of grain size groups: first, silt and very
507 fine sand, and second, fine sand and medium sand. Our data support the alternative hypothesis that
508 both the silt and very fine sand groups have a different age distribution when each is compared to
509 the fine sand and medium sand groups.

510 *Crystal fan detritus characterization*

511 The densities of the crystal fan detrital light and heavy fractions are shown in **Table 2**, as
512 well as the relative proportions that fall into the light and heavy fractions are shown in **Table 2**.
513 The Rainstorm crystal fan detritus is extremely unusual: the majority of the sediment, by both
514 volume and mass, was separated into the heavy fraction during heavy liquids separation using
515 methyl-iodide. A more typical split between the heavy and light fractions of a sedimentary
516 sandstone would be on the order of 1% or less.

517 The density of the light fraction ($\rho = 2.62 \text{ g/cm}^3$) is similar to that of quartz ($\rho = 2.65 \text{ g/cm}^3$)
518 and feldspar minerals ($\rho = 2.55 \text{ to } 2.78 \text{ g/cm}^3$). The density of the heavy fraction ($\rho = 3.19 \text{ g/cm}^3$)
519 is slightly less than that of methylene iodide ($\rho = 3.32 \text{ g/cm}^3$), the heavy liquid used to separate the
520 detrital minerals. It is possible that the methylene iodide, which had been previously used and
521 recycled in an active mineral separation lab, was incompletely reclaimed following mixing with
522 ethanol, and so the heavy liquid used in this separation was less dense than reported by the
523 manufacturer. It is also possible that grain-grain interactions during settling of this unusual
524 sediment enriched in heavy minerals resulted in sediment clumping or dragging, with some light

525 grains sinking to the bottom, or that some air remained in the pore spaces of sediment during
526 displacement with water, leading to a too-low density measurement.

527 Grain size data for both the heavy and light crystal fan detrital fractions are shown in **Table**
528 **2** and in **Figure 8**. The heavy fraction D₅₀ grain size (43.1 μm) is much finer than the light fraction
529 D₅₀ grain size (64.9 μm). The predicted heavy grain size, based on the four size-density
530 relationships used in **Figure 4** (equal-sized heavy mineral grains, “grain shielded” small heavy
531 mineral grains, water settling, and air settling) and the measured density of the heavy fraction, is
532 shown in **Table 3** along with the difference between these calculated grain sizes and the measured
533 grain size. The heavy fraction D₅₀ grain size is finer than predicted by all size-density relationships
534 predicted here, but is closest to the “grain shielding” case using the size-density relationship
535 quantified in Garzanti et al. (2008) from measurement of a modern aeolian sand.

536 *Modeling size-provenance relationships in the Ediacaran-Terreneuvian Death Valley succession*

537 Modeled results shown in **Figure 9** show the outcome of size-dependent selection of grain
538 sizes drawn from this succession. As the grain size of the modeled deposit increases, the relative
539 size and contribution of Grenville-aged grains to the resulting spectra also do. At a depositional
540 grain size range of 69 ± 16 μm, between 2.4% and 20.3% (95% confidence interval) of the total
541 spectra came from Grenville-aged (between 1 and 1.2 Ga) sources. 5.4% of the total crystal fan
542 spectra comes from Grenville-aged sources, falling within this confidence interval. At a
543 depositional grain size range of 154 ± 50 μm, between 19.7% and 55.3% (95% confidence interval)
544 of the total spectra came from Grenville-aged (between 1 and 1.2 Ga) sources. 72.4% of the Wood
545 Canyon Formation sample comes from Grenville-aged sources, well outside the 95% confidence
546 interval identified here.

547 **DISCUSSION**

548 *Process in an idealized sedimentary system*

549 Results from our model of source mixing and transport-dependent grain size sorting and
550 deposition highlight that transport process may generate offsets between the detrital zircon and
551 bulk sediment histories in a sample. Such decoupling has been observed between quartz and zircon
552 in an ancient basin (Augustsson et al., 2019). This new model provides a framework for exploring
553 the causes and consequences of such decoupling.

554 **Base case**

555 In this case, with grain size held equal between the two sources, the resulting detrital zircon
556 spectra match the quartz sources well. They are a faithful and quantitative record of sediment
557 source regardless of transport process on average, although randomness within individual spectra
558 can still lead to discrepancies between the true and the measured provenance record. Some of the
559 parameters held constant in this simulation—including consistent zircon fertility and constant
560 grain size between source regions—may often not be equal between source regions in natural
561 settings.

562 **Case 1**

563 In this case, an initial difference in the average detrital zircon size of two populations
564 resulted in a wide range of resultant age spectra. The resulting spectra cover a range of possibilities,
565 from mostly Source 2 to all Source 1, depending on the transport process, and therefore the size-
566 density relationship of the deposit. The ratios of source areas based on zircons do not always
567 quantitatively capture the bulk sediment record.

568 **Case 2**

569 As in Case 1, an initial difference between detrital zircon populations results in a range of
570 spectra due to size-sorting by transport process. The spectra range from mostly Source 1 grains to
571 mostly Source 2 grains. The ratios of source areas based on zircons do not always quantitatively
572 capture the bulk sediment record.

573 Because of the grain size-dependency of the size-density relationship, changes in grain size
574 due to transport process are smaller at these smaller sediment sizes. One consequence of this is
575 that even very small differences between detrital zircon population sizes can affect the outcome of
576 the spectra, if the grain sizes are very close to the depositional sizes for zircon. The other
577 consequence is that if grains are far outside the depositional window (for example, in this case a
578 zircon source with an average size much larger than the 30-40 micron grains that are deposited),
579 they will not be represented in the deposit.

580 **Process in provenance records**

581 These results demonstrate that transport process can impart a signature to the detrital zircon
582 record of sediment if there are systematic differences in grain size between zircon populations.
583 The properties of the transport medium, as well as grain-grain and grain-bed interactions, play a
584 role in defining the size-density relationship of mineral grains in a deposit.

585 Each case explores how transport process could impart different signatures relative to the
586 same bulk quartz grain size. Size-density relationships are unlikely to remain constant through a
587 succession. For example, strata may oscillate between a size-density relationship controlled
588 primarily by settling to a size-density relationship affected by selective entrainment, as observed
589 by Steidtmann and Haywood (1982). Grain size also varies within a stratigraphic succession. In
590 natural sediments, both bulk sediment size and size-density relationships change through a
591 succession, mixing and matching different size-density relationships across grain sizes. In the
592 context of this simple model, this might be analogized by moving from the Case 1 “grain shielding”
593 spectra measured in a sandstone deposited from traction to a Case 2 “settling through water”
594 siltstone layer. The synthetic spectra in these two cases differ in the relative contribution of zircons
595 from different source areas but are identical in bulk sediment provenance. The complexity of real
596 spectra will be greater than the simple examples shown here, as they typically contain more than
597 two sources, and the zircon grain size distributions of real zircons are likely more complex than
598 the generalized normal distributions used here.

599 Analysis of modern rivers suggests that the presence of a population, rather than the size of
600 that population, is more significant in the interpretation of detrital zircon spectra (Link et al., 2005).
601 Our study supports this conclusion, suggesting that differences in grain size may affect the size of
602 a population without necessarily requiring a change in provenance. However, we also note that
603 under some conditions, detrital zircons from a particular source may not be deposited in a unit at
604 all, depending on size characteristics.

605 *The Ediacaran-Terreneuvian succession of Death Valley*

606 **Detrital zircon size**

607 Analysis of our dataset show that it is susceptible to hydrodynamic sorting (**Figure 7,**
608 **Table 4**). Across all grains considered here, Grenville-aged grains are the largest on average
609 (median long axis size of 160 μm). Other detrital zircon grains are somewhat smaller (median long
610 axis size of 122 μm for grains between 1.2 and 1.5 Ga and 112.7 μm for grains between 1.5 and
611 1.9 Ga) and the oldest grains (older than 1.9 Ga) have the smallest median long axis size within
612 the dataset (103.2 μm) (**Figure 6**). There is substantial overlap in sizes between these age ranges.
613 Large and small zircons occur within each age bin, and the age of an individual grain cannot be
614 derived from its size alone. However, an overall trend of average grain size decreasing with grain

615 age is clear. A similar trend is noted within other datasets collected on other continents (Lawrence
616 et al., 2011; Yang et al., 2012).

617 One hypothesis consistent with these data is that older zircon grains are subject to more
618 sedimentary cycles, and erosion and abrasion results in a smaller average size. Another hypothesis
619 is that there is been time-bound variation in the sizes to which zircons grow through Earth history.
620 This suggestion is consistent with the observation that Zr concentrations have increased in Earth's
621 continental igneous rocks through time and that overall zircon abundance has increased over the
622 last 4 Ga (Keller et al., 2017). If zircon size is linked to Zr concentration and abundance, which
623 seems probable based on observations of the high-Zr Grenville-aged granites that also produce
624 abundant, large zircons (Moecher and Samson, 2006; Liu et al., 2017; Samson et al., 2018), then
625 secular variation in average igneous zircon size could result. This systematic variation could then
626 impact the detrital zircon record. Limited available grain size data for detrital and igneous zircon
627 grains across Earth history makes further assessment of this hypothesis challenging.

628 **Grenville sources**

629 **Figure 6** shows that Grenville-aged grains are large relative to other detrital zircon
630 populations. This result is consistent with reports that Grenville zircons are large (Moecher and
631 Samson, 2006) and with observations that both Grenville rocks in particular, and igneous rocks
632 connected to the assembly of Rodinia in general, are Zr-rich (Liu et al., 2017; Samson et al., 2018).
633 One impact of this is that as simulated spectra are generated for coarser and coarser samples, the
634 relative contribution or peak associated with Grenville-aged sources increases (**Figure 9**). To some
635 degree, changing grain size, rather than any change in sediment provenance, may be responsible
636 for changing the contribution of Grenville-aged zircons between samples. This is perhaps best
637 illustrated by the relatively fine-grained Rainstorm samples, which have minor Grenville-aged
638 peaks and are the finest samples studied. Although the lack of a prominent Grenville-aged peak in
639 these samples may truly reflect a lack of Grenville-aged sources, our analysis shows that this is a
640 non-unique explanation. **Figure 9** demonstrates that synthetic spectra randomly chosen from the
641 suite of zircons analyzed through the succession with a similar depositional grain size will also
642 have minor to absent Grenville-aged peaks. These results are consistent with the Nd isotope
643 analyses done by Schoenborn et al. (2012) which suggests a Yavapai-Matzatzal source for the
644 Rainstorm Member. As in that study, these data cannot definitively rule out a cryptic Grenville
645 source (Schoenborn et al., 2012); a Grenville source may not be represented in the detrital zircon

646 record because of sediment sorting. The minor Grenville-aged peaks in both samples above and
647 below this unit, combined with the potential for a “missing” Grenville peak in the Rainstorm
648 crystal fans controlled by grain size, suggest that Grenville sources could have been minor but
649 persistent through the interval.

650 The prominent Grenville-aged peak in the Wood Canyon Formation noted by Stewart et
651 al. (2001) was the special focus of Muhbauer et al. (2017), which found no evidence for grain size
652 sorting exerting a strong control on the presence of the Grenville-aged peak. We note that the
653 Wood Canyon Formation zircons are the coarsest population examined in this study. As seen in
654 **Figure 9**, synthetic spectra with a coarser grain size will develop more prominent Grenville-aged
655 peaks. However, none of these synthetic spectra—even the coarsest, which takes the mean and
656 standard deviation of its depositional size from the Wood Canyon Formation sample—reaches the
657 same magnitude Grenville-aged peak. This suggests that, although grain size may indeed
658 contribute to the substantial Grenville-aged peak seen in the Wood Canyon Formation, grain size
659 alone does not account for it. In another example, grain size alone is also unlikely to account for
660 the minimal or absent Grenville-aged peak seen in the upper Stirling Quartzite or lowermost
661 Johnnie Formation; synthetic spectra coarser than 100 μm have clear Grenville-aged peaks (**Figure**
662 **9**). These data support the conclusion that shifts in the sandstones may represent genuine shifts in
663 provenance (Schoenborn et al., 2012; Muhlbauer et al., 2017), though we note that fine-grained
664 samples from these units are not included in this study. It is possible that size influence would be
665 apparent in a finer-grained sample suite from these units.

666 These results support the case that change in detrital zircon spectra cannot be *prima facie*
667 understood as evidence for changes in provenance (Lawrence et al., 2011; Ibañez-Mejía et al.,
668 2018). In some cases, gaps or peaks in detrital zircon age spectra may be understood as either a
669 change in provenance or a shift in grain size or a shift in sedimentary processes. Considering both
670 age and grain size together can either highlight the non-unique interpretations of a detrital record
671 or bolster the case for true changes in provenance.

672 **Process and provenance in Rainstorm Member crystal fans**

673 Previous studies have investigated the effect of size sorting on detrital zircon spectra in both
674 modern and ancient examples (Hietpas et al., 2011b; Lawrence et al., 2011; Slama and Kosler,
675 2012; Augustsson et al., 2017; Muhlbauer et al., 2017; Ibañez-Mejía et al., 2018; Malkowski et
676 al., 2019; Leary et al., 2020). This study integrates detrital zircon data through a succession in both

677 fine and coarse-grained material. As described above, we find evidence that grain size can exert
678 control on the detrital zircon spectra in our dataset, in particular for the size of the Grenville-aged
679 peak in fine-grained samples. We also characterize how the grain size of detrital zircons within a
680 sediment can depend—both in an absolute sense, and relative to the average bulk sediment grain
681 size—on the transport processes involved. Changes in transport process, and the type of grain-
682 grain or grain-bed interactions, can yield changes in the detrital zircon grain size within a sample
683 by affecting the size-density distribution of the sediment. Here, we investigate the size-density
684 distribution of detrital sediment within the Rainstorm crystal fans to further elucidate
685 environmental conditions during their deposition.

686 Rainstorm Member zircons and sediments are, on average, finer than other units within this
687 succession—consistent with Rainstorm Member deposition during an interval of maximum
688 flooding (Summa, 1993; Bergmann et al., 2011, 2013). Examining the measured grain size offsets
689 between the heavy and light detrital fractions of Rainstorm crystal strata is puzzling (Table 3). The
690 heavy fraction is finer than the fine fraction, as expected—but it is much finer than would be
691 predicted for either settling in air or in water. It is also much finer than would be predicted by the
692 grain size-density relationship quantified by Garzanti et al. (2009) from observation of aeolian
693 dunes, showing that selective entrainment in the Rainstorm yielded a different size-density
694 relationship with a larger size change at lower densities. Entrainment equivalence between mineral
695 grains has been expressed as a function of grain size, mineral and fluid density, shape, stacking
696 pattern, and bed roughness (Komar, 2007) although the relationship remains largely empirically
697 untested. The subaqueous depositional setting of the Rainstorm Member, and resulting greater
698 density contrast between heavy and light minerals compared to subaerial settings, seems to have
699 played a role in defining a more extreme size-density relationship compared to the modern aeolian
700 dunes studied by Garzanti et al. (2008).

701 Sedimentary structures within the Rainstorm Member can provide information about flow
702 regime and bed roughness during its deposition. The Rainstorm Member contains special
703 hummocky cross-stratification and amalgamated scours interpreted as evidence for intermittent,
704 high-energy storm events, as well as abundant storm-generated edgewise conglomerates (Summa,
705 1993; Pruss et al., 2008; Bergmann et al., 2013). Both discontinuous laminae within scours and
706 intraclast beds form from periodic erosion, reworking, and deposition of sediment. This suggests
707 that extensive reworking of sediment occurred episodically within the Rainstorm Member. The

708 formation of heavy mineral-rich placer sands within the swash zone of beaches during storm events
709 has been documented in modern settings (Woolsey et al., 1975; Komar and Wang, 1984)

710 The centimeter-scale crystal fans growing up from the sediment-water interface within the
711 Rainstorm Member could act as baffles for sediment, increasing bed roughness and shielding fine
712 grains from entrainment. Crystal fans are often found with mineral grains in inter-blade fill (note
713 gray, heavy mineral-rich zones in interstitial spaces of crystal fan in **Figure 3**) (Summa, 1993;
714 Pruss et al., 2008; Bergmann et al., 2013). Entrainment and settling of sediment as it passes over
715 a rough bed leads to the formation of heavy mineral-enriched deposits in the interstices of gravel
716 beds in some modern placers (Slingerland and Smith, 1986; Day and Fletcher, 1991); possibly the
717 interstices of crystal fans protruding from the sediment-water interface could play a similar role.
718 The importance of the crystal fans as sedimentary baffles in forming this heavy-mineral enriched
719 deposit could be assessed by evaluating the heavy mineral enrichment, if any, in Rainstorm
720 sandstones. Because the crystal fans often nucleate on and grow from the heavy mineral-rich
721 deposits—thus postdating them—they may play a secondary role in its formation, if they play a
722 role at all.

723 Rapid cementation of detrital mineral grains in a carbonate-precipitating environment could
724 enhance selective entrainment by cementing grains in place. This early cementation would impact
725 smaller grains, regardless of their density, to a greater degree than larger grains. Carbonate
726 cementation of siliciclastic deposits is ubiquitous through the Rainstorm Member (Summa, 1993).
727 In a study of amalgamated scours within the Rainstorm Member, Summa (1993) suggested that a
728 high degree of sediment cohesion, potentially generated by carbonate cementation, was required
729 to allow the high-angle scour walls, indicating that cementation occurred rapidly following
730 deposition.

731 Thus, selective entrainment due to flow characteristics and bed roughness, and potentially
732 enhanced by rapid carbonate cementation, is consistent with the observed size-density relationship
733 in Rainstorm crystal fan detritus. The observed enrichment in heavy minerals seen in these deposits
734 is also consistent with winnowing of sediment and removal of low-density grains by selective
735 entrainment. Further supporting the role of selective entrainment is the provenance data from the
736 unit, which is typical for the succession—demonstrated by both Nd (Schoenborn et al., 2012) and
737 detrital zircon data to be consistent with expectations for a Laurentian passive margin, and wholly
738 consistent with provenance throughout the succession. The unusual character of the siliciclastic

739 material contained in these strata does not appear to be the result of an unusual sediment source,
740 but of a set of transport processes operating on a siliciclastic deposit to yield an unusual deposit.
741 Winnowing and selective entrainment, combined with low sediment supply, are consistent with
742 other work on the origin of these deposits (Bergmann et al., 2013). In the Rainstorm crystal fan
743 strata, grain shielding effects led to the deposition of an especially fine-grained set of detrital zircon
744 grains. These grain-size effects may have contributed to the minimal Grenville-aged peak seen in
745 these units.

746 Previous work suggested an aeolian source for the heavy mineral-rich silt-sized detritus within
747 Rainstorm Member crystal fans because of the grain size, grain mineralogy and depositional style
748 (Bergmann et al., 2013). The observed size-density relationships in these minerals do not conform
749 with predictions for grains settling through air (**Table 3**) (Bagheri and Bonadonna, 2016; Raffaele
750 et al., 2020). However, sedimentological evidence indicates extensive, storm-influenced
751 reworking of Rainstorm sediment (Summa, 1993; Pruss et al., 2008; Bergmann et al., 2013).
752 Because the size-density relationships of sediments reflect their most recent environment (Hand,
753 1967), any initial air-settling size-density relationship could be overprinted following deposition
754 and reworking in a submarine environment. The winnowing of light minerals by wind from beach
755 dunes is implicated in the formation of some modern beach placer deposits (Woolsey et al., 1975)
756 and may have influenced the Rainstorm Member. The lack of a size-density relationship consistent
757 with settling through air neither rules out nor requires an aeolian influence on siliciclastic detritus
758 the Rainstorm Member.

759 **The Rainstorm Member and the Shuram excursion**

760 One of the motivations for study of the Rainstorm Member is understanding the
761 sedimentological and environmental context of strata that record Earth's most negative carbon
762 isotope excursion. This study has investigated the Rainstorm Member's provenance using detrital
763 zircon and quantified the size-density relationship of siliciclastic detritus within carbonate crystal
764 fan-bearing units. Our data suggest that the provenance of the Rainstorm Member, despite the
765 unusual, heavy mineral-enriched deposits found within, are entirely consistent with provenance
766 throughout the succession. This indicates that the formation of these deposits was the result of
767 sediment transport processes occurring alongside deposition, rather than erosion of a distinct
768 source not represented elsewhere in the succession. Sedimentological evidence for frequent and
769 energetic storm activity indicates that sediment reworking was ongoing through deposition of the

770 unit (Summa, 1993; Pruss et al., 2008). Climate can also influence the composition of clastic
771 deposits (Johnsson, 1993). Available data suggest that the Johnnie Formation did not undergo cold-
772 weather weathering and that it represents moderate to intense weathering of granitoid sources
773 (Schoenborn and Fedo, 2011). Weathering products and elemental compositions were not
774 measured as part of this study.

775 To date, no comparable heavy mineral suite has been reported in other Shuram excursion-
776 bearing strata. In the course of this study, a sample of carbonate-cemented siltstone from the
777 Shuram Formation of Oman was disaggregated through dissolution of the carbonate matrix and
778 separated using heavy liquids (**Data Repository**). The Shuram Formation siltstone shares some
779 characteristics with Rainstorm Member detritus: a coarse silt grain size, carbonate cement,
780 abundant mica, and ample sedimentological evidence for storm action (McCarron, 1999; Le
781 Guerroué et al., 2006b, 2006a; Grotzinger et al., 2011; Bergmann, 2013). In both Oman
782 (McCarron, 1999; Le Guerroué et al., 2006b, 2006a; Bergmann, 2013) and Death Valley (Corsetti
783 and Kaufman, 2003; Corsetti et al., 2004; Kaufman et al., 2007; Pruss et al., 2008; Bergmann,
784 2013), the carbonate strata recording the Shuram excursion include oolite and edgewise
785 conglomerate interbedded with purple or red siltstone. Despite these similarities, the Shuram
786 Formation siltstone was not enriched in heavy minerals. Only about a dozen grains were found in
787 the heavy fraction following separation by heavy liquids and no size-density relationship was
788 measured.

789 This difference between the Shuram Formation and the Rainstorm Member could be
790 explained, in part, by differences in sediment delivery between the two locations. Limited sediment
791 supply is linked to the formation of heavy mineral-enriched deposits in the Rainstorm Member
792 (Bergmann et al., 2013). The Rainstorm Member is 50-100 meters thick (Stewart, 1970; Summa,
793 1993; Bergmann et al., 2013); the strata in Oman that contain the Shuram excursion (Shuram and
794 Buah formations) are 500-600 meters thick in outcrop (McCarron, 1999). Assuming a synchronous
795 excursion in Oman and Death Valley, this implies depositional rates in Oman were more than an
796 order of magnitude greater, and that sediment supply was higher and potentially incompatible with
797 formation of heavy mineral-enriched deposits by winnowing and selective entrainment in a
798 sediment-starved setting. These observations of finer grain sizes, decreased sediment supply, and
799 extensive sediment reworking and winnowing are consistent with interpretations of Rainstorm
800 Member deposition during an interval of maximum flooding (Summa, 1993; Bergmann et al.,

2011, 2013). Other similarities between the Rainstorm Member and other Shuram excursion-bearing successions, including the abundance of siltstone (often carbonate-cemented) and evidence for storm activity (McCarron, 1999; Le Guerroué et al., 2006a, 2006b; Grotzinger et al., 2011; Bergmann, 2013; Husson et al., 2015), remain intriguing.

CONCLUSION

Drawing on insights from draws on insights from process sedimentology, we demonstrate that sediment transport processes, from simple settling through fluid to selective entrainment, may lead to differential biases in detrital zircon records. Recognition that detrital zircon grains are not perfect records of provenance, but transported grains subject to sedimentary processes, encourages greater complexity and care in interpretations of provenance from these records. In the specific case study explored here, we find evidence that detrital zircon grains from the Ediacaran-Terreneuvian succession of Death Valley are susceptible to hydrodynamic fractionation. The effects of hydrodynamic fractionation are especially relevant for the relative contribution of Grenville-aged (1.0-1.2 Ga) zircon grains which are, on average, larger than older grains. The Rainstorm Member, which is geochemically, sedimentologically, and mineralogically unusual, nevertheless has a typical sediment provenance compared to other units in the succession, with substantial input from Yavapai-Mazatzal sources and a potentially cryptic Grenville source due to its fine grain size. The size-density relationship of detrital minerals and heavy mineral enrichment within carbonate strata of the Rainstorm Member are consistent with selective entrainment and winnowing of sediment during storms and low sediment supply during deposition.

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Data Availability

All data and code related to this project are available at <https://github.com/mcantine/rainstorm-zircons>.

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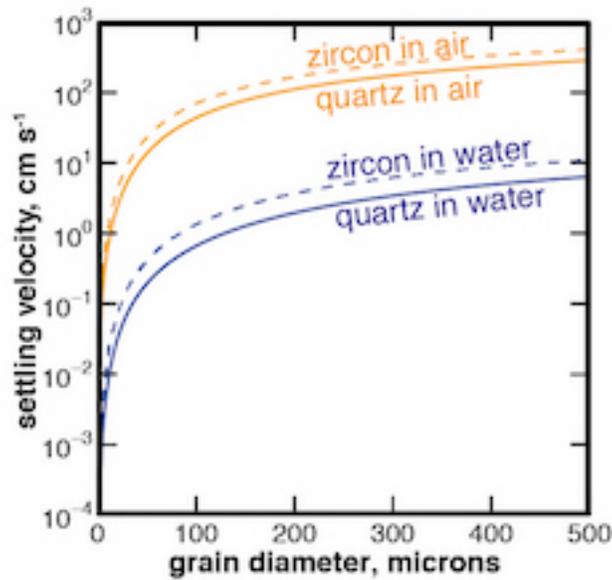
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1383 **Figures and Figure Captions**

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1387 **Figure 1.** The settling velocities of quartz and zircon grains in water and air are shown as a
1388 function of grain diameter based on empirical measurements (Dietrich 1982, Bagheri and
1389 Bonadonna 2016, Raffaele et al. 2020). Due to its greater density, zircon grains settle more
1390 rapidly than quartz grains of the same size, so zircon grains with the same settling velocity as
1391 quartz grains are smaller. The settling velocities of quartz and zircon grains in air are more rapid
1392 because air is less viscous and dense than water. Relative to the same quartz grain size, the
1393 hydrodynamically equivalent zircon in air will be larger than the hydrodynamically equivalent
1394 zircon in water.

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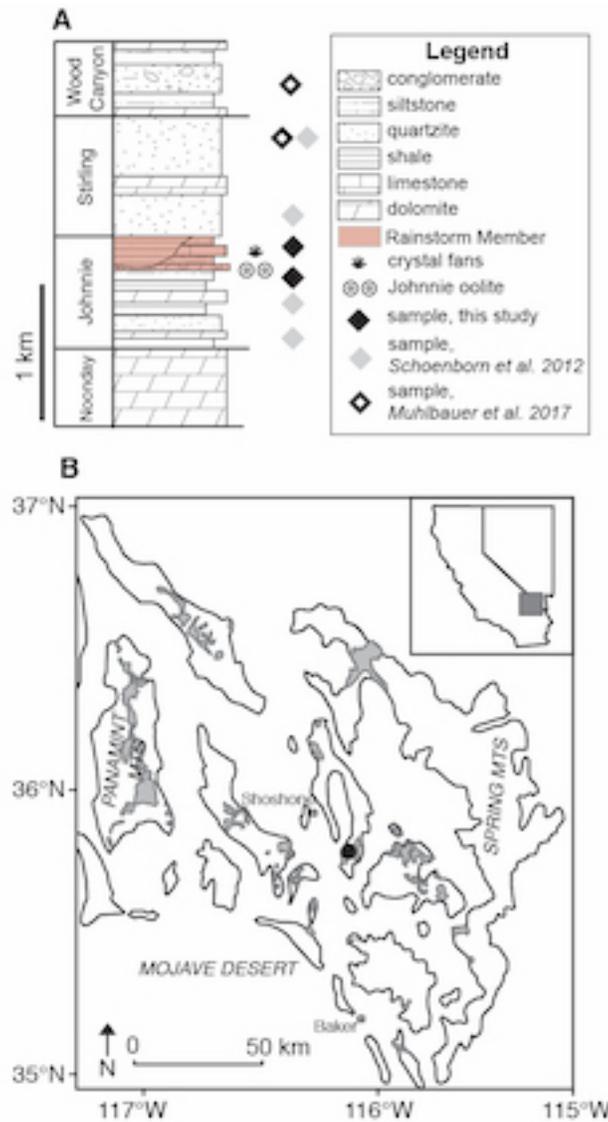
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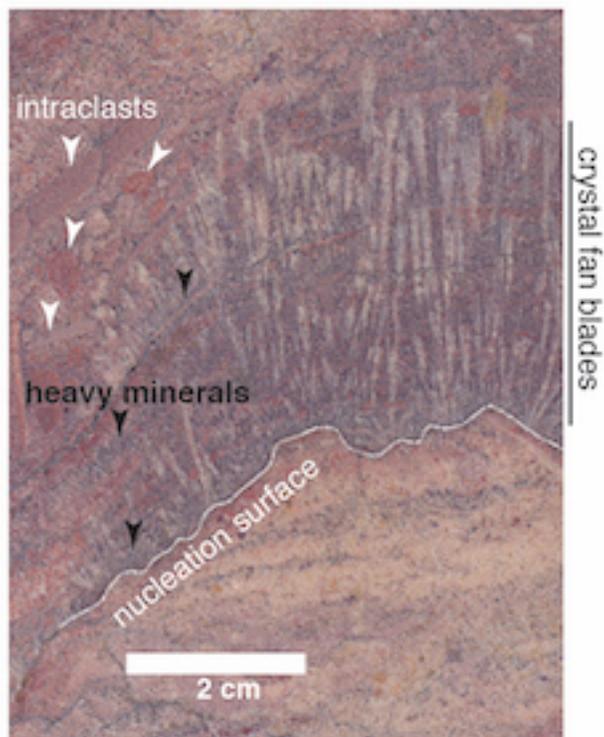
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Figure 2. A. The Edicaran to Terreneuvian succession of Death Valley, Nevada and California, USA, is shown. The lowermost Noonday Dolomite is considered earliest Ediacaran near its base (Prave 1999, Petterson et al. 2015) and the Wood Canyon Formation straddles the Ediacaran-Cambrian boundary (Corsetti and Hagadorn 2000). The approximate stratigraphic position of samples from this study and previous studies (Schoenborn et al. 2012, Muhlbauer et al. 2017) are shown as diamonds. The Rainstorm Member, the particular focus of this study, is highlighted in pink.

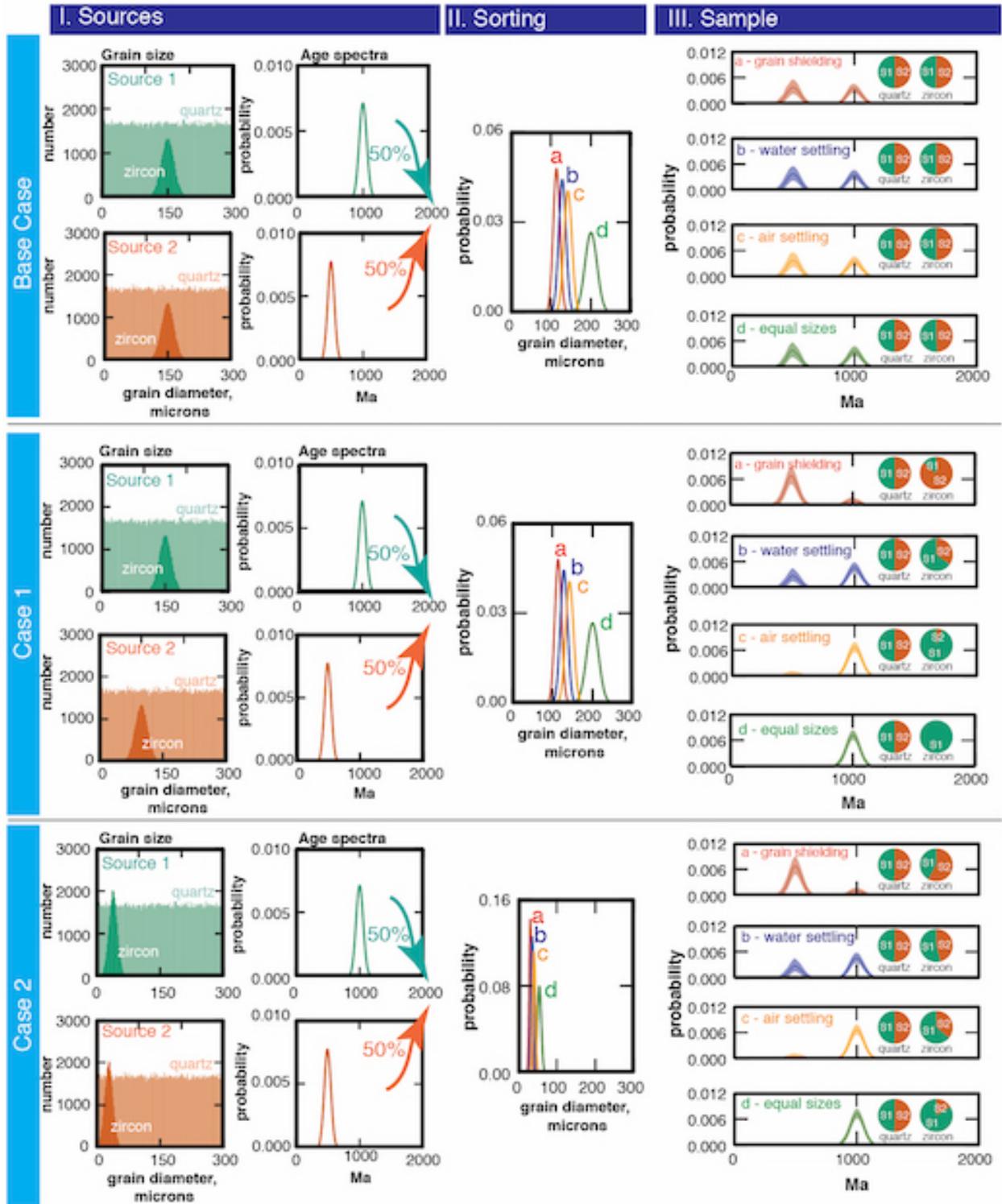
1412 B. The aeral extent of the Johnnie Formation, of which the Rainstorm Member is the uppermost
1413 unit, is shown in gray. The Southern Nopah Range, where Rainstorm samples were collected for
1414 this study, is marked with a black dot. Modified after Bergmann et al. 2011, figure 2.



1415
1416 **Figure 3.** A representative crystal fan-bearing limestone bed from the Rainstorm Member is
1417 shown in a cut and polished slab. Formerly aragonite (Pruss et al. 2008) crystal fans nucleate on
1418 a scoured seafloor surface, traced with a dotted white line. Heavy mineral-enriched clastic
1419 detritus is abundant in the sample, visible as dark gray to black zones; a few examples are
1420 indicated with black arrowheads. Carbonate intraclasts are also abundant, and a few examples
1421 are indicated with white arrowheads. Modified after Bergmann et al. 2013, Figure 5A.

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1434 **Figure 4.** The process-dependent, hydrodynamic sorting of different zircon populations in an
1435 idealized sedimentary system is illustrated here in three cases. The grain size and age spectra of
1436 two source populations, and the proportion in which these populations are mixed, are shown in
1437 column I. Column II shows four different depositional sizes informed by four different size-
1438 density relationships between deposited quartz and zircon, from the “grain shielding” case in
1439 which deposited zircon grains are smaller than predicted by settling relationships, to the “equal
1440 sizes” case in which zircon and quartz grains have the same size. Size-density relationships from
1441 the settling of sediment in water and air are intermediate to these two cases. The resulting
1442 synthetic age spectra are shown in column III (95% confidence interval envelope), along with the
1443 proportions of grains from different sources.

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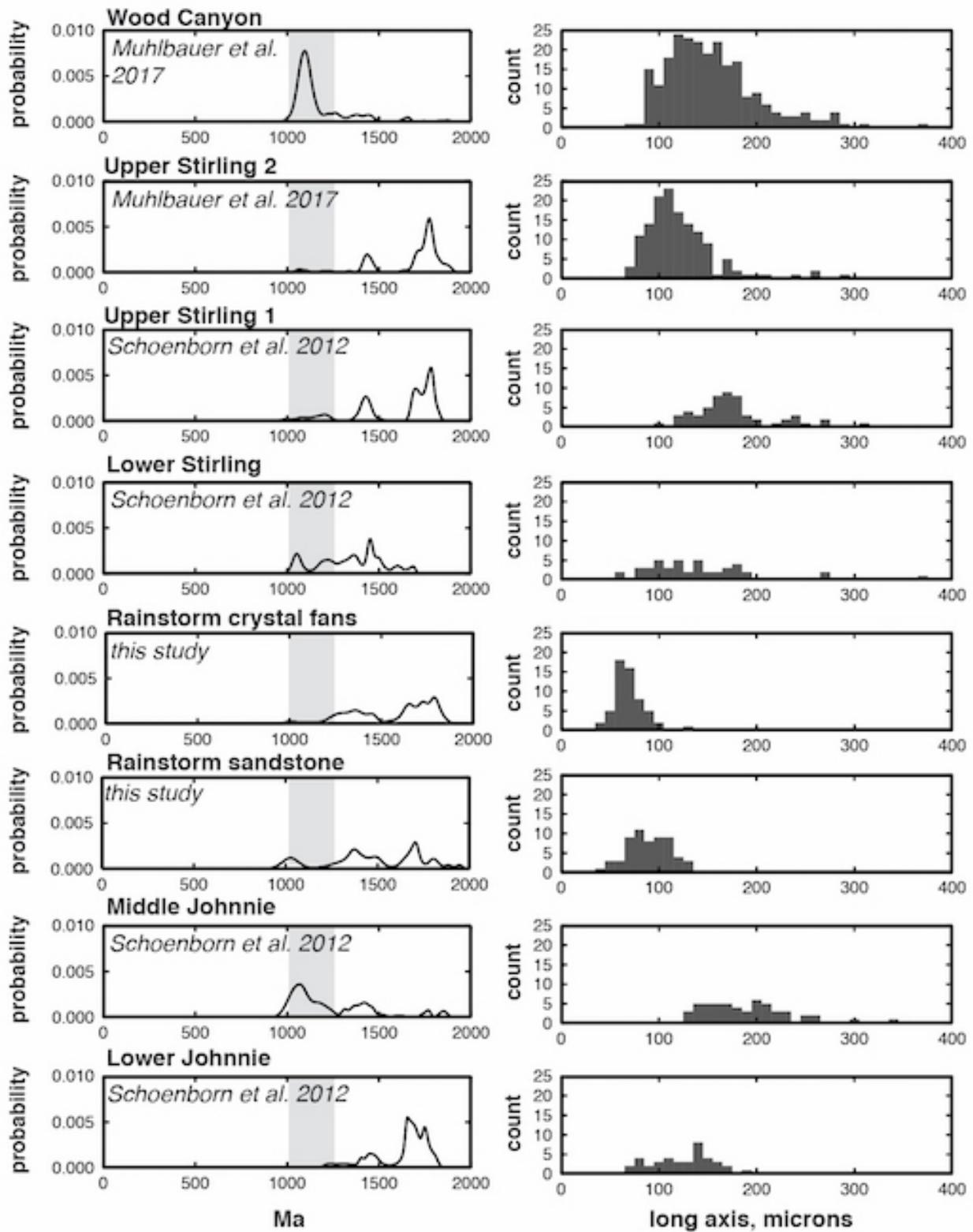
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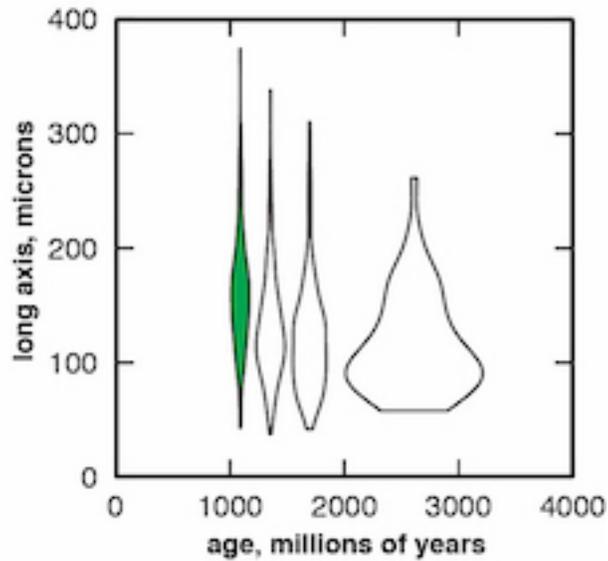
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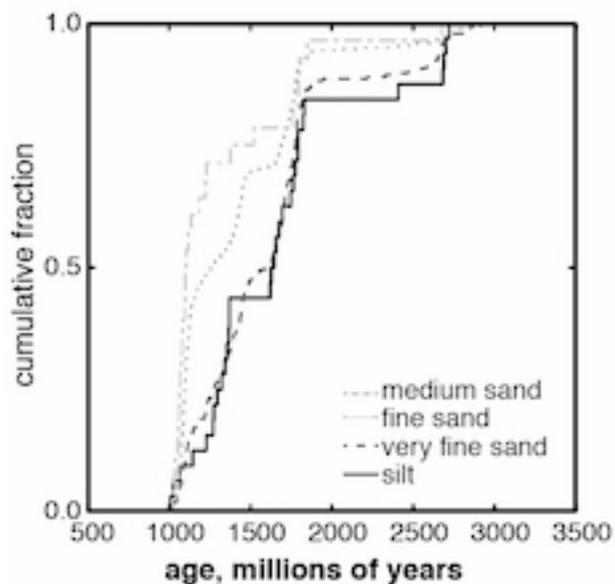
1469 **Figure 5.** Measured age spectra and size distributions for samples included in this study.
1470 Grenville age (1.0-1.2 Ga) ranges are highlighted in gray. Samples from the Rainstorm Member
1471 are finer-grained than other samples.

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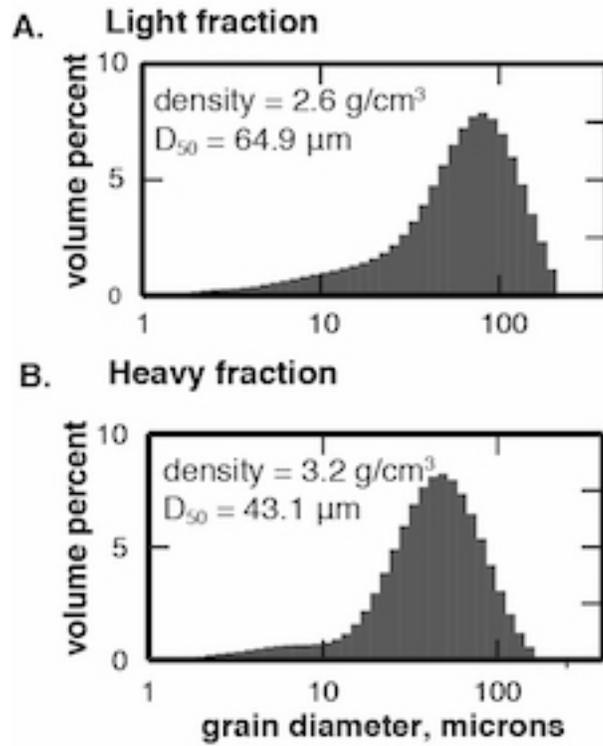


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1476 **Figure 6.** This violin plot describes size variation within and between age populations from
1477 detrital zircons measured by this study and others (Schoenborn et al. 2012, Muhlbauer et al.
1478 2017). Grenville-aged grains (1.0-1.2 Ga, highlighted in green) are larger on average than their
1479 older counterparts, though all age populations contain larger and small grains.

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 1491 **Figure 7.** These empirical cumulative distribution function curves describe the fraction of grains
 1492 and their ages in a detrital zircon dataset measured by this study and others (Schoenborn et al,
 1493 Muhlbauer et al. 2017). The silt and very fine sand-sized grain populations are similar to ach
 1494 other, as are the fine and medium-sand sized populations. The largest difference is for Grenville-
 1495 aged (1.0-1.2 Ga) grains, which make up more than 50% of the fine and medium sand sized
 1496 grains but less than 50% of the silt and very fine sand sized-grains.



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1498 **Figure 8.** LPSA data for both the light and heavy fractions of crystal fan-associated detritus are
1499 shown.

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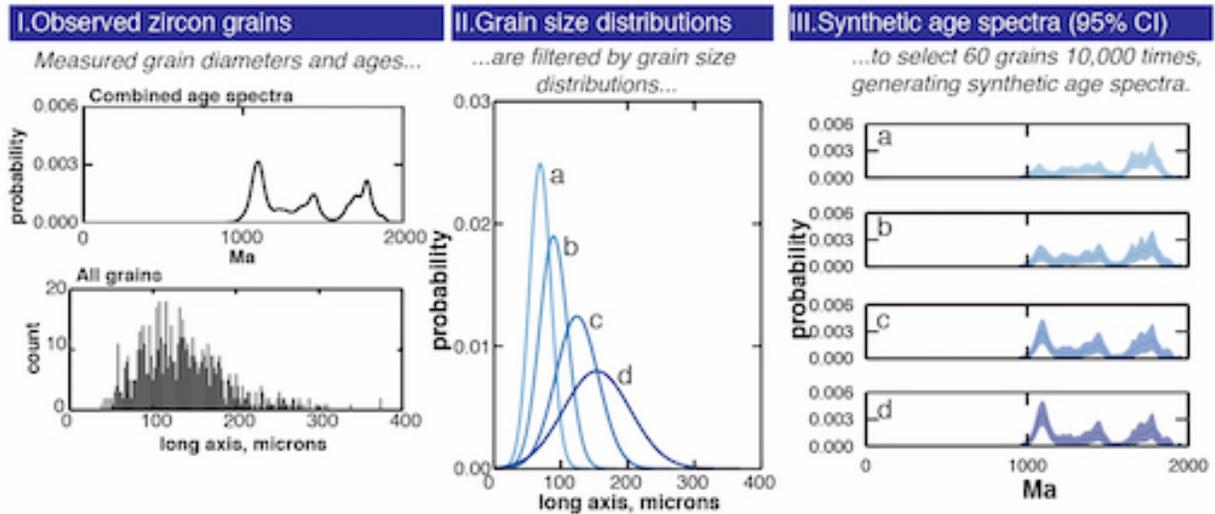
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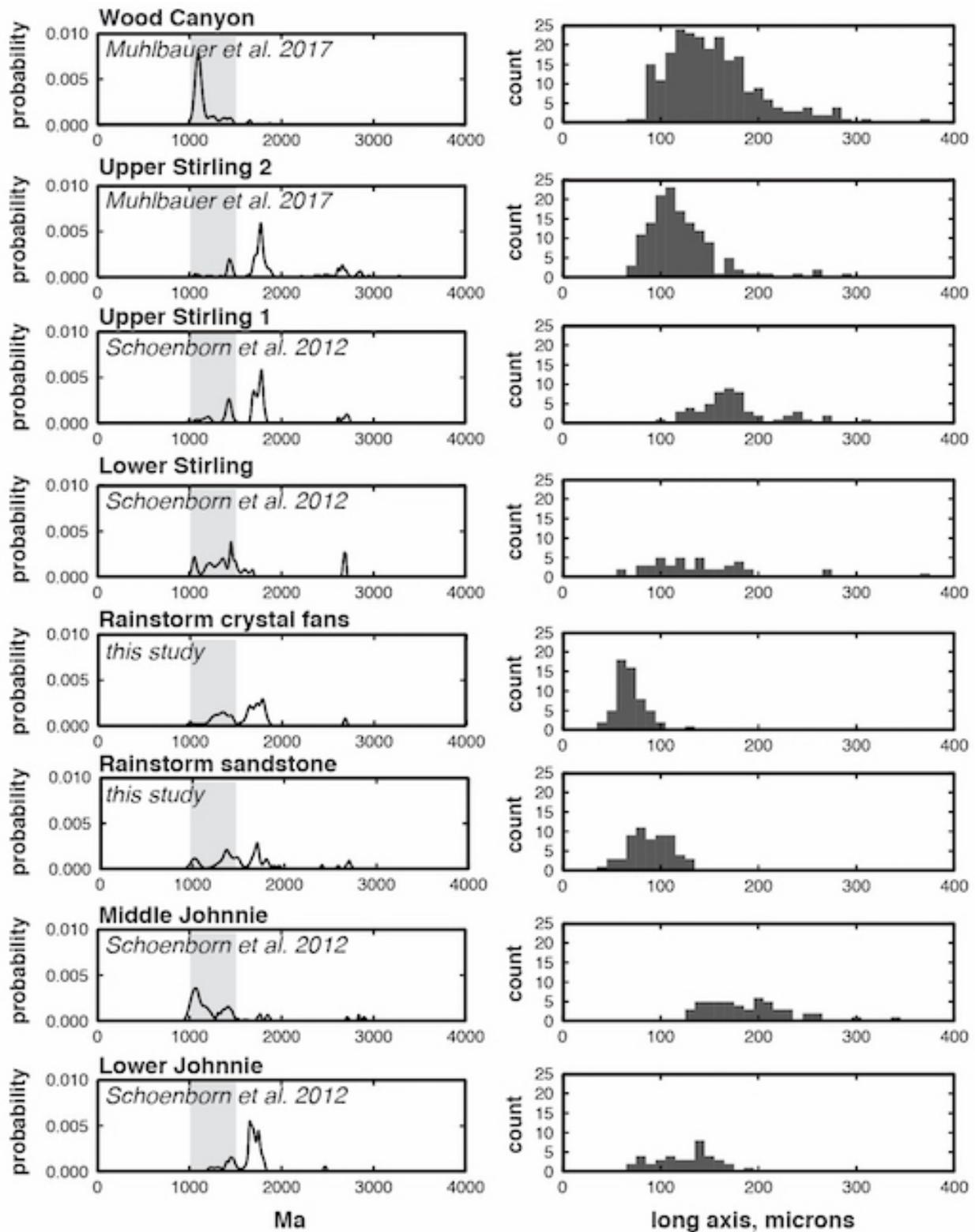
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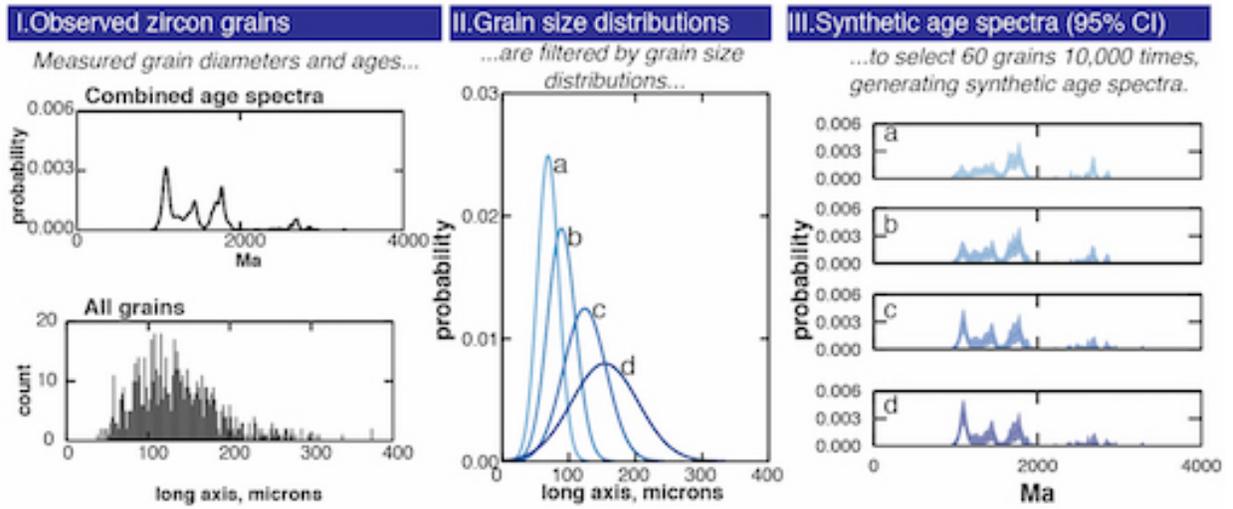
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 1507 **Figure 9.** All measured concordant zircon grains with a known size from this study and others
 1508 (Schoenborn et al, Muhlbauer et al. 2017) are shown in column I. These grains are pooled
 1509 together and 60 grains are chosen 10,000 times to generate a series of synthetic age spectra. The
 1510 grains are selected used size-based probability distribution functions shown in column II. The
 1511 95% confidence interval of the resulting synthetic age spectra are shown in column III. As the
 1512 synthetic samples grow coarser (from a to d), the prominence of the Grenville-aged peak also
 1513 grows.



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1515 **Figure S1.** This figure shows the same data as Fig. 5 in the main text but includes grains older

1516 than 2000 Ma.



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1518 **Figure S2.** This figure shows the same data as Fig. 9 in the main text but includes grains older

1519 than 2000 Ma.

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